

Fitting B/C cosmic-ray data in the AMS-02 era

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Work based on :

Fitting B/C cosmic-ray data in the AMS-02 era: a cookbook

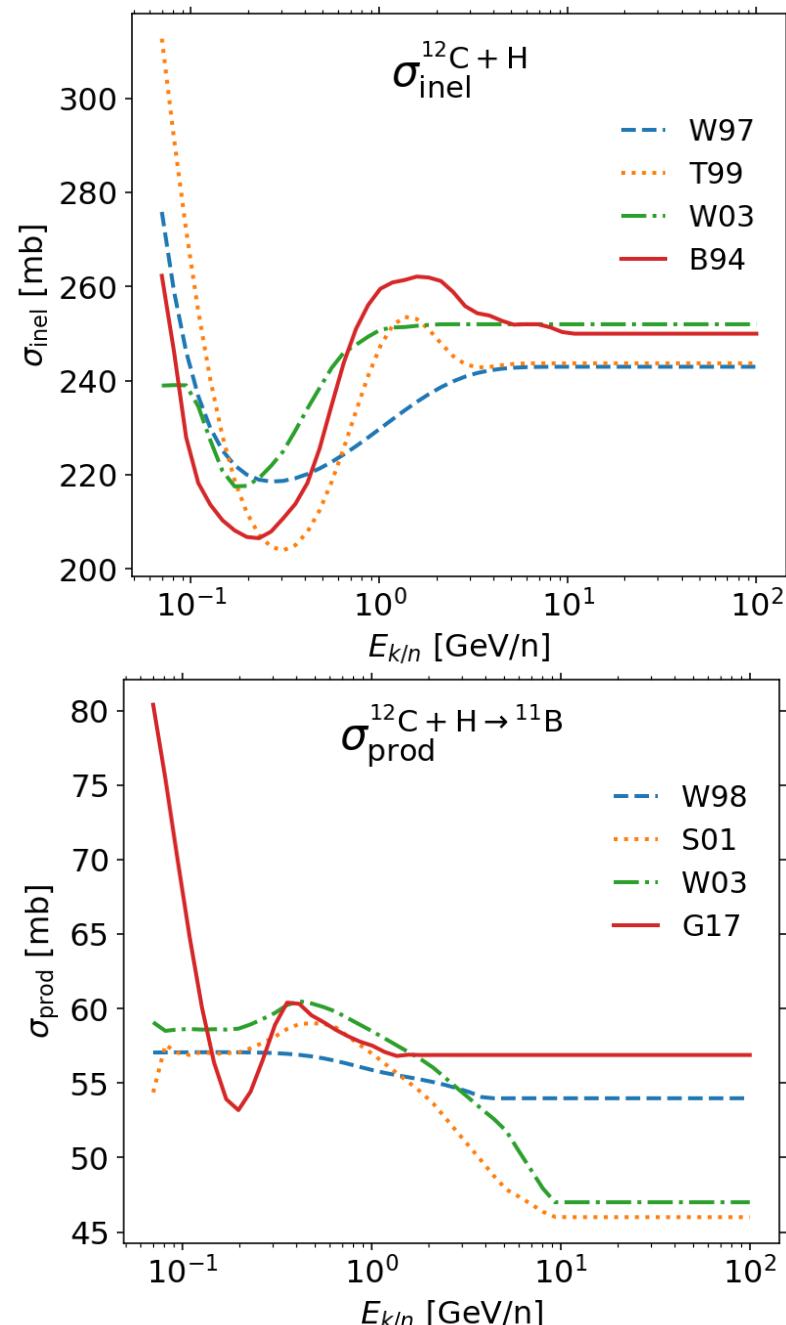
Astronomy & Astrophysics, 627, A158 (2019)

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- New generation of experiments (AMS-02, ...): percent-level precision, systematic-dominated data.
 - Methods used to constrain models with these data should be updated:
 - improved model precision:
 - Boundary condition
 - Stability of numerical solution
 - Handling of cross-section uncertainties
 - Handling of systematics errors from experimental data
- } This presentation

Nuclear Cross Sections

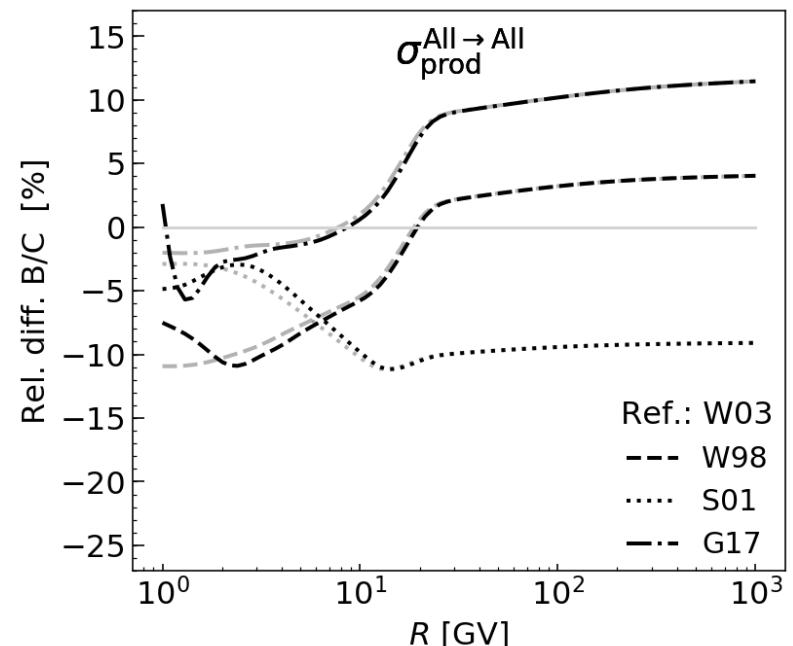
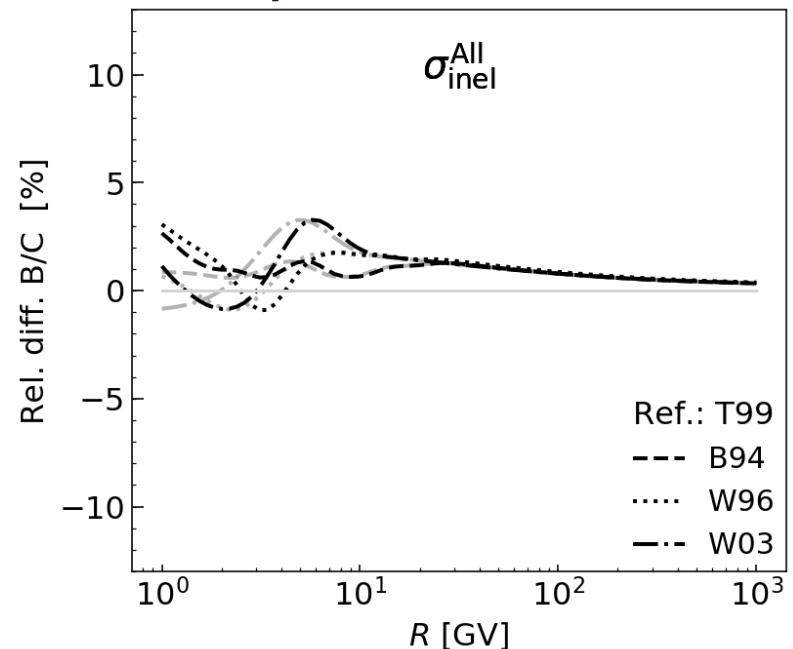
- Major ingredient in the modelling of GCR nuclei:
 - Inelastic cross sections: sink term for both primaries and secondaries.
 - Production cross sections: source term for secondaries.
- Several parameterisations of the reactions, based on experimental data, available:
 - Inelastic: Barashenkov [B94], Wellisch [W97], Tripathi [T99], Weber [W03]
 - Production: Weber [W98], Soutoul 01 [S01], Weber [W03], Galprop [G17]



Nuclear Cross Sections

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 - Impact on B/C ratio larger than data precision

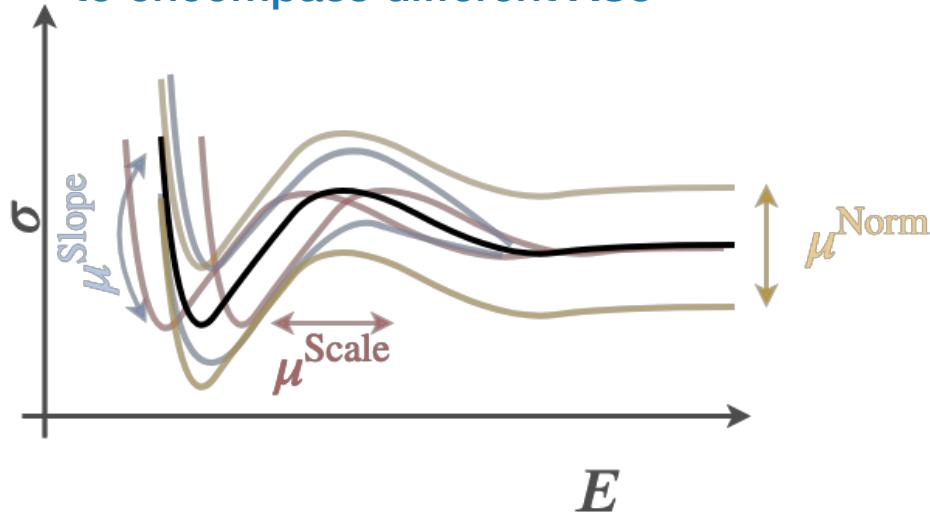
Impact on B/C ratio:



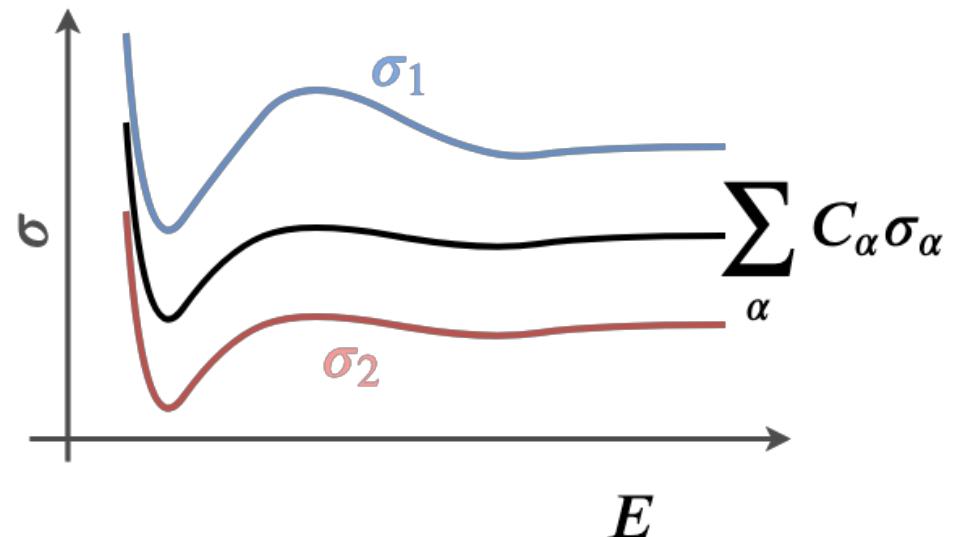
Nuclear Cross Sections

2 approaches to implement cross-section systematics:

- Norm Scale Slope (NSS)
 - Simple transformations from reference XS.
 - 3 Nuisance parameters with gaussian distribution and with ranges chosen to encompass different XSs



- Linear Combination (LC)
 - Linear combination of available XSs.
 - Nuisance parameters C_α with a flat distribution.
 - Can fully recover each XSs parametrisation



- Cross Sections considered (not the all network):
 - $\sigma_{\text{Inel.}}$ for B, C, O
 - $\sigma_{\text{Prod.}}$ for C \rightarrow $^{10, 11}\text{B}$ and O \rightarrow $^{10, 11}\text{B}$,
- In total: 14-20 nuisance parameters used to handle cross-section systematics

Propagation Models

- 2 configurations :

- Model A (diffusion + reacceleration + convection aka QUANT):

$$K(R) = \beta^{\eta} K_0 \left(\frac{R}{1 \text{ GV}} \right)^{\delta} + \text{reacceleration } (V_a) + \text{convection } (V_c)$$

5 free parameters

- Model B (pure diffusion aka SLIM):

$$K(R) = \beta K_0 \left(\frac{R}{1 \text{ GV}} \right)^{\delta} \left(1 + \left(\frac{R_l}{R} \right)^{(\delta + \delta_l)/s_l} \right)^{s_l}$$

4 free parameters

- Mock Data (1000) generated with statistical errors around reference models for A and B.
- Fit of the mock data, two cases considered:
 - Unbiased: $\sigma_{\text{Inel., Prod.}}(\text{Fit}) = \sigma_{\text{Inel., Prod.}}(\text{Gen.})$ (Inel.: W97, Prod.: G17)
 - Biased: $\sigma_{\text{Inel., Prod.}}(\text{Fit})$ (Inel.: T99, Prod.: W03) $\neq \sigma_{\text{Inel., Prod.}}(\text{Gen.})$



Unbiased

$$\sigma(\text{Gen.}) = \sigma(\text{Fit})$$

No nuisance (black):

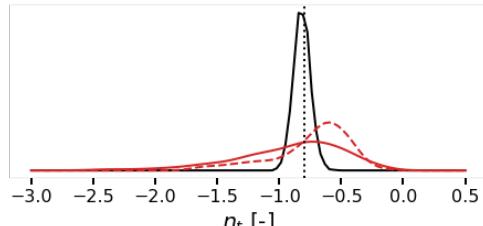
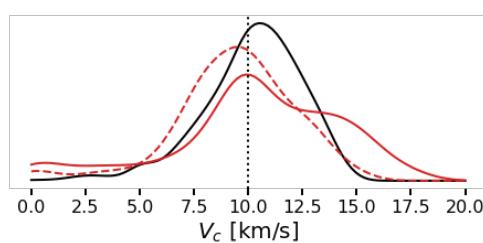
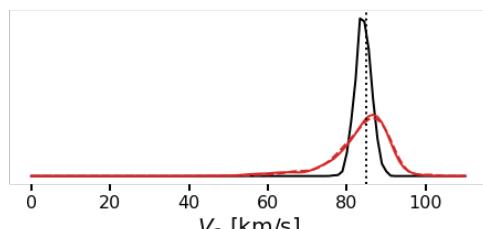
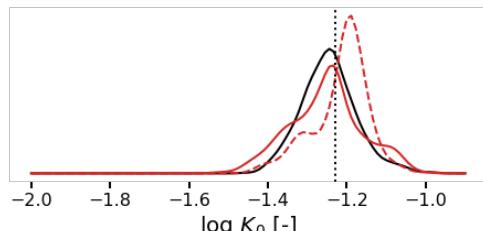
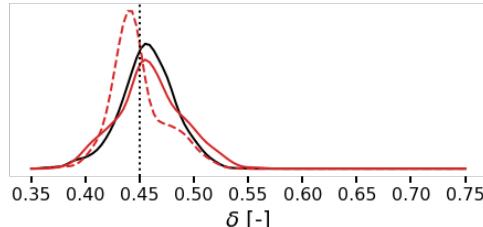
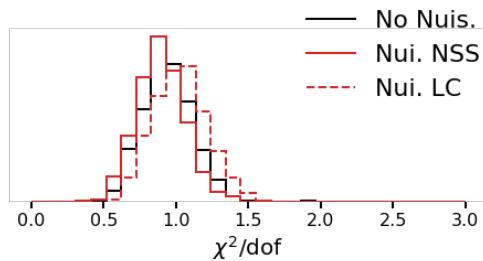
- $\chi^2/\text{dof} \sim 1$
- no bias w.r.t parameters values of the models (vertical dashed lines)

With nuisances NSS (–) and

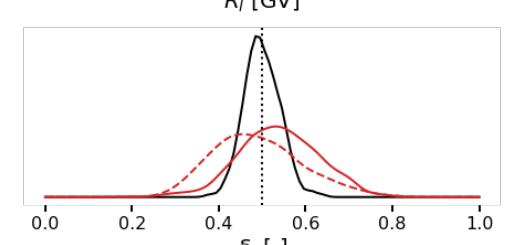
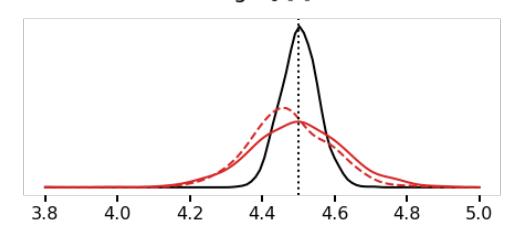
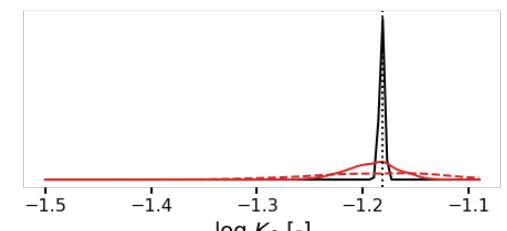
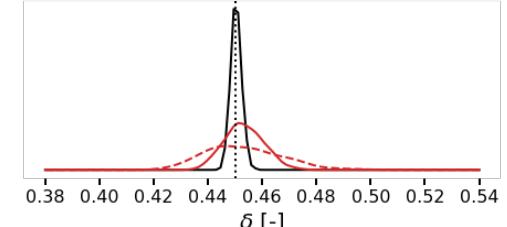
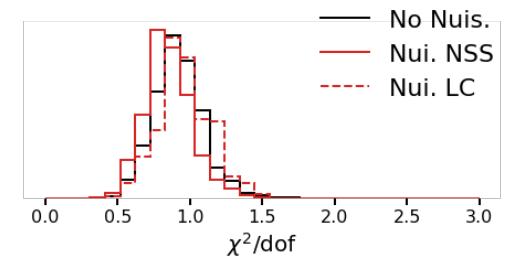
LC (--):

- $\chi^2/\text{dof} \sim 1$
- Larger errors (systematic from XS now included in the errors from the fit)

MODEL A



MODEL B



Biased

$$\sigma(\text{Gen.}) \neq \sigma(\text{Fit})$$

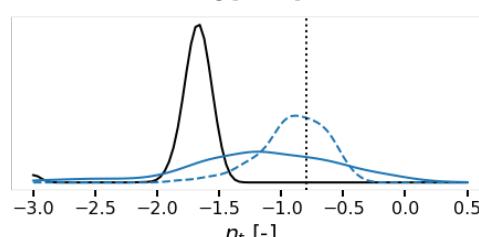
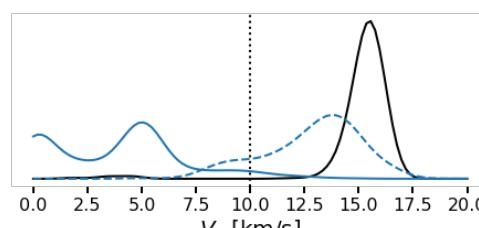
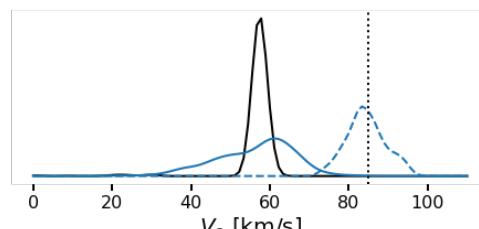
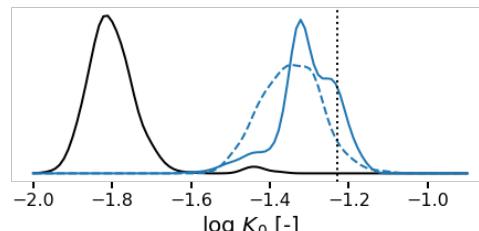
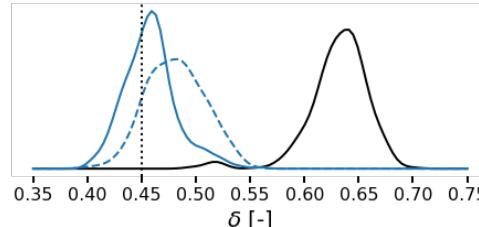
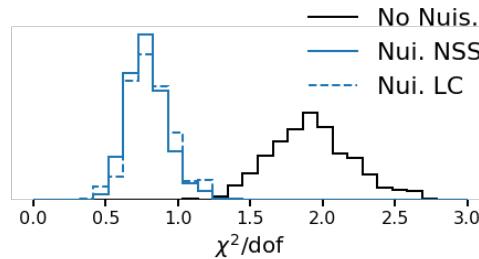
No nuisance (black):

- Large χ^2/dof
- Biased best fits w.r.t to parameters values (vertical dashed lines)

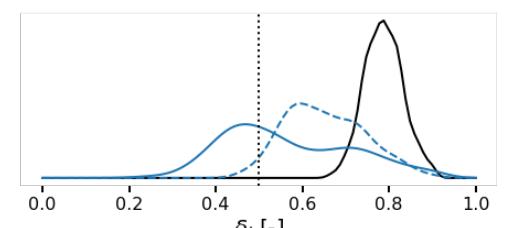
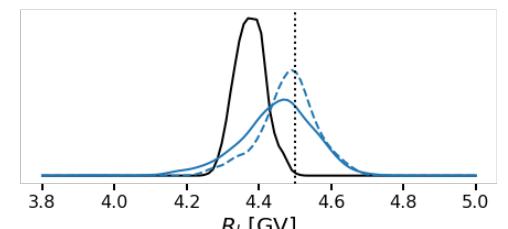
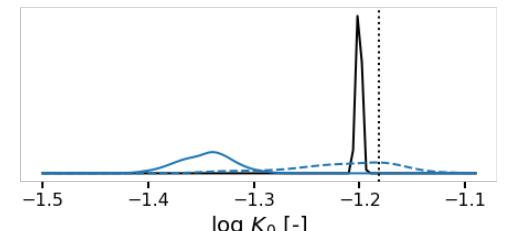
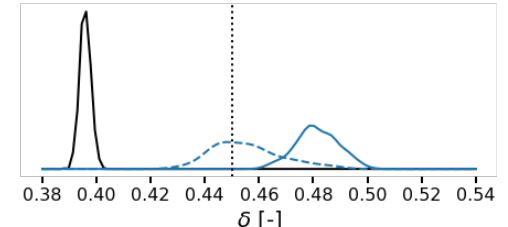
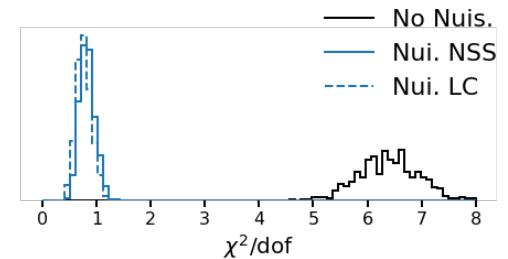
With nuisances NSS (—) and LC (--):

- good χ^2/dof recovered
- Reduced biases on the parameter distributions
- Better results for LC than for NSS as expected.

MODEL A



MODEL B



Systematic errors on Data

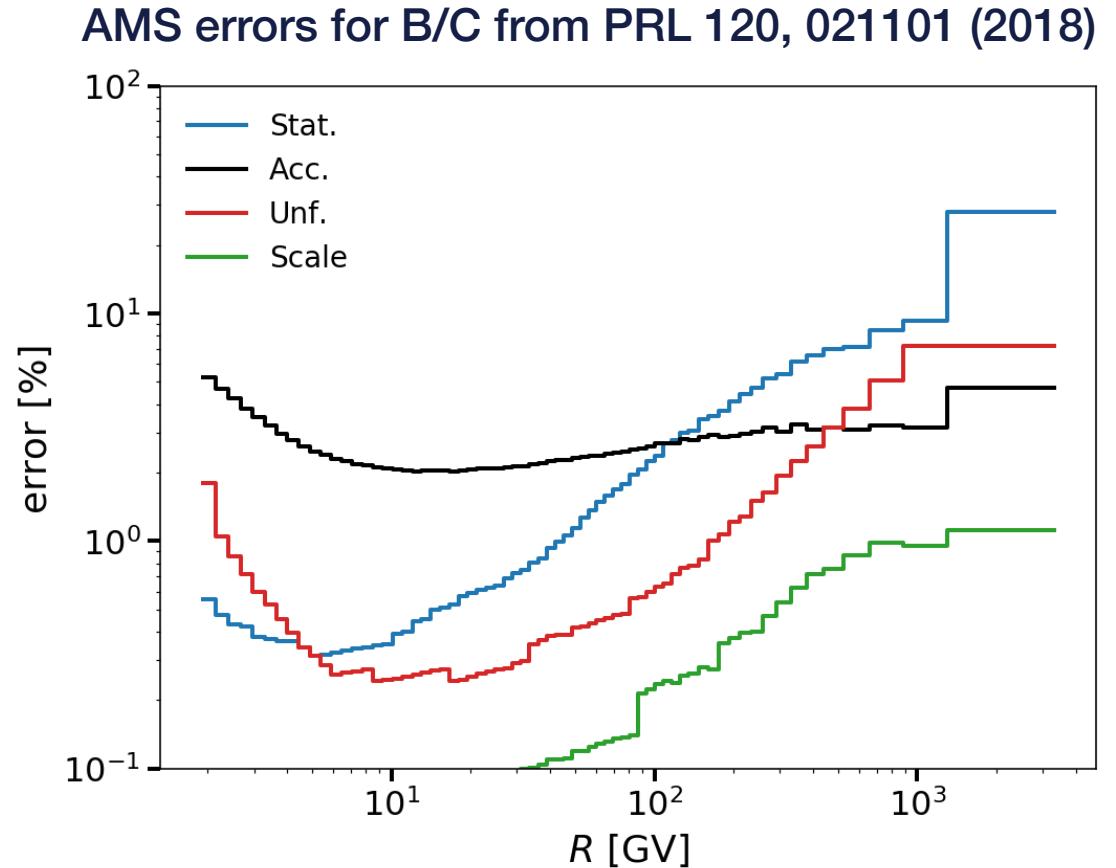
AMS02 B/C:

- Dominated by systematic Acceptance error up to 100 GV
- No covariance matrix provided to account for bin-to-bin correlation of systematic errors.
- Model guessed for the covariance matrices ($\alpha = \text{Stat.}, \text{Acc.}, \text{Unf.}, \text{Scale}$):

$$(C_{\text{rel}}^{\alpha})_{ij} = \sigma_i^{\alpha} \sigma_j^{\alpha} \exp \left(-\frac{1}{2} \frac{(\log_{10}(R_i/R_j))^2}{(l_{\rho}^{\alpha})^2} \right)$$

where:

- σ_i^{α} is the error from AMS
- l_{ρ}^{α} is the correlation length in unit of decade of rigidity.



- $l_{\rho}^{\text{Stat.}} = 0$
- $l_{\rho}^{\text{Unf.}} = 0.5$
- $l_{\rho}^{\text{Scale}} = \infty$
- $l_{\rho}^{\text{Acc.}} = 0.01\dots 3$

How the choice of $l_{\rho}^{\text{Acc.}}$ impact the best fit values ?

Systematic errors on Data

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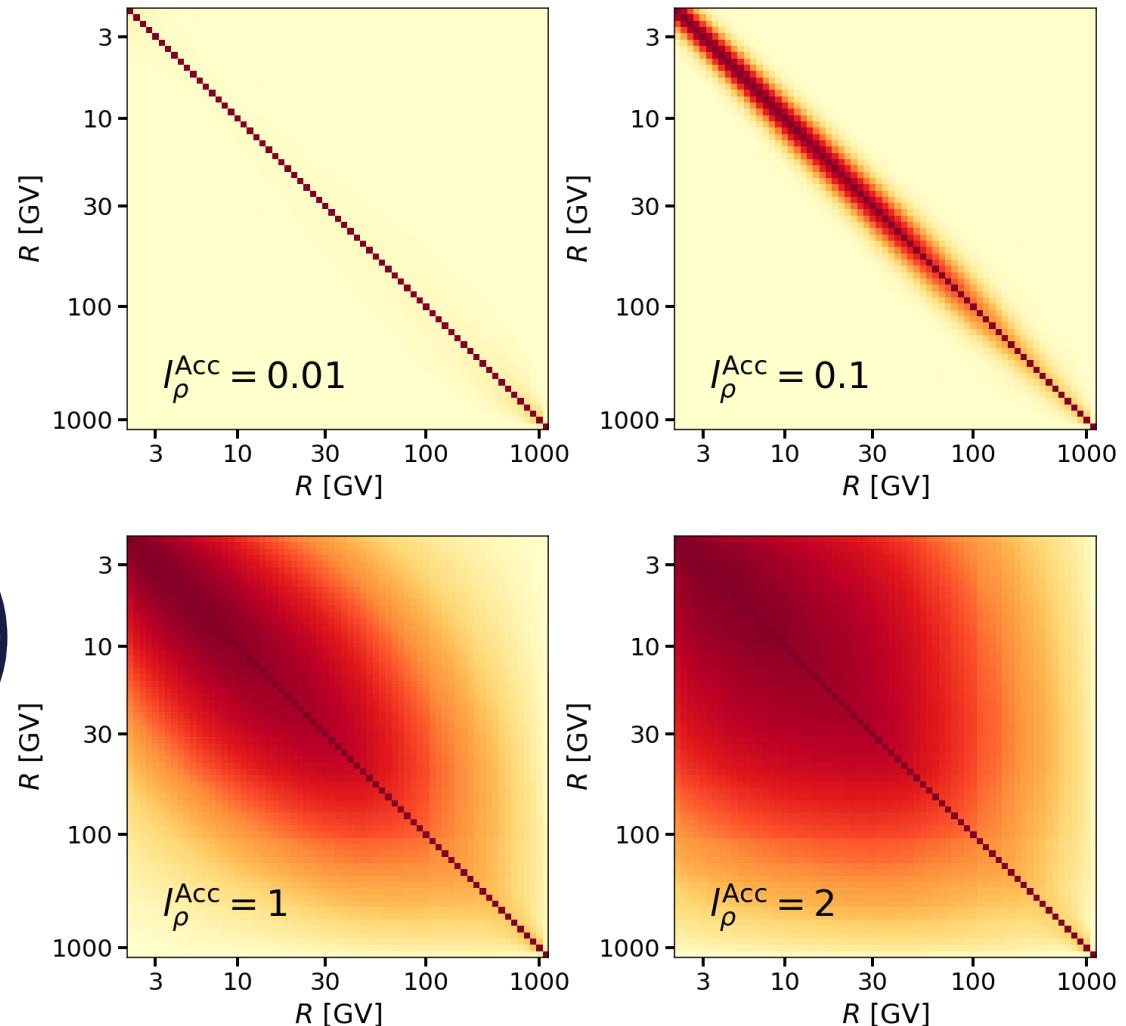
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Total correlation matrix

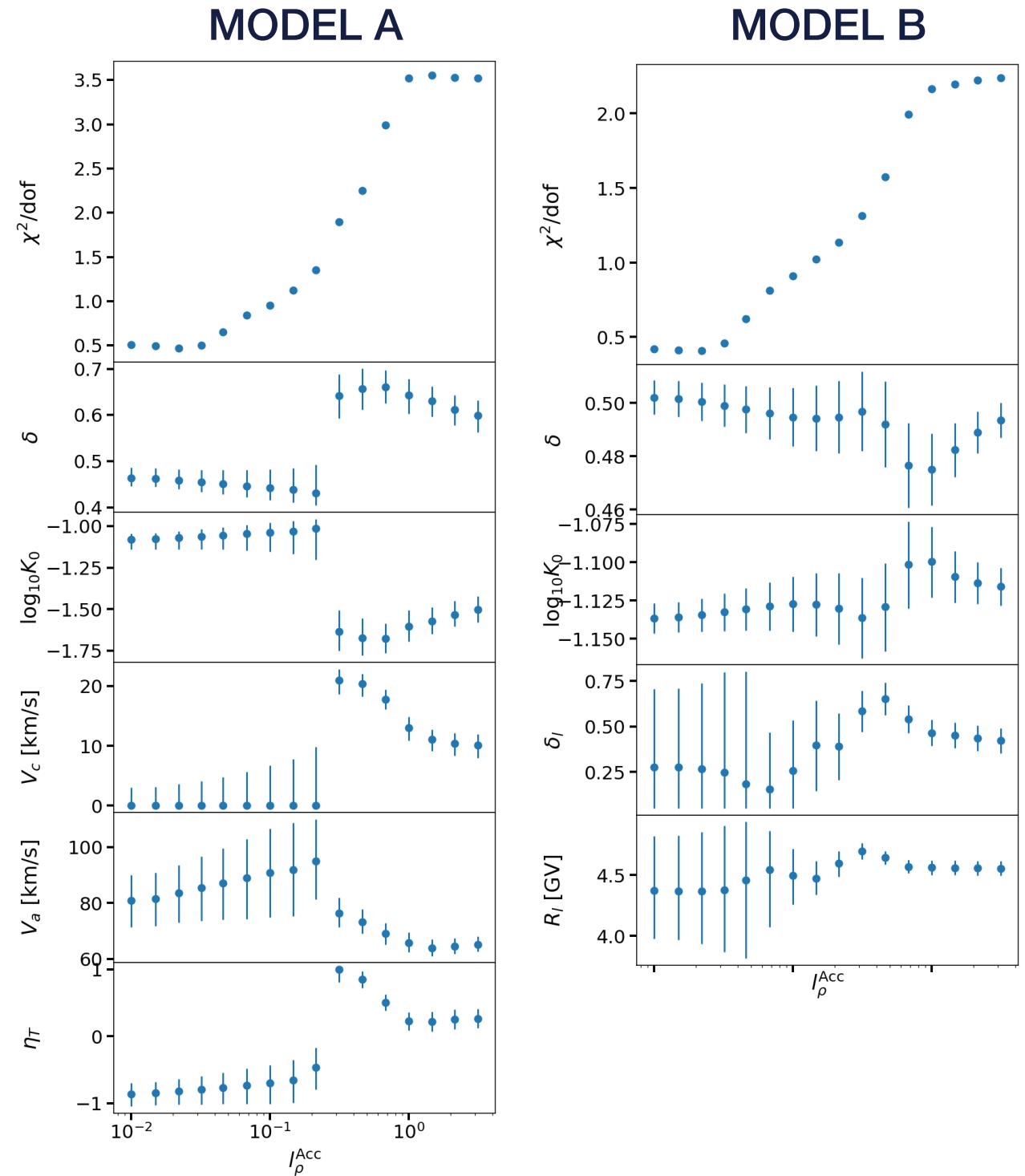


AMS-02

B/C fit

AMS B/C fit results as a function of $l_\rho^{\text{Acc.}}$:

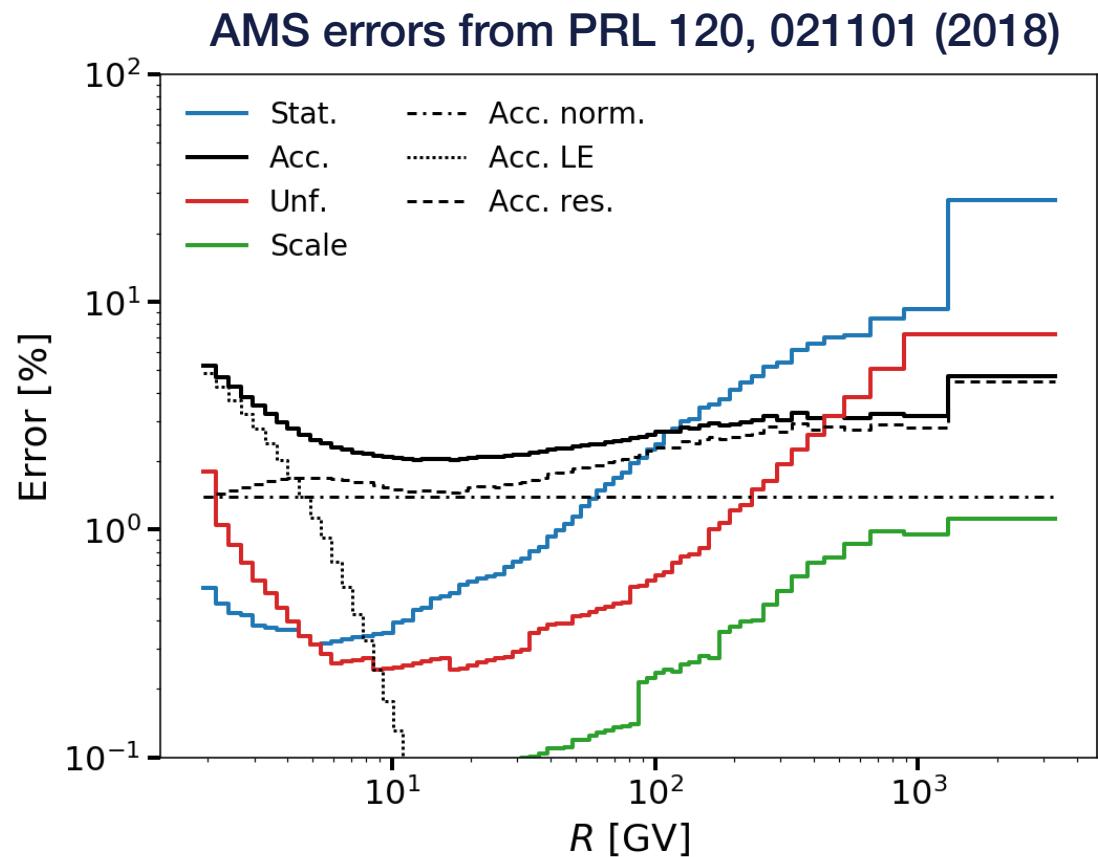
- χ^2/dof from 0.5 to 3.5
- Large best-fit values and errors dependence:
 - Model A: 2 different regimes for low $l_\rho^{\text{Acc.}}$ and high $l_\rho^{\text{Acc.}}$
 - Model B: Less dependence for values. Still large dependence for errors.



Systematic errors on Data

Acceptance error handling is critical:

- Sum of different contribution, may have different correlation lengths.
- Split acc. error into 3 contributions:
 - Acc. Norm: Flat normalisation error, related to systematic on survival probability : $l_\rho \sim 1.0$
 - Acc. LE: Low Energy error : $l_\rho \sim .3$
 - Acc. res. : Residual error data/ MC corrections, ...
 $l_\rho = 0.01\dots 3$

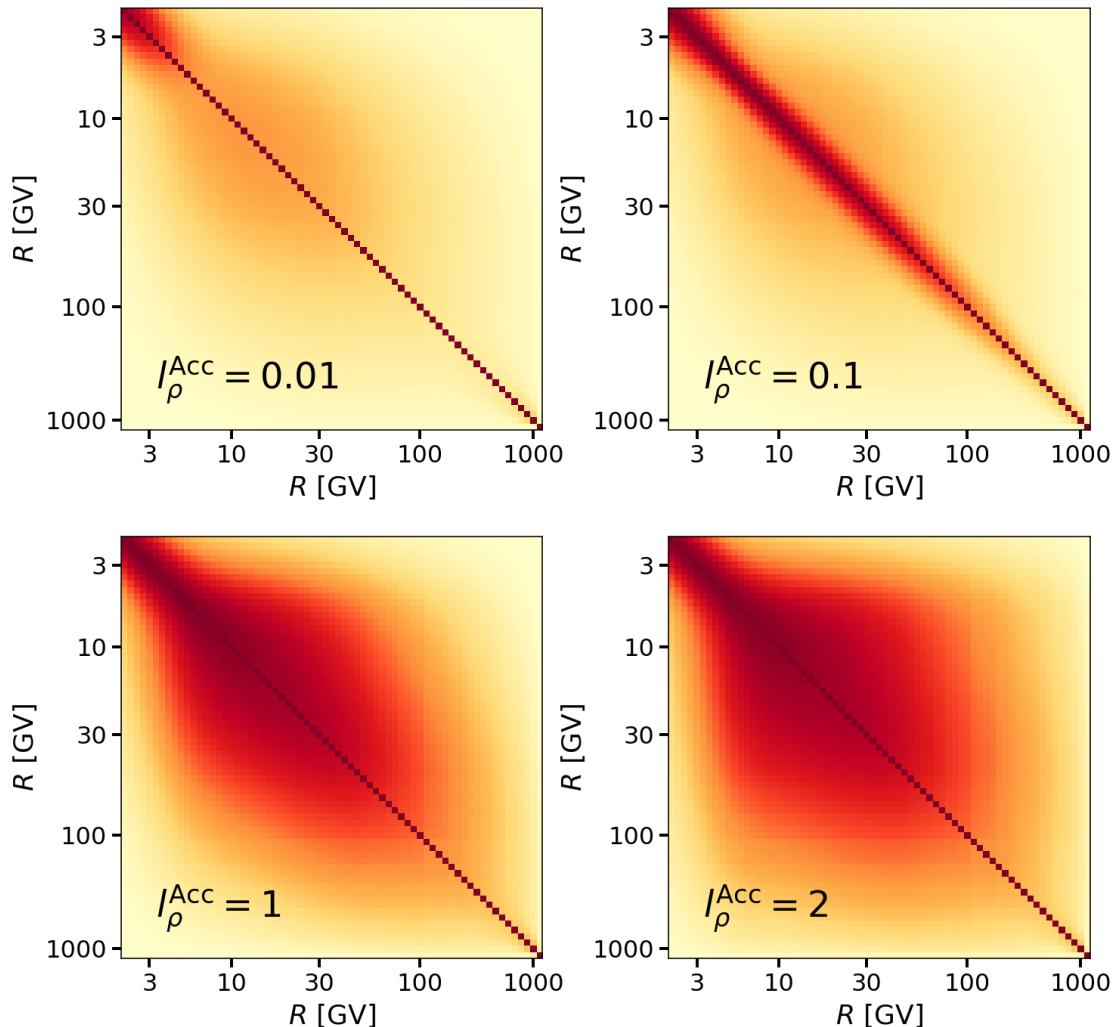


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Total correlation matrix

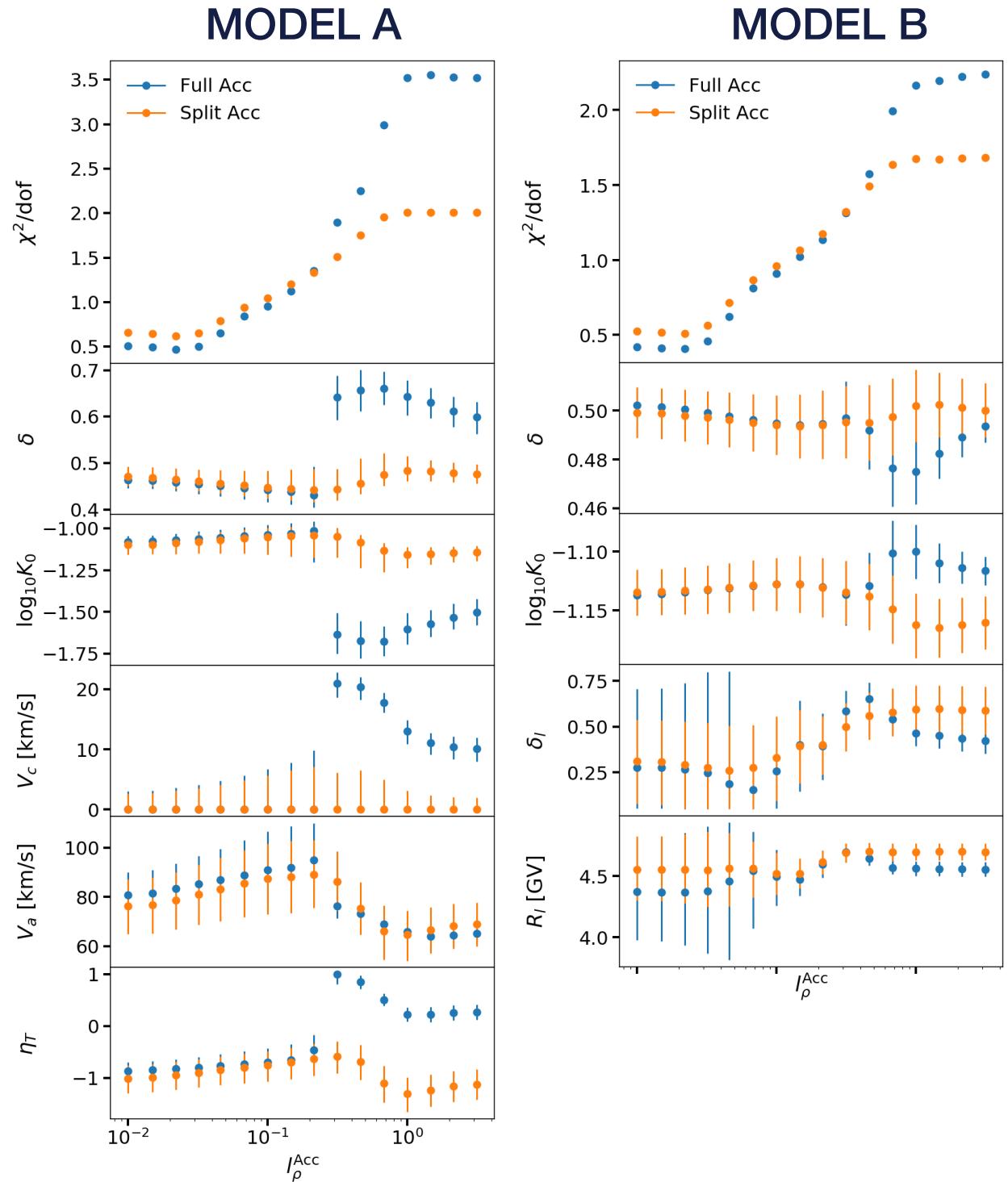


AMS-02

B/C fit

AMS B/C fit results as a function of $l_p^{\text{Acc. res.}}$:

- χ^2/dof from 0.7 to 2
- Best-fit values dependence reduced. More stable results for both Model A and B.
- $l_p^{\text{Acc.}} = 0.1$ gives $\chi^2/\text{dof} \simeq 1$, stable results and conservative errors.



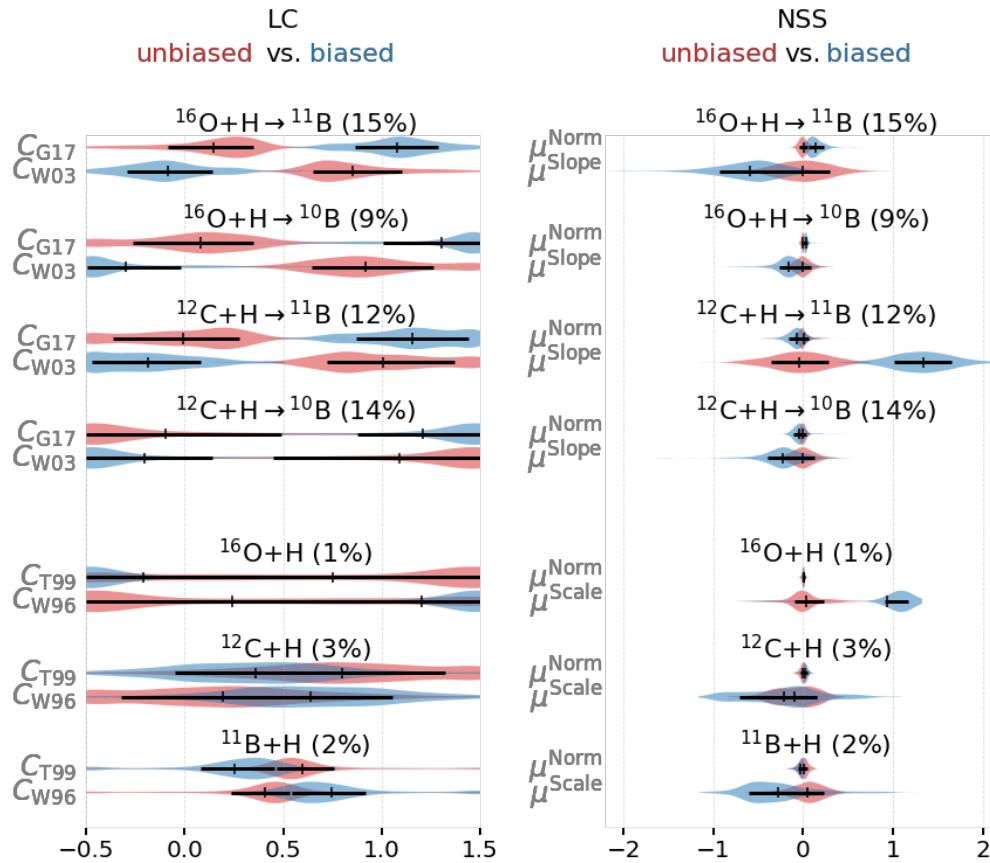
Conclusions

- Production and inelastic nuclear cross-sections can be implemented in the models with nuisances parameters to propagate ‘uncertainties’ and remove biases from wrong cross-sections.
- Handling of systematic on data from AMS-02 requires to model the bin-to-bin correlation for each source of systematics.
- Full description of analyses and results presented here: A&A, 627, A158 (2019)
- All analyses performed with USINE [<https://lpsc.in2p3.fr/usine>]
- Methodology presented here used in:
 - CRD1b (Thursday, July 25) AMS-02 Antiprotons are Consistent with a Secondary Astrophysical Origin, M. Boudeau
 - CRD6a (Monday, July 29): Cosmic ray transport from AMS-02 B/C data: reference parameters and physical interpretation, Y. Génolini

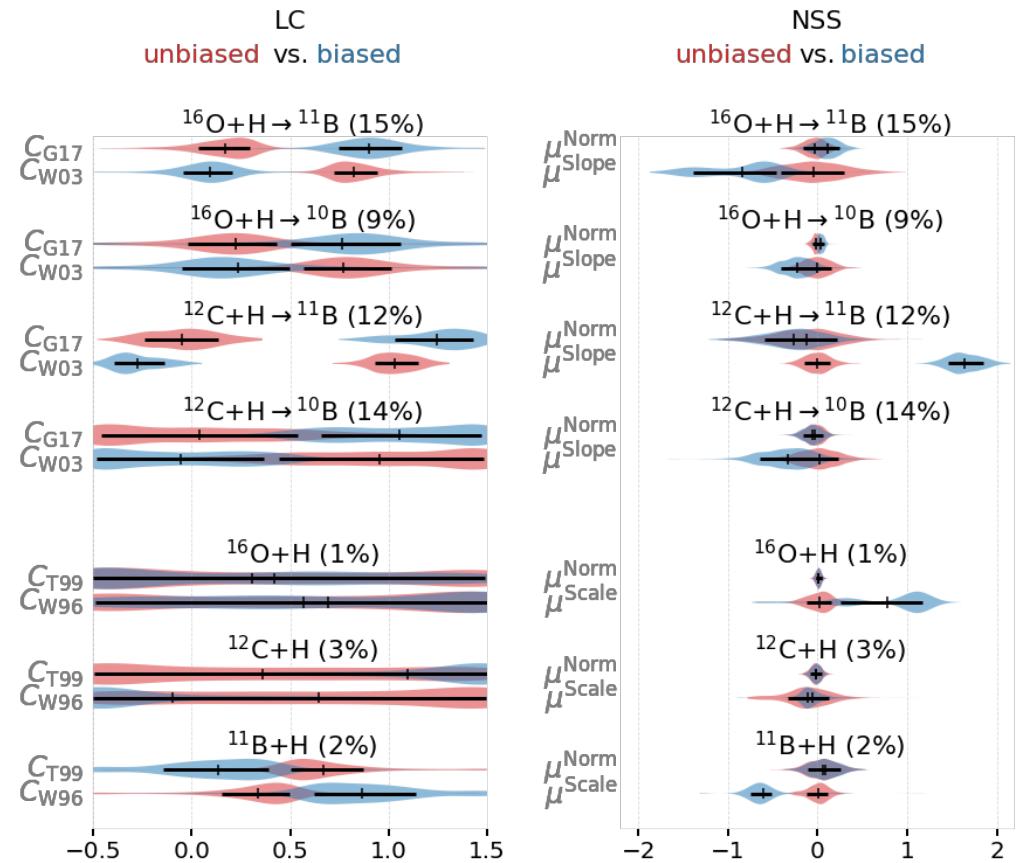
Backup slides

Nuclear Cross-Section

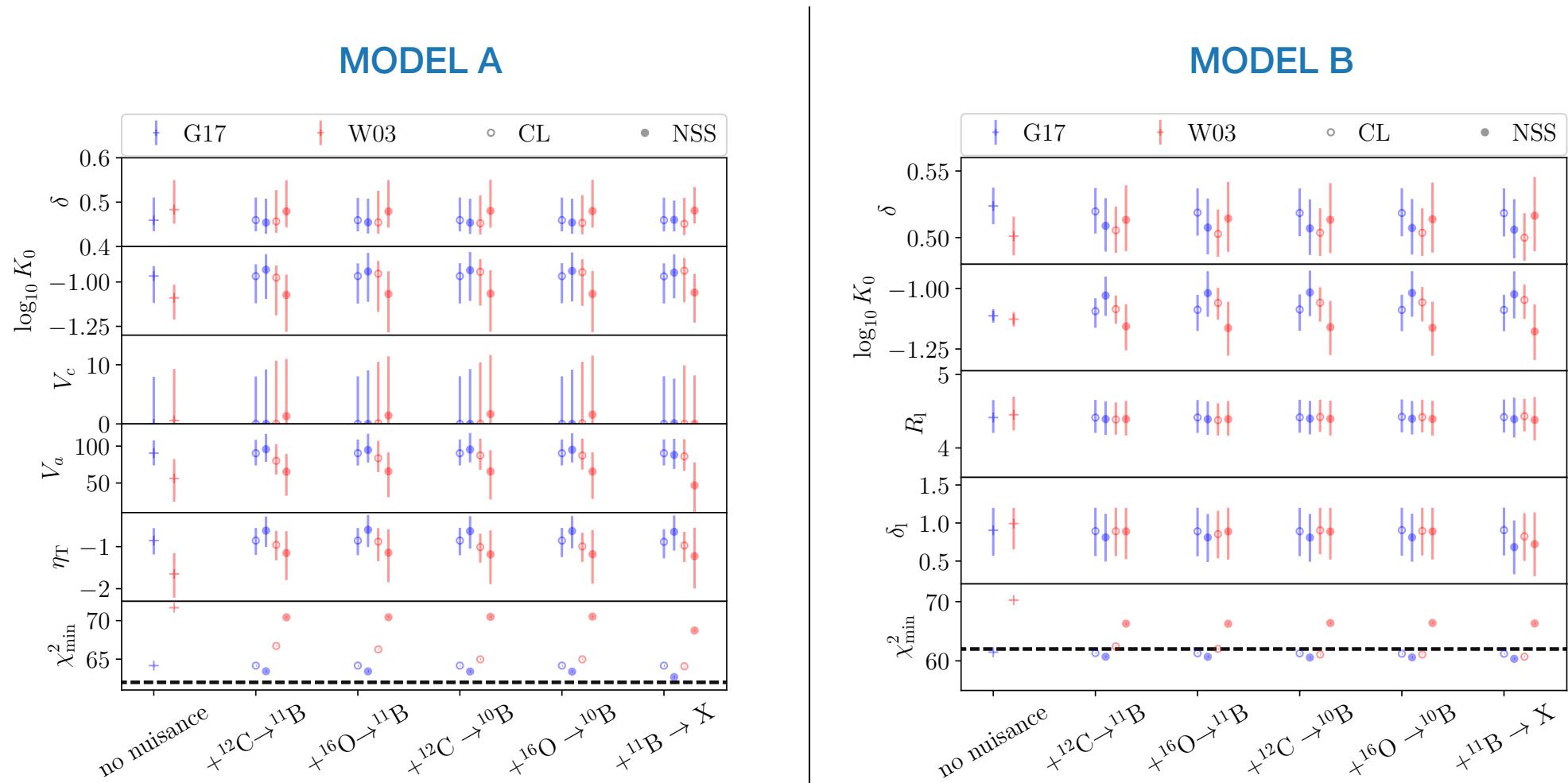
MODEL A



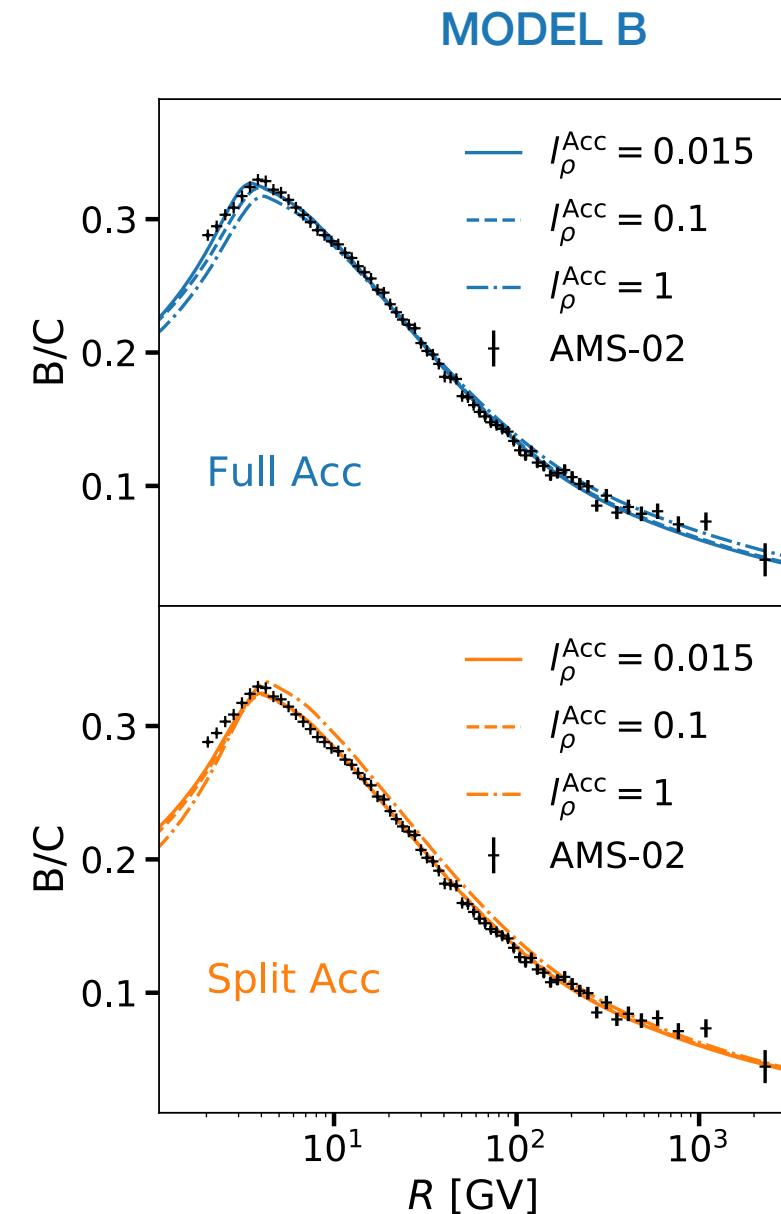
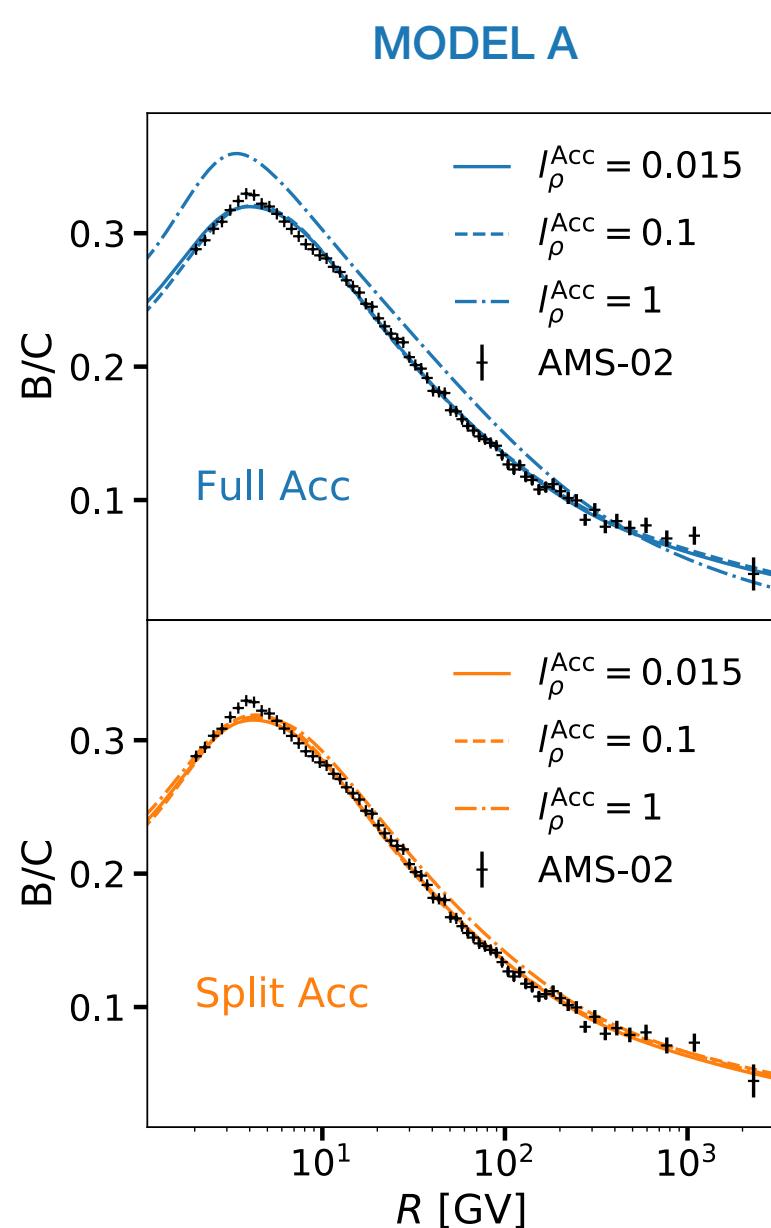
MODEL B



Systematic errors on Data + Nuclear Cross Sections



Nuclear Cross-Section



Propagation Models

- 2 configurations :

- Model A (diff.+reac.+conv. aka QUAINT):

$$K(R) = \beta^{\eta_t} K_0 \left(\frac{R}{1 \text{ GV}} \right)^{\delta} K_{HE}(R)$$



+ reacceleration (V_a) + convection (V_c)

- Model B (pure diffusion aka SLIM):

$$K(R) = \beta K_0 \left(\frac{R}{1 \text{ GV}} \right)^{\delta} \left(1 + \left(\frac{R_l}{R} \right)^{(\delta + \delta_l)/s_l} \right)^{s_l} K_{HE}(R)$$



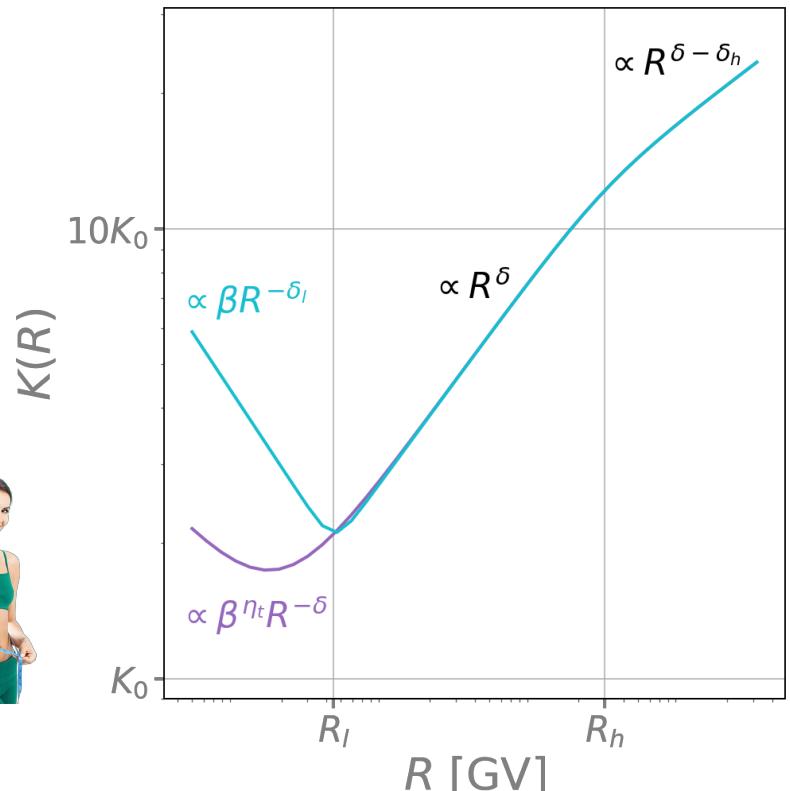
- Where $K_{HE}(R) = (1 + (R/R_h)^{\delta_h/s_h})^{-s_h}$ (fixed)

- Mock Data generated with statistical errors around reference models for A and B.

- Fit of the mock data, two cases considered:

- Unbiased: $\sigma_{\text{Inel., Prod.}}(\text{Fit}) = \sigma_{\text{Inel., Prod.}}(\text{Gen.})$ (Inel.: W97, Prod.: G17)

- Biased: $\sigma_{\text{Inel., Prod.}}(\text{Fit})$ (Inel.: T99, Prod.: W03) $\neq \sigma_{\text{Inel., Prod.}}(\text{Gen.})$



Nuclear Cross Sections

- Major ingredient in the modelling of GCR nuclei:
 - Inelastic cross sections: sink term for both primaries and secondaries.
 - Production cross sections: source term for secondaries.
- Impact on B/C ratio:
 - Inelastic XSs: 3% at ~ 5 GV, decreases to zero at higher R .
 - Production XSs: $\sim 10\%$ above 20 GV
 - Larger than data precision

Impact on B/C ratio:

