

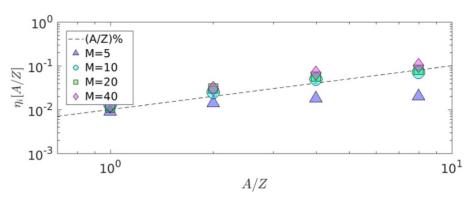


Acceleration of Helium Nuclei at Non-Relativistic Shocks

Cory Cotter, Marwah Roussi, Damiano Caprioli

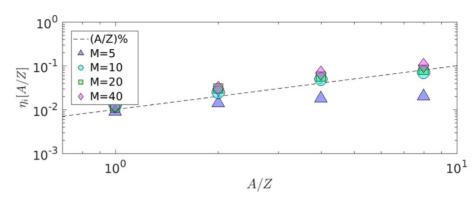
Previous studies

- Hybrid kinetic PIC simulations
- Caprioli et al. (2017)
 - He and other nuclei as test particles
 - Selection rate goes as (A/Q)²



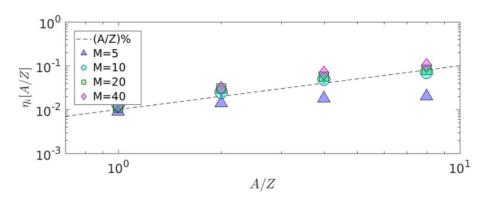
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 - Selection rate goes as (A/Q)
- This work
 - Changes in dynamics



- He should contain a comparable amount of energy to protons
- He and other heavy nuclei are preferentially accelerated

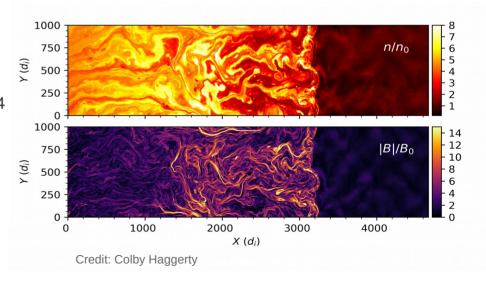
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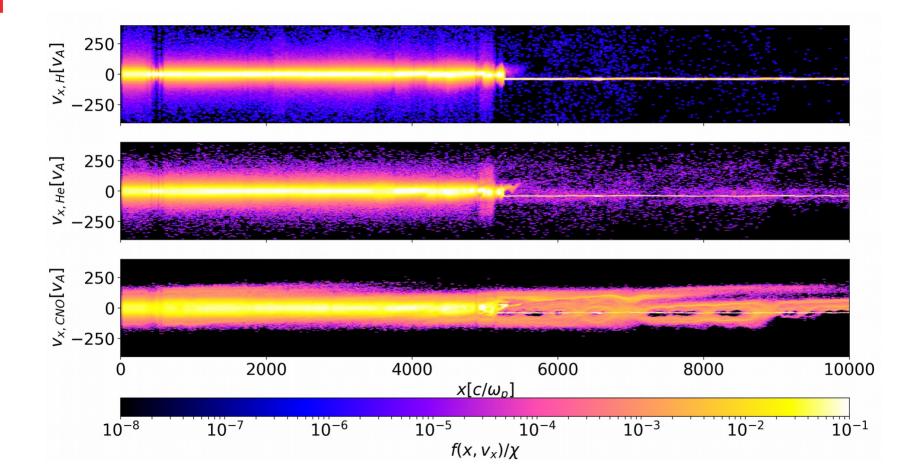
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- Has yet to be studied in detail in simulations
- May alter expected CR spectrum and γ-ray spectrum

Simulation Setup

- 2D hybrid simulation code dHybrid (Gargaté et al. 2007)
- Supersonic flow against reflecting wall
- Mach numbers: 5, 20, 40
- Solar abundances for H, He, CNO
 - $x_{H}=1, x_{He}=0.0963, x_{CNO}=9.54*10^{-4}$

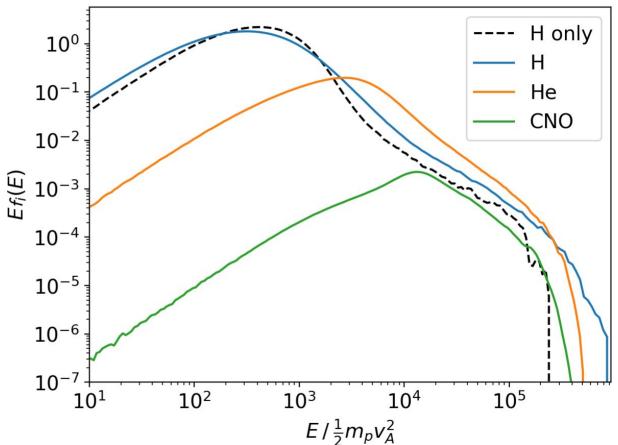


Phase Space



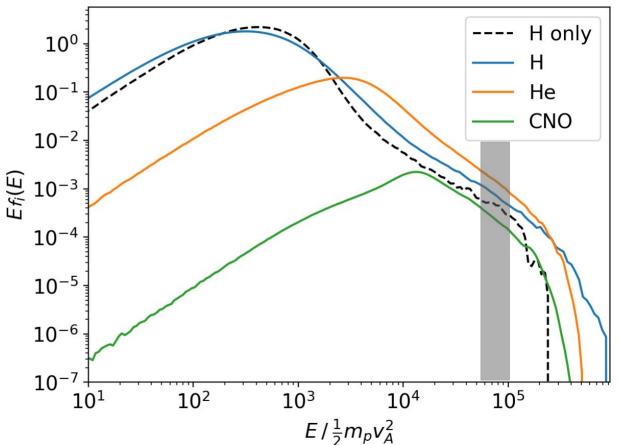
Energy Spectra

- Taken at $t = 520\omega_c$
- M=40
- Max energy increased by ~2-3



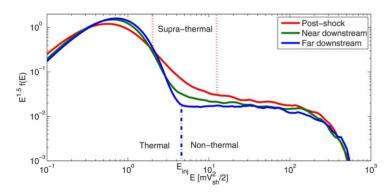
Energy Spectra

- Taken at $t = 520\omega_{c}$
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- Max energy increased by ~2-3
- He contains most energy



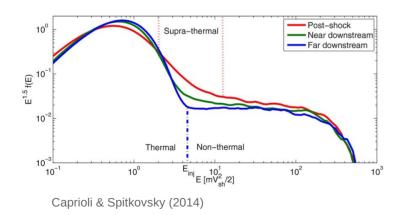
Acceleration Efficiency

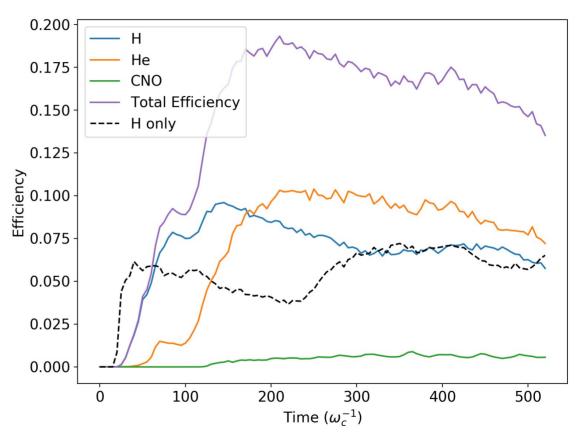
Energy in accelerated particles over total energy in simulation



Acceleration Efficiency

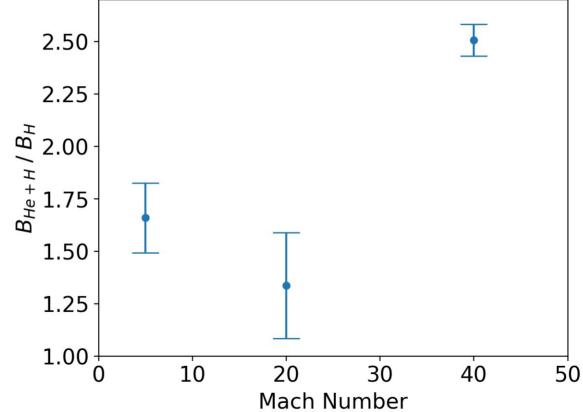
- Energy in accelerated particles over total energy in simulation
- He dominates
- CNO not important





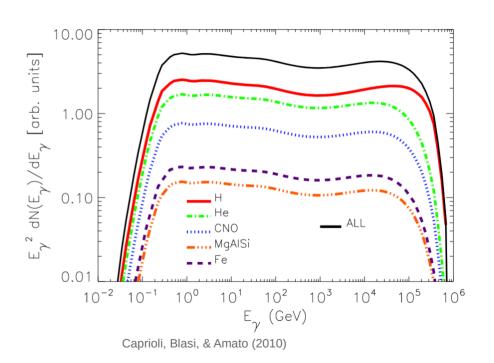
Magnetic Field Amplification

- Average upstream
 B₁ field over 100ω_c
- Most important for higher mach number
- Reduces the tension on galactic sources (Bell et al. 2013, Cardillo et al. 2015)



Gamma Ray / Neutrino Emission

- Production via nuclear interactions with thermal particles
- He becomes the primary source of emission
- Parent protons now have a maximum energy 20 times larger than photons (instead of 10)



Conclusions

- Simulations w/o heavy ions leave out physics at higher mach numbers
- More energy in He than H
- B field and maximum energy of H increase by ~2-3 at M=40
- Gamma ray emission dominated by He
- Cotter, Roussi, & Caprioli (in progress)

B Field spectrum

