

# Cosmic-ray isotope measurements with HELIX

Presented by Nahee Park  
for HELIX Collaboration







# HELIX Collaboration

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● Mark Gebhard, Brandon Kunkler, James Musser, Kelli Michaels, Gerard Visser

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## Northern Kentucky University

● Scott Nutter

## Ohio State University

● Patrick Allison, James J. Beatty, Keith McBride

## Pennsylvania State University

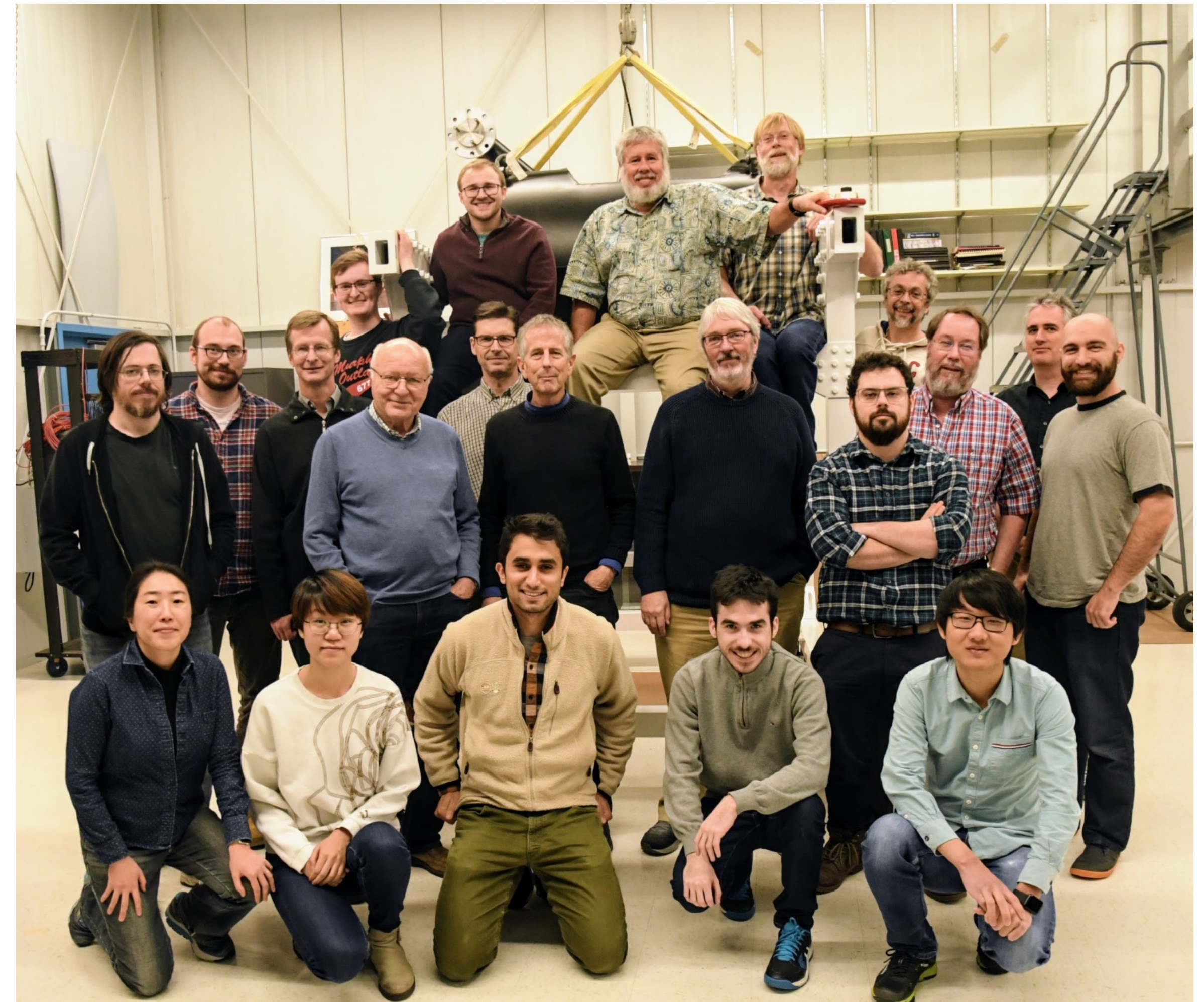
● Yu Chen, Stephane Coutu, Issac Mognet, Monong Yu

## University of Michigan

● Noah Green, Gergory Tarle, Andrew Tomasch

## University of Wisconsin-Madison

● Nahee Park

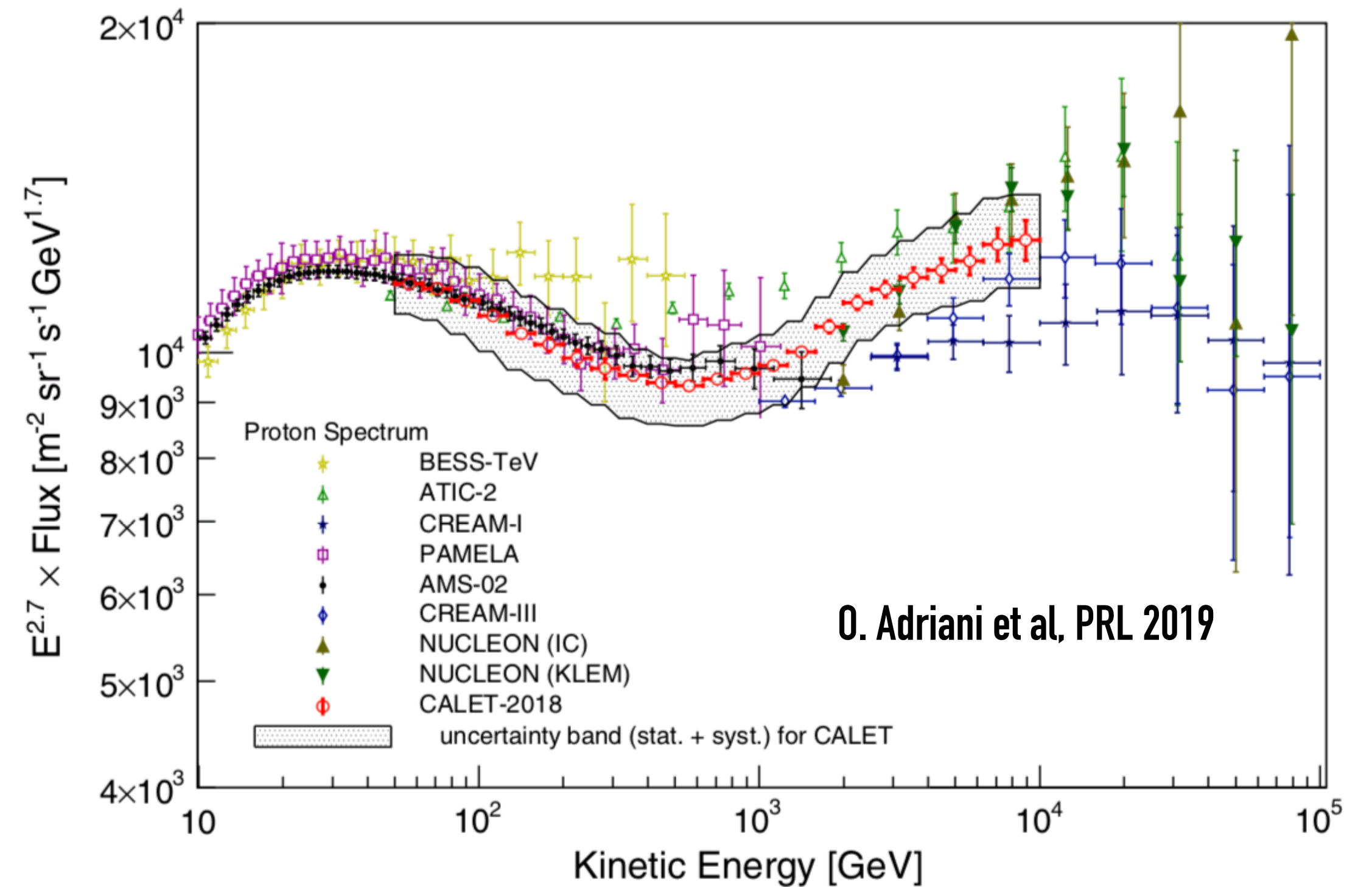
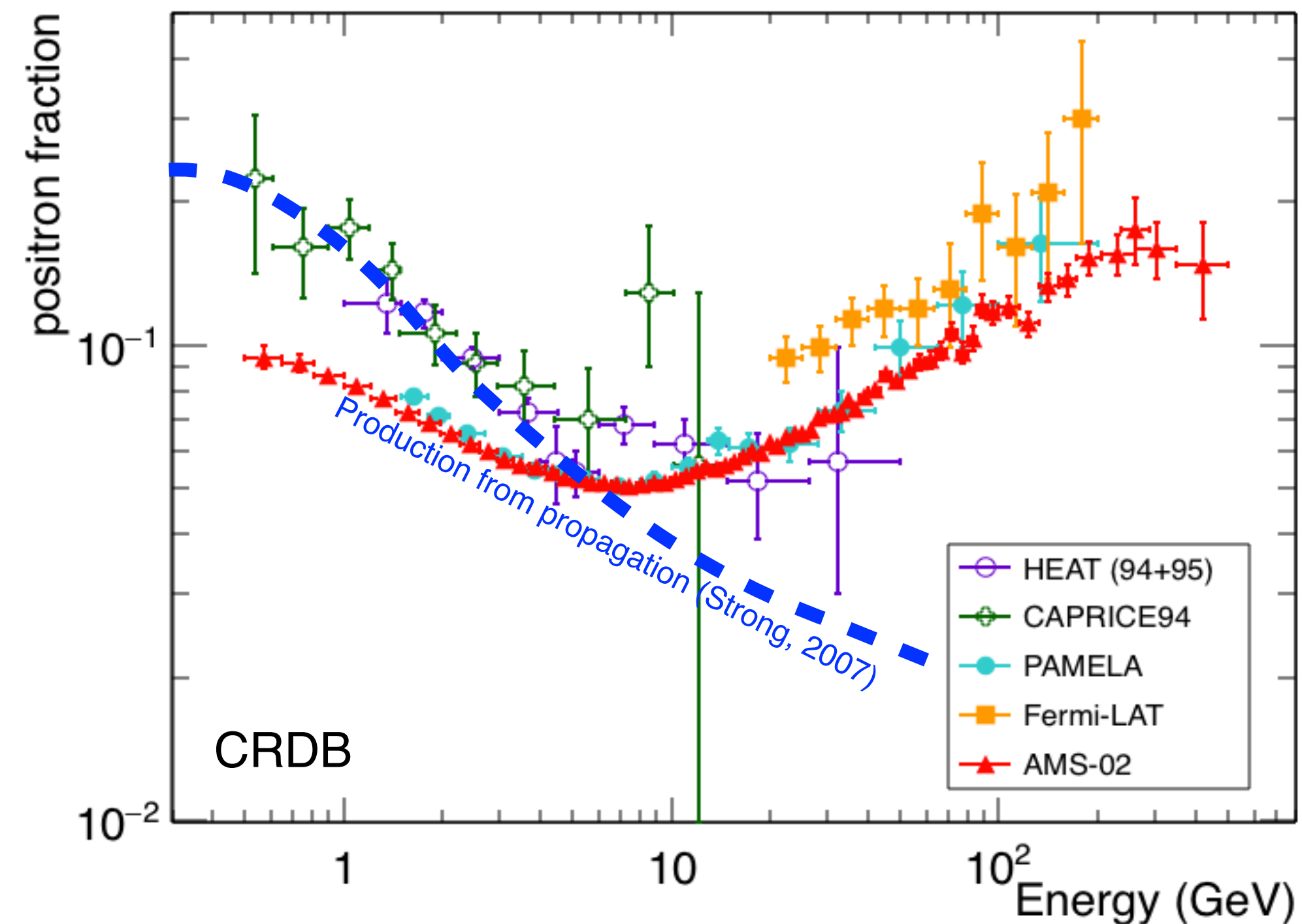




# Recent Updates from Direct Measurement

A new era of precision space-based measurements has brought some real surprises

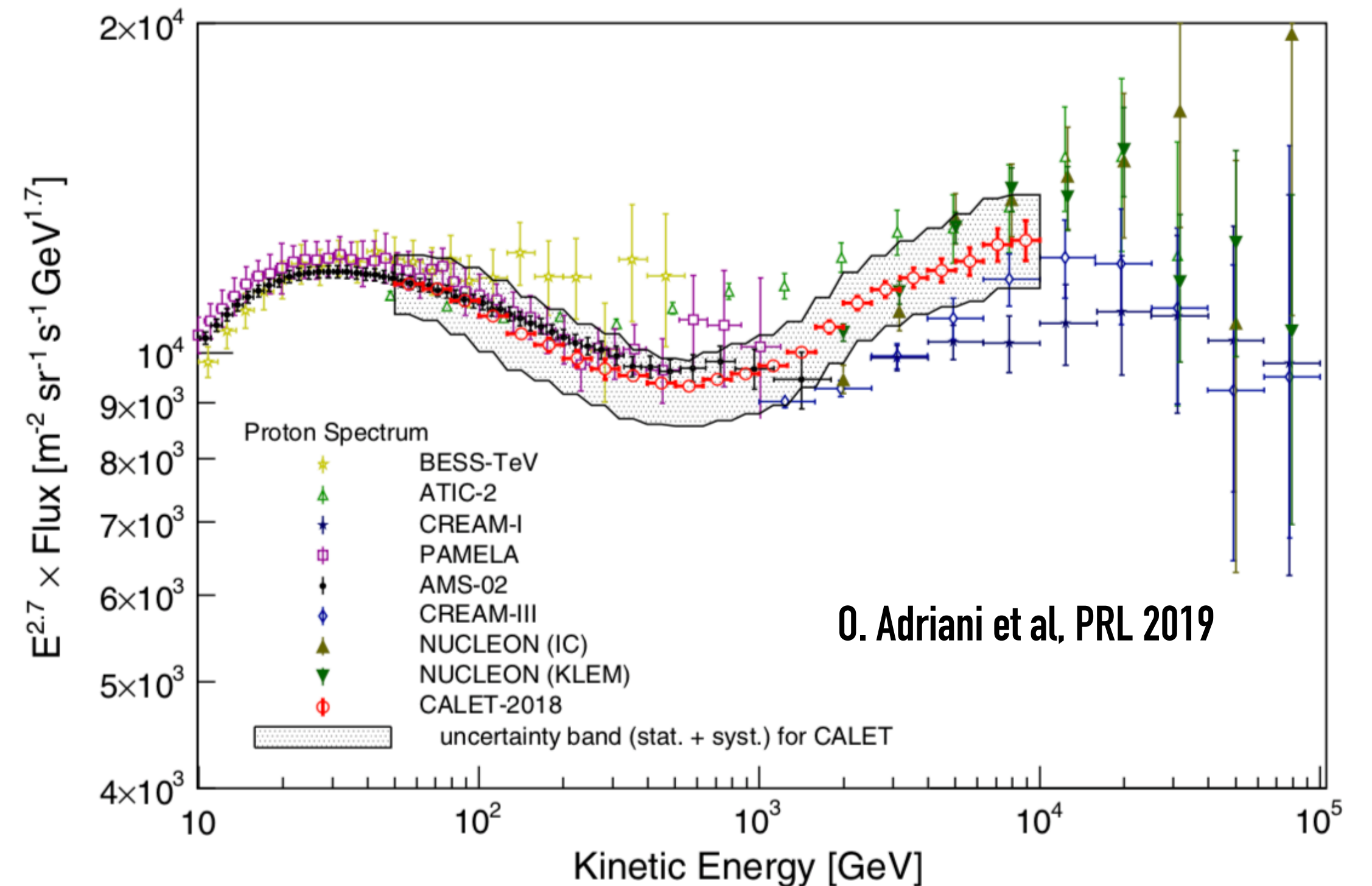
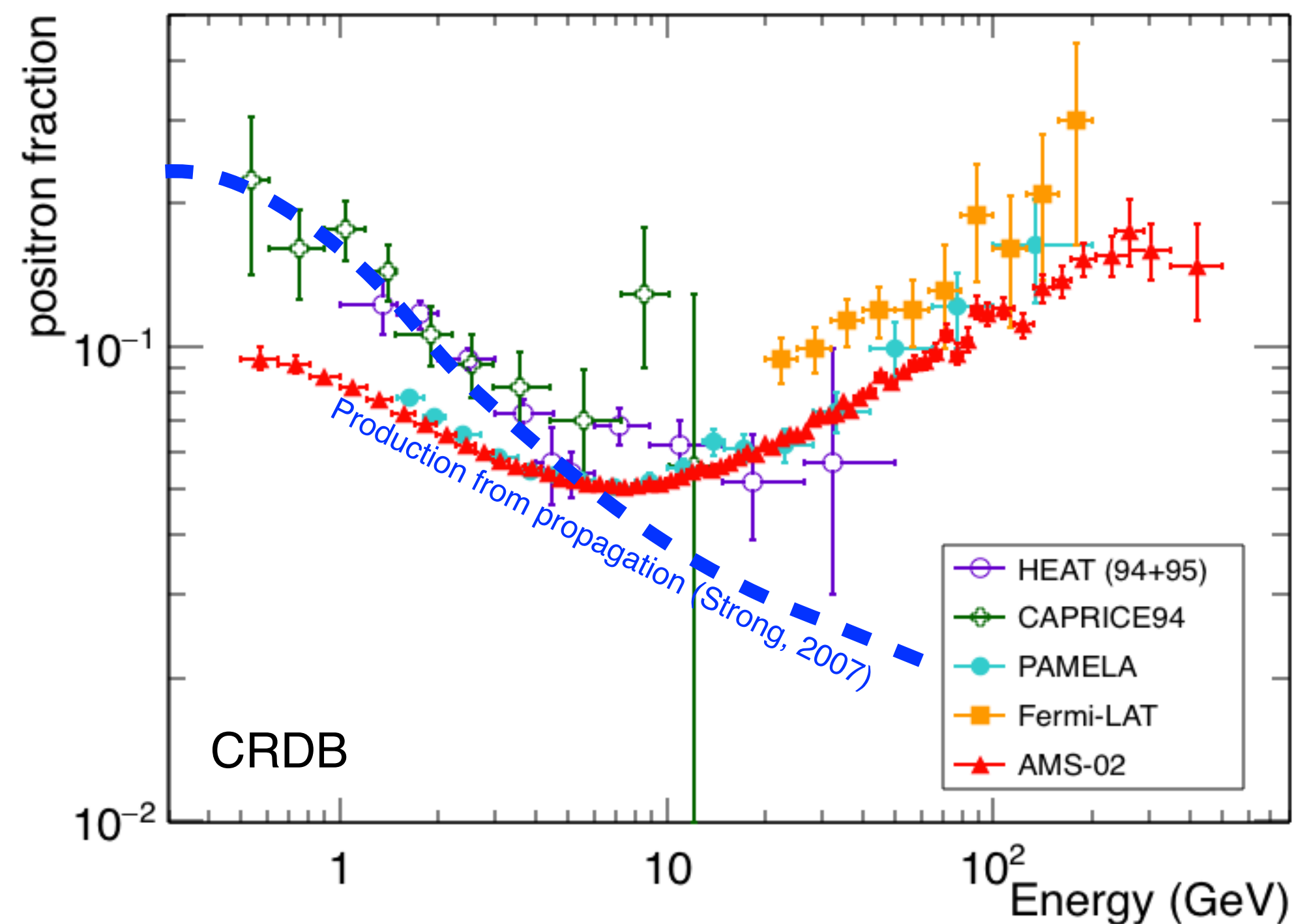
- Rising positron fraction
- Potentially rising anti-proton fraction
- Hardening at  $\sim 300$  GV in the spectra of H, He, Li, C, O, ...
- Different spectral index between proton and helium



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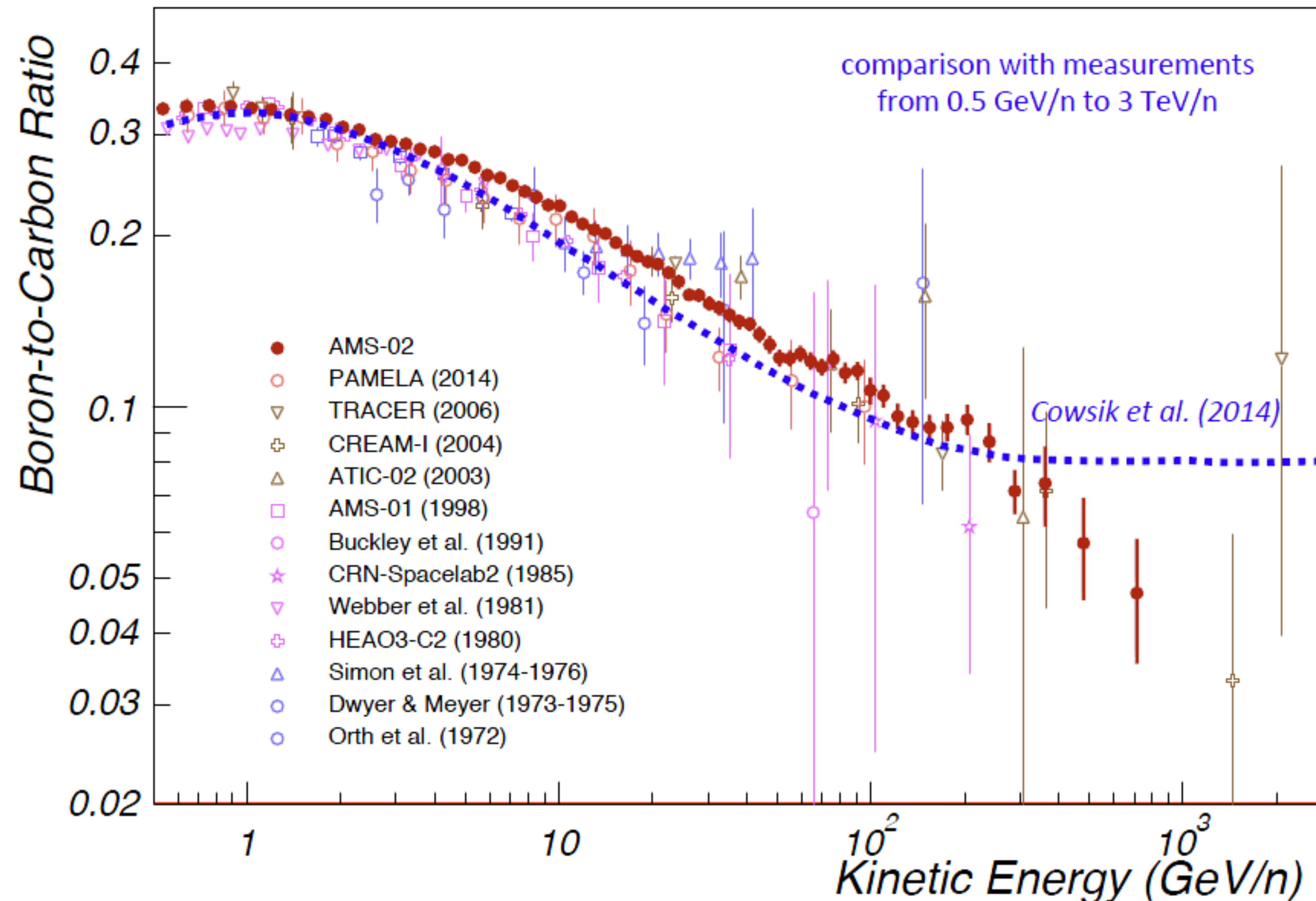
→ *It is critical to understand the propagation!*



# Secondary-to-Primary ratio

**Best measured observable to study the propagation: Secondary-to-primary ratio (e.g. B/C)**

- Sensitive to the amount of matter traversed by the CRs
  - Degeneracy between average amount of matter traversed and average life time





# Propagation Clock Isotope, $^{10}\text{Be}$

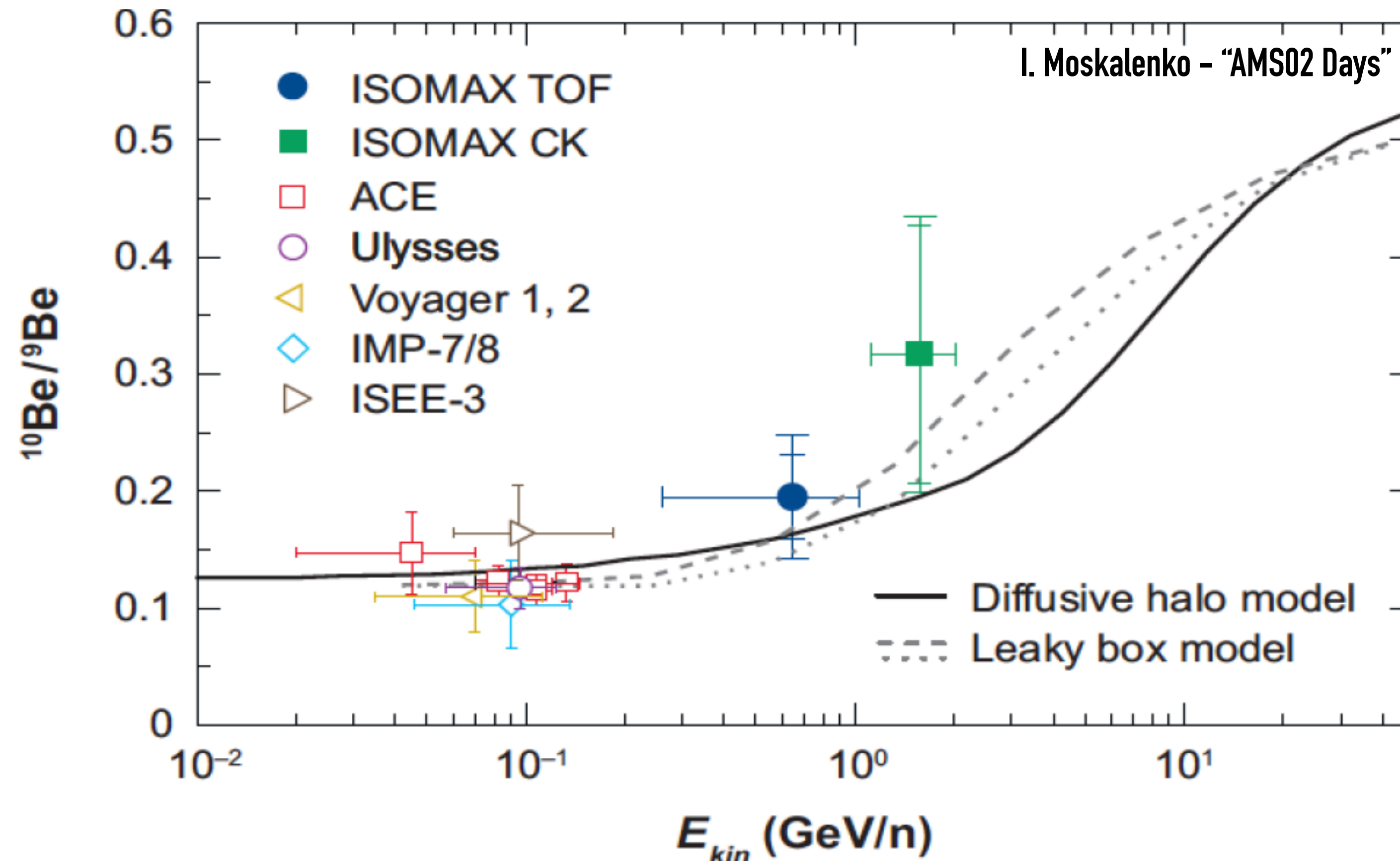
$^{10}\text{Be}$  : Unstable isotope w/ known half life of  $1.5 \times 10^6$  yr

- $^{10}\text{Be}/^9\text{Be}$  ratio provides strong constraints for the propagation models

- Good model discriminating power around 3 GeV/nuc

- Challenging measurements

★ Several good measurements at a few hundred MeV/nuc. Above this, the ISOMAX balloon payload covers up to  $\sim 2$  GeV/nuc





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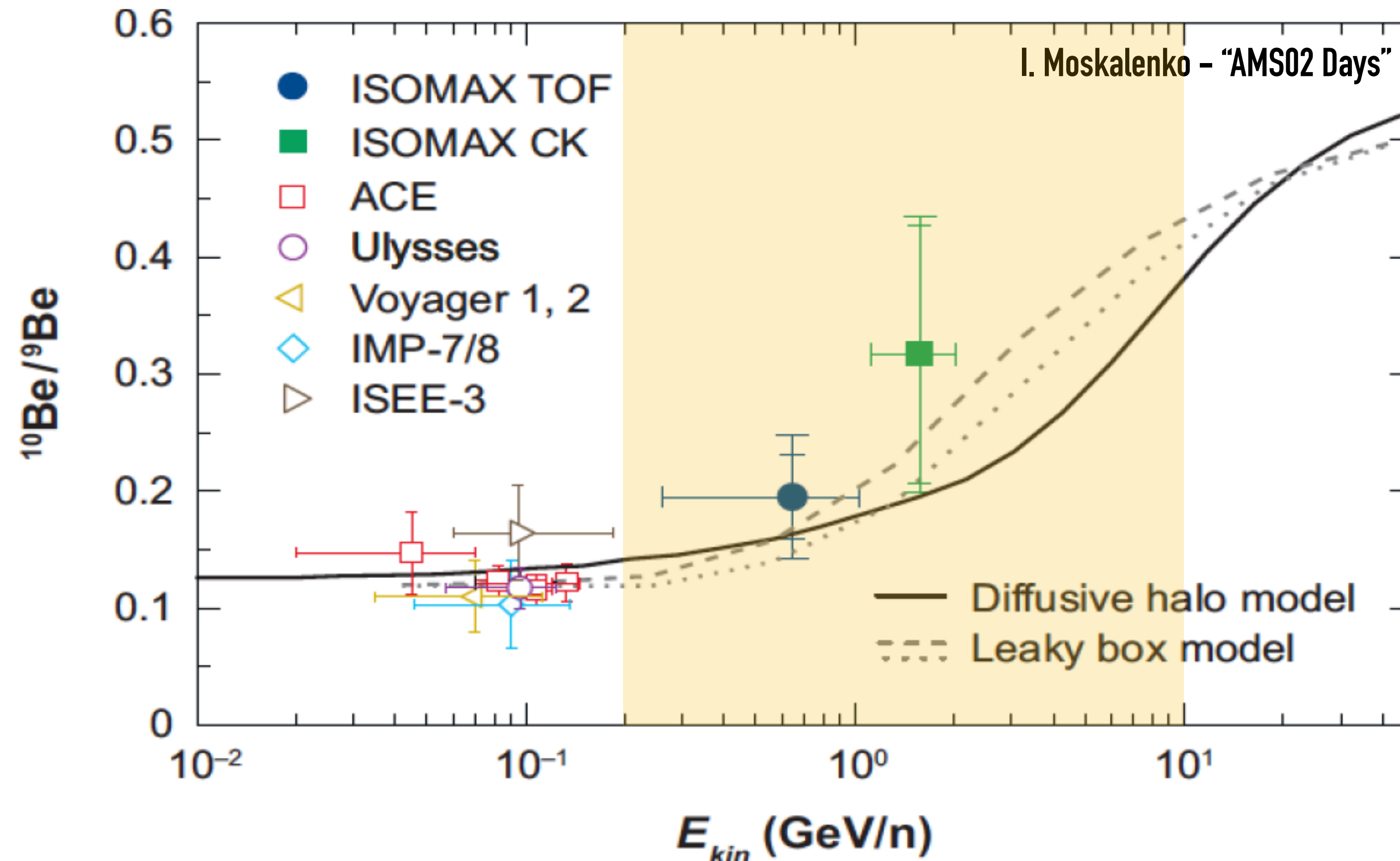
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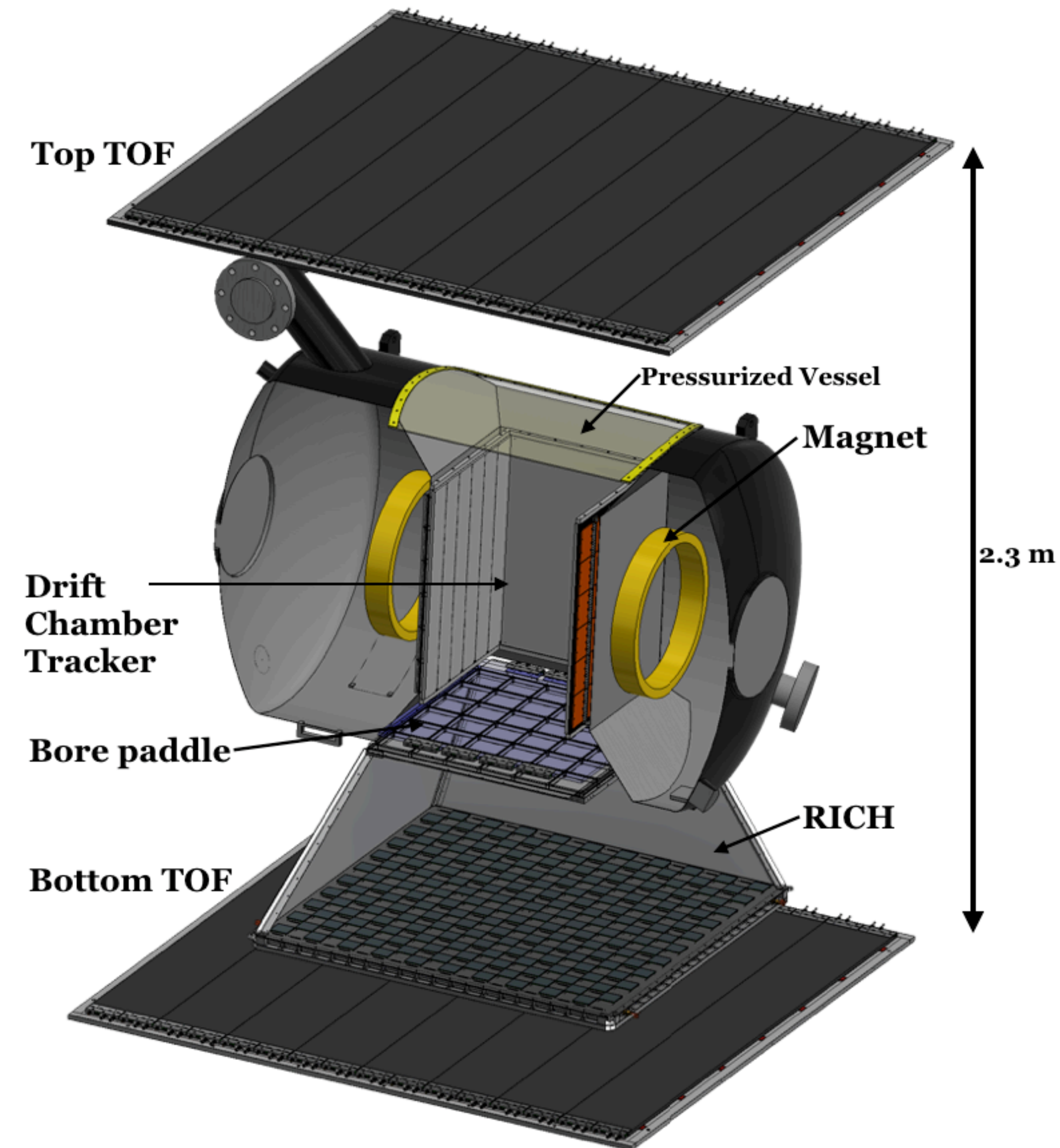


# High Energy Light Isotope eXperiment

A new magnet spectrometer payload to measure  $^{10}\text{Be}/^9\text{Be}$  isotope ratio up to 10 GeV/n

$$m = Ze R \frac{\sqrt{1 - \beta^2}}{\beta}$$

- Two stage approach to cover wider range of energy
- Stage 1 : covers up to  $\sim 3$  GeV/nuc, designed to have a flight in Antarctica with a long duration balloon in 2020



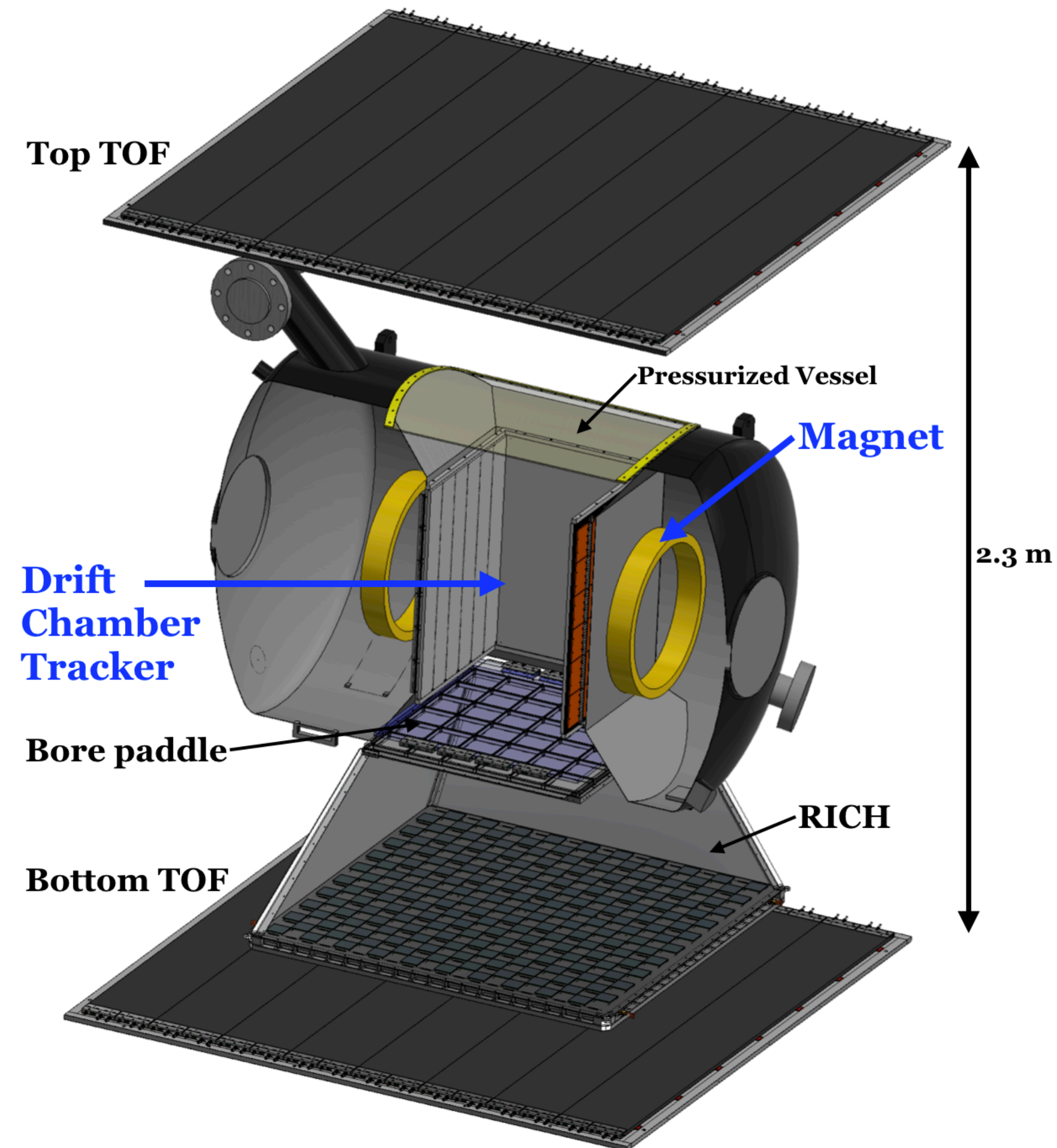


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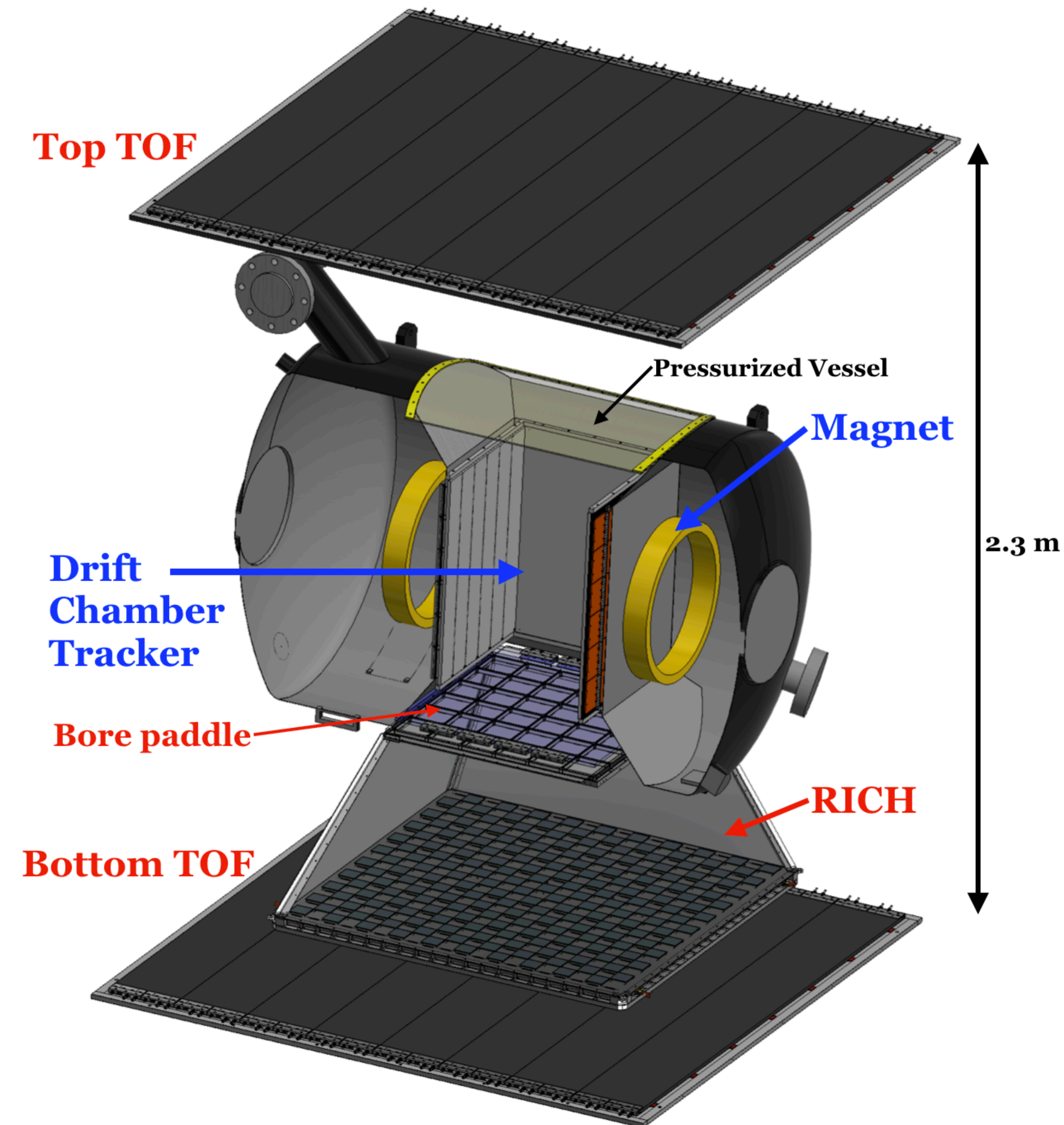


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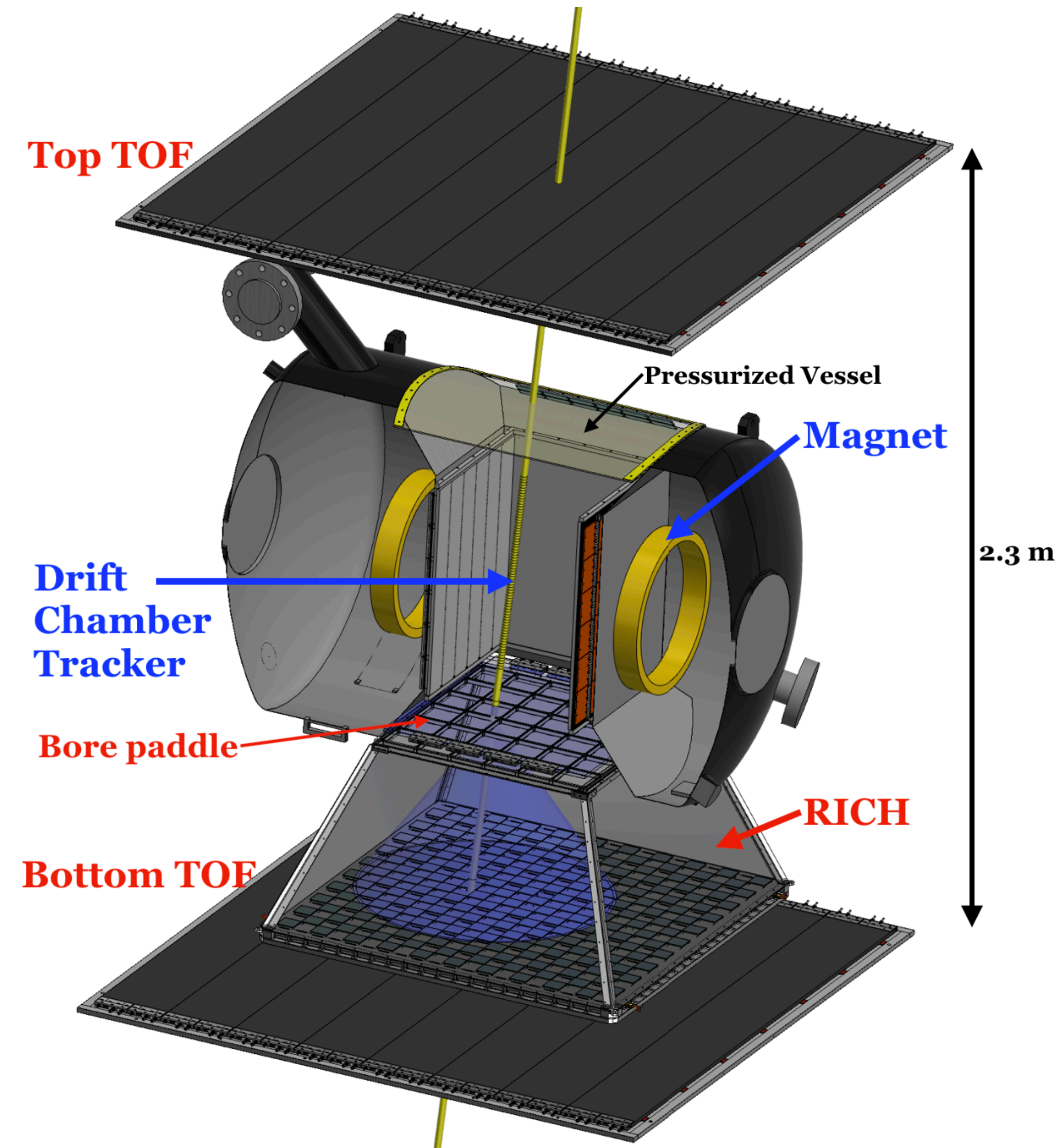


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- Two stage approach to cover wider range of energy
- Stage 1 : covers up to  $\sim 3$  GeV/nuc, designed to have a flight in Antarctica with a long duration balloon in 2020
- Very challenging measurements
  - ★ Mass resolution of few % up to 10 GeV/n
  - ★ Readout within a very strong magnetic field ( HEAT superconducting magnet, B field at the center  $\sim 1$  T )
  - ★ All SiPM readout needs good thermal design
  - ★ Total  $\sim 26\text{k}$  channels for full configuration

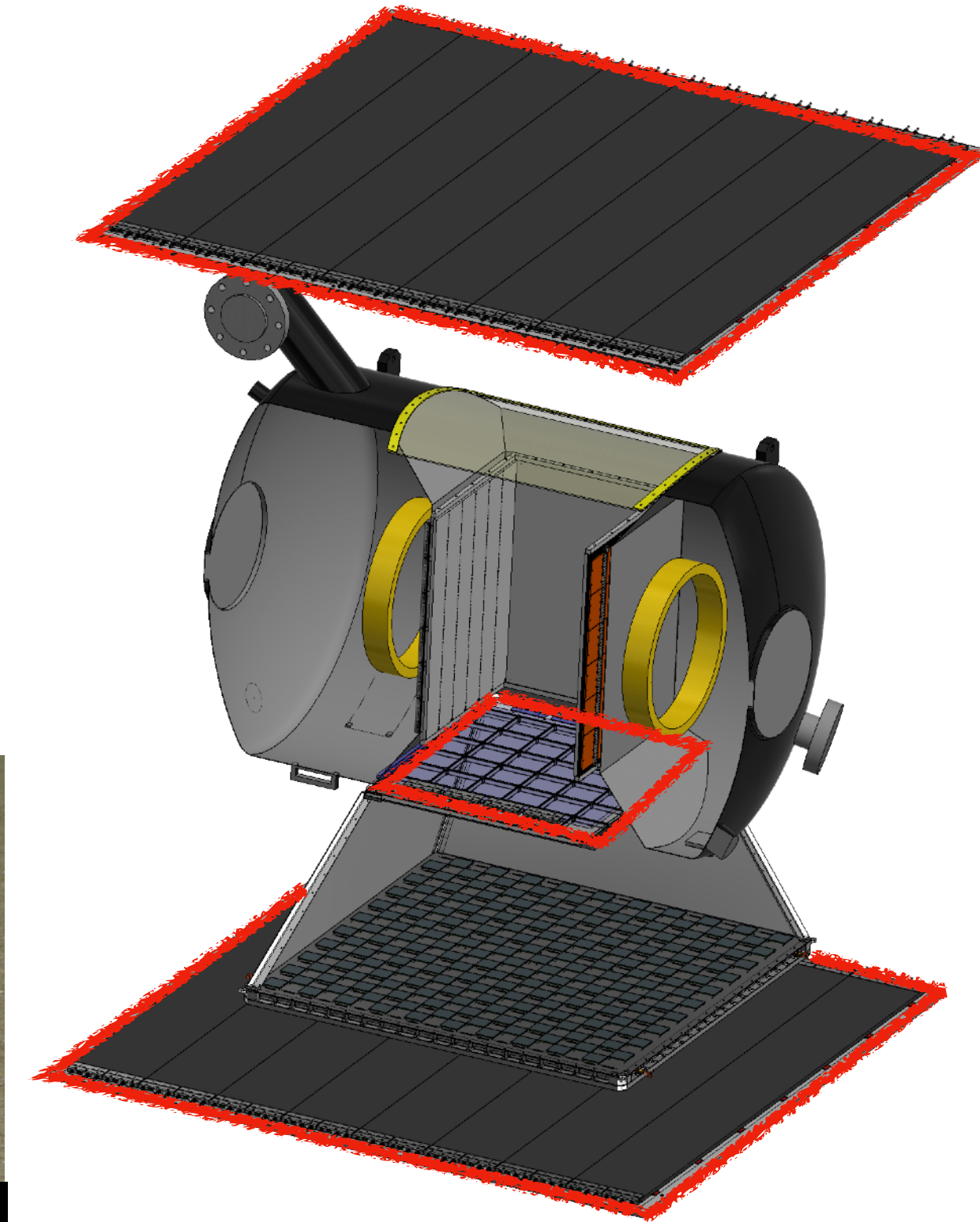
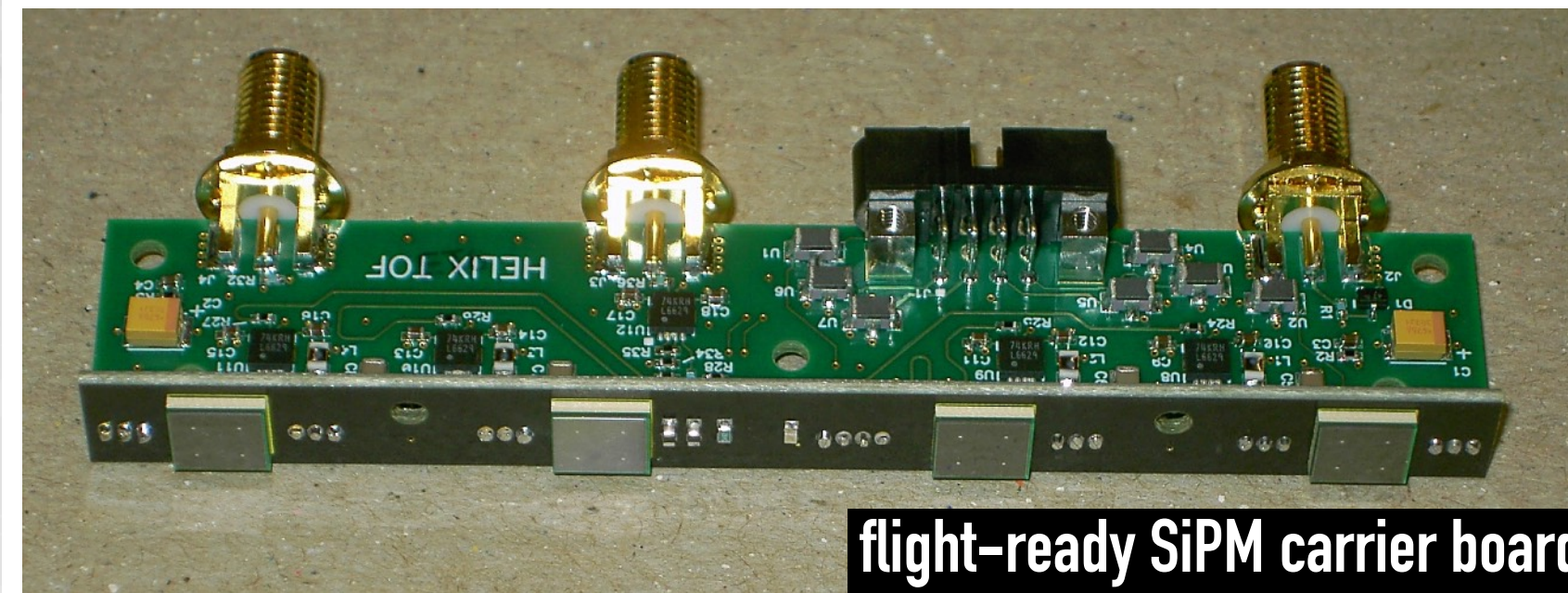
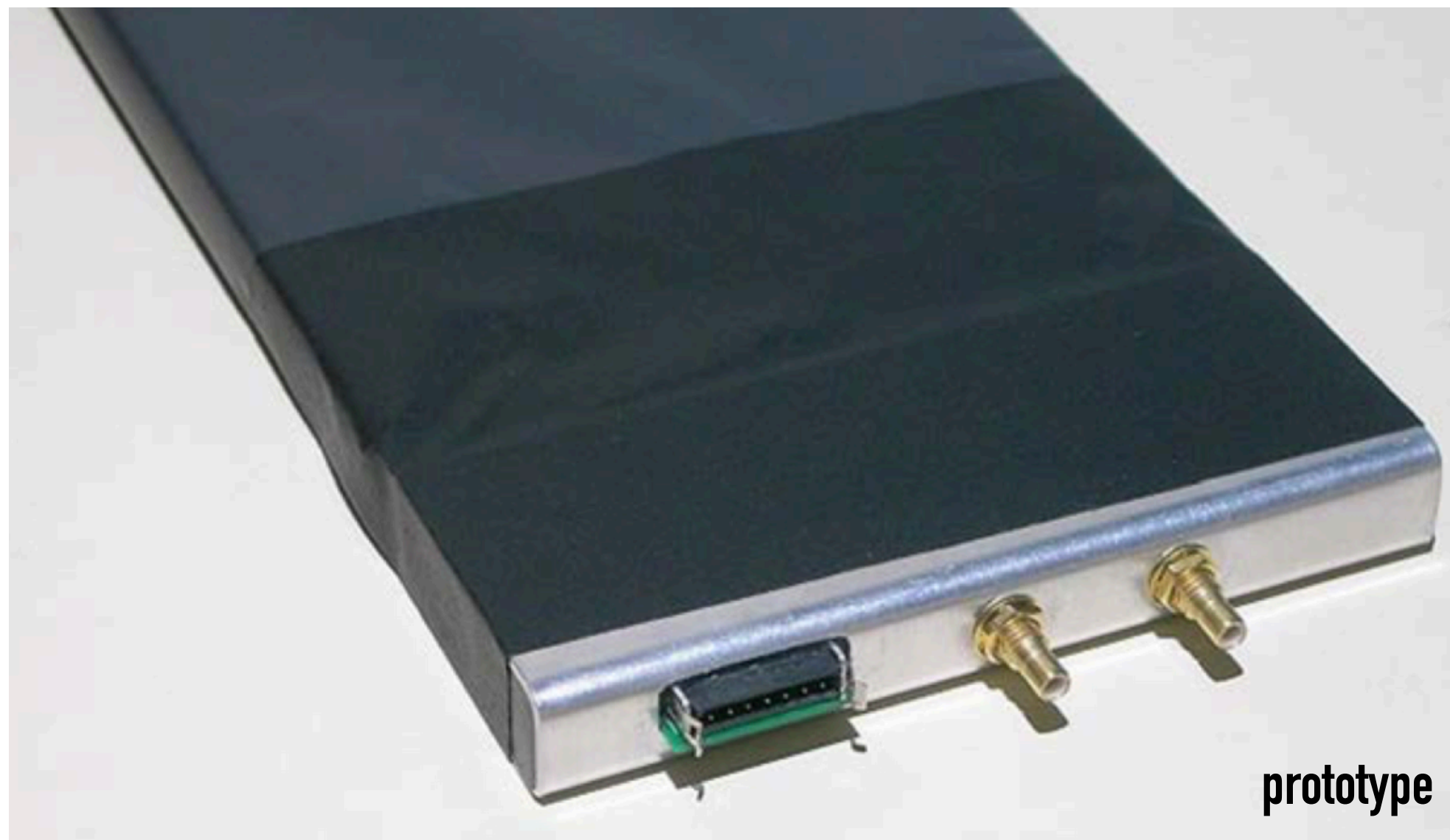




# Time-Of-Flight

**Three layers of 1 cm thickness fast plastic scintillator, L=2.3 m**

- Timing resolution of <50 ps timing resolution for  $Z > 3$
- Combination of three layers of TOF provide system-wide trigger



## Muon test with SiPMs

- Timing resolution of SiPM output was matching with expected performance from the simulation

**Ready for the paddle assembly**

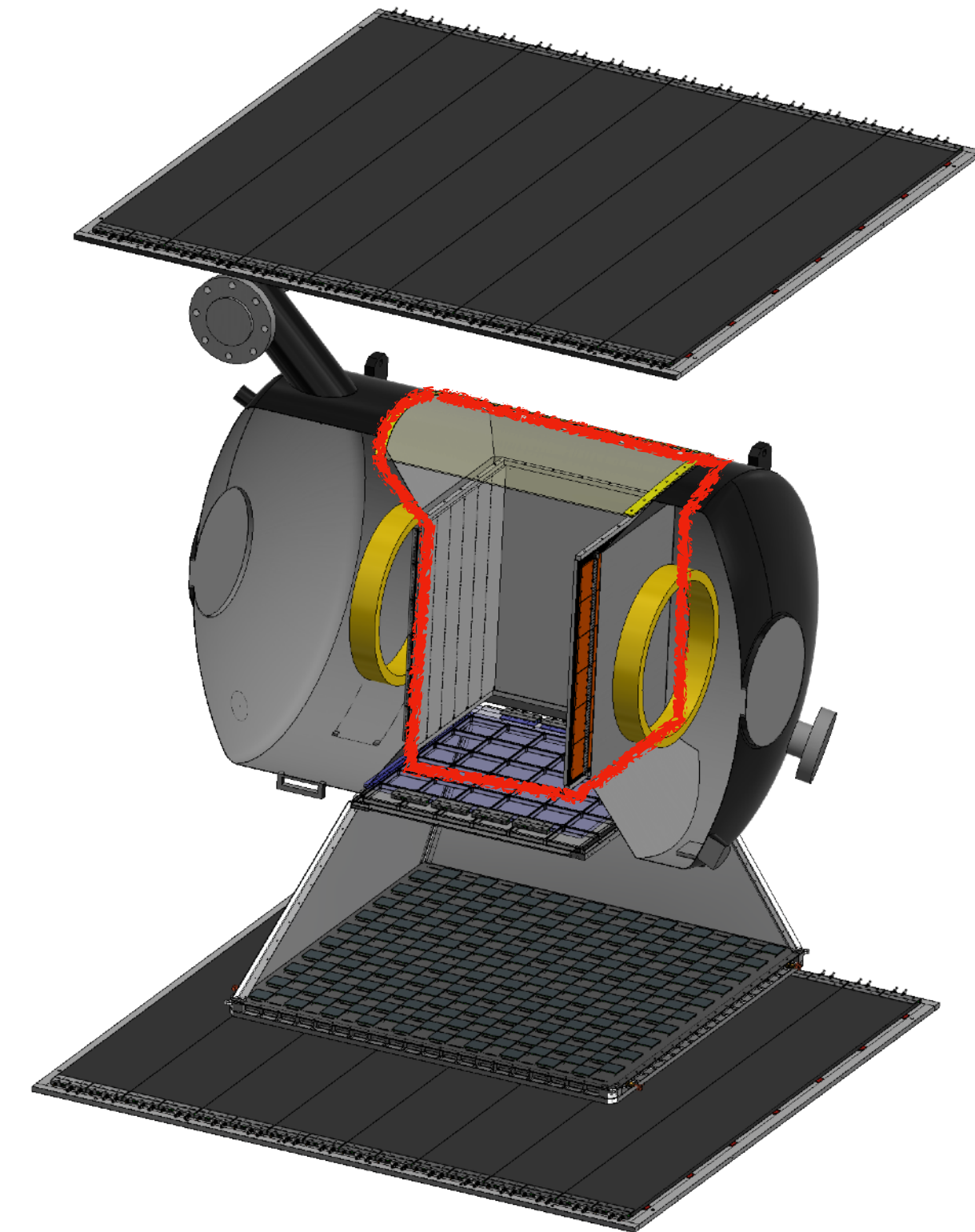
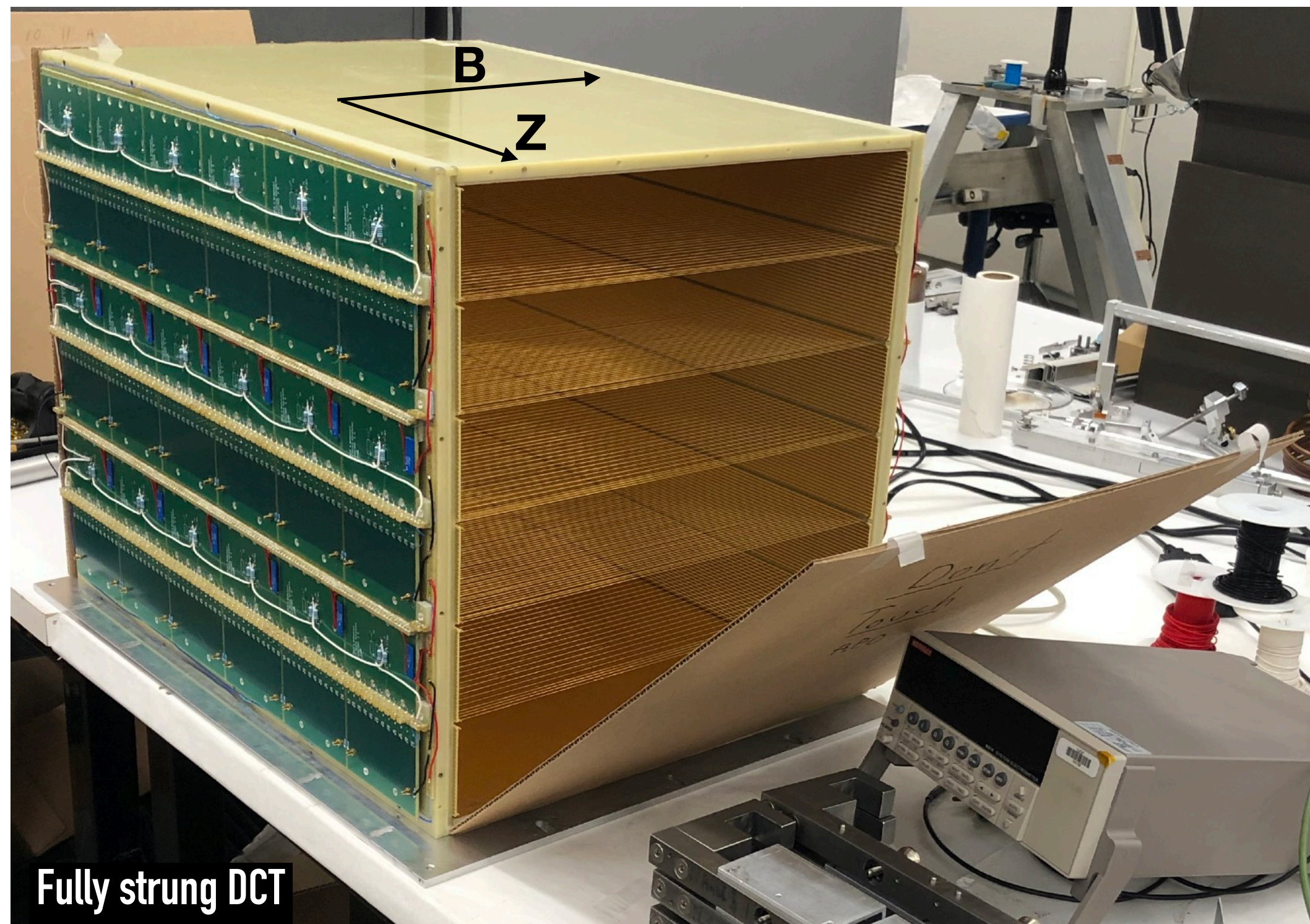


# Drift Chamber Tracker

## Multi-wire drift chamber with drift gas CO<sub>2</sub>

- Spatial resolution of 65  $\mu\text{m}$  for  $Z > 3$
- 72 sense layers, readout with high-speed sampling electronics
- Installed in the bore of magnet within a thin pressure vessel

## Wire fully strung & final test is underway





# Magnet

## 1T Superconducting magnet

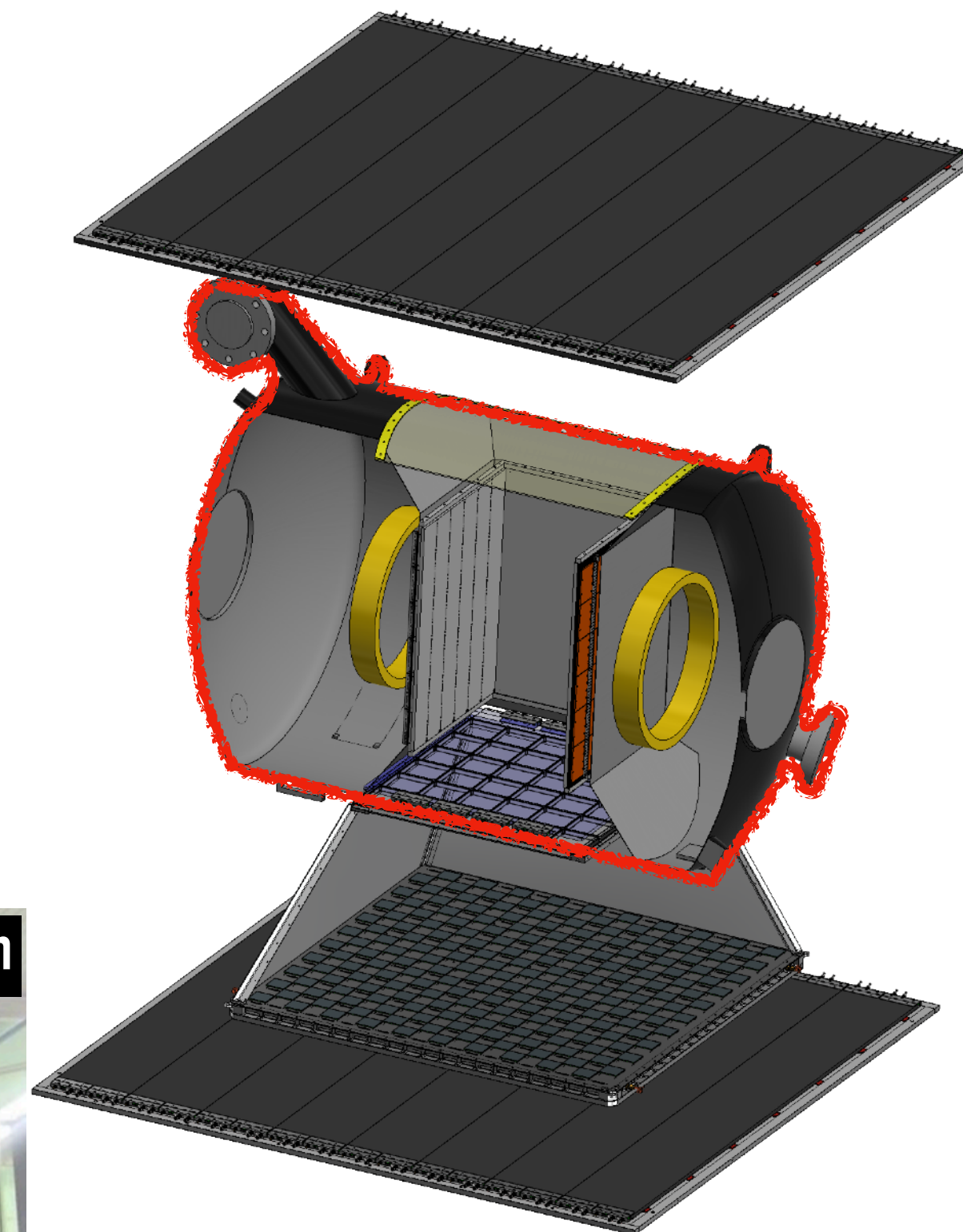
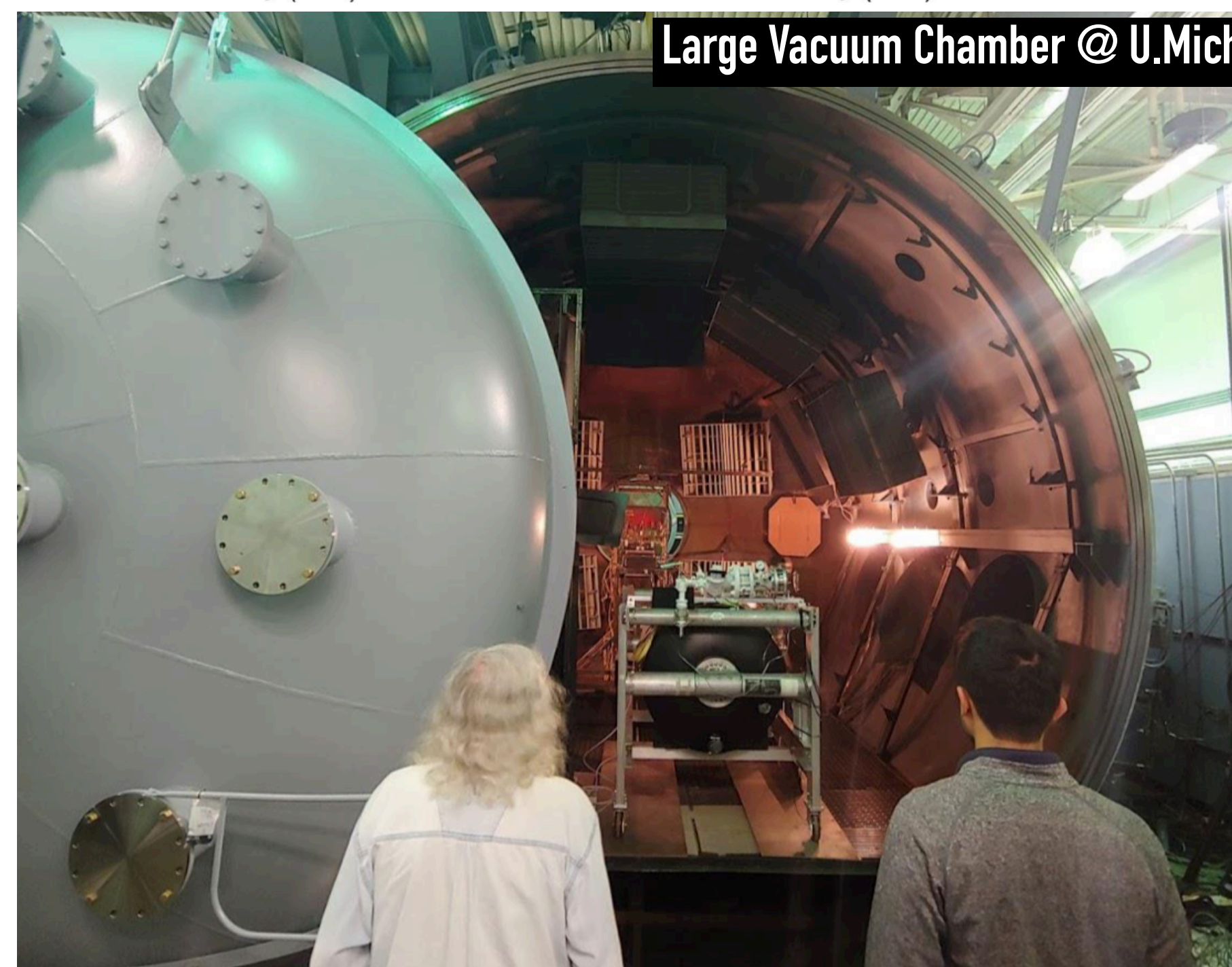
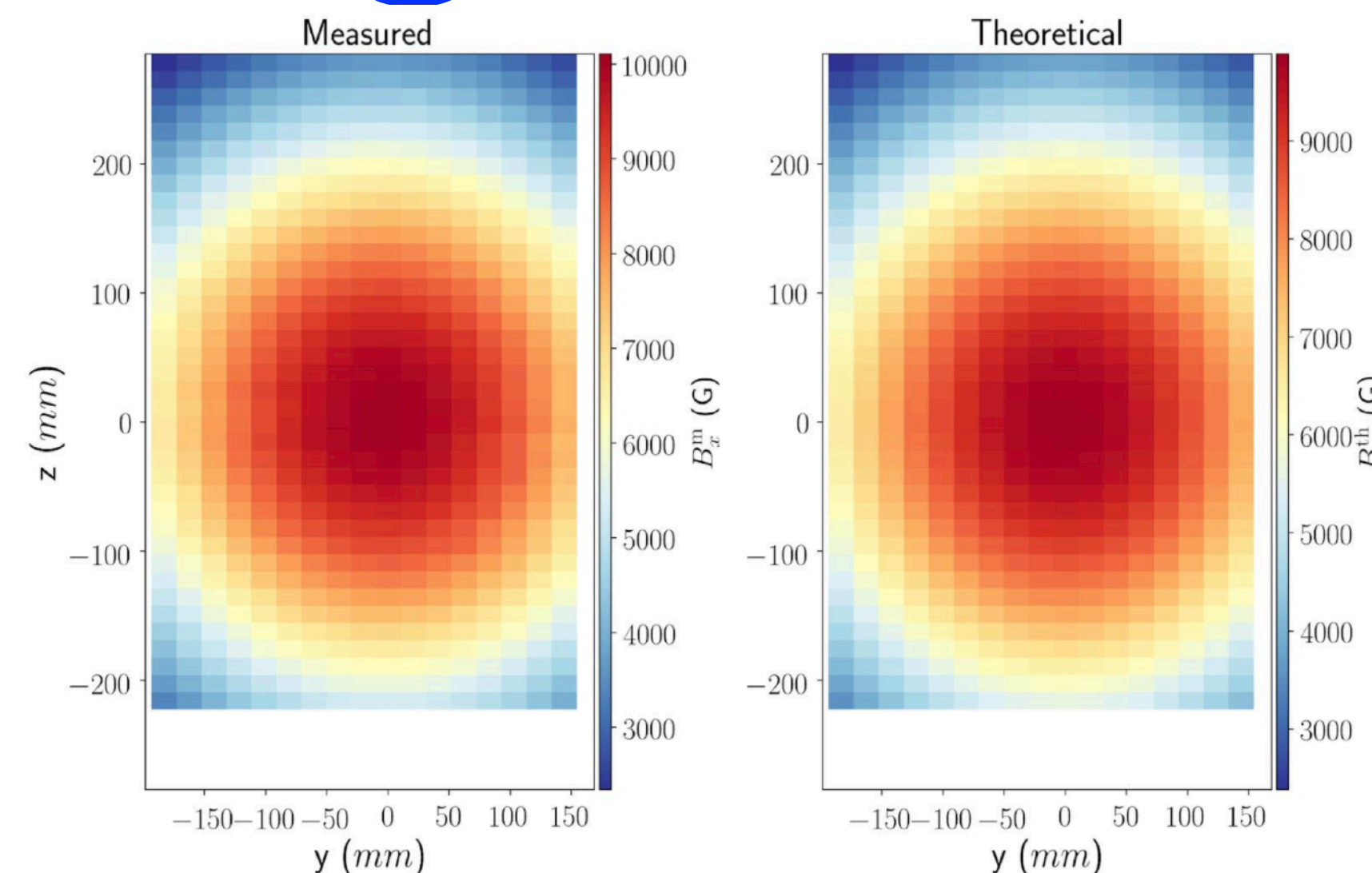
- Hold time : ~7 days
- Reused from the HEAT instrument
  - ★ Refurbished to operate the magnet without pressure vessel
- NbTi coils cooled to ~ 4.2 K

## 2 Successful cool down tests

- Measured detailed 3D magnetic field map
  - ★ Matching well with the theoretical model

## Successful vacuum test at Large Vacuum Chamber

- Successful operation at the flight vacuum condition (5 torr)

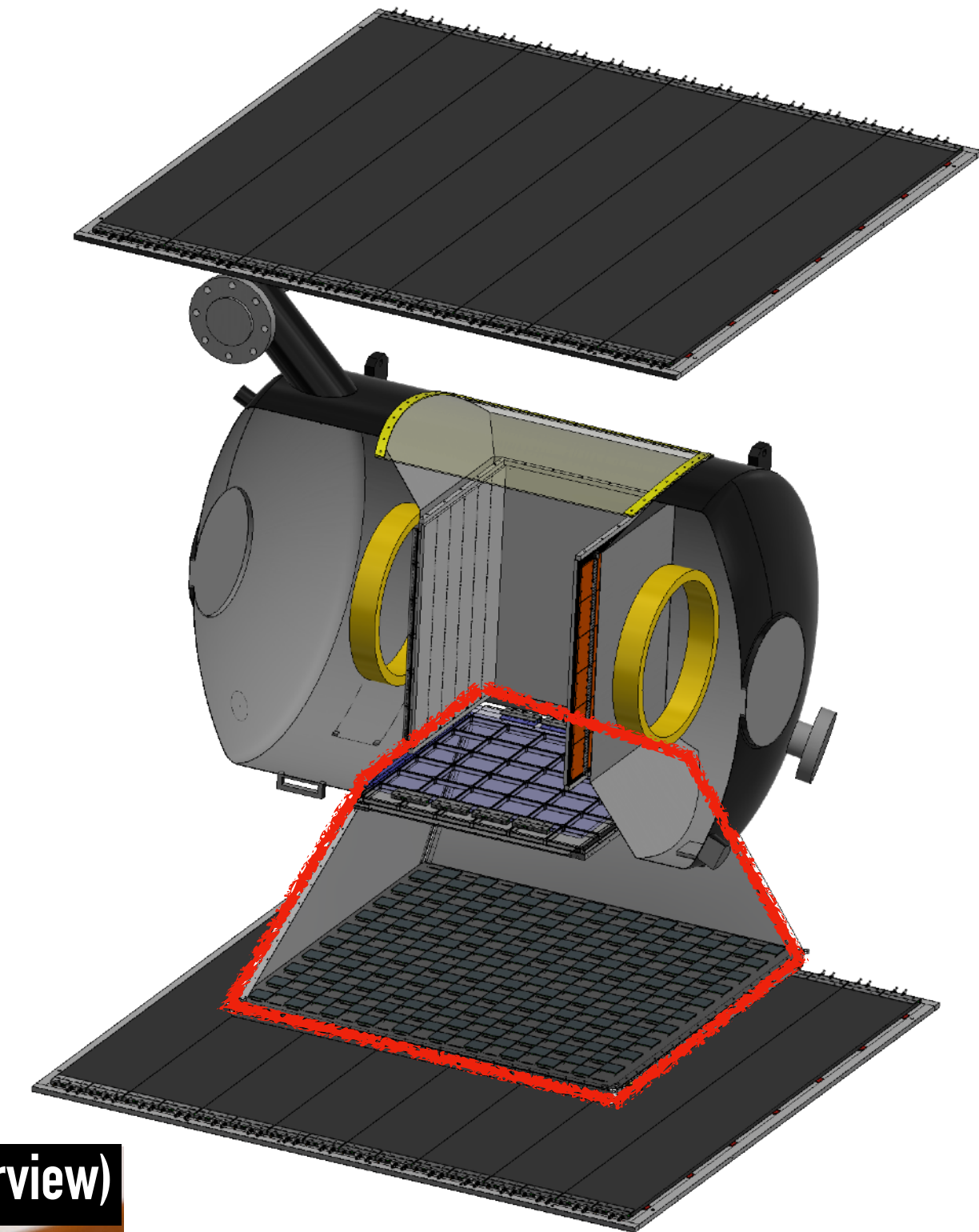




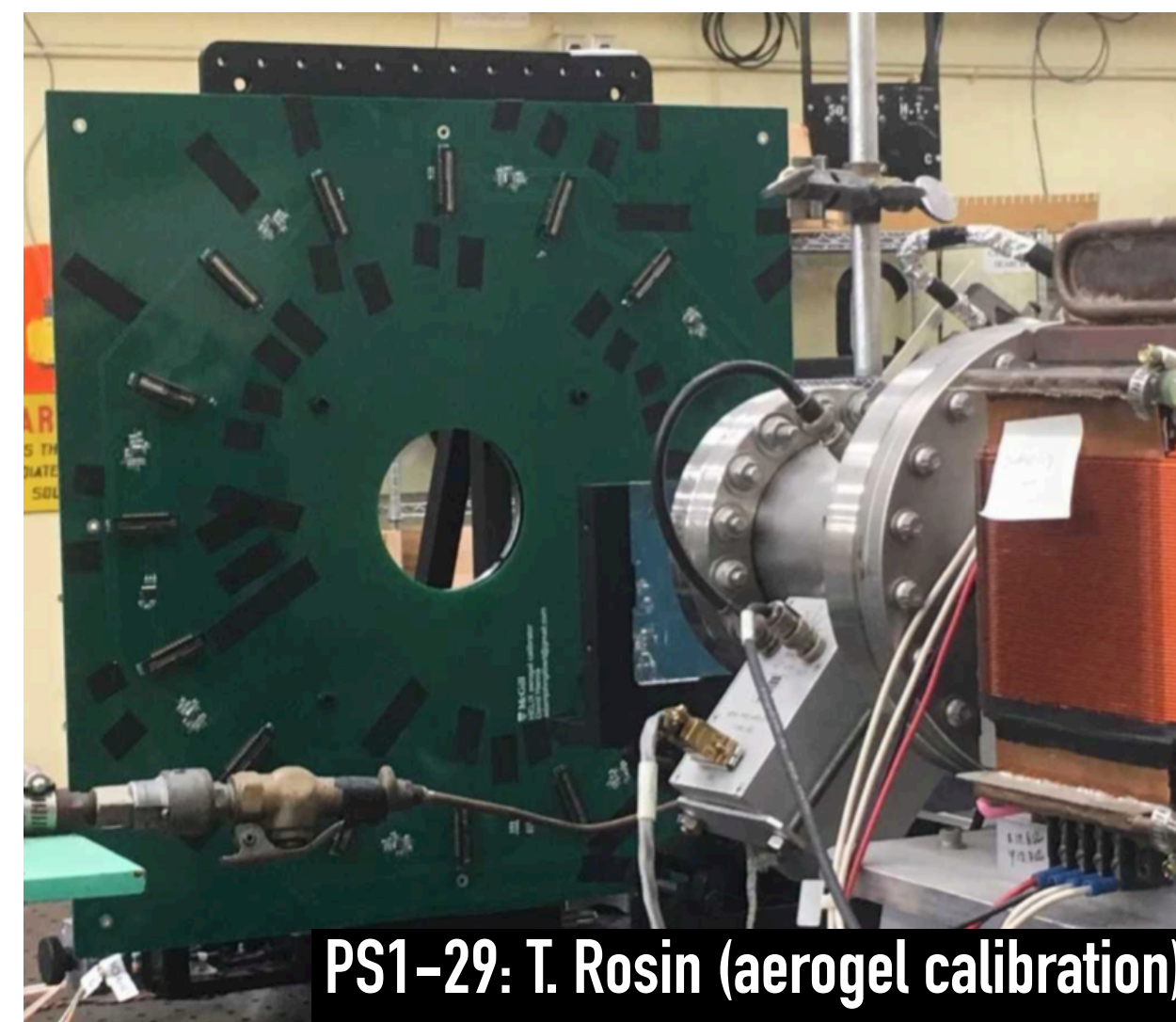
# Ring Imaging Cherenkov Counter

## Proximity-focused RICH with SiPM readout (→PS1-39: I. G. Wisher)

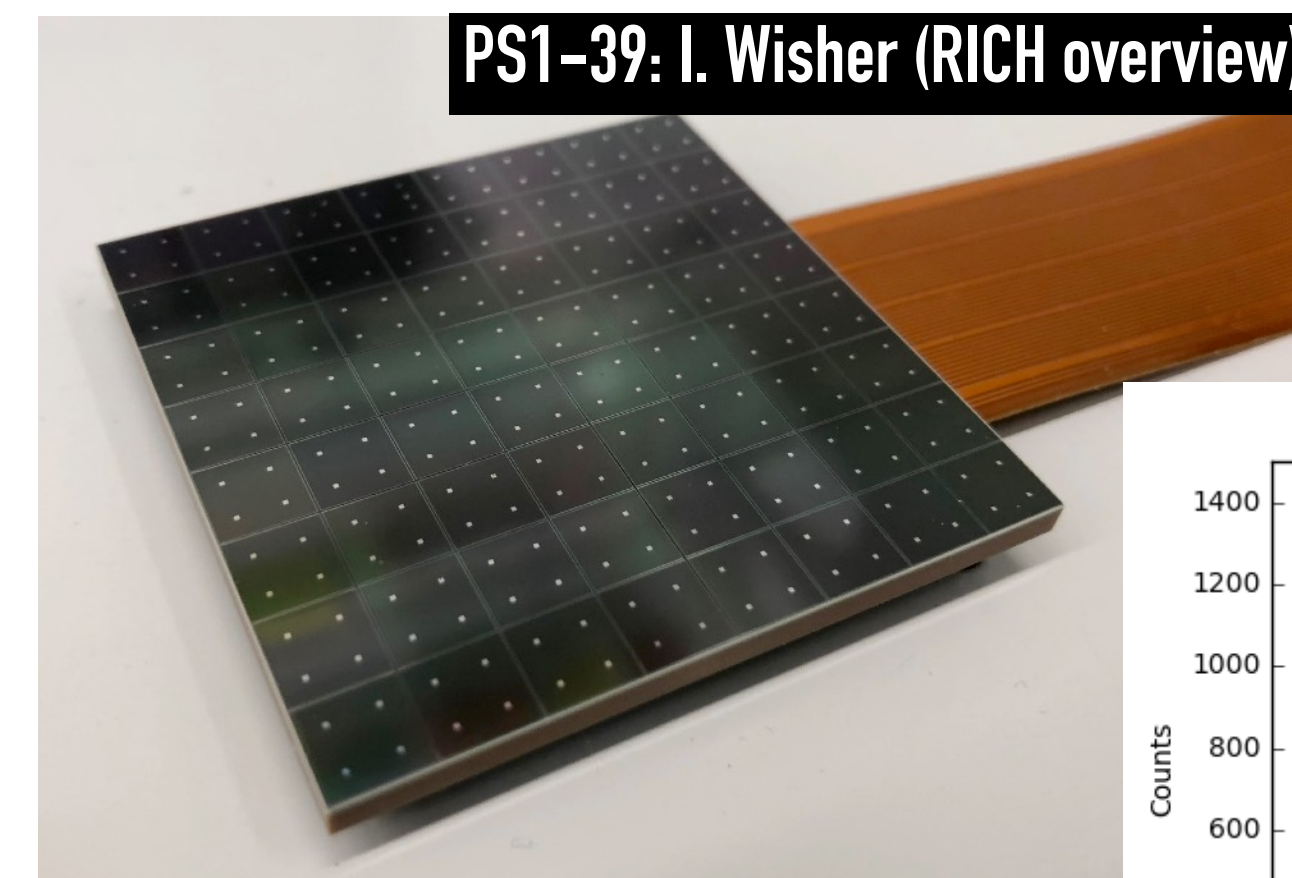
- Velocity resolution of  $\Delta\beta/\beta \sim 1 \times 10^{-3}$  for  $Z > 3$  for  $E > 1$  GeV/n
- Radiator : high refractive index aerogel ( $n \sim 1.15$ )
  - ★ Highly transparent & hydrophobic (→PS1-33: M. Tabata)
  - ★ Refractive index calibration w/ systematic error of  $10^{-4}$  level (→PS1-29: T. Rosin)
- Focal plane
  - ★  $1 \text{ m} \times 1 \text{ m}$  focal plane covered by Hamamatsu SiPM array (half-filled)
    - Single p.e. detectability
    - Thermal plate underneath to provide cool temperature



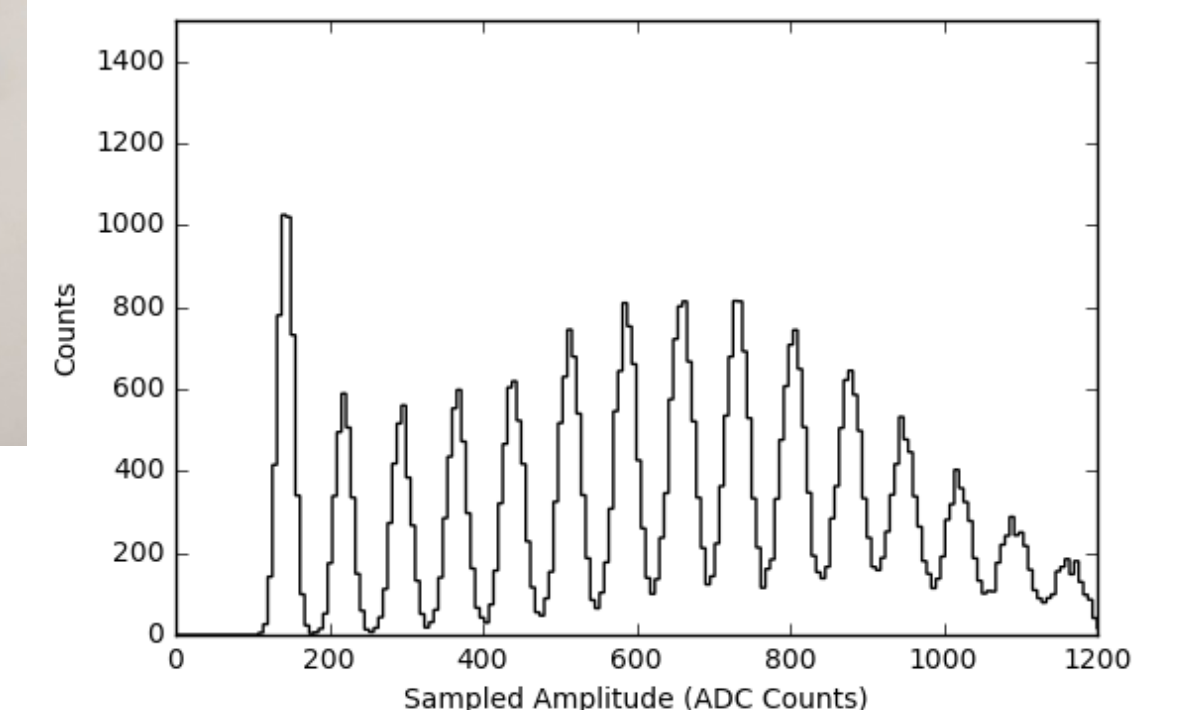
PS1-33: M. Tabata (aerogel fabrication)



PS1-29: T. Rosin (aerogel calibration)



PS1-39: I. Wisher (RICH overview)

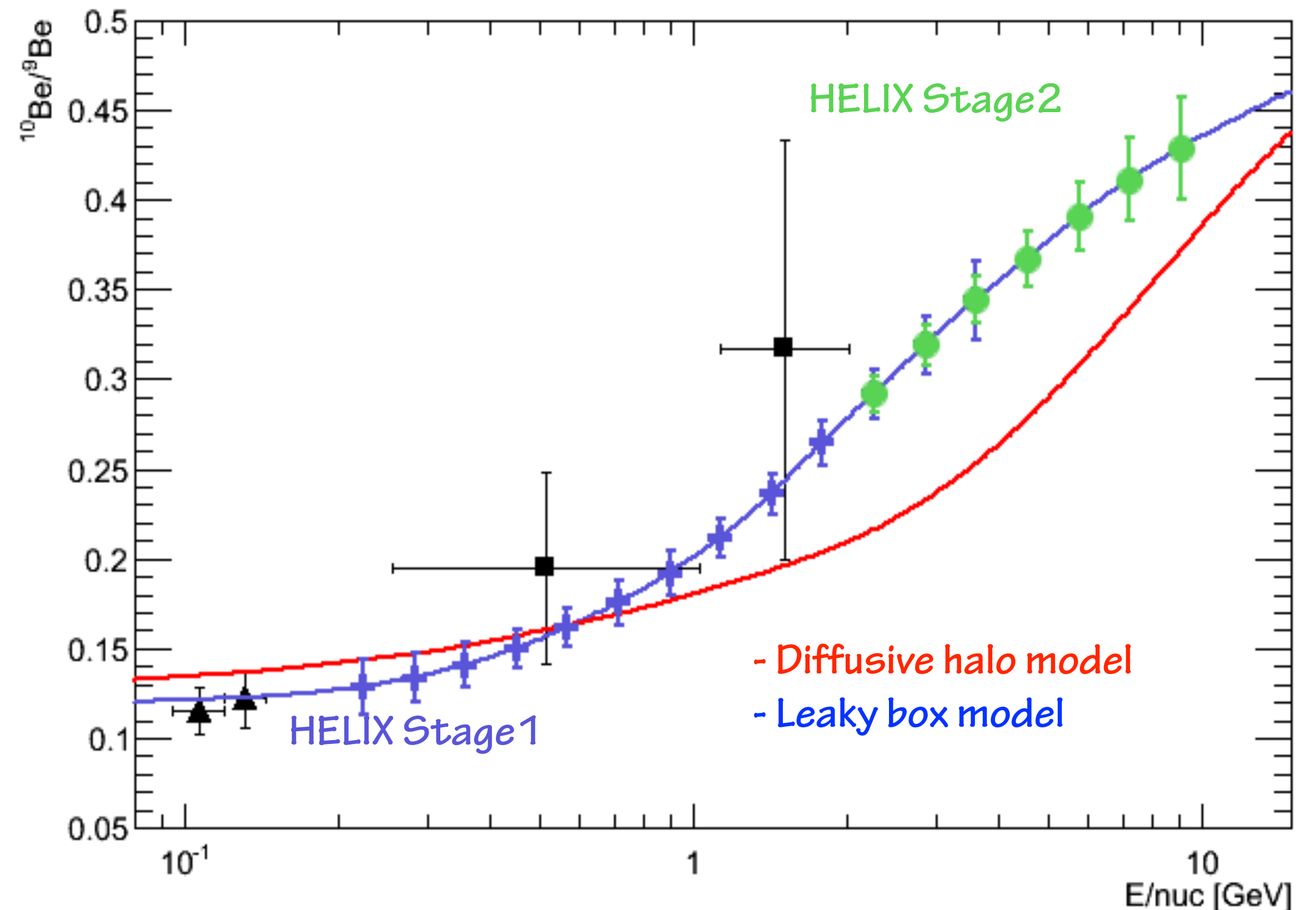
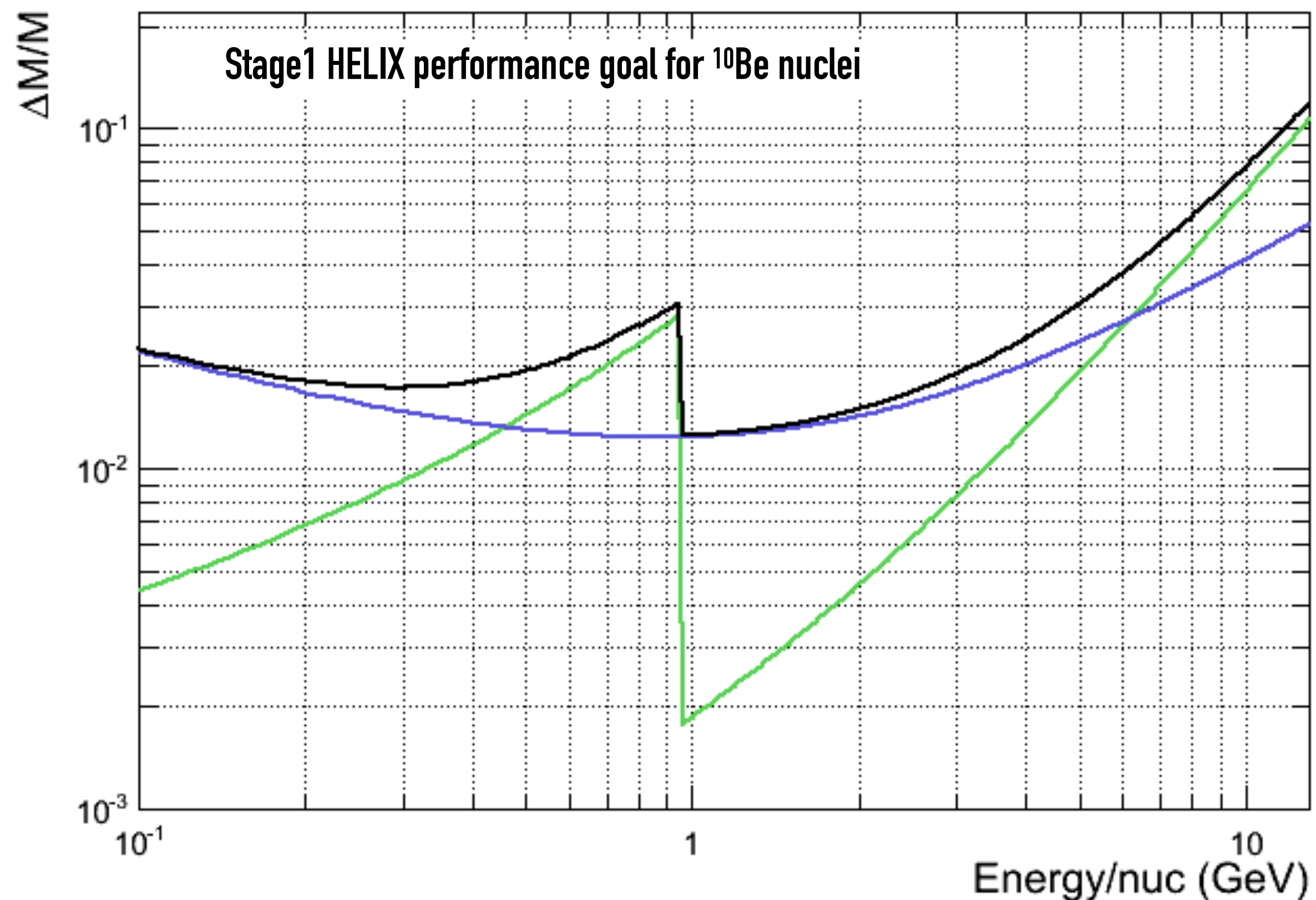




# HELIX Stage1 Performance

**$^{10}\text{Be}/^9\text{Be}$  ratio up to  $\sim 3 \text{ GeV/n}$  with  $\Delta m/m \sim 2.5\%$**

- 7-14 day exposure with  $0.1 \text{ m}^2\text{sr}$  geometry factor
- Measure the charge of CR up to neon ( $Z=10$ )
- Mass resolution of few percentage for light isotopes up to  $3 \text{ GeV/n}$



# Summary

**HELIX is moving forward to be ready for integration test in 2019, and an Antarctica flight in 2020!**

Recent discoveries of new features of CRs require better understanding of CR propagation.

Measurement of propagation clock isotope, such as  $^{10}\text{Be}$  can provide essential data.

HELIX is a balloon-borne experiment designed to measure  $^{10}\text{Be}$  from 0.2 GeV to beyond 3 GeV/n.

All of the sub detectors are under construction.

