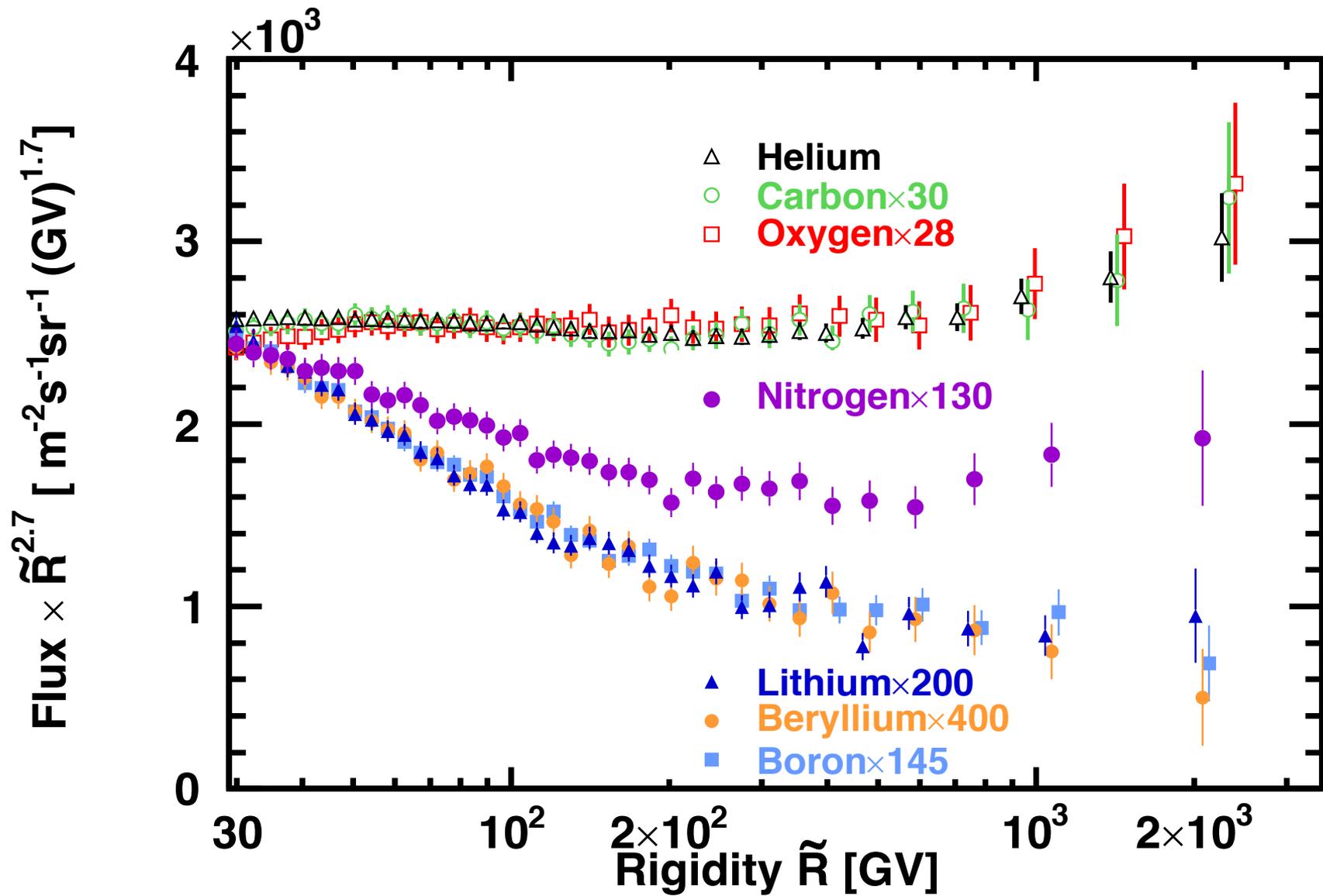


# Measurements of Neon, Magnesium, Silicon and Sulfur Fluxes in Cosmic Rays with AMS



***Q. Yan / MIT on behalf of the AMS collaboration***

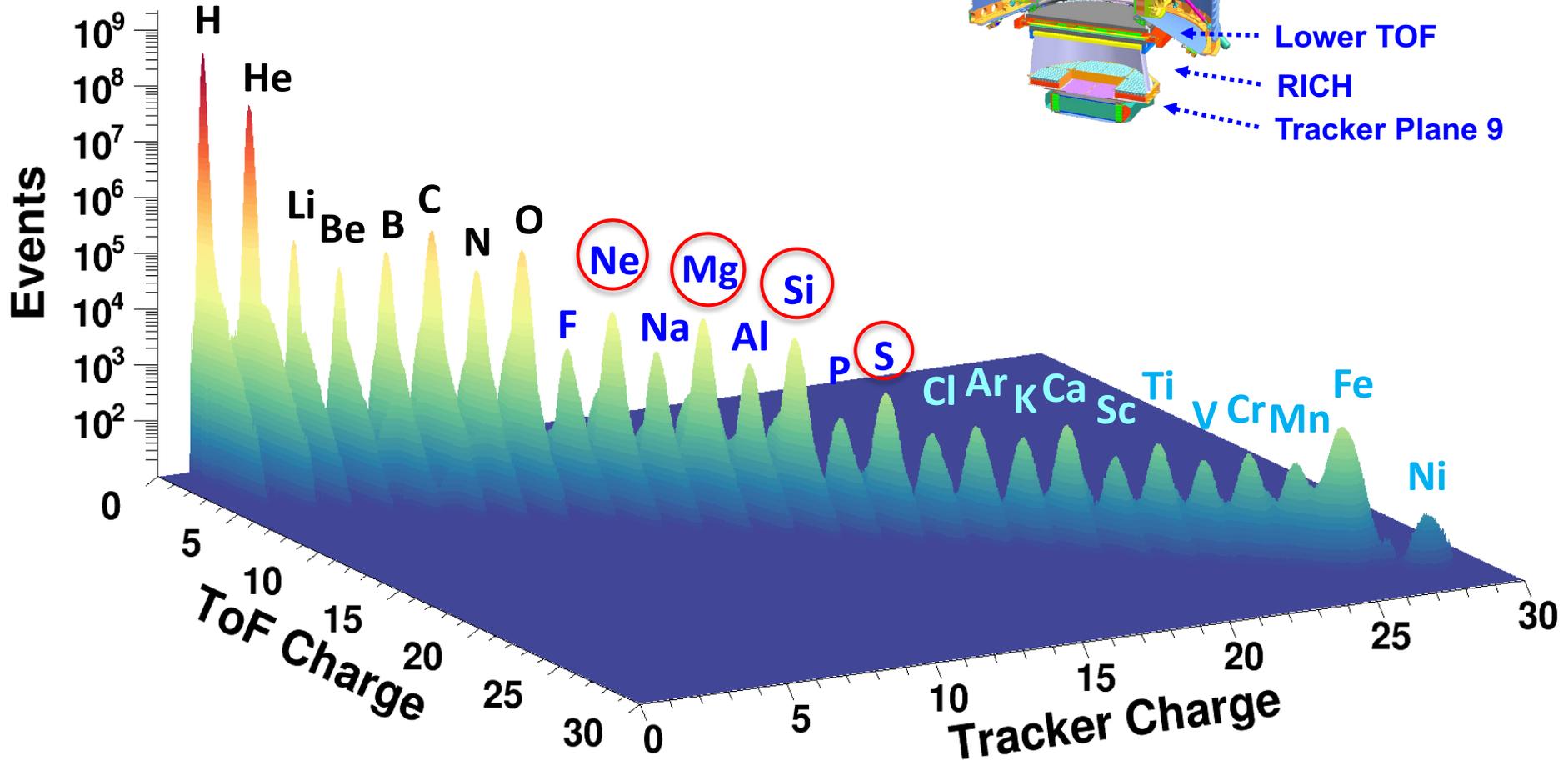
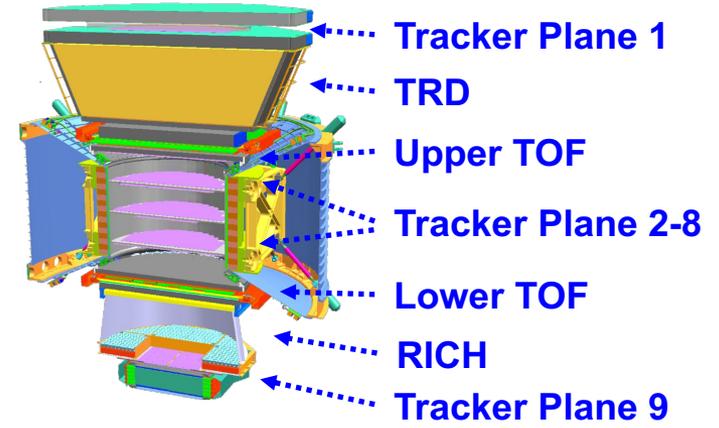
**36<sup>th</sup> International Cosmic Ray Conference, Madison, USA**



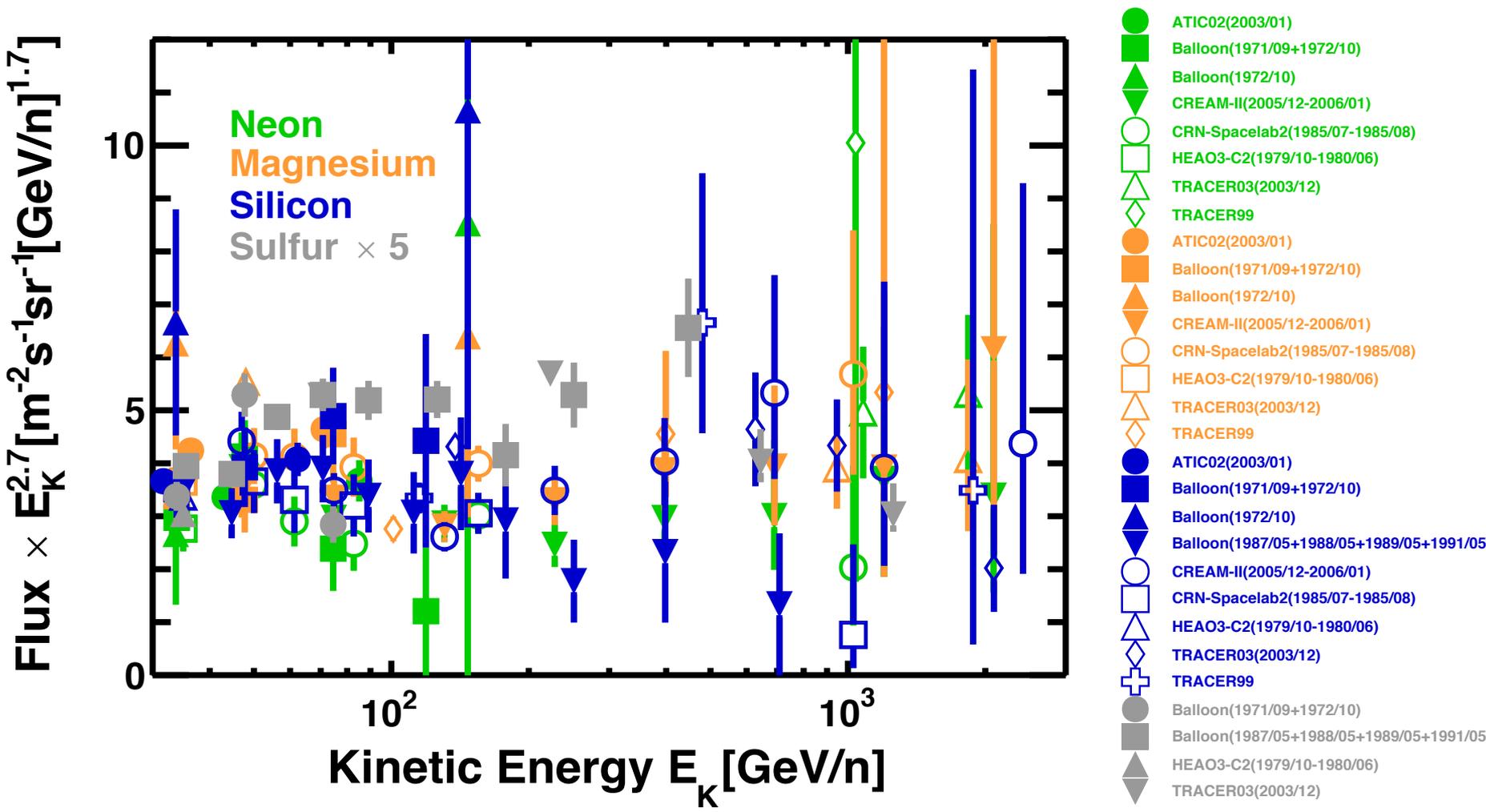
- [1] M. Aguilar *et al.*, Phys. Rev. Lett. **119**, 251101 (2017).  
 [2] M. Aguilar *et al.*, Phys. Rev. Lett. **120**, 021101 (2018).  
 [3] M. Aguilar *et al.*, Phys. Rev. Lett. **121**, 051103 (2018).

# Primary Cosmic Rays Ne (Z=10), Mg (Z=12), Si (Z=14) and S (Z=16)

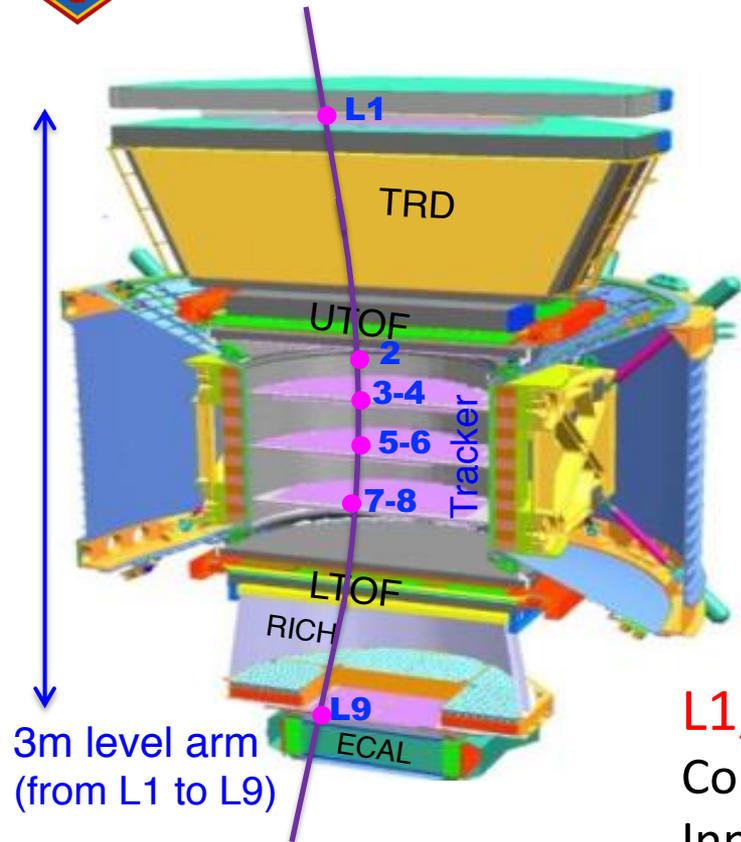
AMS has seven instruments which independently measure Cosmic Nuclei charge



# Measurements of Primary Cosmic Rays Ne, Mg, Si and S Before AMS



# AMS Nuclei Measurements (Ne-S)



## Tracker + Magnet

Rigidity (Momentum/Charge)

Bending Coordinate Resolution  $\approx 6-8 \mu\text{m}$

MDR  $\approx 3 \text{ TV}$

## TOF (4 Layers)

Velocity and Direction

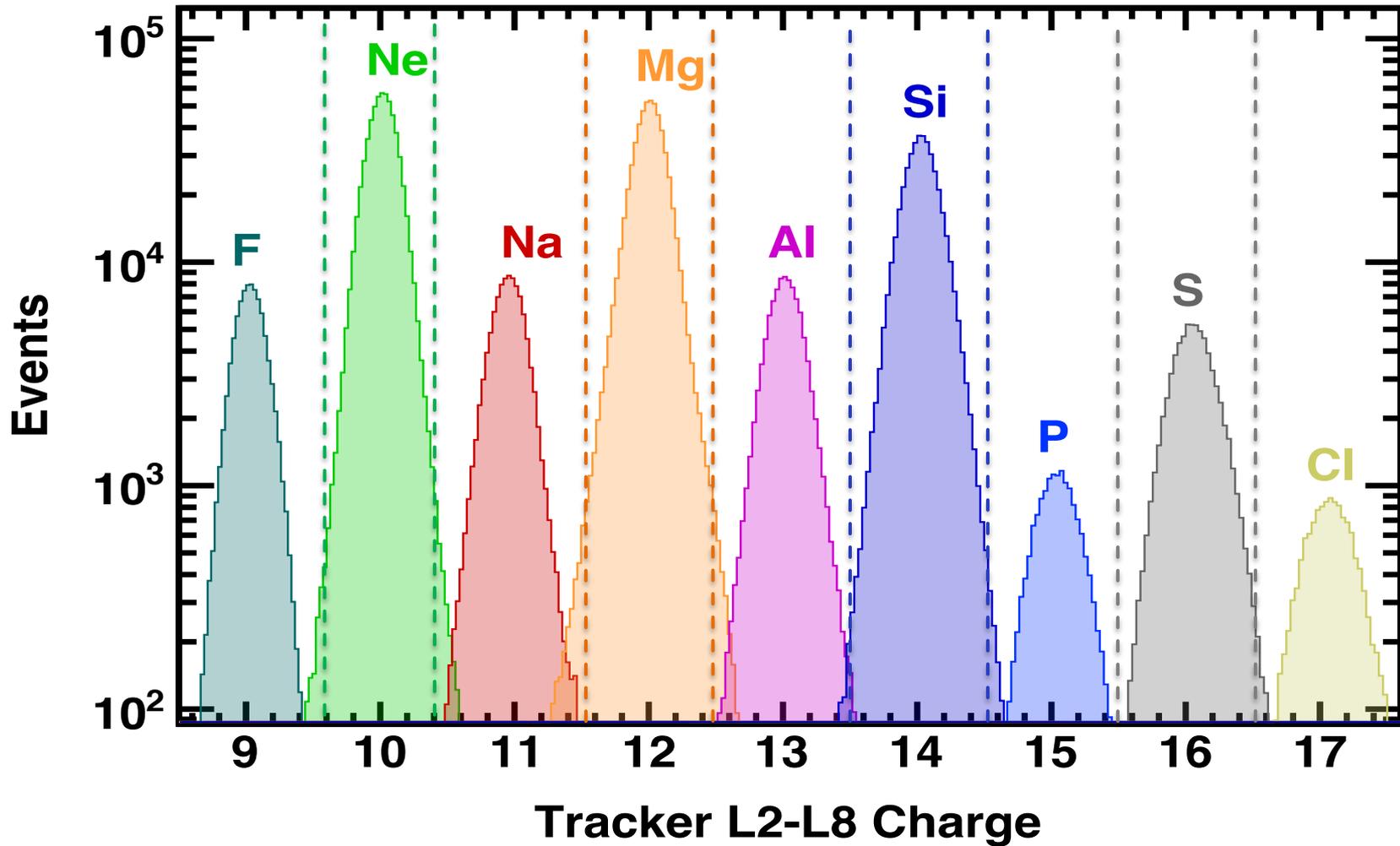
$\Delta\beta/\beta^2 \approx 1\%$

**L1, UTOF, Inner Tracker (L2-L8), LTOF\* and L9\***

Consistent Charge Along Particle Trajectory

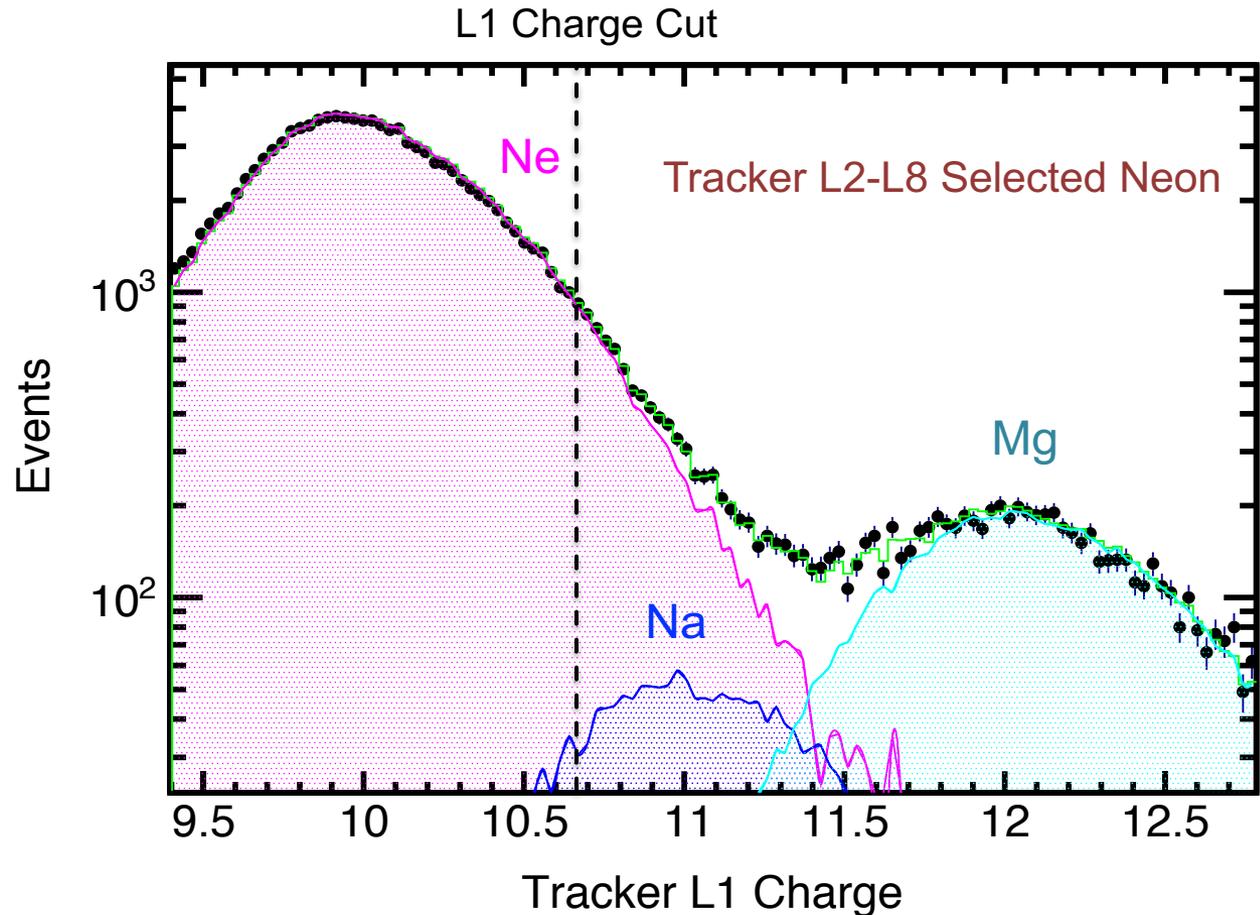
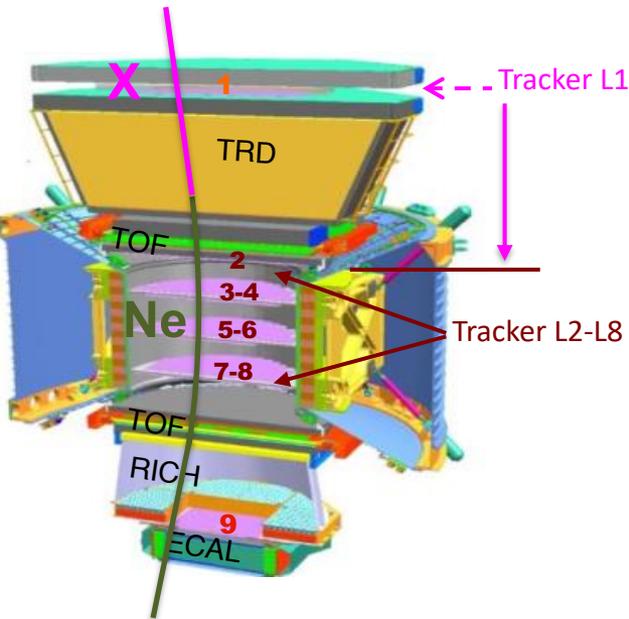
Inner Tracker only Resolution  $\Delta Z \approx 0.13-0.19$

# Ne, Mg, Si and S Events Selection

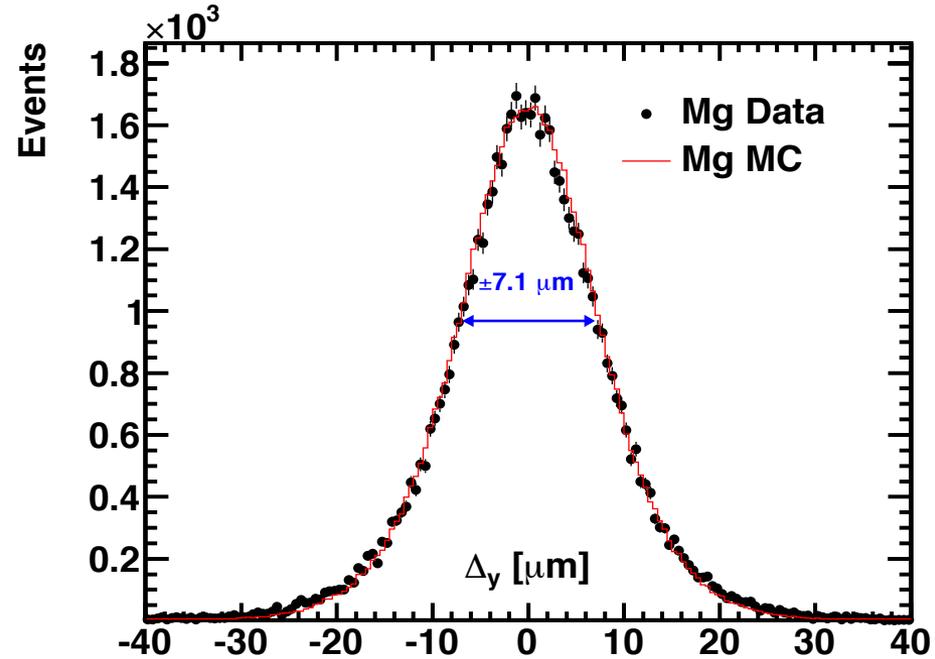
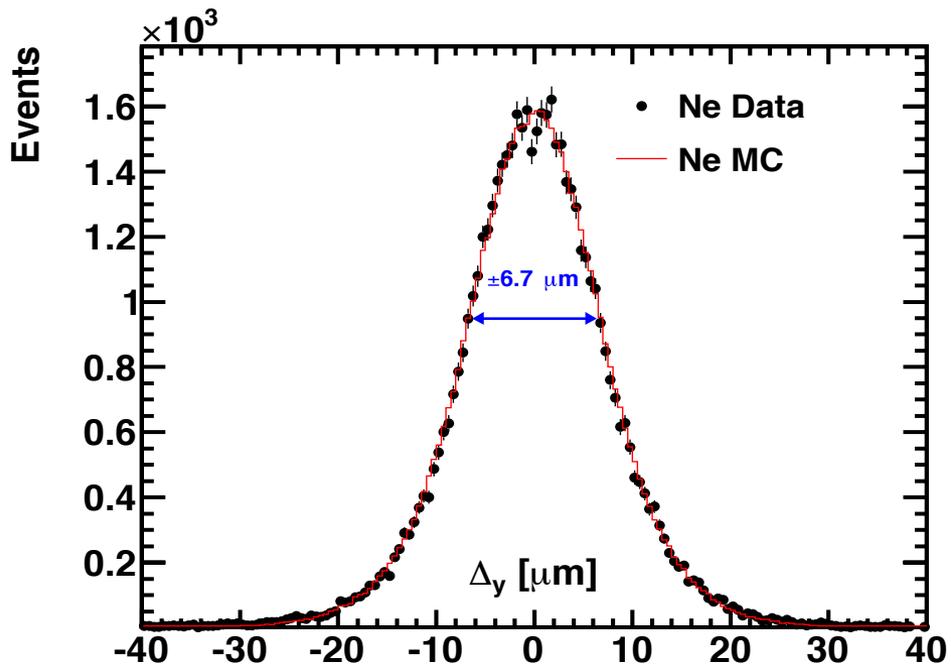
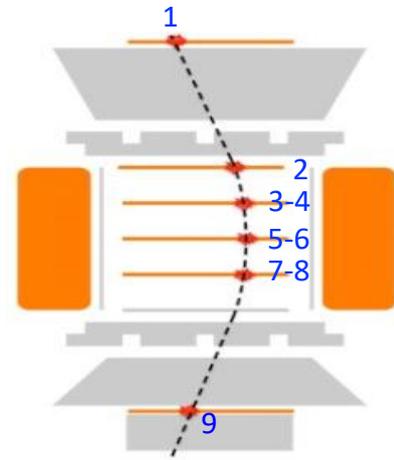


# Background

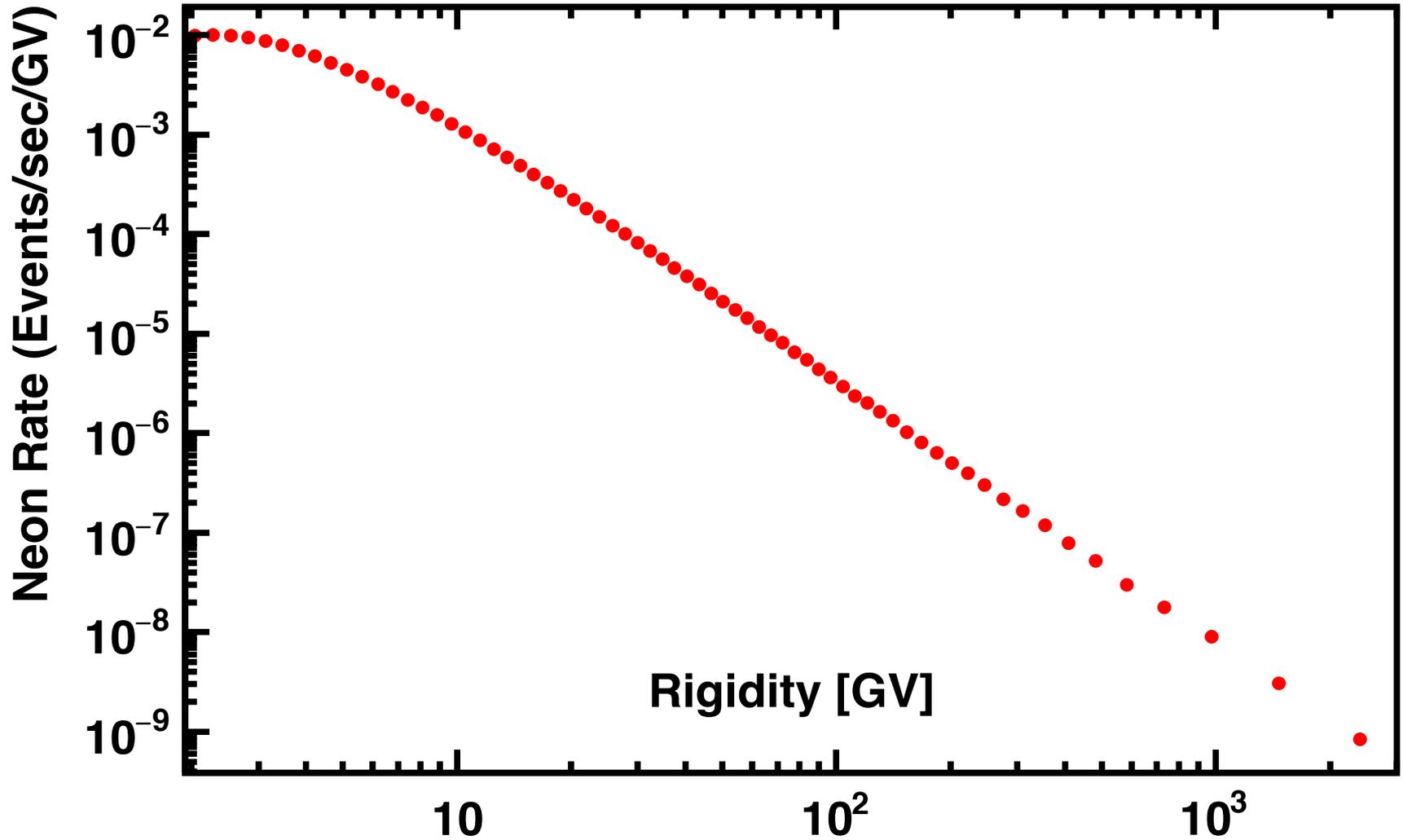
The background resulting from heavier nuclei which interact **between tracker L1 and L2** (for example: Mg,Na ...->Ne) can be subtracted by L1 charge measurement. After the selection, the background is **<0.2%** in the entire rigidity range for Ne, Mg, Si and S.



# Tracker Coordinate Resolution Data and MC



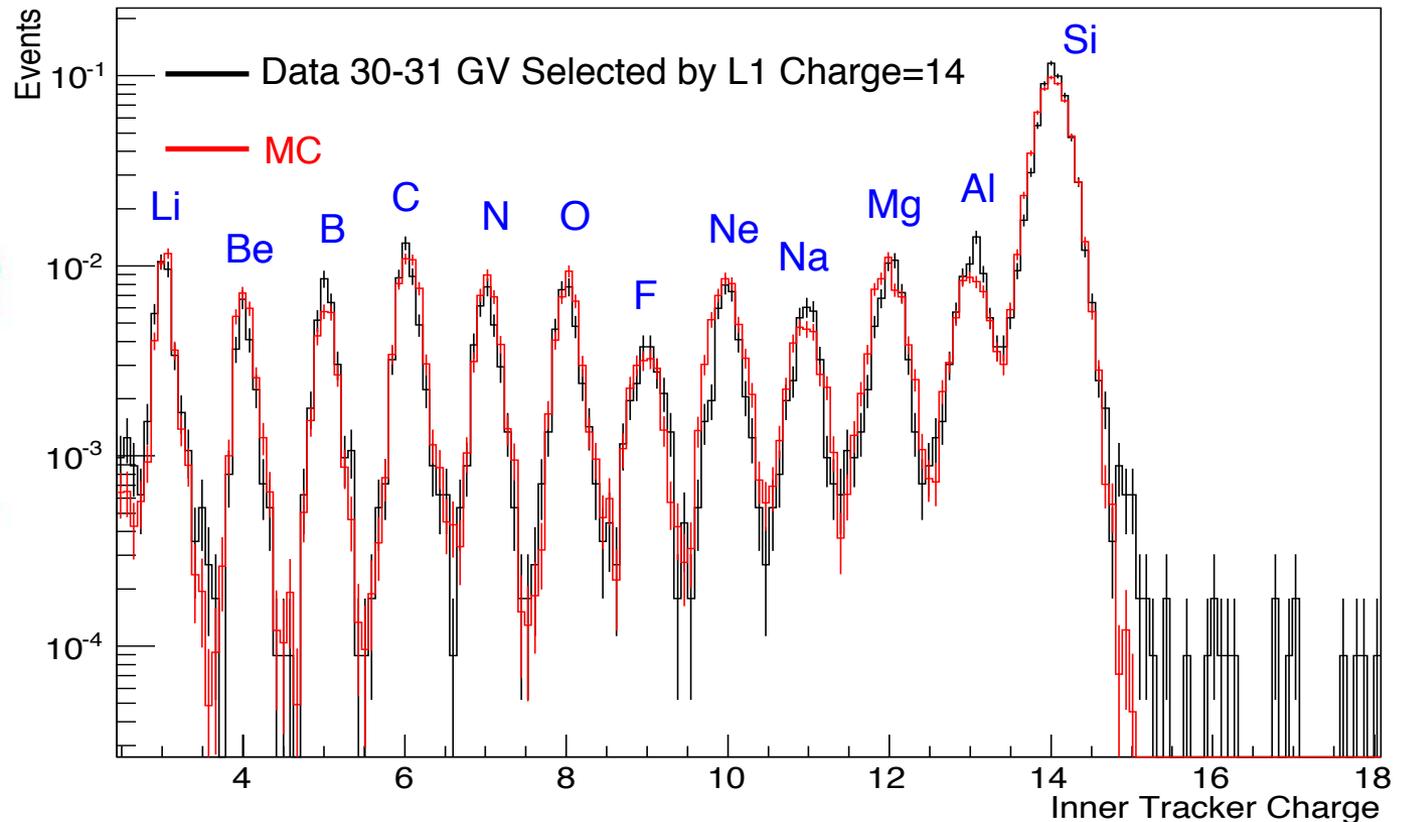
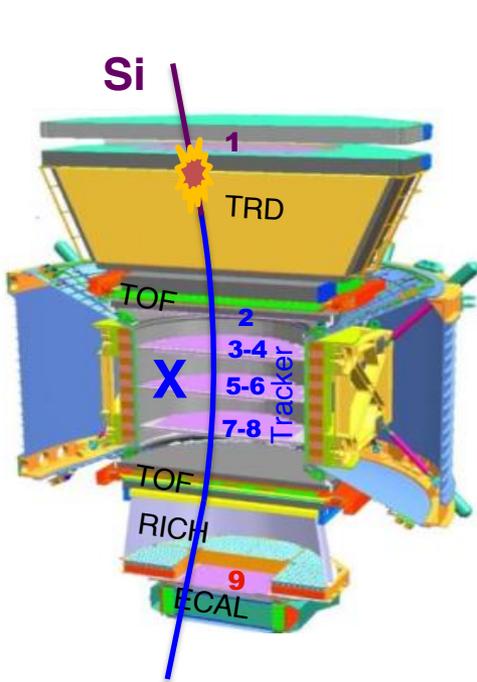
# Neon Rate (Events/T/ $\Delta R$ )



# Nucleus Inelastic Interaction and Breaking-Up Channels Distribution (Silicon)

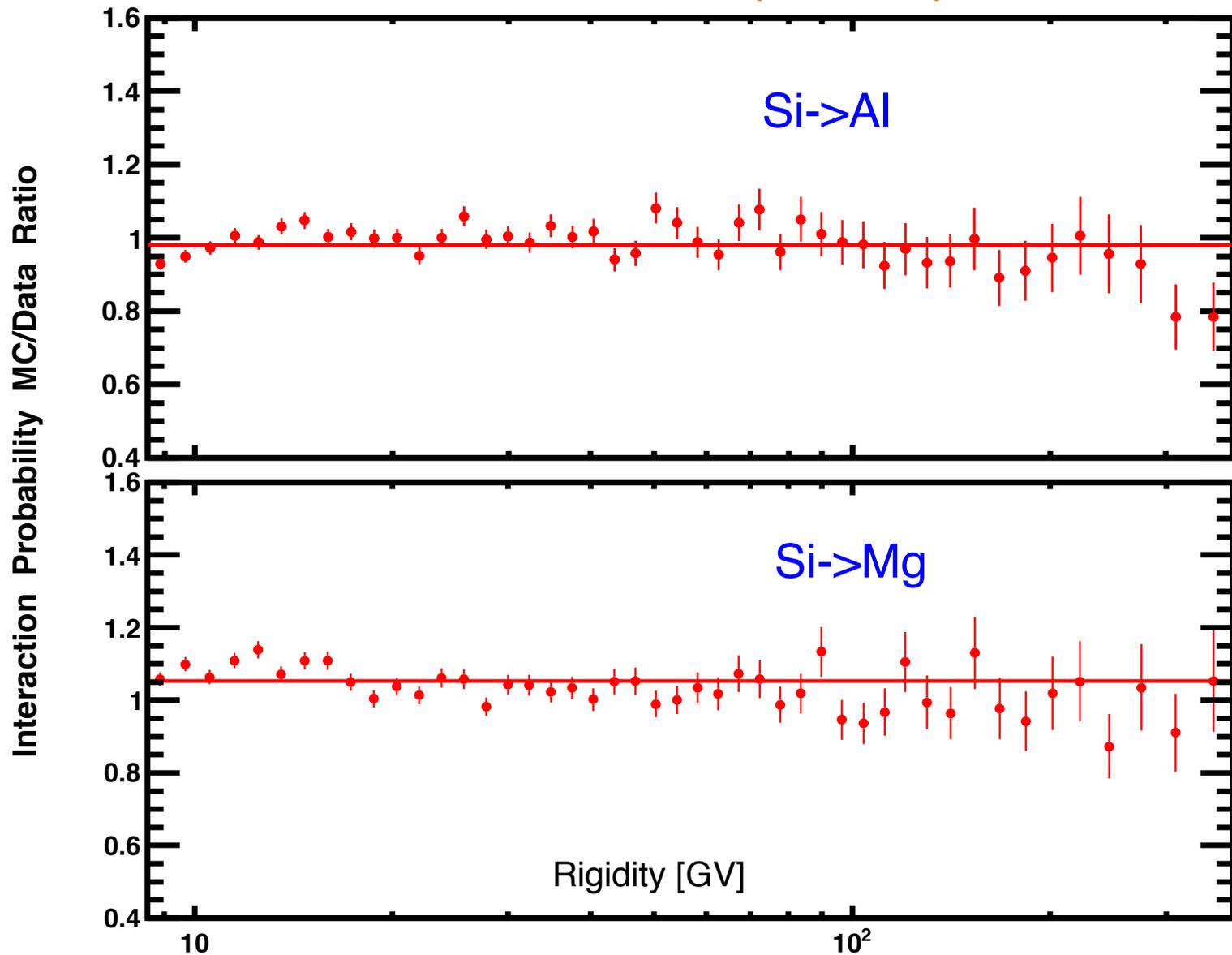
Statistics in “horizontal” runs are not sufficient for high Z survival probability evaluation, so different method was used

- Select primary nuclei by L1 charge
- Obtain survival probability by comparing charge measured with inner tracker

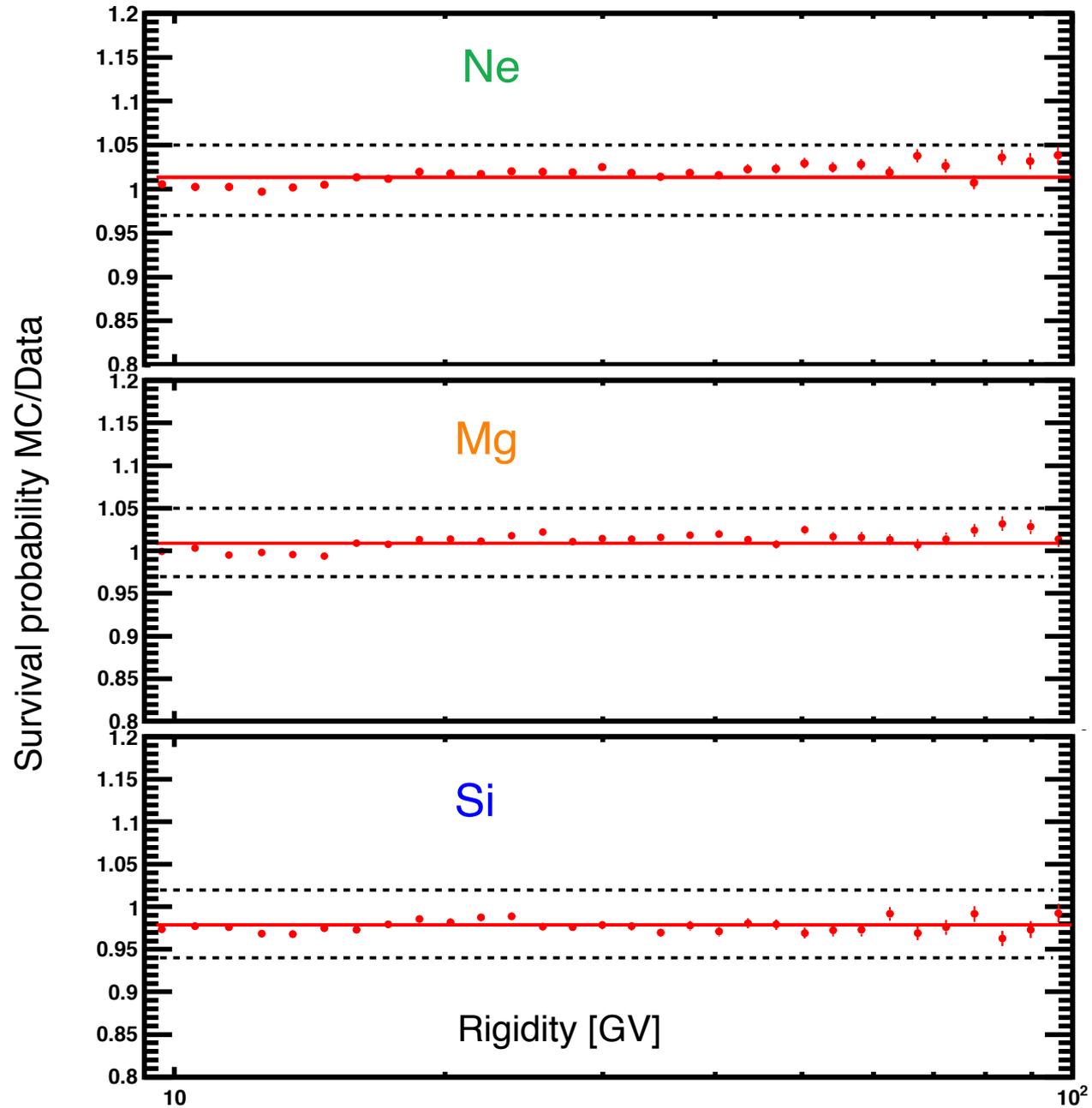
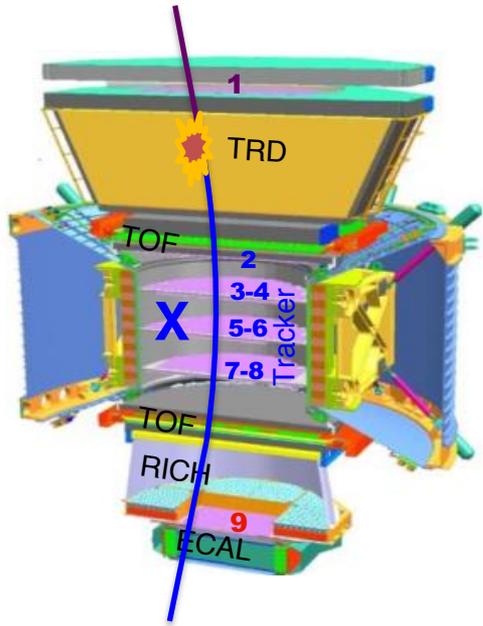


# Nucleus Interaction Breaking-Up Channels

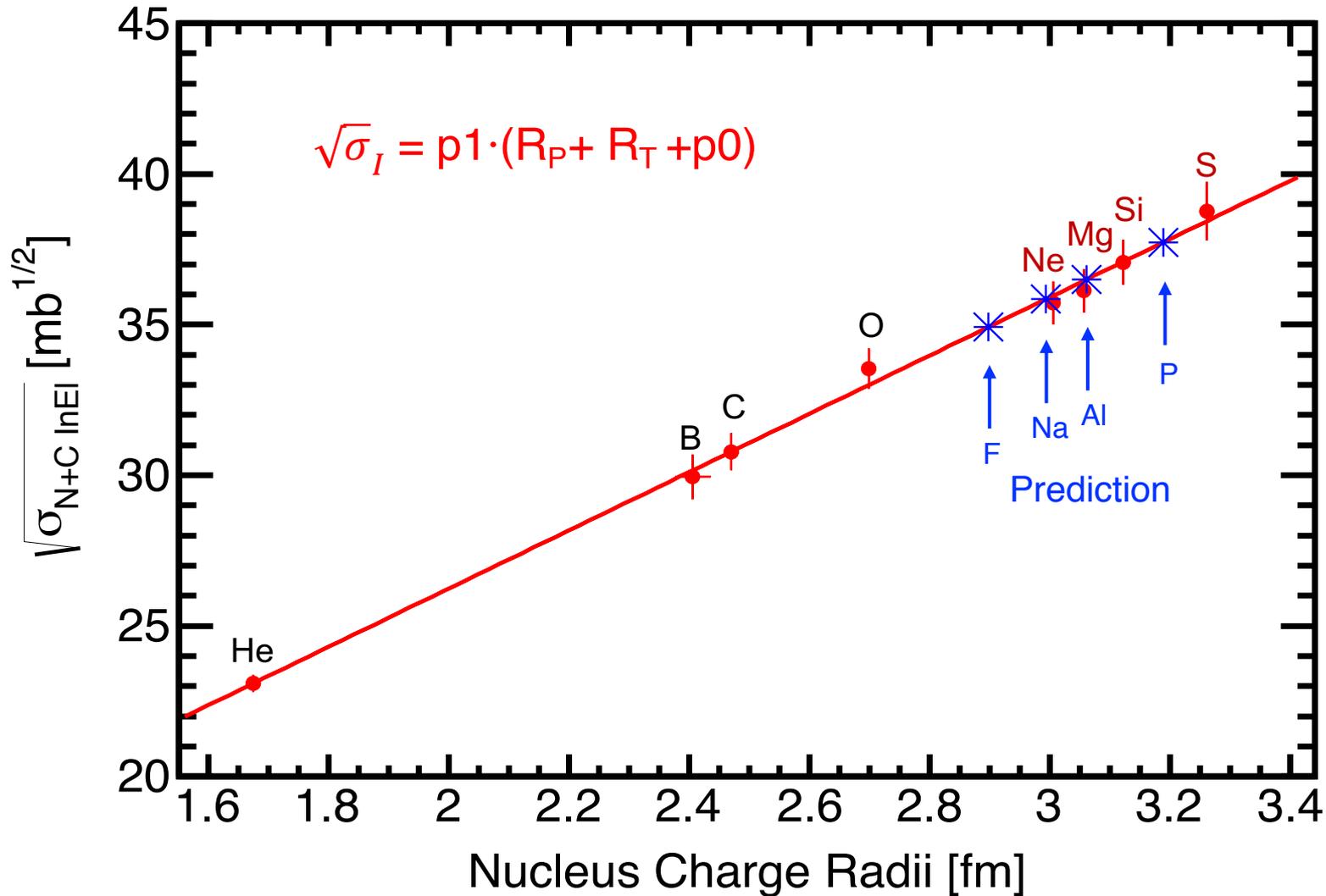
## MC to Data Ratio (Silicon)



# Nuclear Inelastic Interaction Survival Probability MC/Data

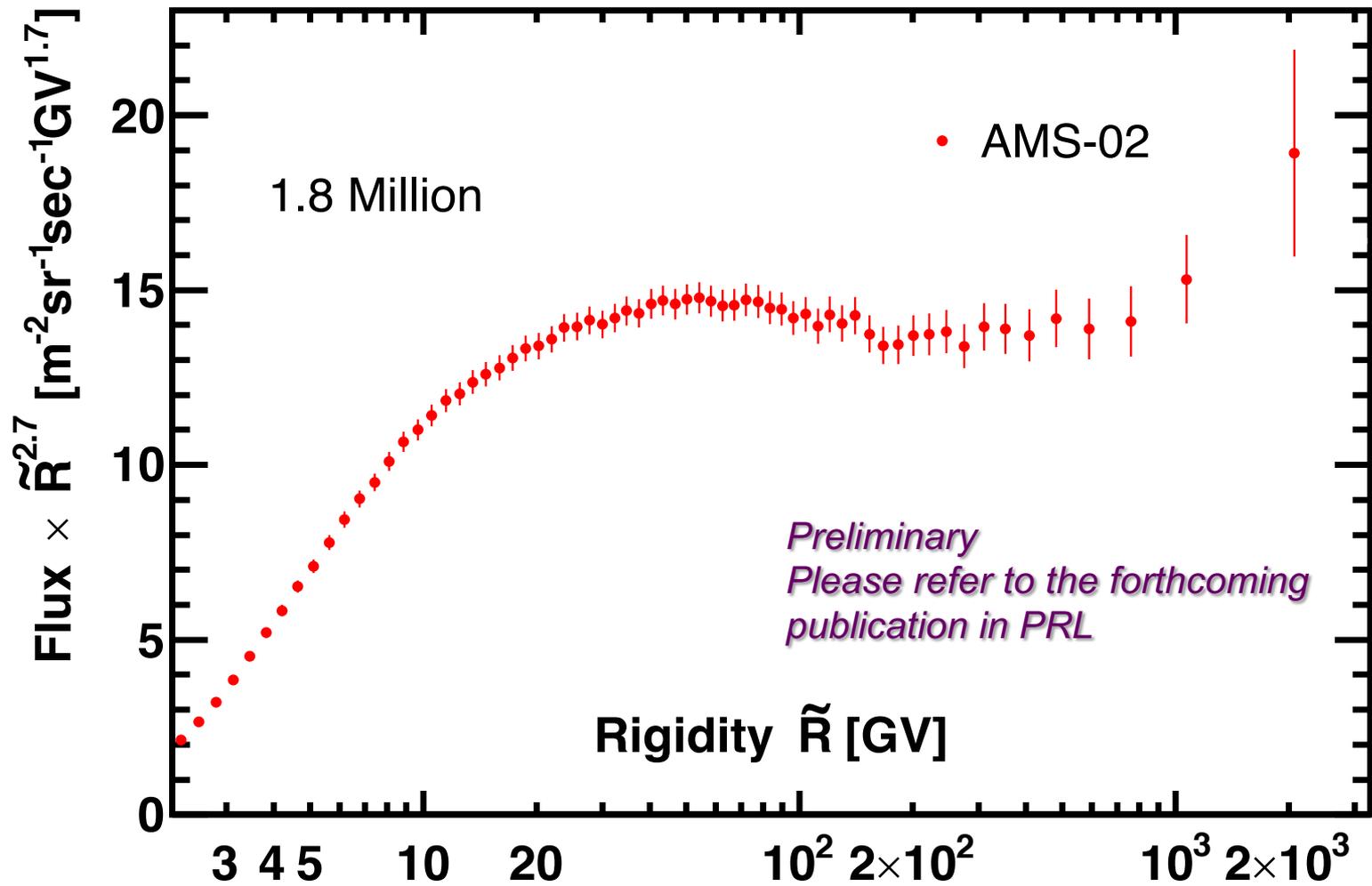


# AMS Nucleus + C Inelastic Cross Section Measurements (5-100 GV Rigidity)

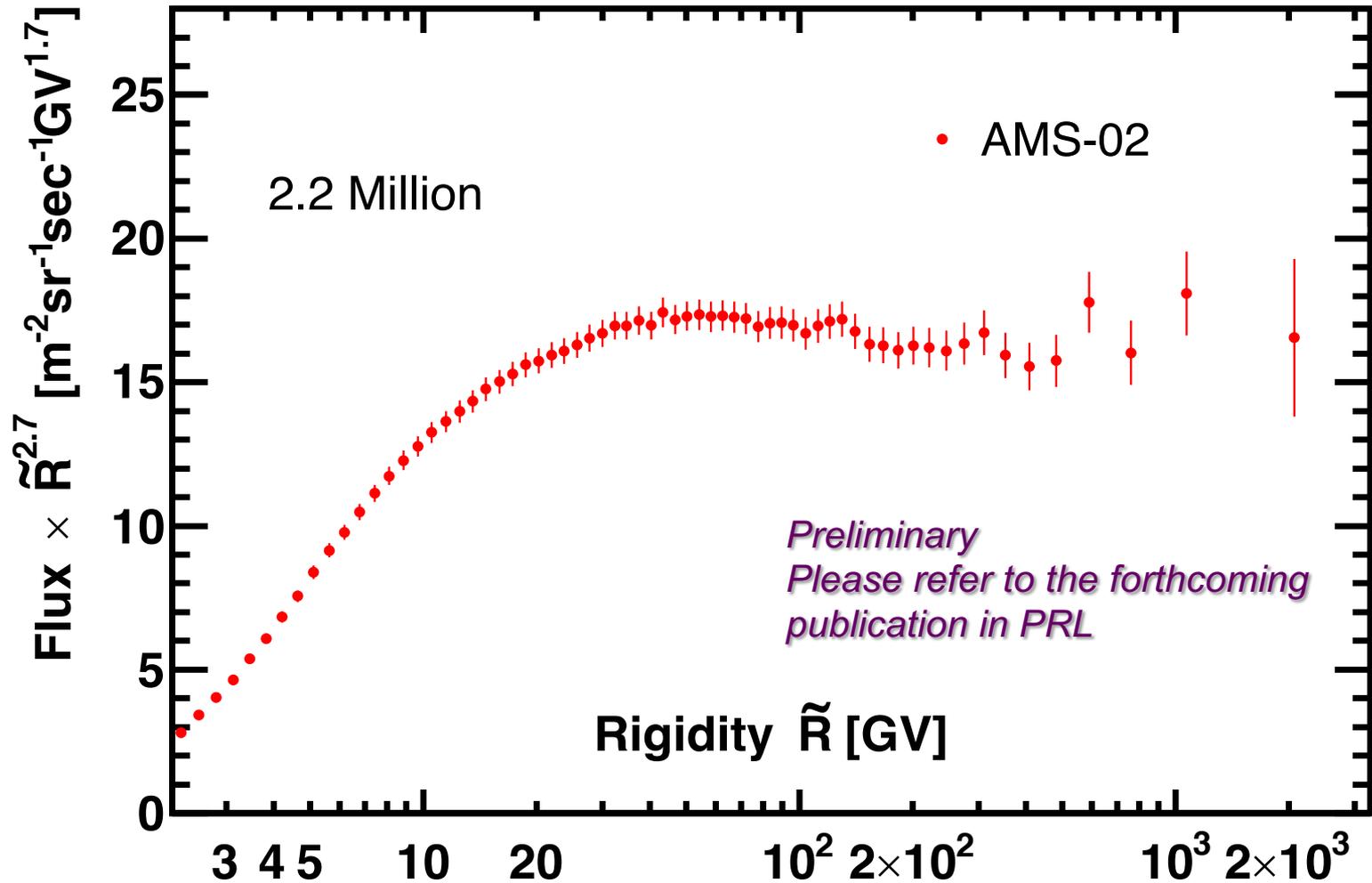


I. Angeli and K. P. Marinova. "Table of experimental nuclear ground state charge radii: An update." *Atomic Data and Nuclear Data Tables* 99.1 (2013): 69-95

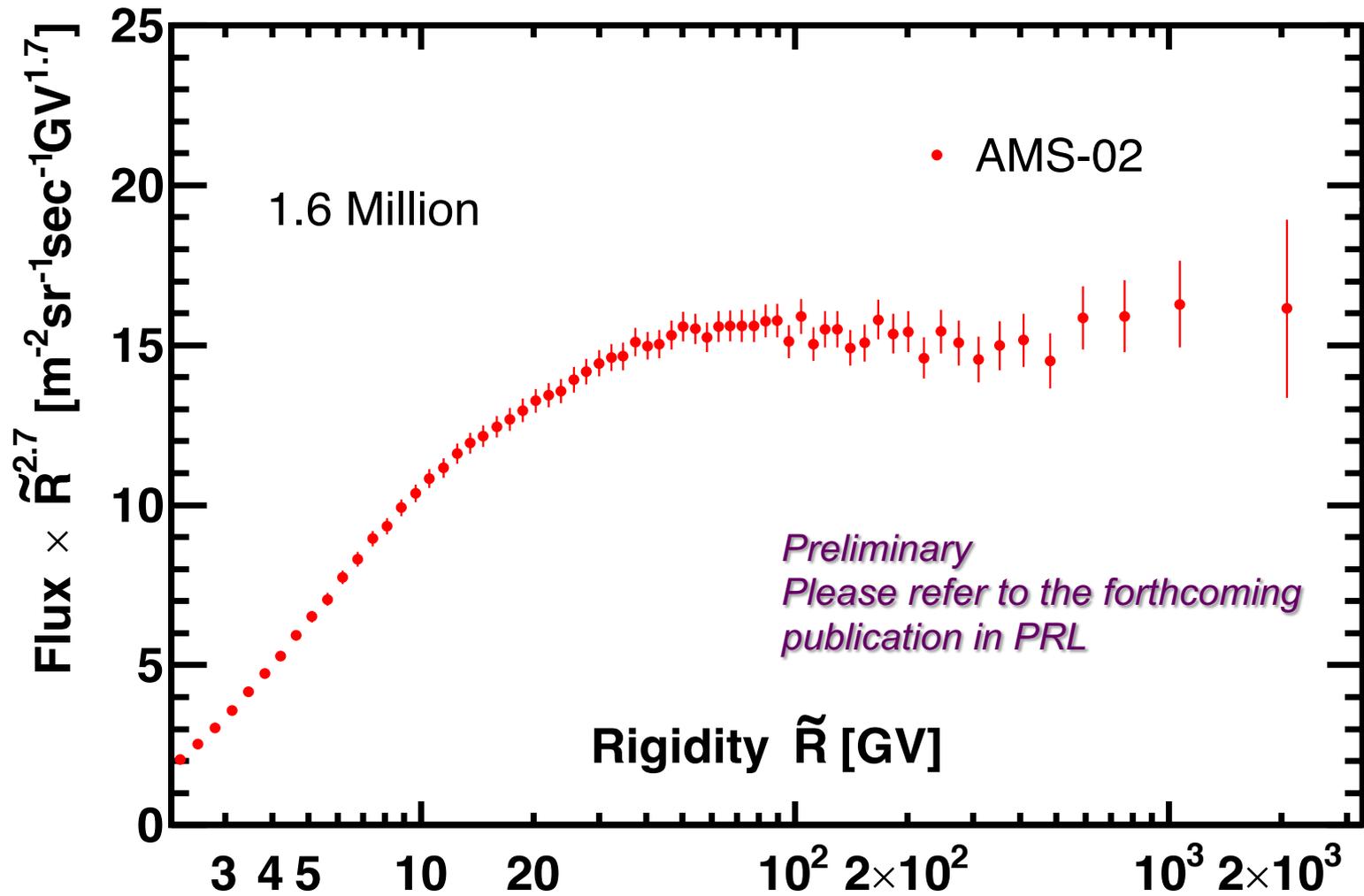
# Neon Flux



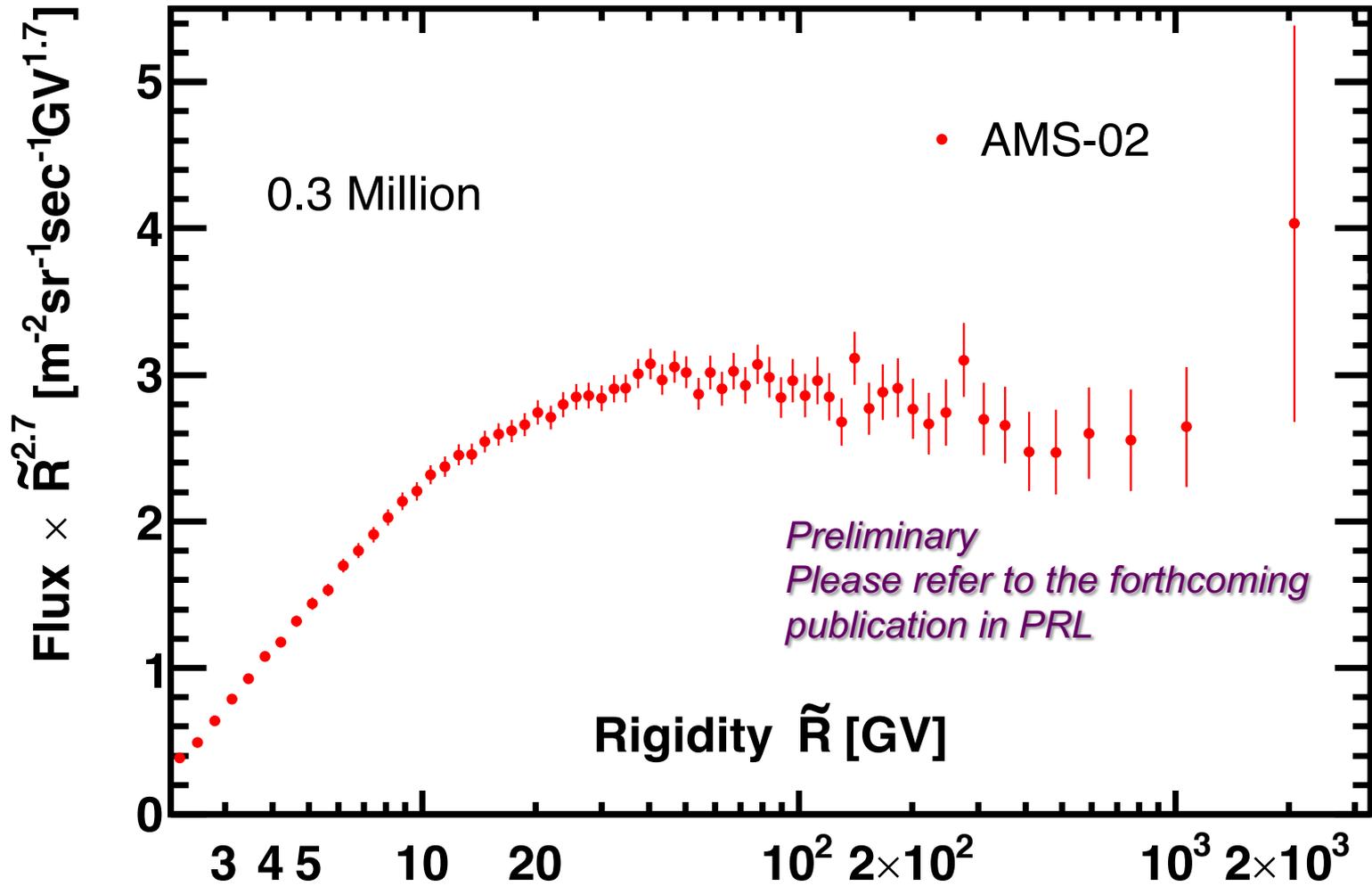
# Magnesium Flux



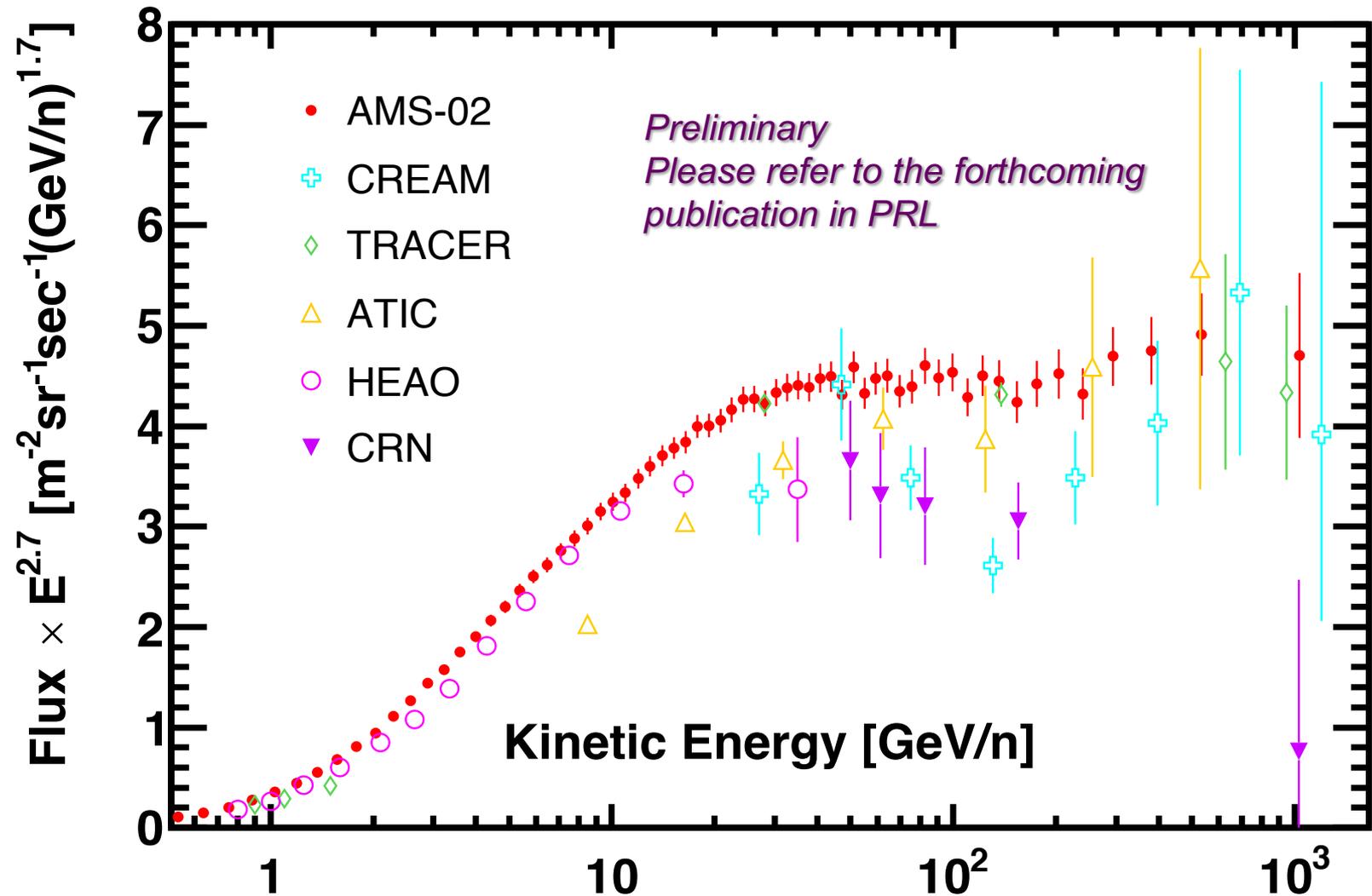
# Silicon Flux



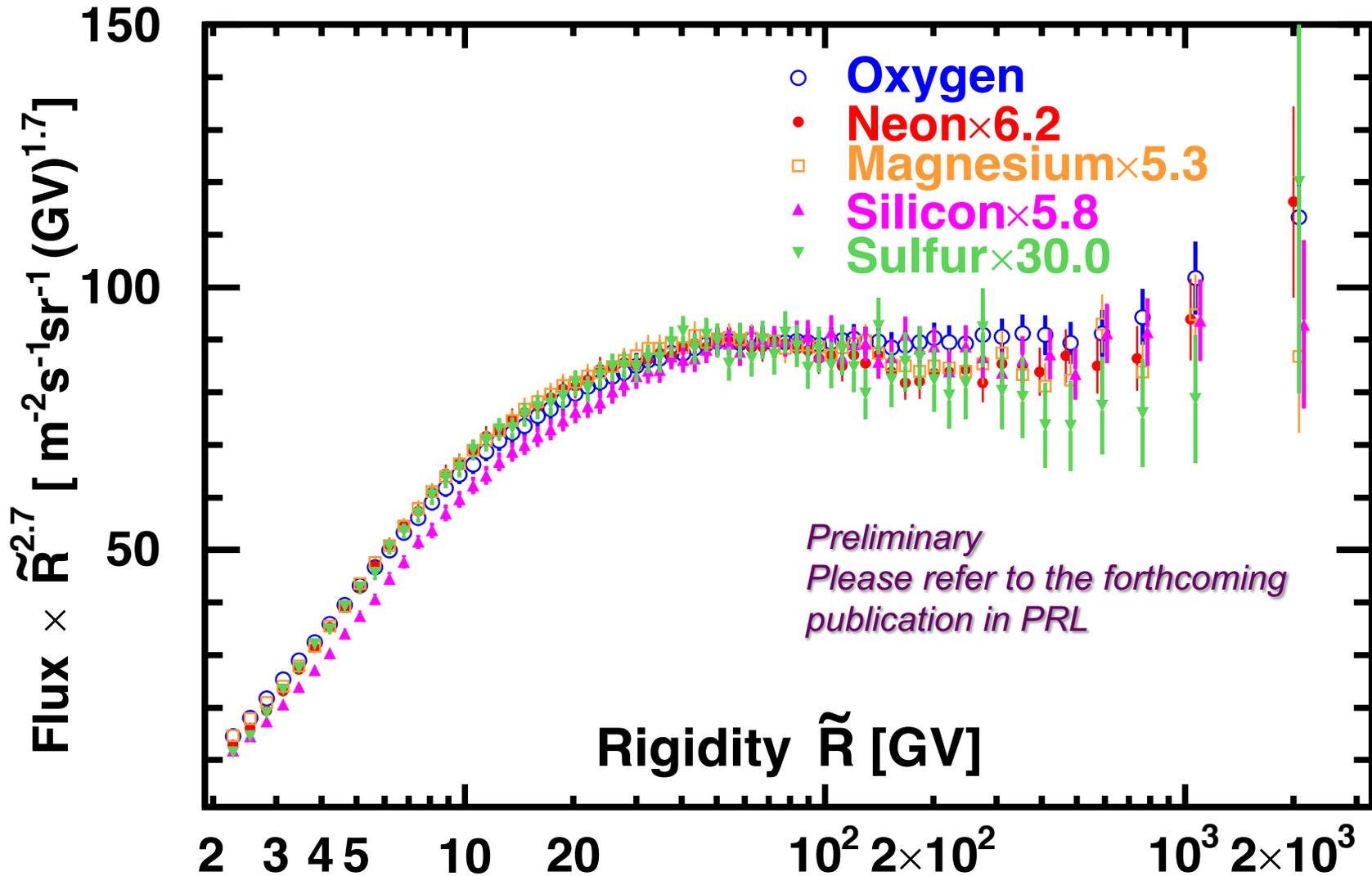
# Sulfur Flux



# Compared with Other Experiments (Silicon Flux)

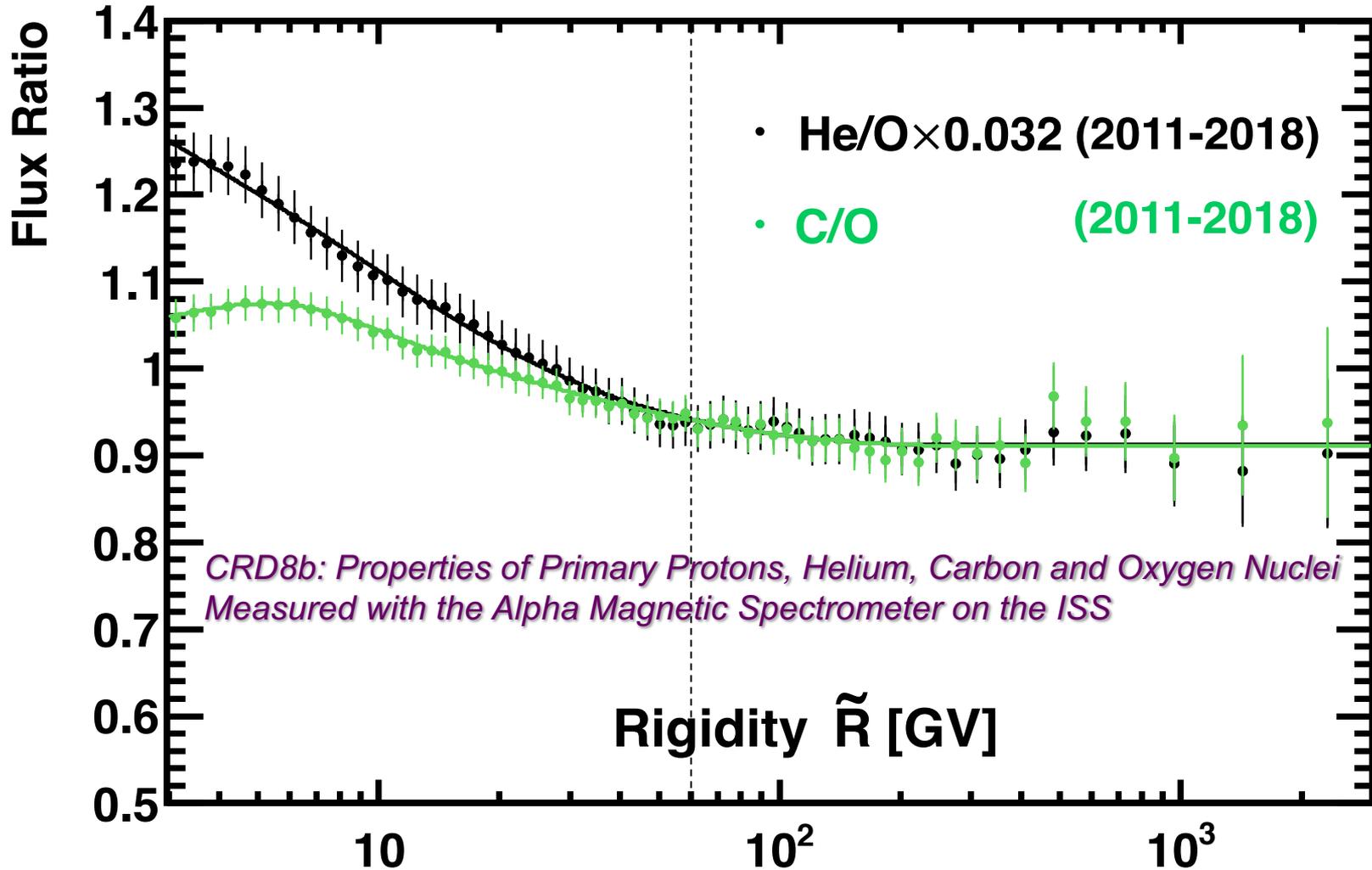


# O, Ne, Mg, Si and S Fluxes



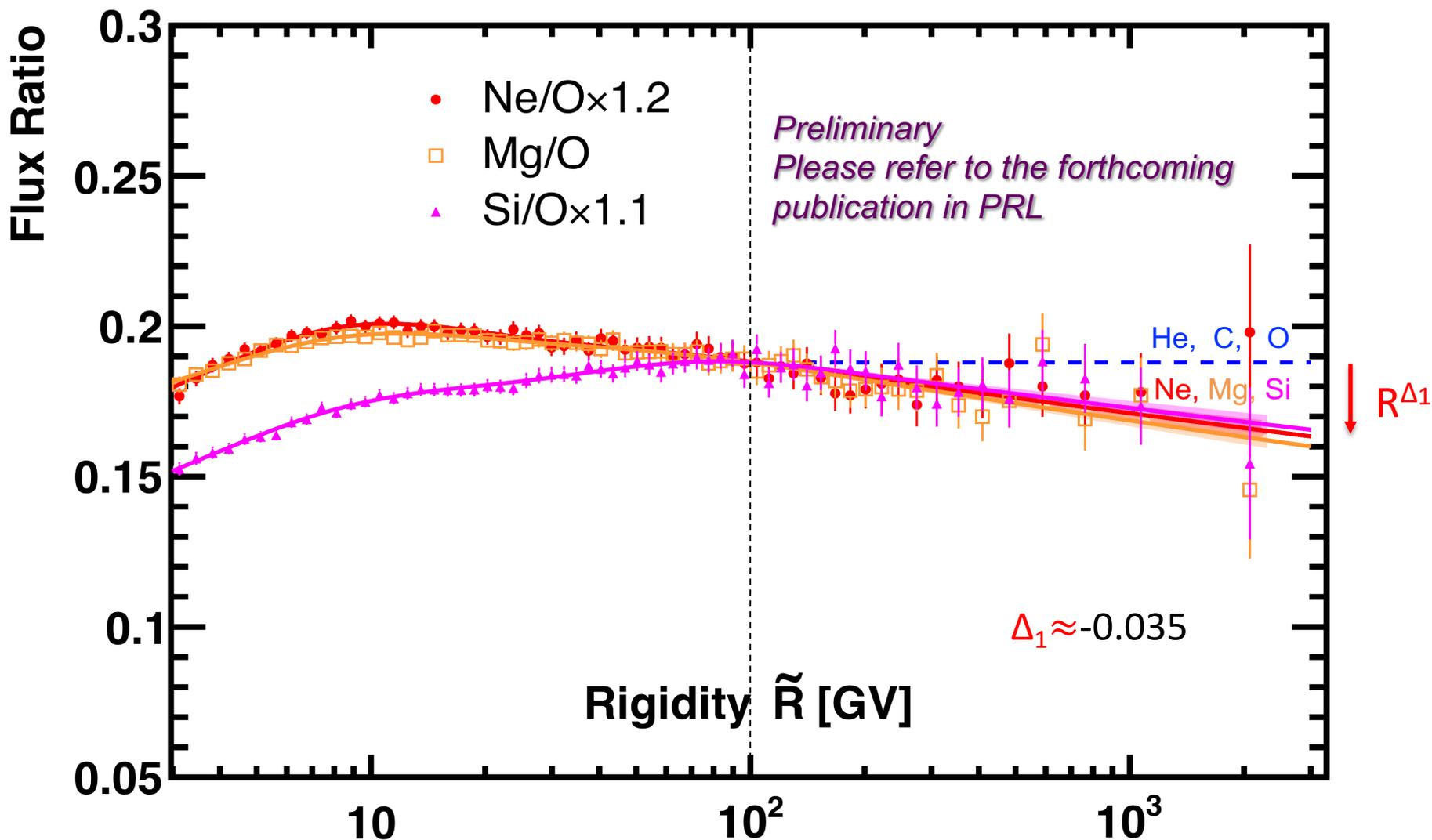
# Light Elements Flux Ratio to Oxygen (He/O, C/O)

Above  $\sim 60$  GV, the He, C and O spectra have identical rigidity dependence



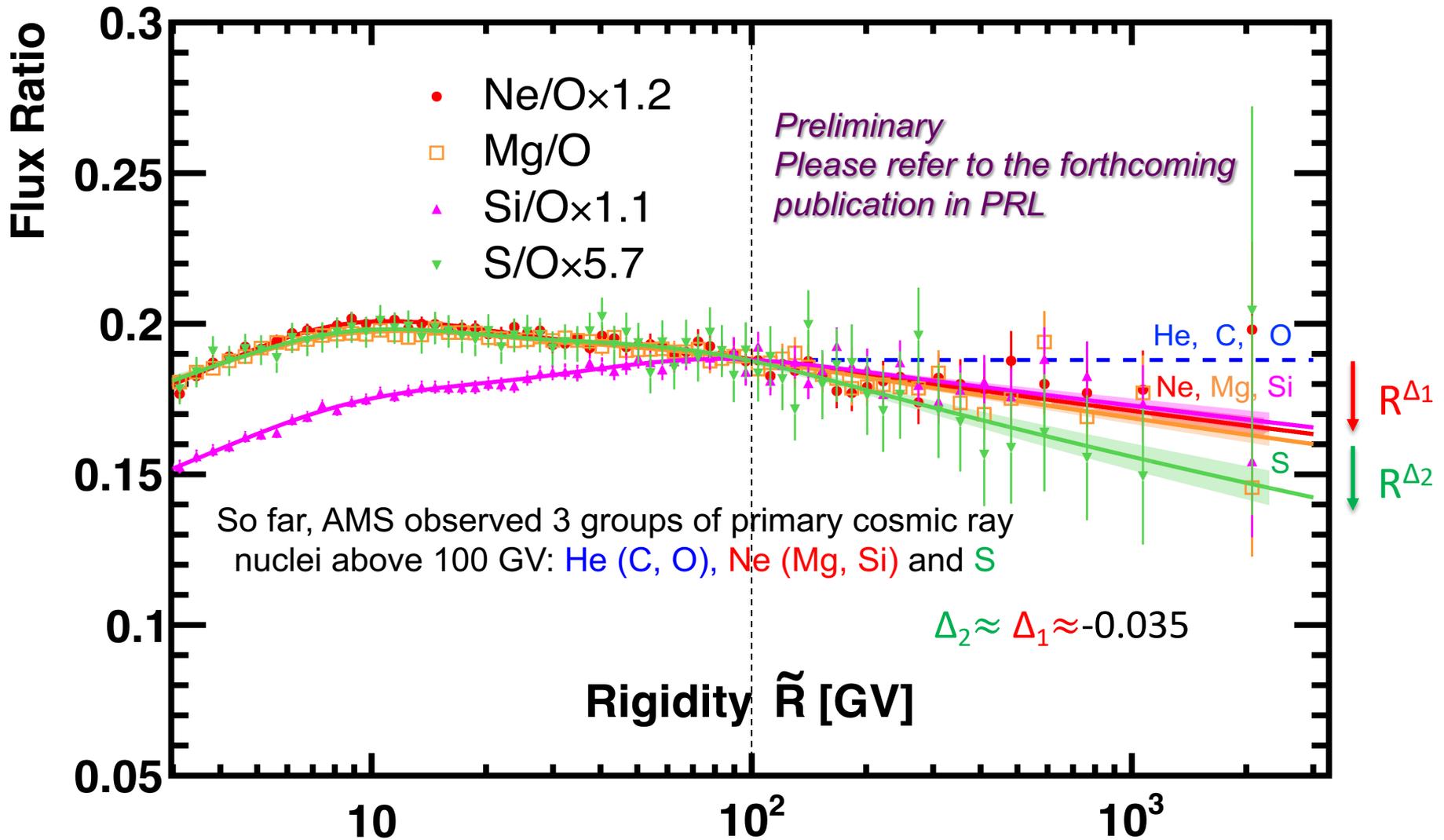
The new AMS result (2011-2018) is consistent with earlier AMS PRL result (2011-2016) “M. Aguilar *et al.*, Phys. Rev. Lett., **119**, 251101 (2017)” but with improved accuracy

# Heavy Elements Flux Ratio to Oxygen (Ne/O, Mg/O, Si/O)



Above 100 GV, the spectra rigidity behavior of Ne, Mg, and Si is identical, but it differ from O (He, C) by a single power law ( $R^{\Delta_1}$ ) with spectral index  $\Delta_1 \approx -0.035$

# Flux Ratio to Oxygen (Ne/O, Mg/O, Si/O and S/O)



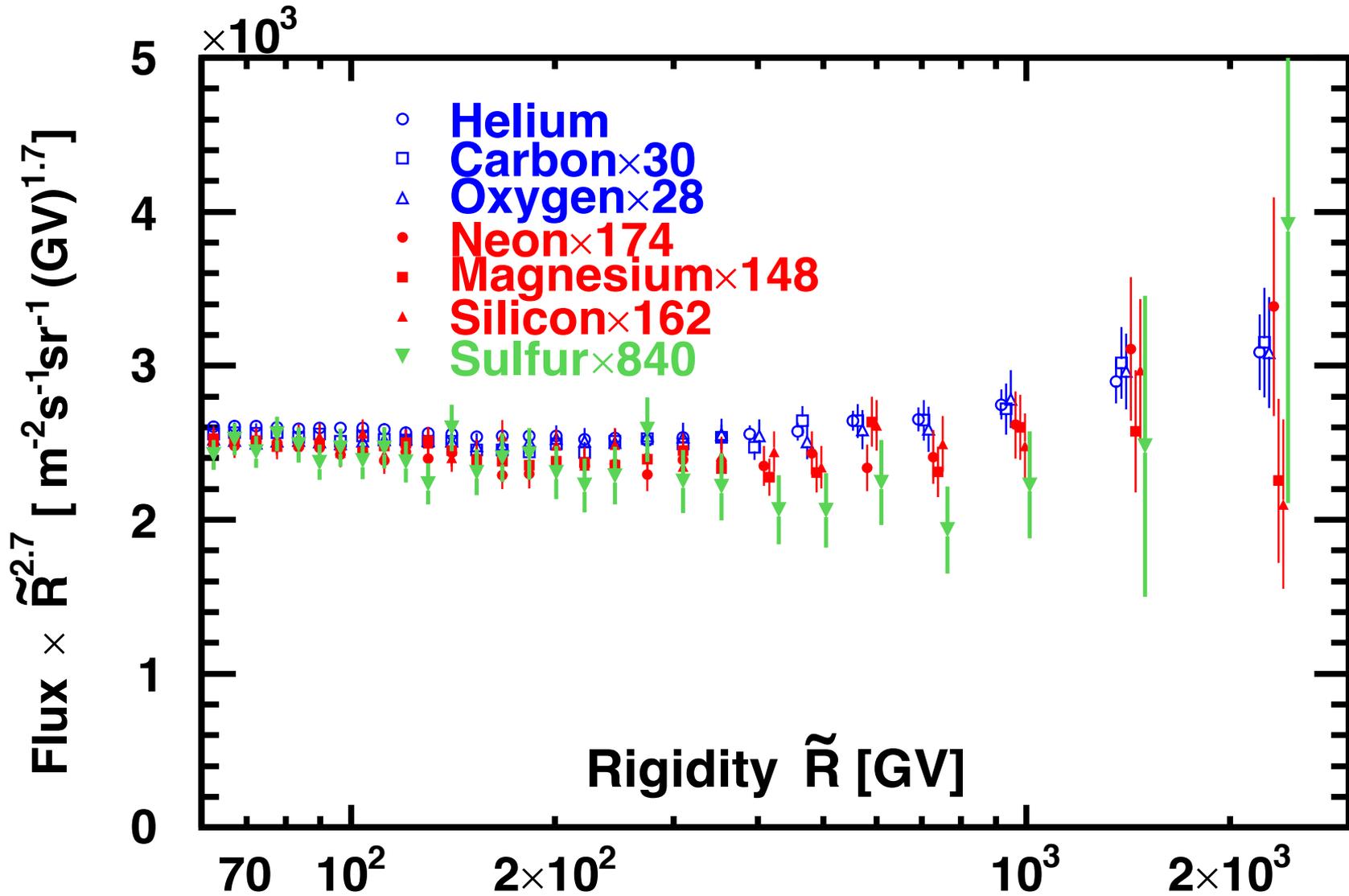
Above 100 GV, the spectrum rigidity behavior of S is different from Ne (Mg, Si) by another single power law ( $R^{\Delta_2}$ ) with spectral index  $\Delta_2 \approx \Delta_1 \approx -0.035$

# Summary

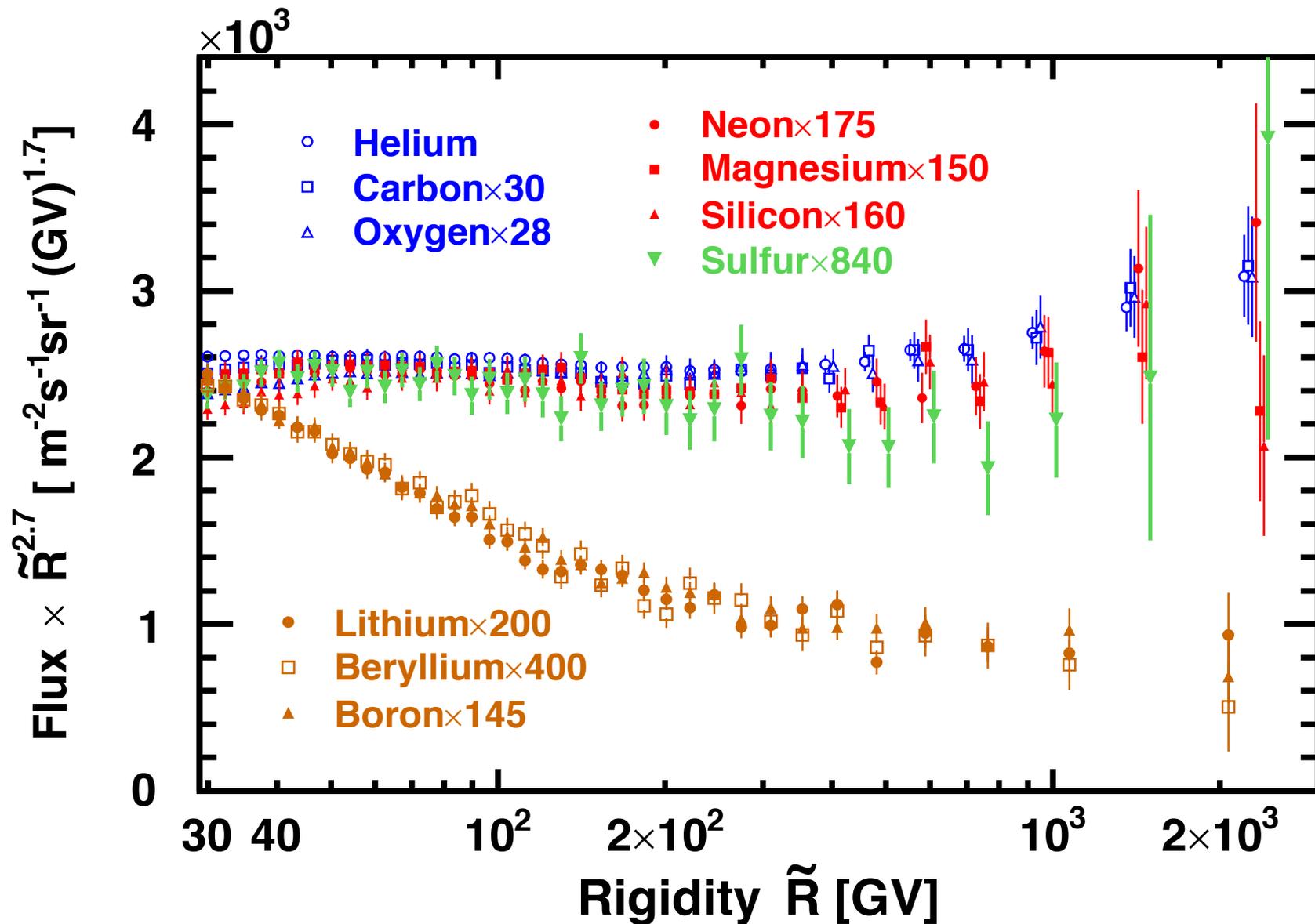
- Precision measurements of Primary Cosmic Rays Neon, Magnesium, Silicon and Sulfur ( $Z=10,12,14,16$ ) fluxes from 2 GV to 3 TV based on 7 years (2011-2018) AMS data have been presented.
- From Oxygen ( $Z=8$ ) to Neon ( $Z=10$ ), the flux drops by a factor of  $\sim 5$ . The abundance of the elements Neon, Magnesium, and Silicon ( $Z=10,12,14$ ) is in the same order of magnitude ( $<20\%$  difference). Above 100 GV, the flux Rigidity behaviors of Neon, Magnesium, and Silicon are identical, while they are all softened than Oxygen by a single power law ( $R^{\Delta_1}$ ) with a spectral index  $\Delta_1 \approx -0.035$ .
- From Silicon ( $Z=14$ ) to Sulfur ( $Z=16$ ), the flux decreases by another factor of  $\sim 5$ . And the flux of Sulfur is softened than Silicon by another single power law ( $R^{\Delta_2}$ ) with spectral index  $\Delta_2 \approx \Delta_1$  above 100GV.
- The flux of Iron ( $Z=26$ ) ?



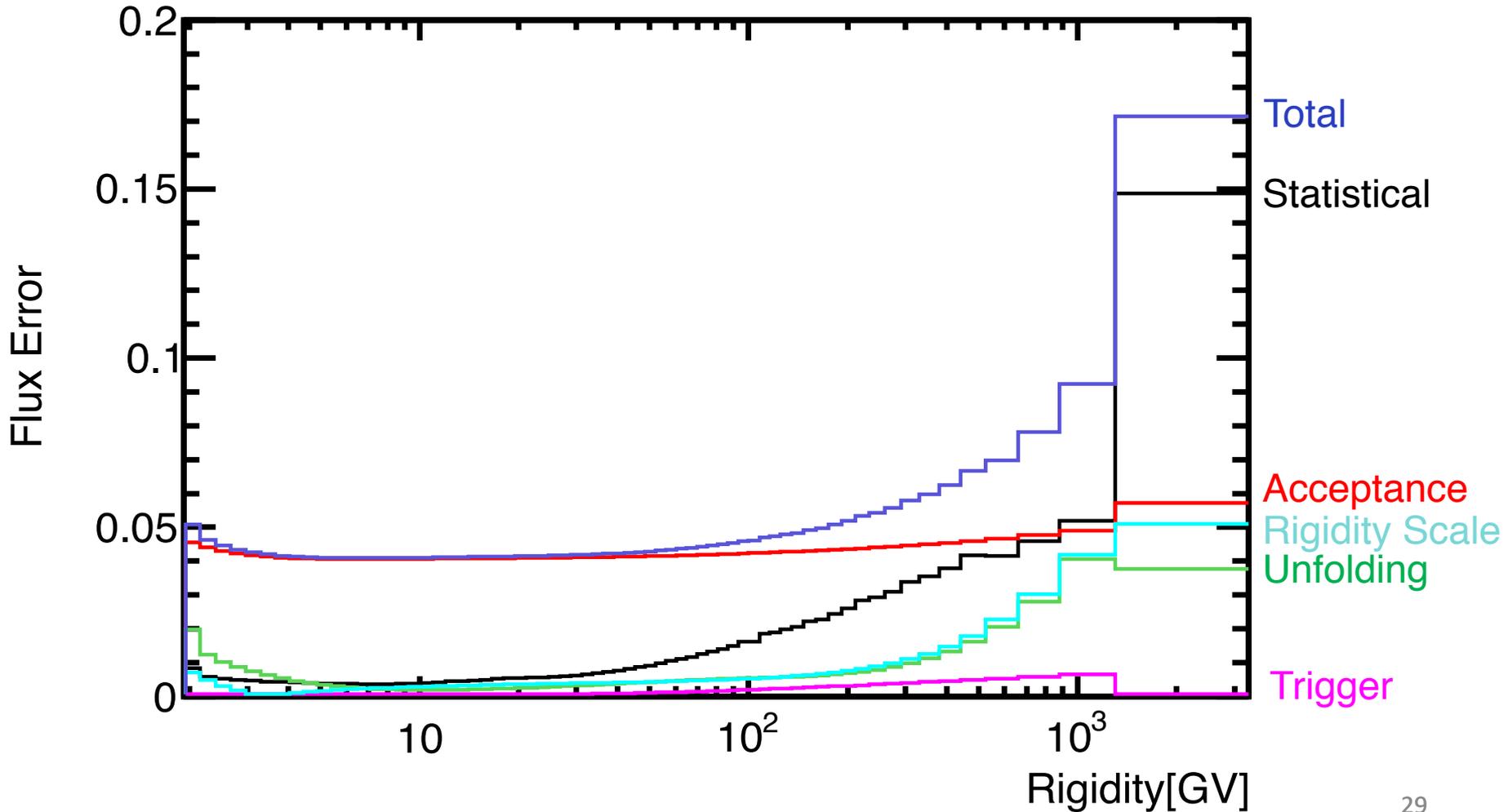
# He, C, O, Ne, Mg, Si and S Fluxes



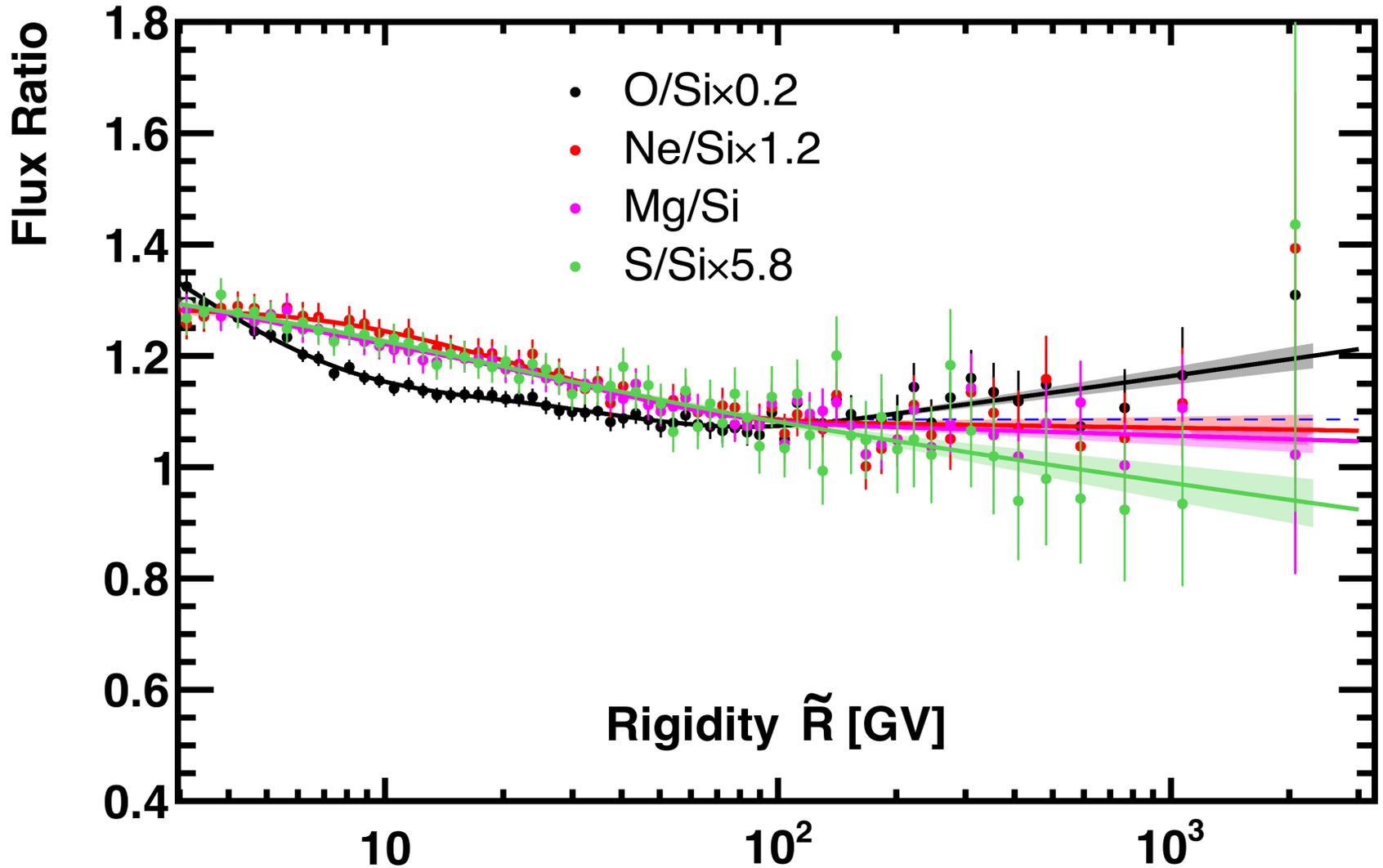
# He-S Fluxes



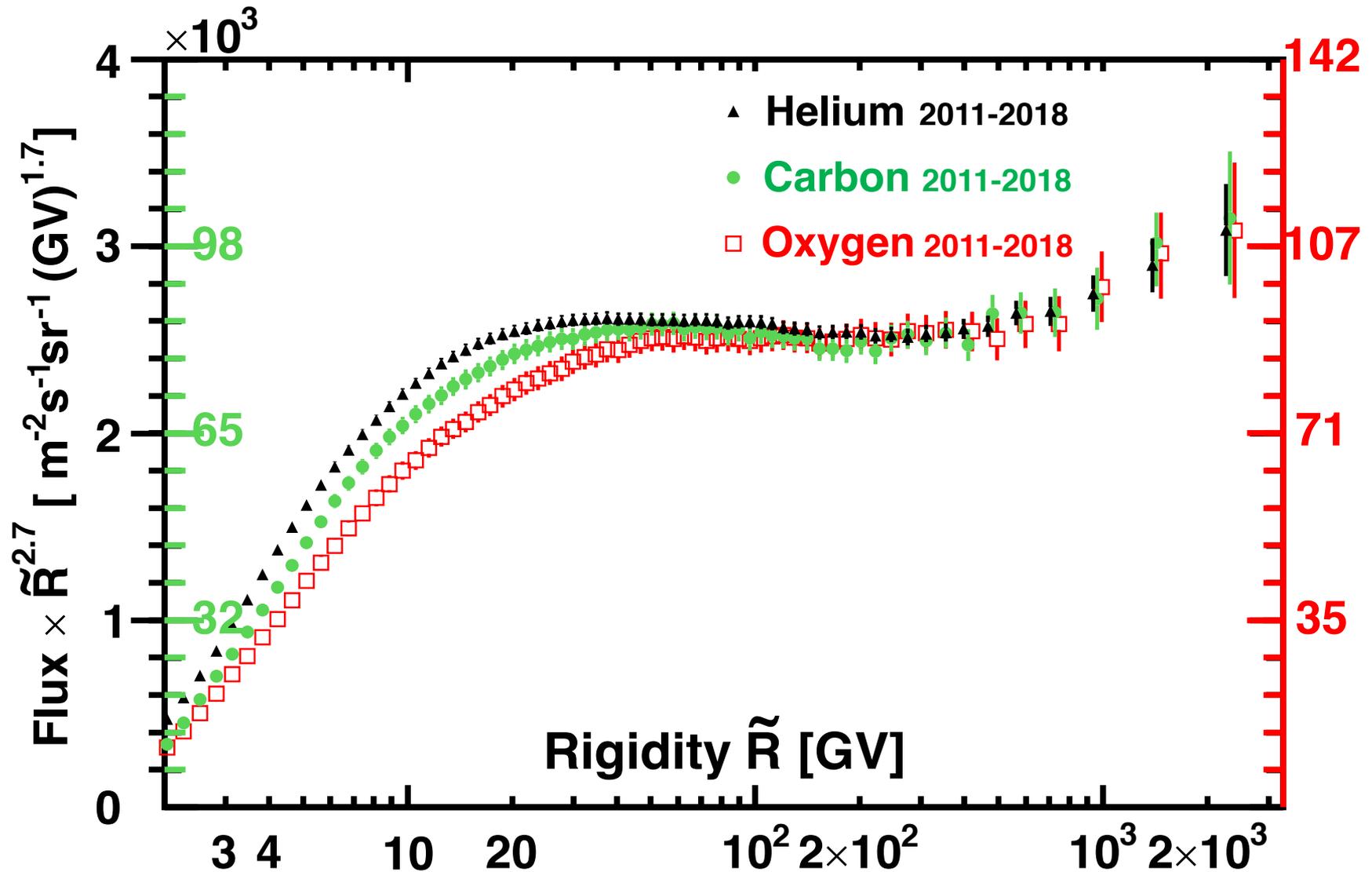
# Error Break Down (Silicon Flux)



# Flux Ratio to Silicon



# Latest AMS Measurements of He, C and O spectra



The new AMS result (2011-2018) is consistent with earlier AMS PRL result (2011-2016) “M. Aguilar *et al.*, Phys. Rev. Lett., **119**, 251101 (2017)” but with improved accuracy

# The Primary and Secondary Components of Carbon Flux

