

Status of Cosmic-ray Antinuclei Searches



ICRC July 2019

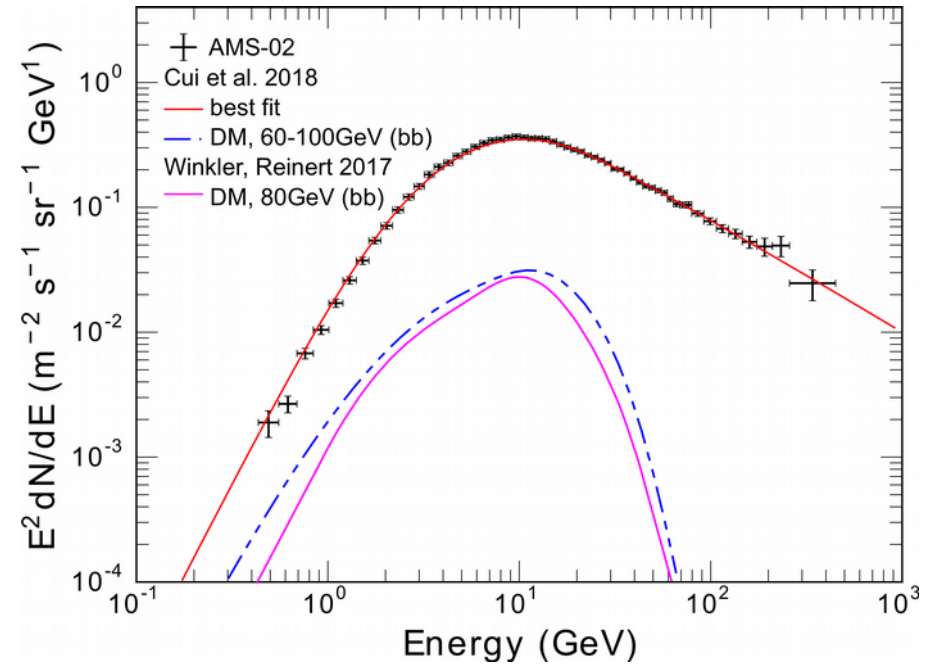
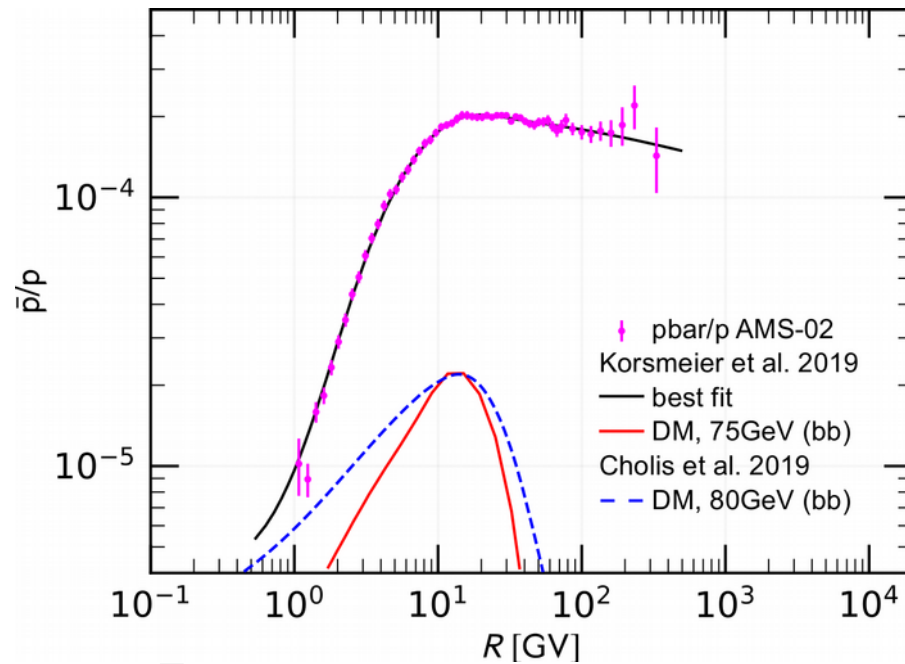
Philip von Doetinchem - philipvd@hawaii.edu

Department of Physics & Astronomy - University of Hawaii at Manoa



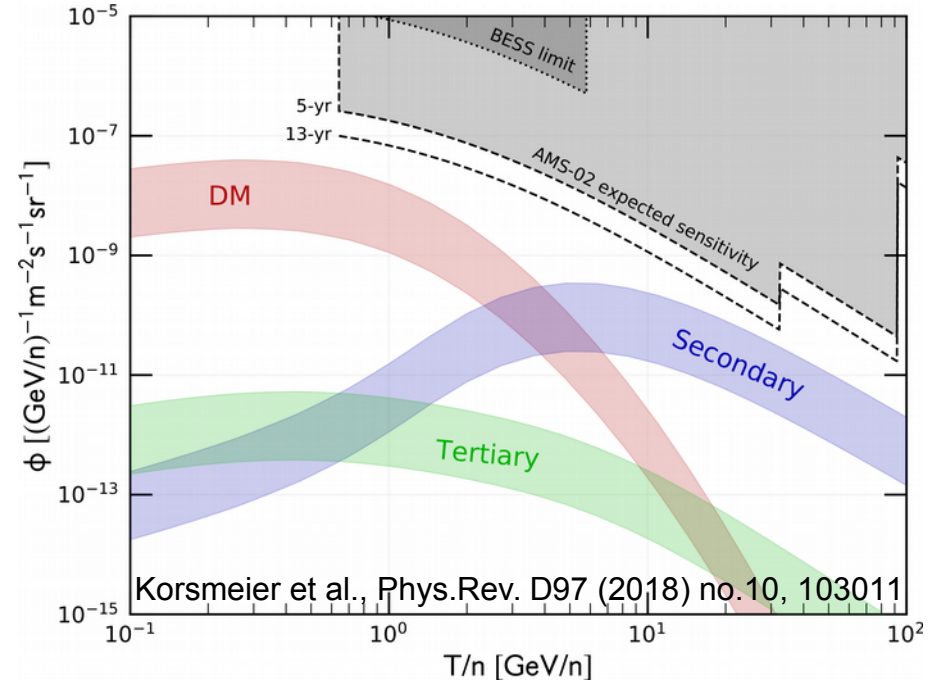
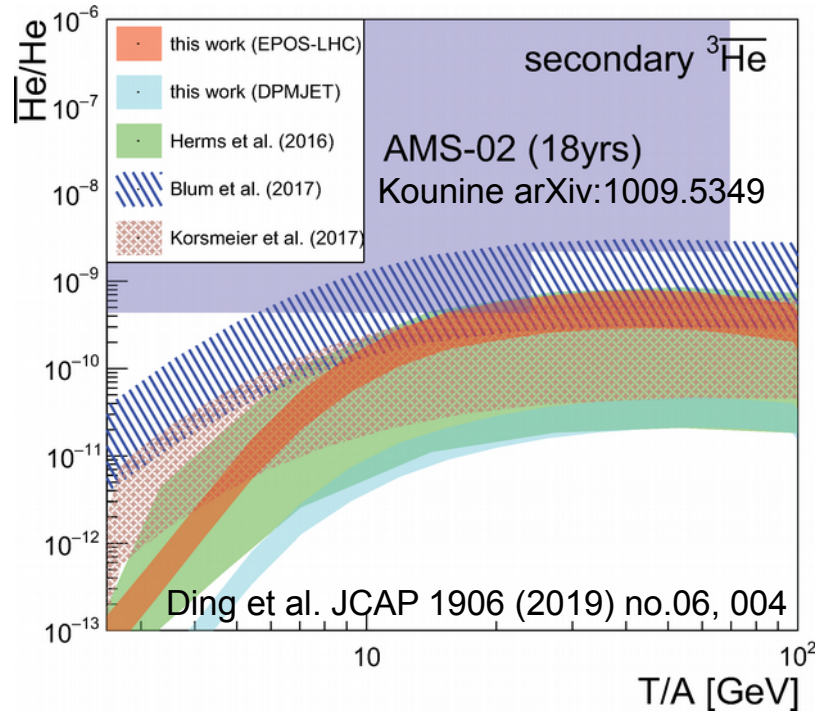
- This talk is a review of some key results of the:
2nd Cosmic-ray Antideuteron Workshop
(UCLA, March 27-29, 2019, 45 participants)
<https://indico.phys.hawaii.edu/e/dbar19>
(organizers: Rene Ong, Mirko Boezio, Kerstin Perez, PvD)
- Measurement of cosmic-ray antinuclei (antiproton, antideuteron, antihelium) is an exciting way to search for new physical phenomena:
 - experimental cosmic-ray search updates
 - potential primary sources of antinuclei
 - uncertainties of production and propagation in our Galaxy
 - path forward

Low-energy Antiproton Excess



- A small \bar{p} excess in AMS-02 data ([Bertucci, 7/27, 11:30am](#); [Tang, 7/25, 1:30pm](#)) above secondary background predictions at $R \sim 10$ GV was found in various studies (e.g., [Hooper, 7/31, 6pm](#))
→ significance level unclear, but defines interesting target for \bar{d} searches
- This excess is in agreement with explaining the γ -ray excess at the Galactic Center with a signal from dark matter annihilation
- New analysis ([Boudaud, 7/25, 1:45pm](#)) with different technique for error correlation matrix does not need additional contribution

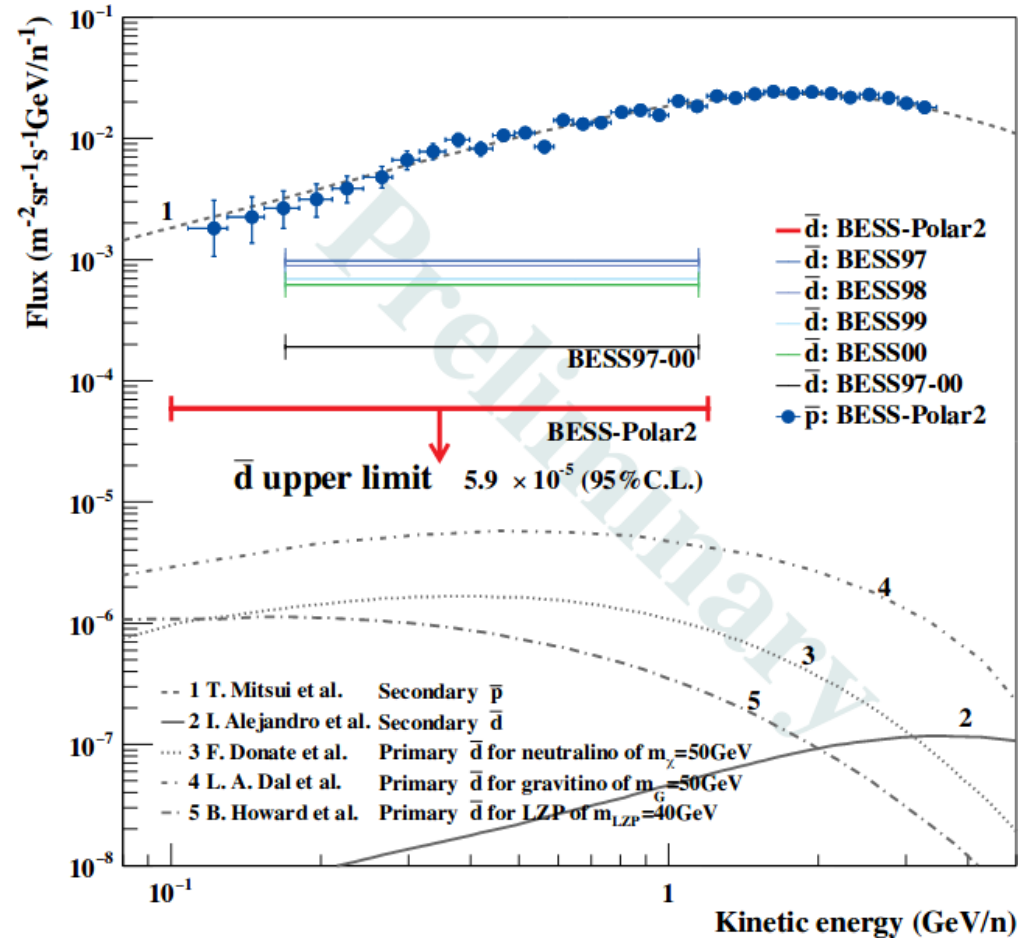
Antihelium



- AMS-02 reported that antihelium candidate events have been observed (Ting, CERN Colloquium, 5/24/18) → interpretations are ongoing
- key constraints: no explanation should overproduce antiprotons and antideuterons?
 - Secondary astrophysical background**: uncertainties in antihelium formation models leave room to explain higher energy antihelium events
 - Dark matter**: difficult to construct a model that does not overproduce antiprotons
 - Nearby antistar**: at distance of ~ 1 pc, profound impact on Big Bang Nucleosynthesis (Poulin et al., Phys.Rev. D99 (2019) no.2, 023016)

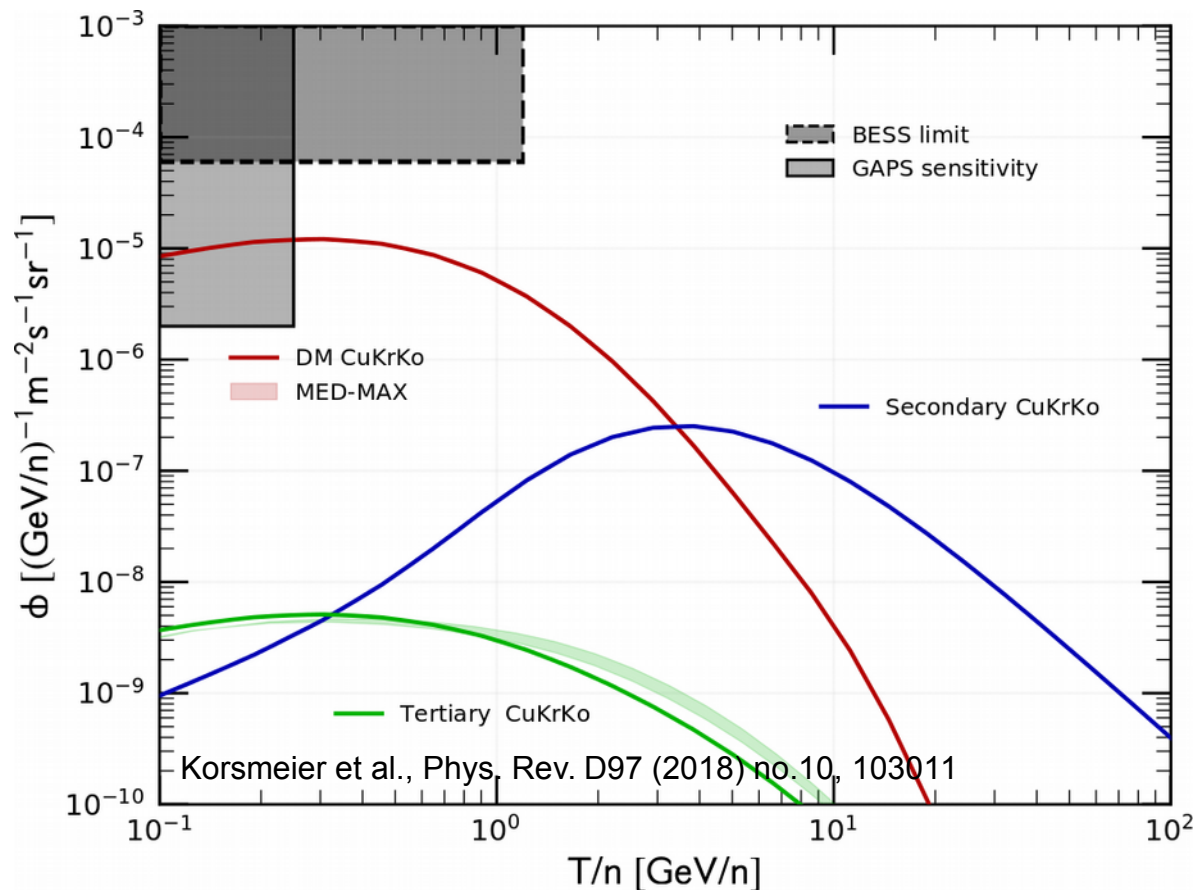
New BESS Antideuteron Results

- Antiproton, antideuteron, antihelium need to be explained in one coherent picture:
 - Antiproton and antihelium both constrain antideuterons
- New analysis from BESS-Polar II flight (2007, 24.5 days):
 - Extended to low-energy range for antiprotons
 - New antideuteron limit (analysis is ongoing)
 - Details: Sakai, 7/25, 2:15pm



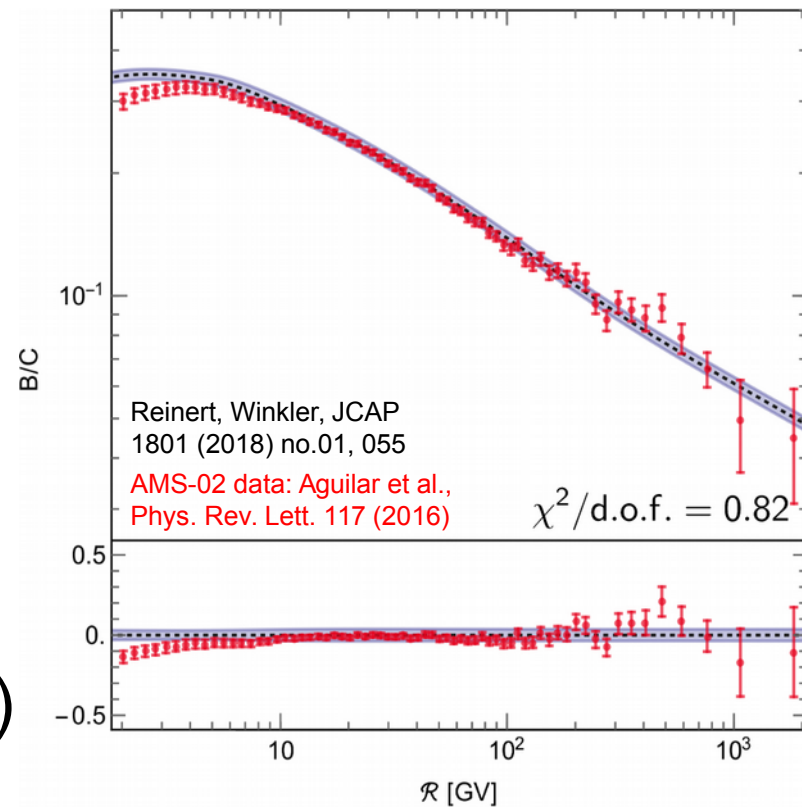
Antideuteron from Dark Matter

- γ -rays from Galactic center and antiproton excess fitted with DM
→ use the same parameters to make prediction for antideuterons:
 - Coalescence momentum from ALICE data
 - Flux propagation with GALPROP
- Dark matter hint from cosmic-ray antiprotons can produce antideuteron signal within experimental reach



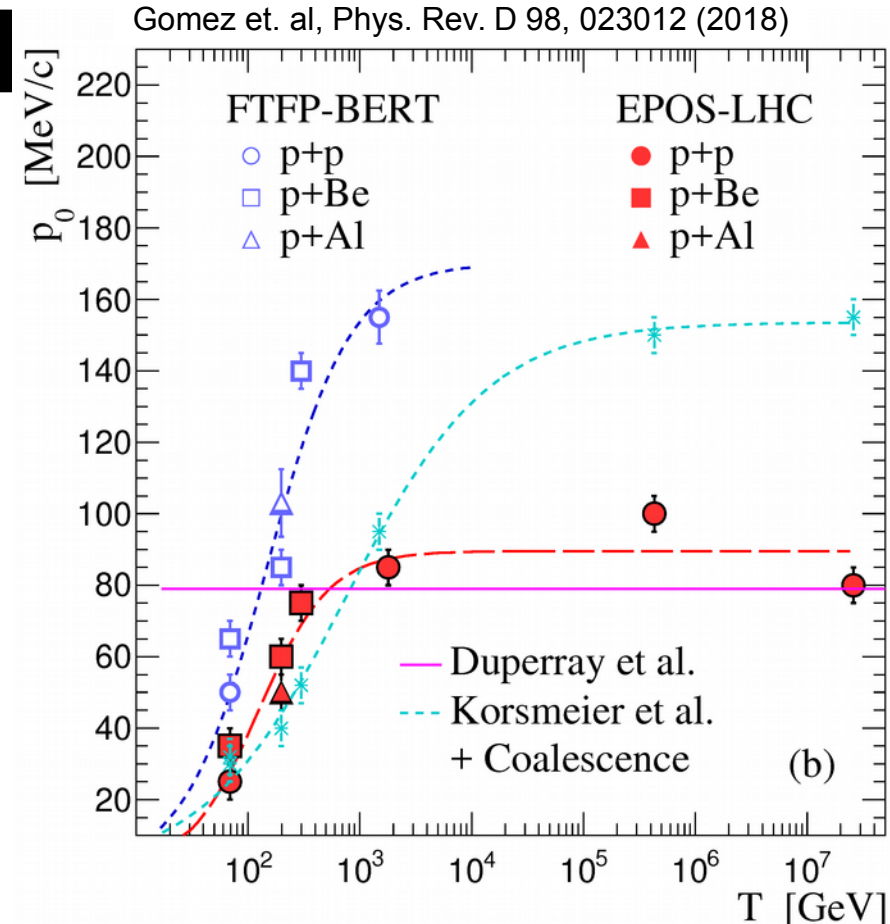
Propagation Uncertainties

- An important constraint for antinuclei flux from dark matter annihilations is the Galactic halo size, which directly scales the observable flux
- Fits of cosmic-ray nuclei data are very important to constrain cosmic-ray propagation models (e.g., Li/C, Li/O, Be/C, Be/O, B/C, B/O)
- Measurement of relevant primary cosmic ray and interstellar medium cross sections is important



Antinuclei Formation

- **Antinuclei formation process breaks the degeneracy of antinuclei with antiprotons**
- **Coalescence:**
 - \bar{d} can be formed by an \bar{p} - \bar{n} pair if relative momentum is small compared to coalescence momentum p_0
 - Hadronic generators not really tuned for antiparticle production
 - hadronic generators do not model coalescence
- **Thermal model:**
 - Antinuclei directly formed at hadronization stage
 - tuned hadronic generator (**Pierog (EPOS) 7/27, 2:45pm**) can produce antideuterons directly from quarks
- **Quantum mechanical approach**
 - Enhanced antihelium formation due to interference of the wave functions of the constituents (Blum et al., Phys. Rev. D96 (2017) no.10, 103021)
 - Wigner function based coalescence model (Kachelriess et al., arXiv: 1905.01192)



- **more high statistics data needed to constrain antinuclei formation models**

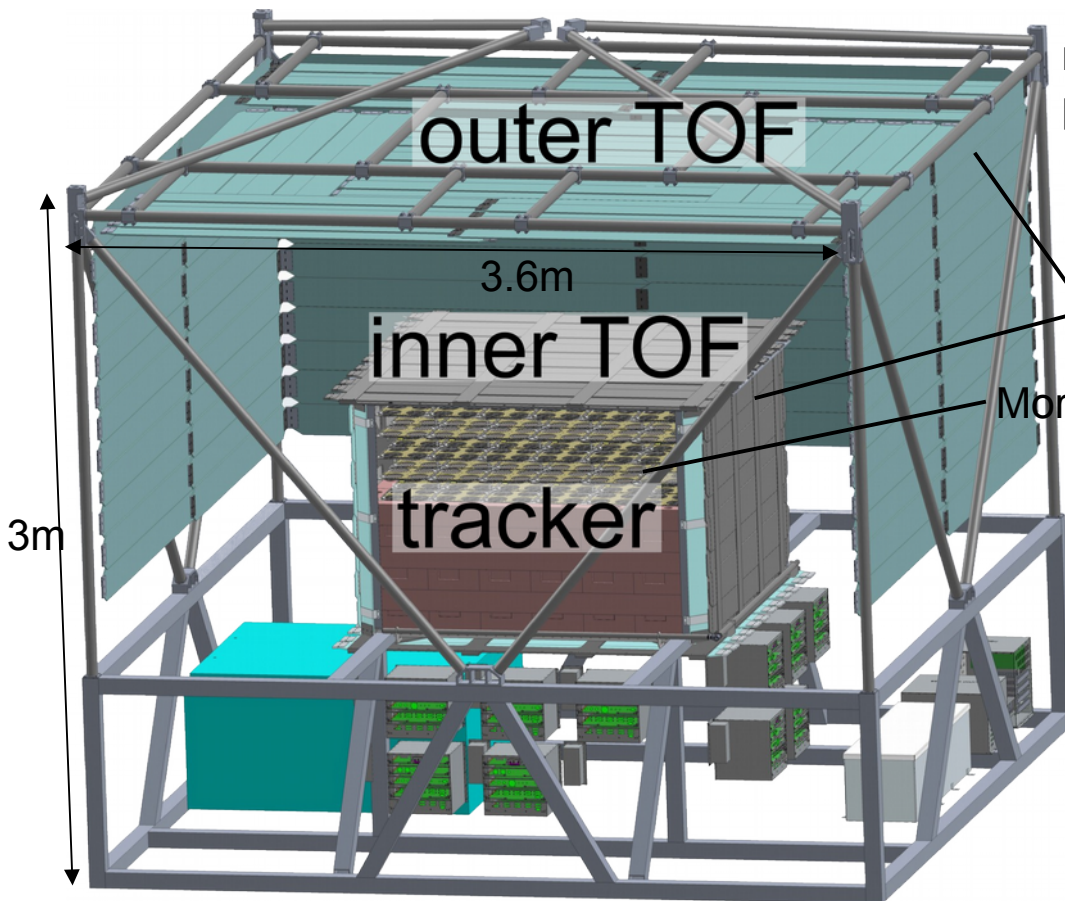
GAPS Experiment

Columbia U, UCSD
UCLA, UCB,
U Hawaii, MIT



mass: ~2,500kg
power: ~1.5kW

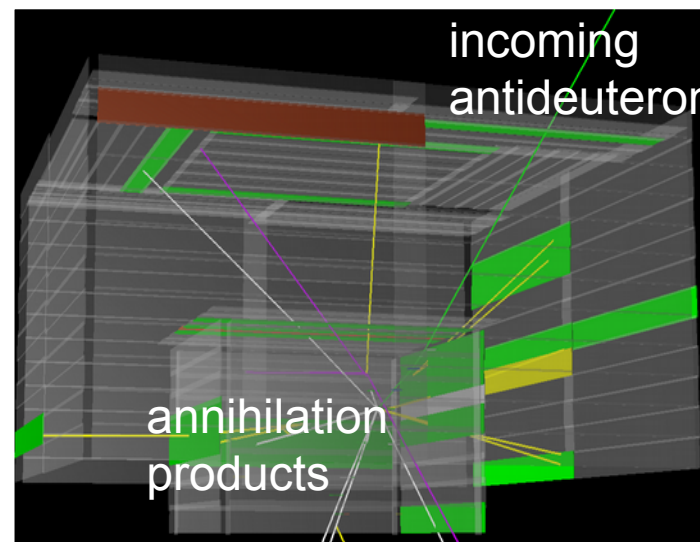
Details: Bird, 7/25, 2:45pm
Posters: Munini, Quinn, Scotti



TOF with SiPM readout

More than 1,000 Si(Li) wafers

- the **General AntiParticle Spectrometer** is designed for low-energy antinuclei
- GAPS is under construction → first Long Duration Balloon flights from Antarctica flight 2021**



Conclusion

- Cosmic-ray antideuteron workshop was in-fact a cosmic-ray antinuclei workshop
→ Antinuclei species need to be explained together
- Antideuteron and antihelium formation are not well understood
- Cross section measurements need to be conducted on for the interpretation
- Thanks a lot to all $\bar{d}19$ workshop participants
→ this review is not complete and a more detailed paper is coming up