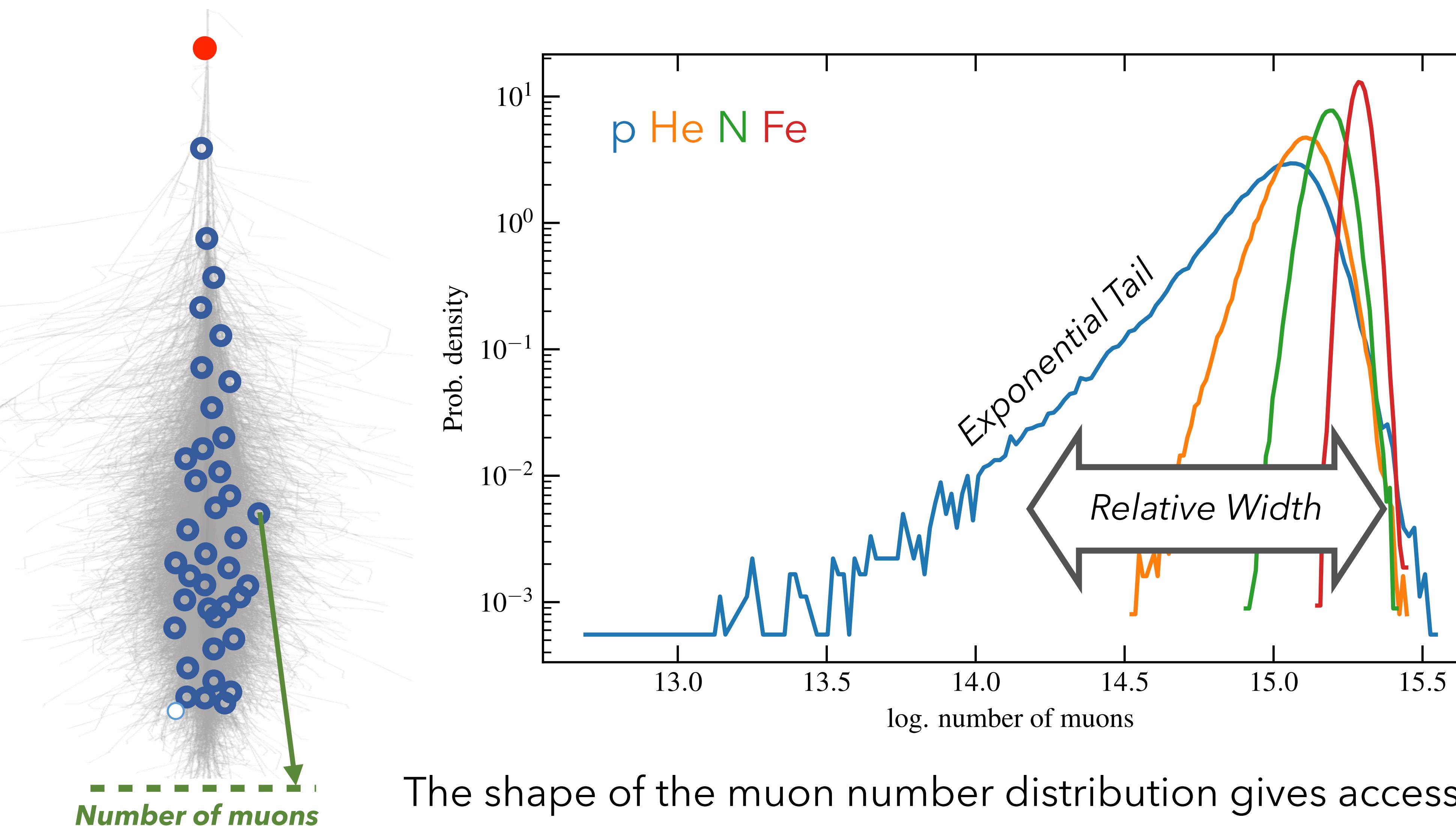


Probing the high-energy spectrum of neutral pions in ultra-high-energy proton-air interaction

Lorenzo Cazon, Ruben Conceição, Felix Riehn, Miguel Martins

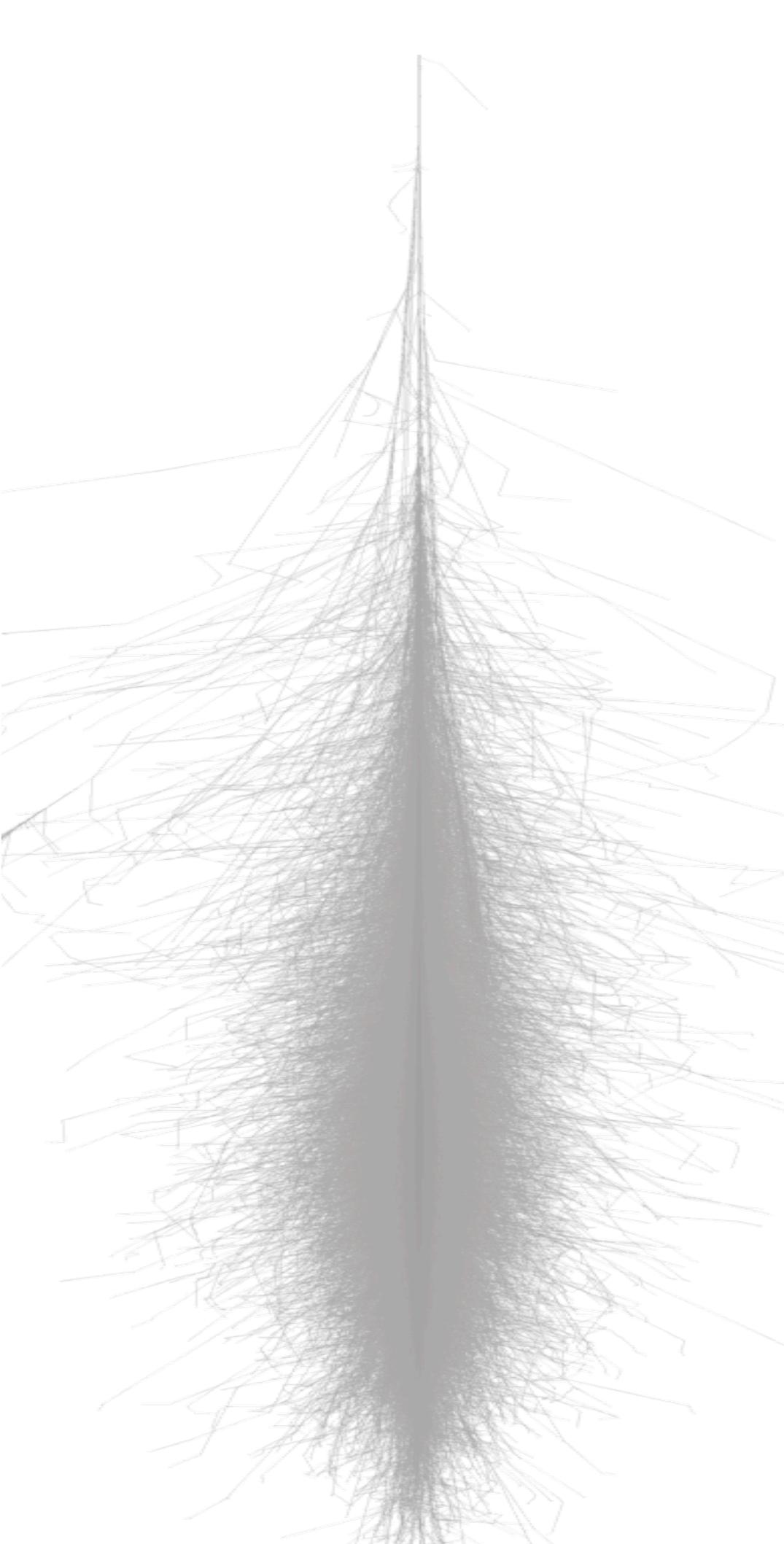


Conclusion

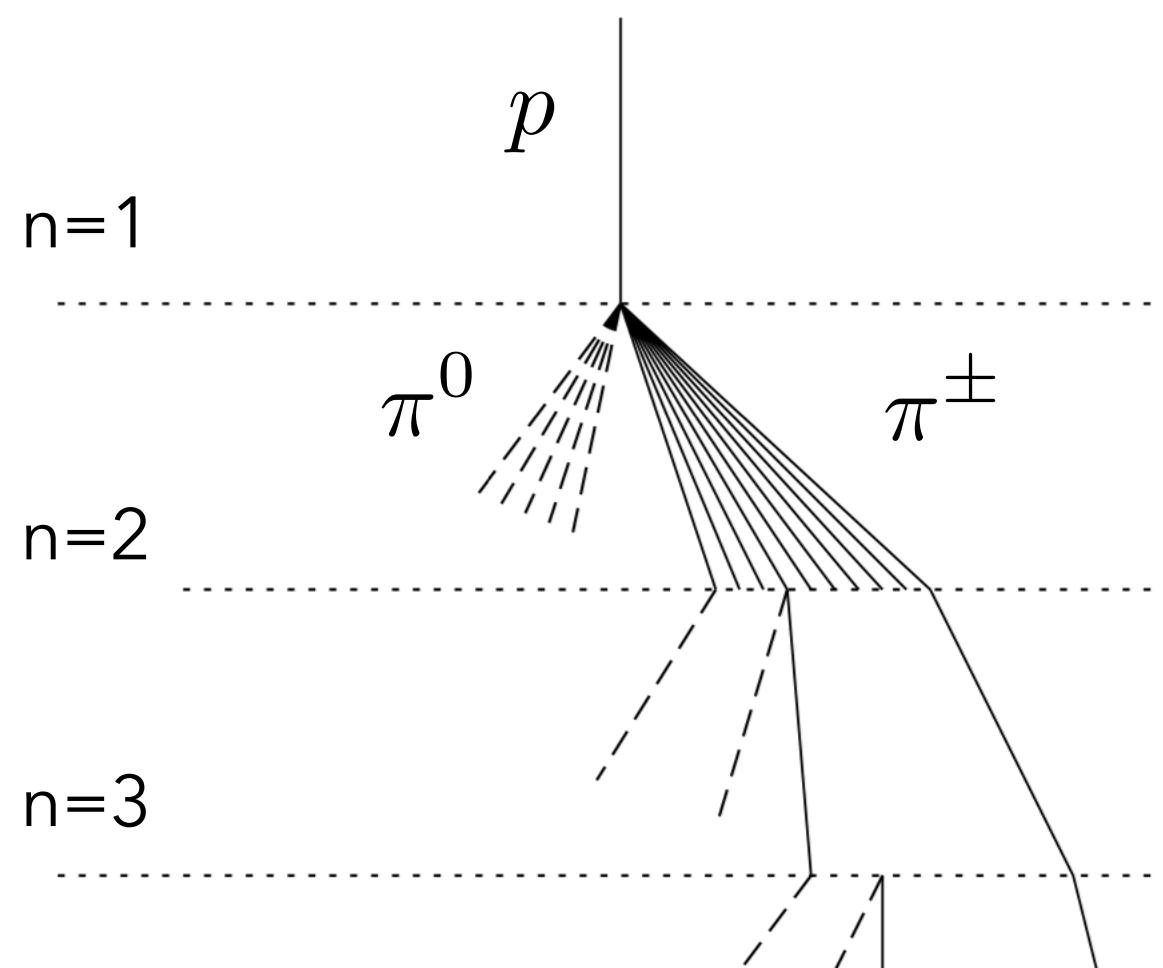


The shape of the muon number distribution gives access to the properties of the **FIRST hadronic interaction**

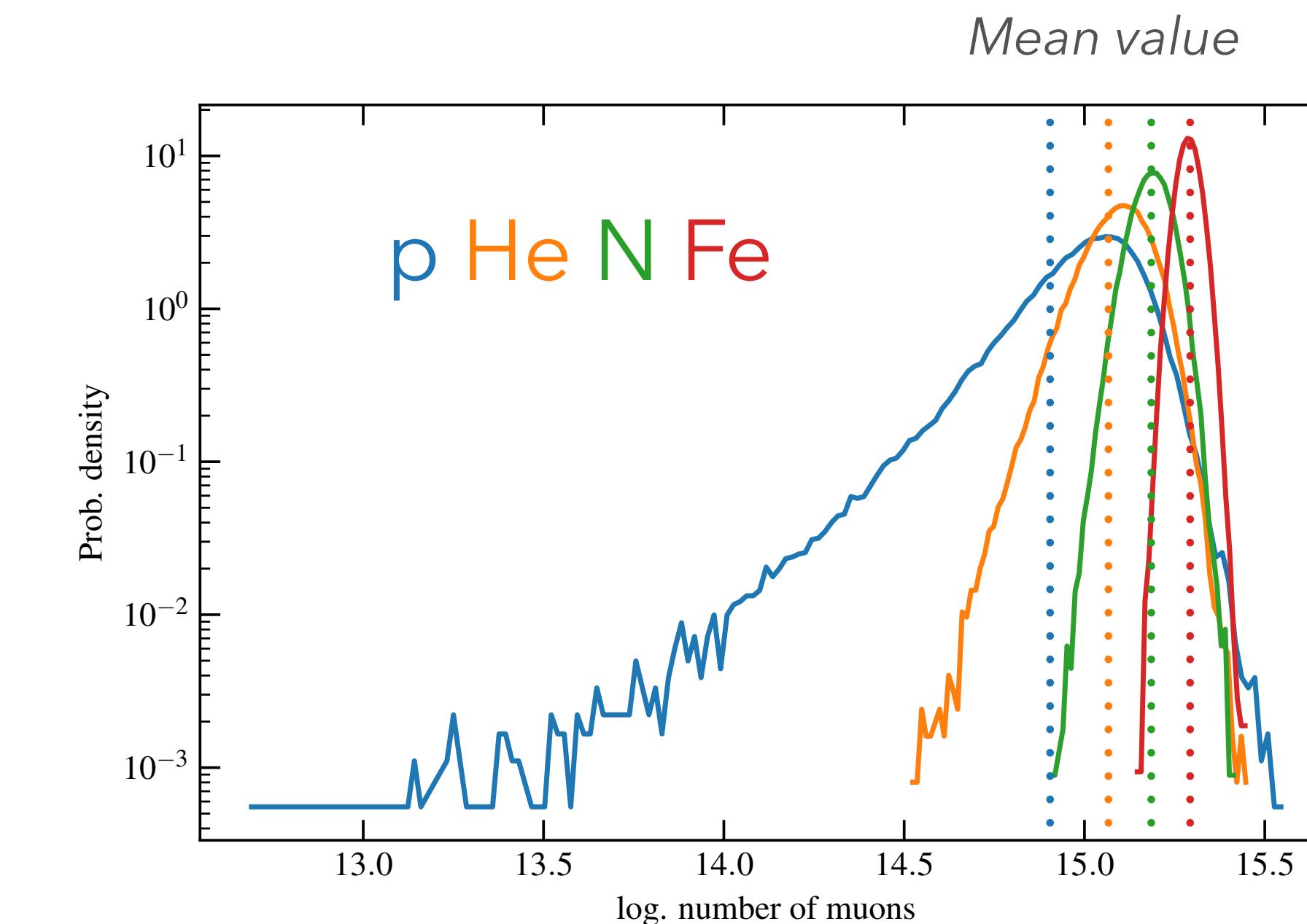
Origin of the fluctuations of the muon number distribution



Heitler-Matthews model



$$\langle N_\mu \rangle \sim E^\beta$$



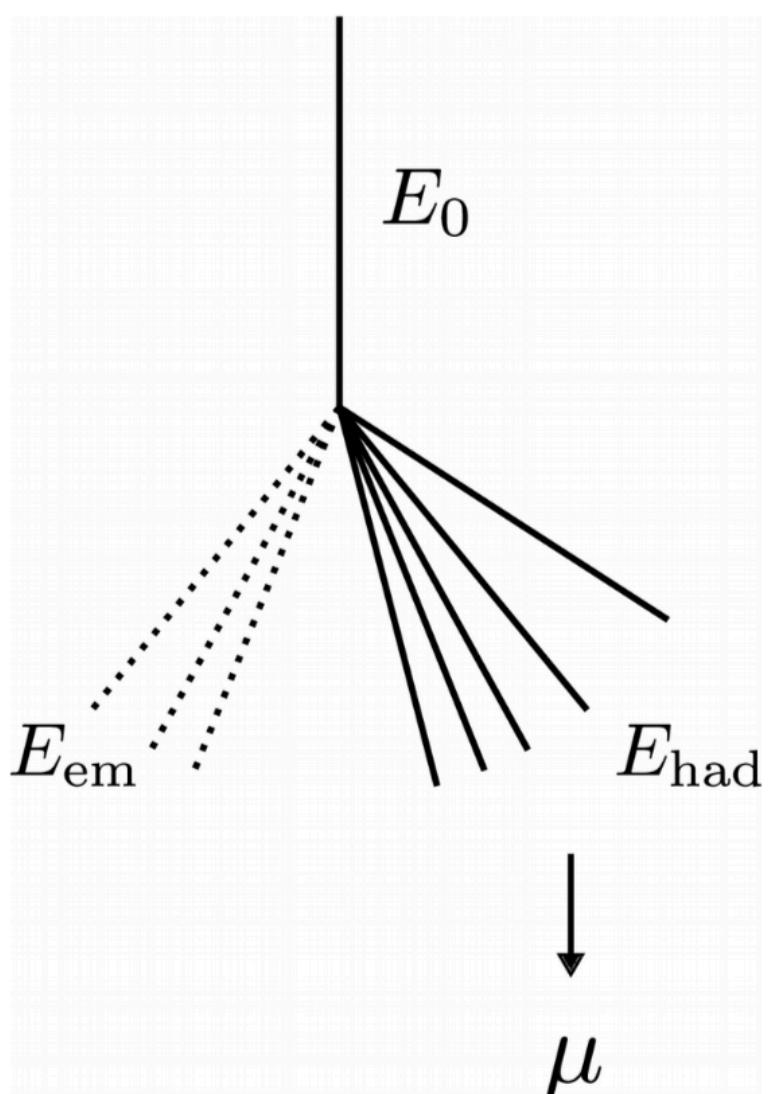
How to explain the higher moments
of the N_μ distribution?

Origin of the fluctuations of the muon number distribution

Heitler-Matthews model

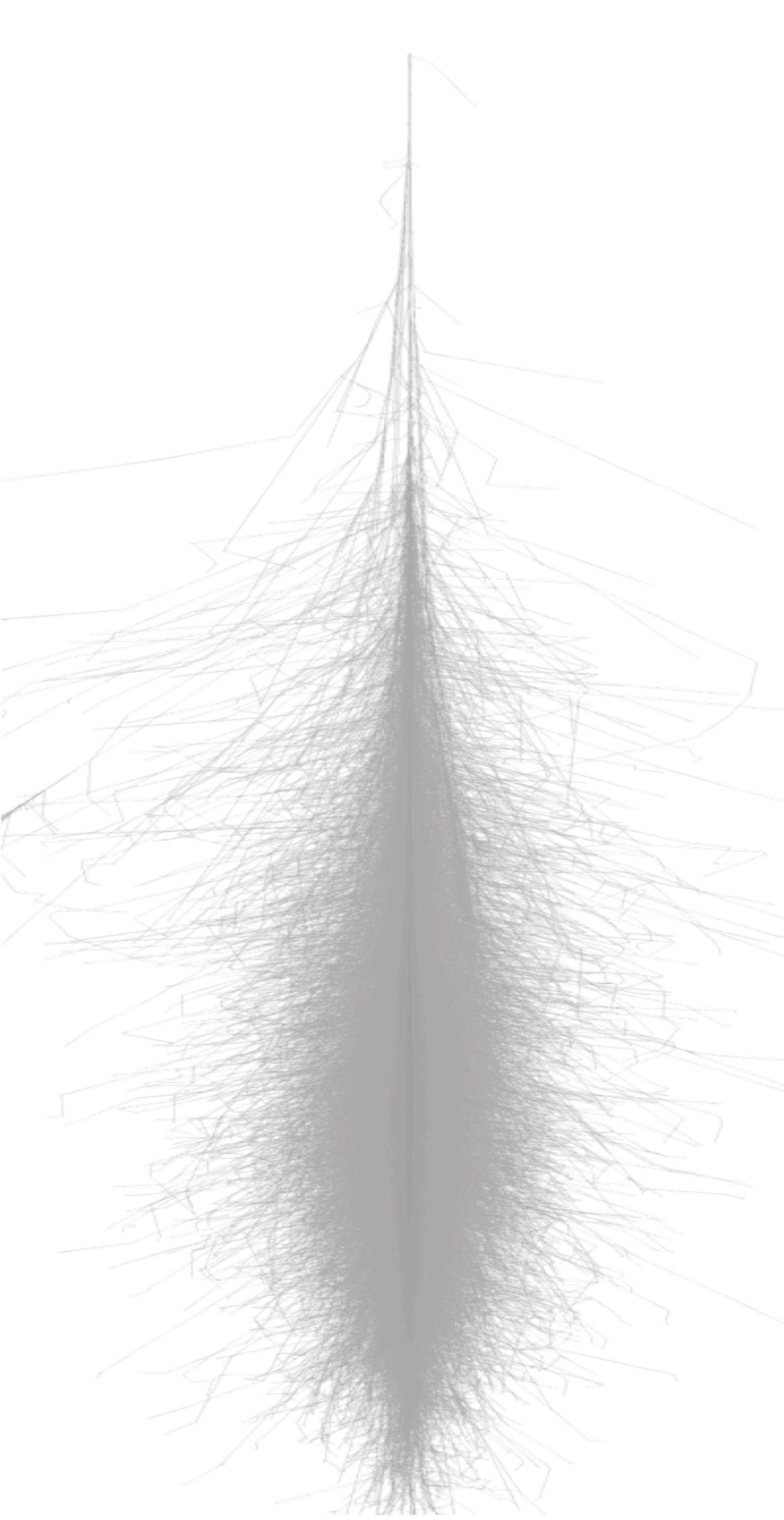
$$\langle N_\mu \rangle \sim E^\beta$$

New model



$$N_\mu \sim E_{\text{had}}$$

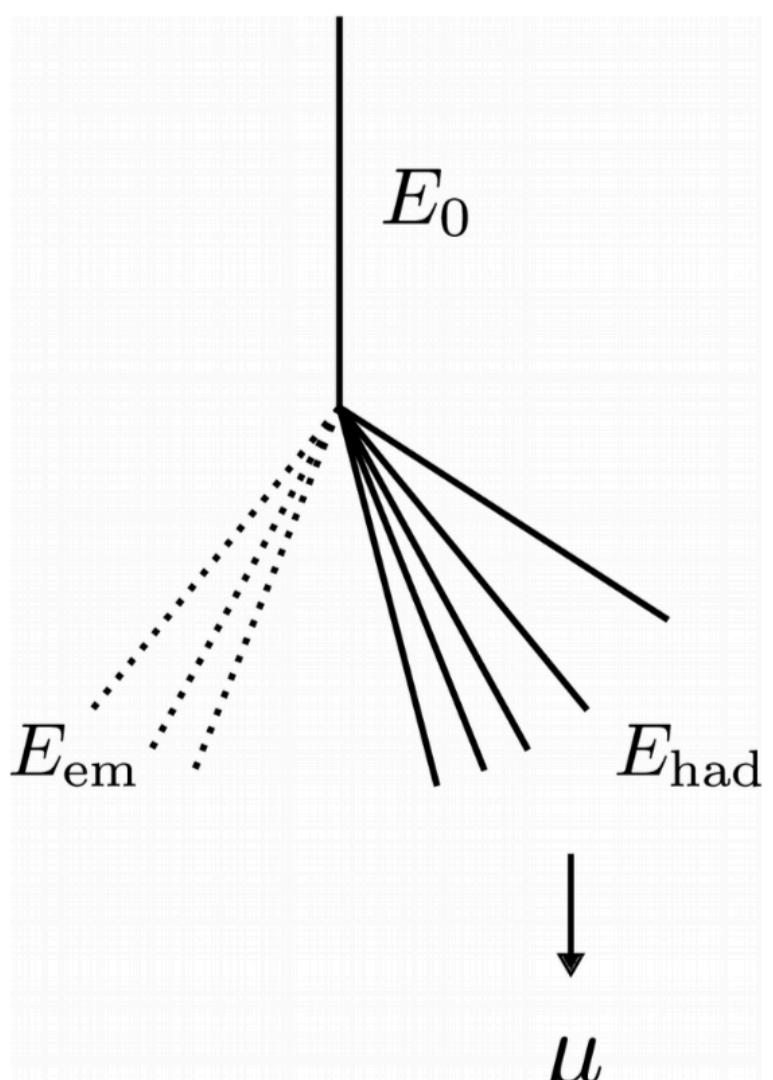
Origin of the fluctuations of the muon number distribution



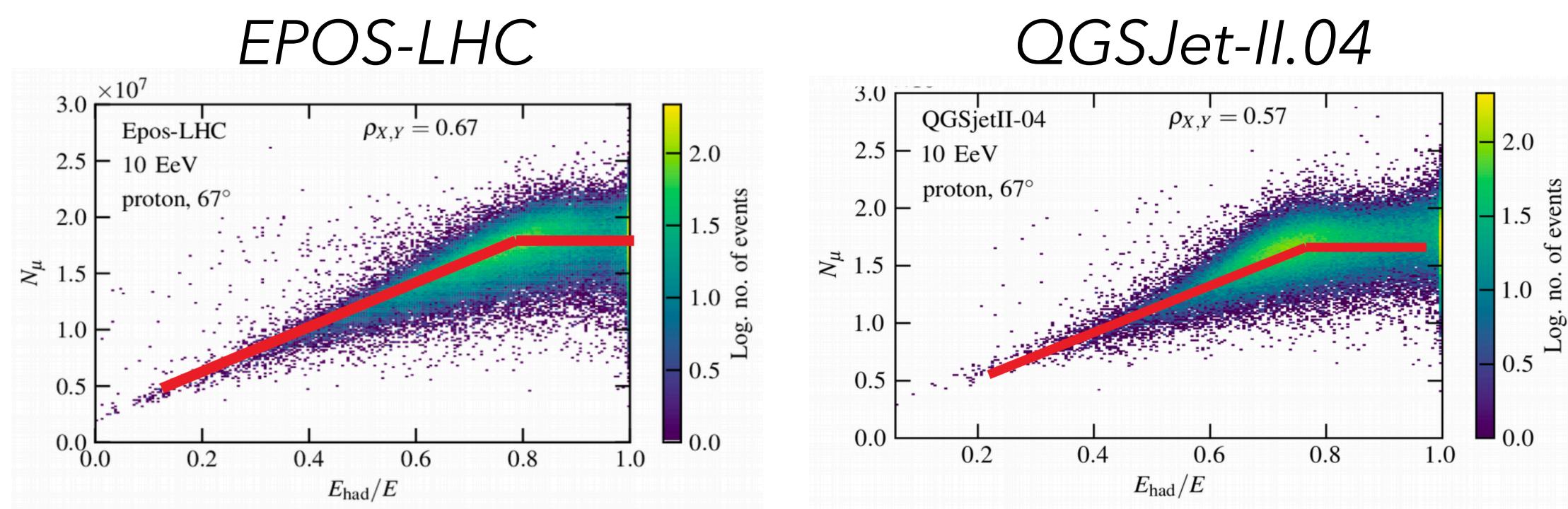
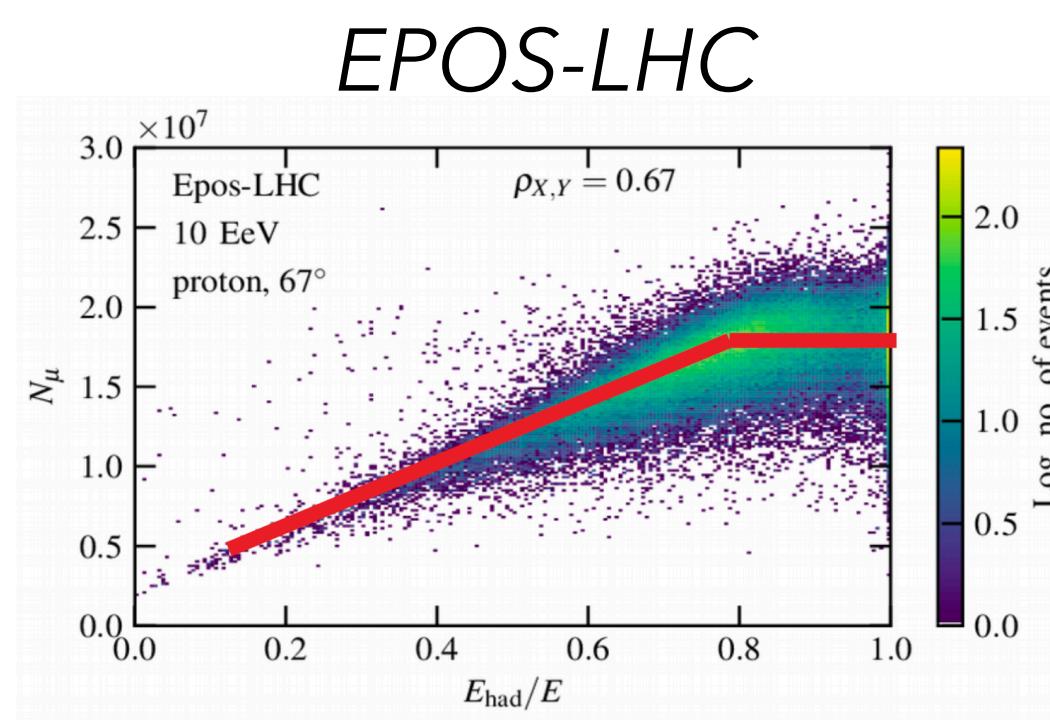
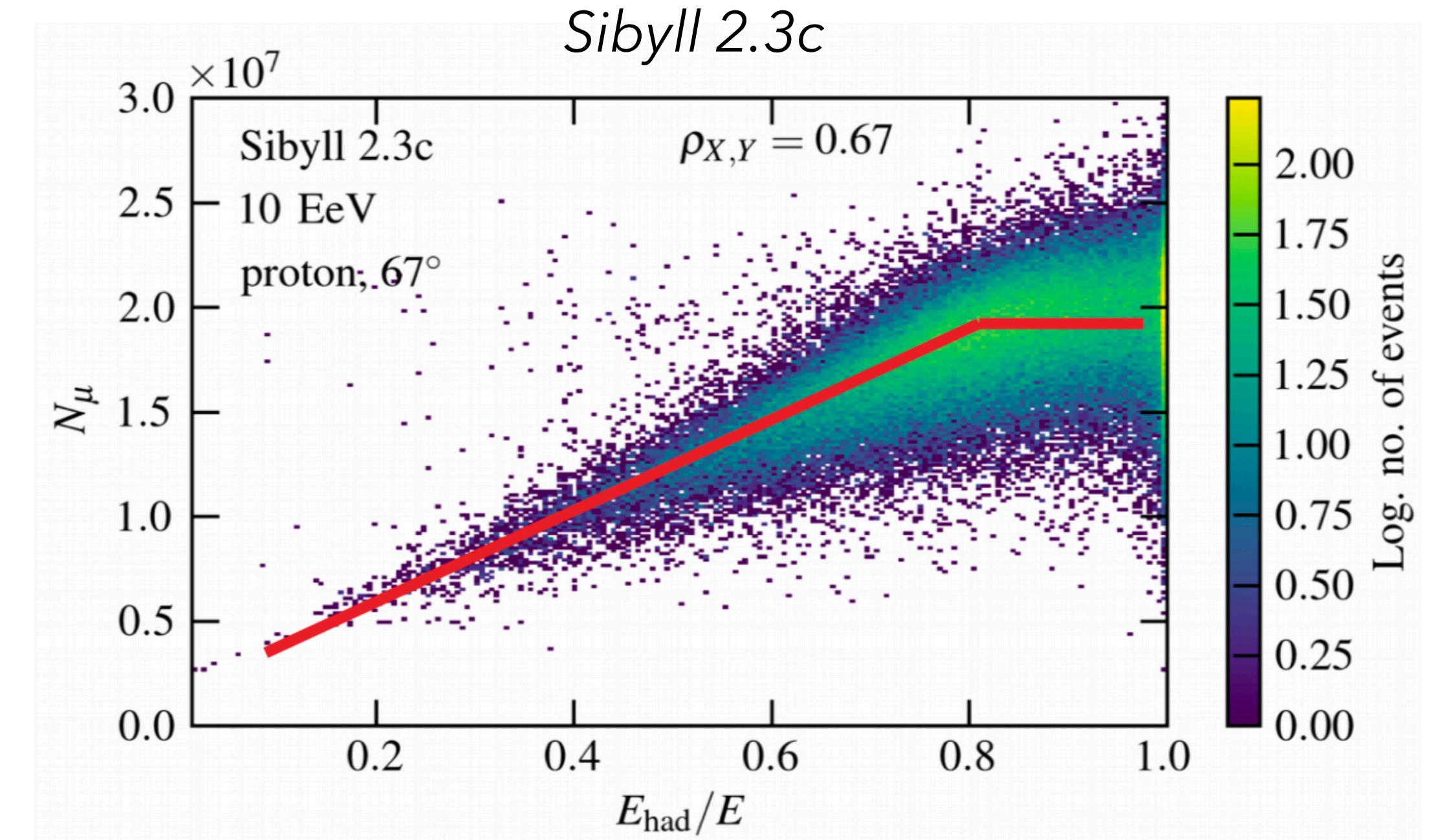
Heitler-Matthews model

$$\langle N_\mu \rangle \sim E^\beta$$

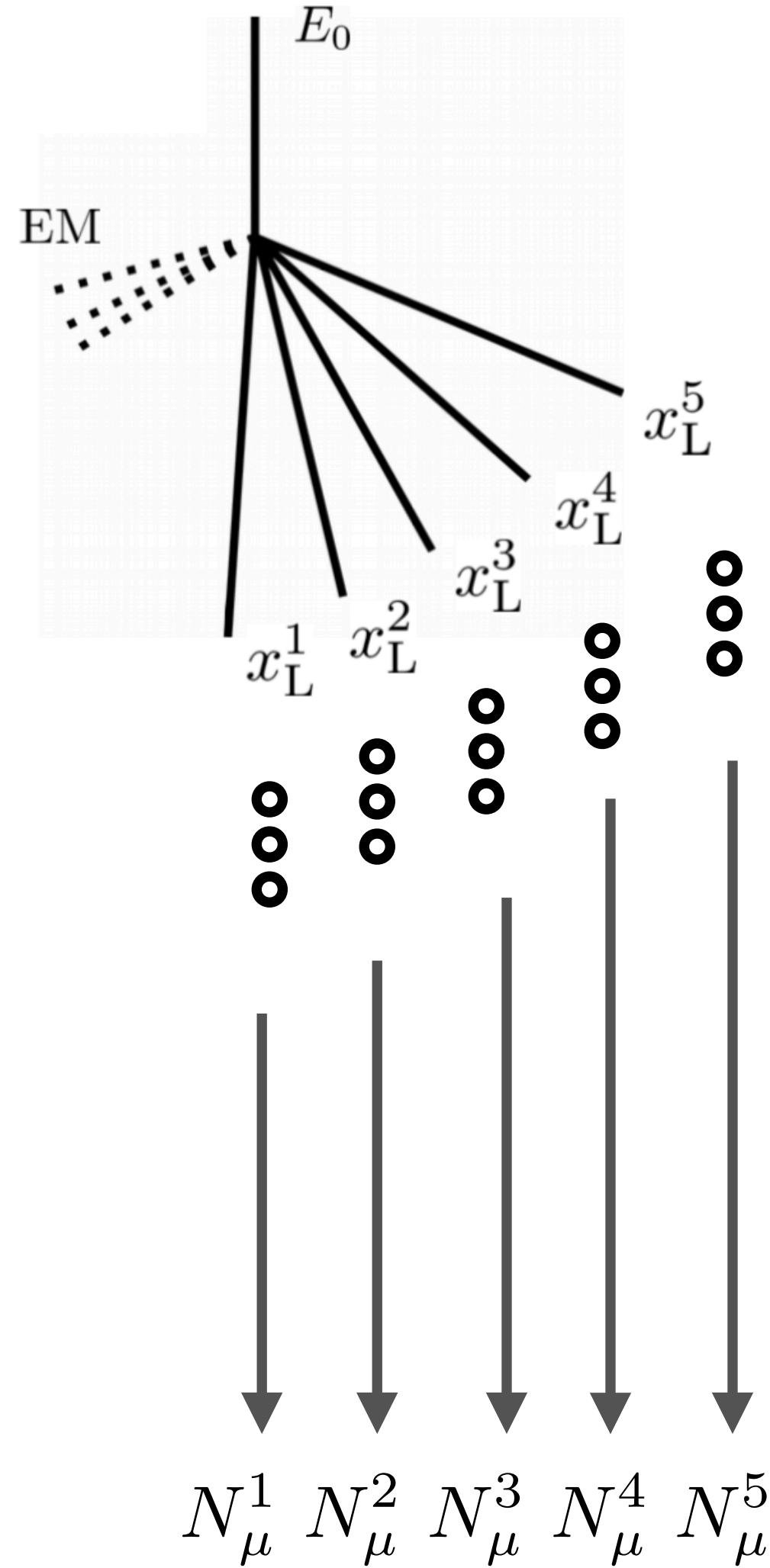
New model



$$N_\mu \sim E_{\text{had}}$$



Origin of the fluctuations of the muon number distribution



New model

$$x_L^i = E_i/E_0$$

$$E_{\text{had}} = E_0 \sum_i x_L^i$$

Since

$$N_\mu = \sum_i N_\mu^i \approx \sum_i \langle N_\mu(E_i) \rangle$$

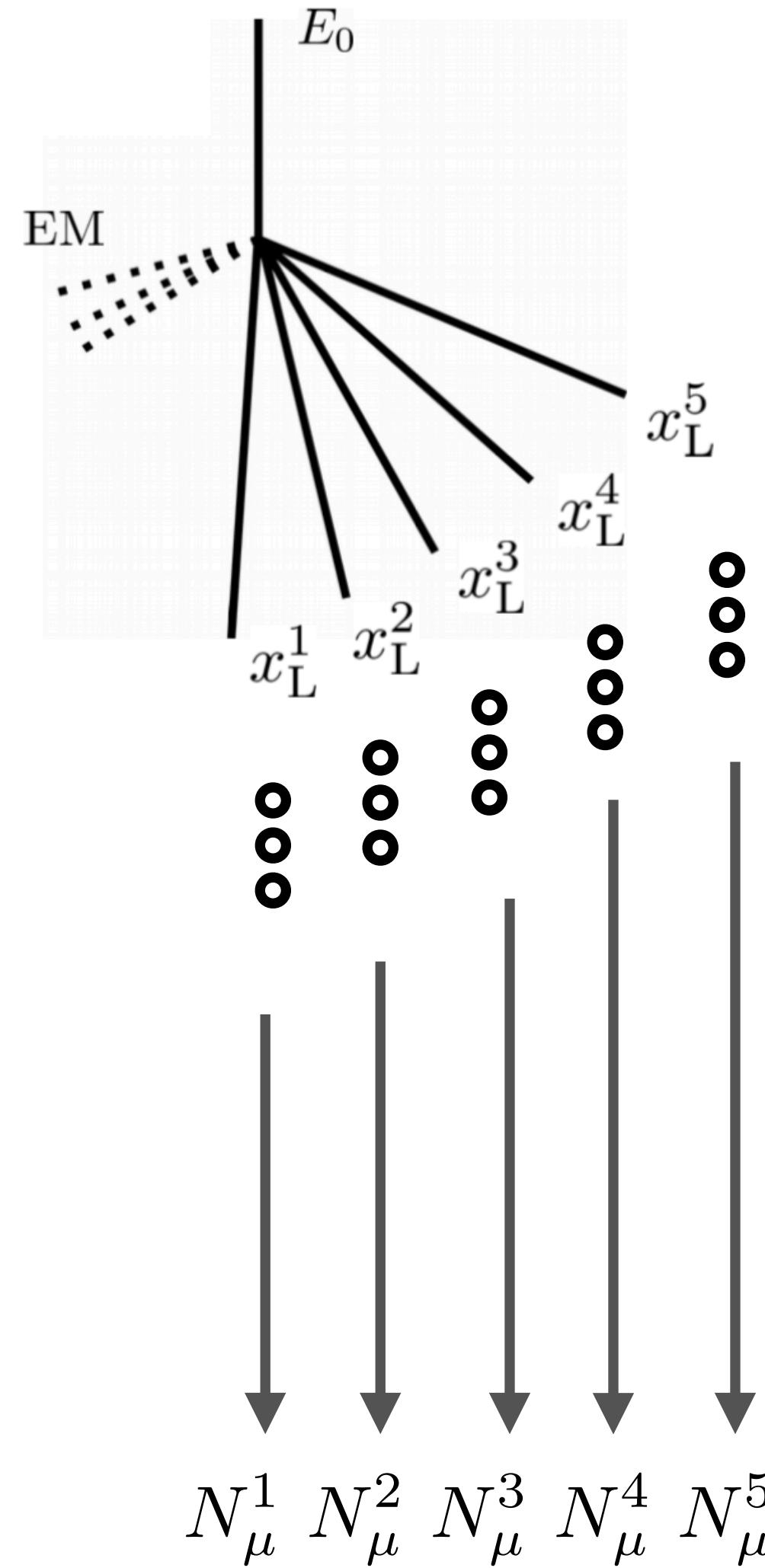
And

$$\langle N_\mu \rangle \sim E^\beta$$

Then

$$N_\mu = a \sum_i (x_L^i)^\beta = a \alpha_1$$

Origin of the fluctuations of the muon number distribution



New model

$$x_L^i = E_i/E_0$$

$$E_{\text{had}} = E_0 \sum_i x_L^i$$

Since

$$N_\mu = \sum_i N_\mu^i \approx \sum_i \langle N_\mu(E_i) \rangle$$

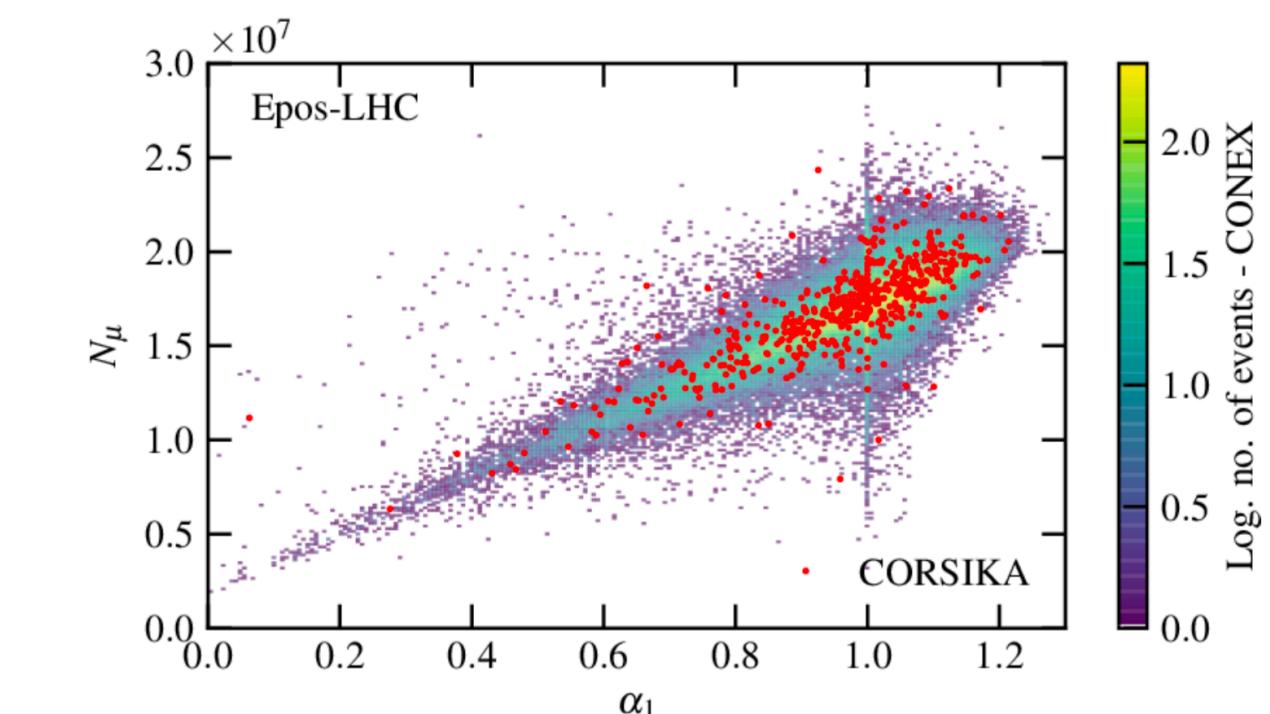
And

$$\langle N_\mu \rangle \sim E^\beta$$

Then

$$N_\mu = a \sum_i (x_L^i)^\beta = a \alpha_1$$

$$\alpha_1 \equiv \sum_i^{m_{\text{had}}} (x_L^i)^\beta$$



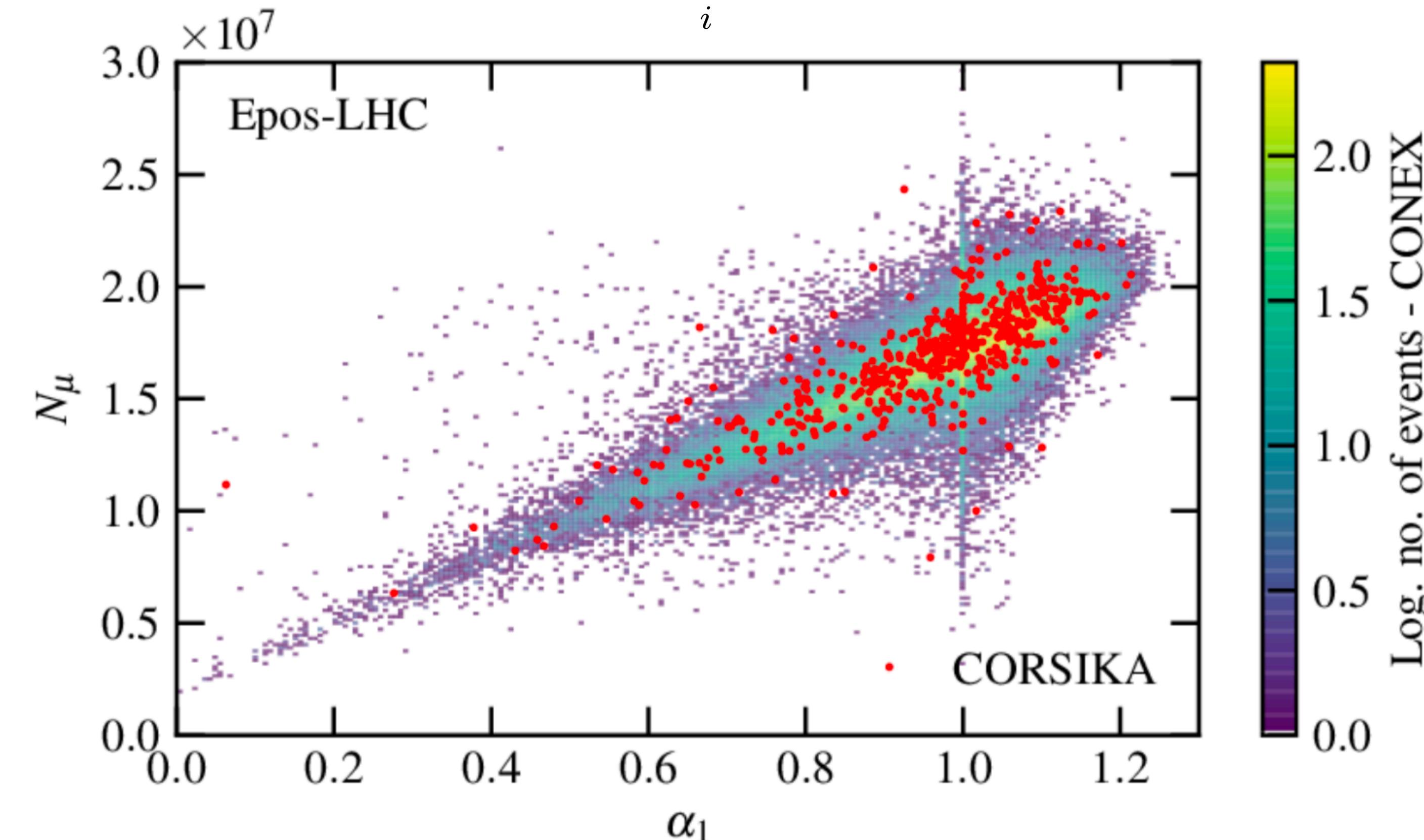
$$\sigma(\alpha) \rightarrow 70\% \sigma(N_\mu)$$

β is a parameter related with the first
Interaction multiplicity

Origin of the fluctuations of the muon number distribution

$$\alpha_1 \equiv \sum_i^{m_{had}} (x_L^i)^\beta$$

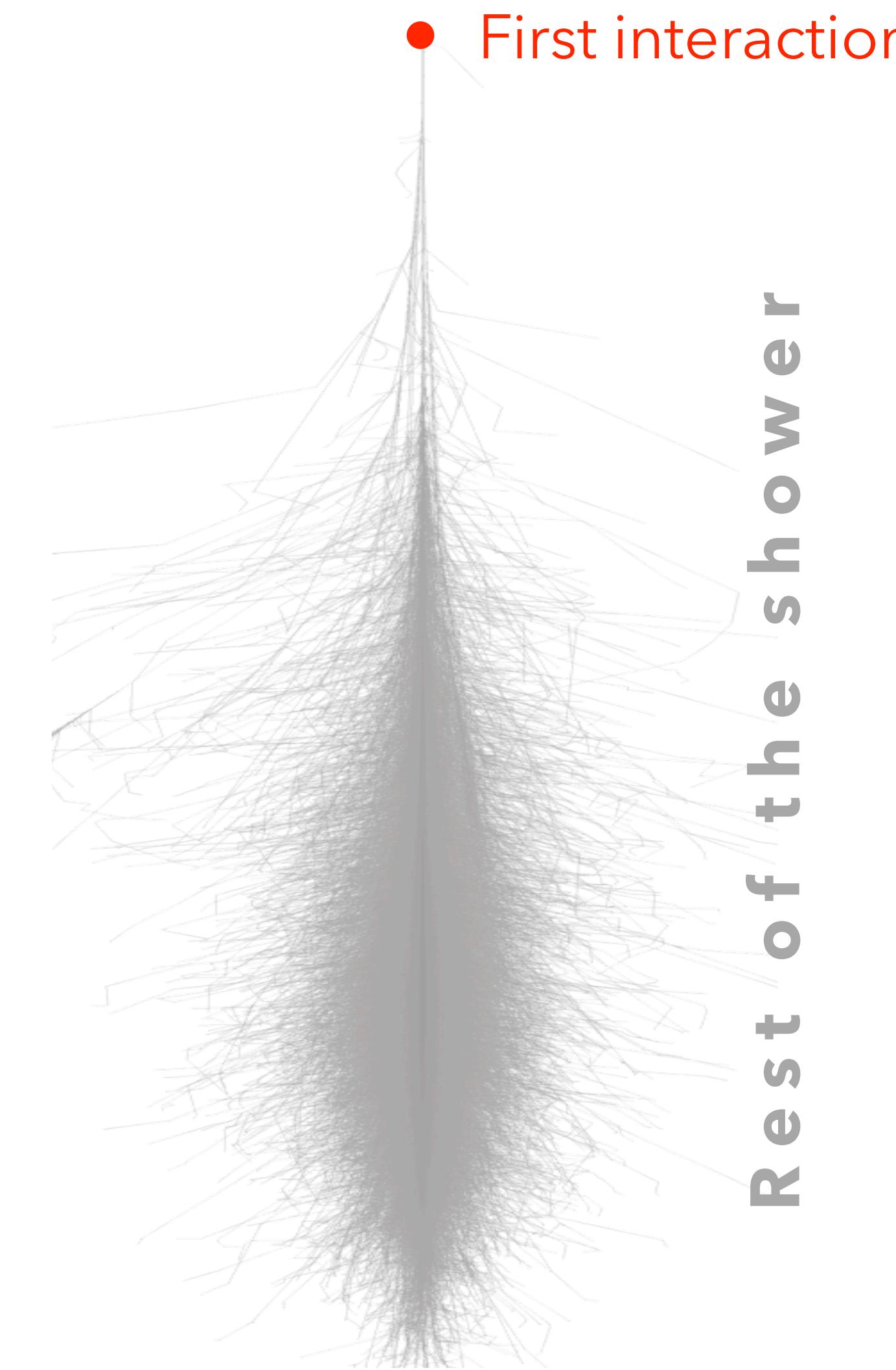
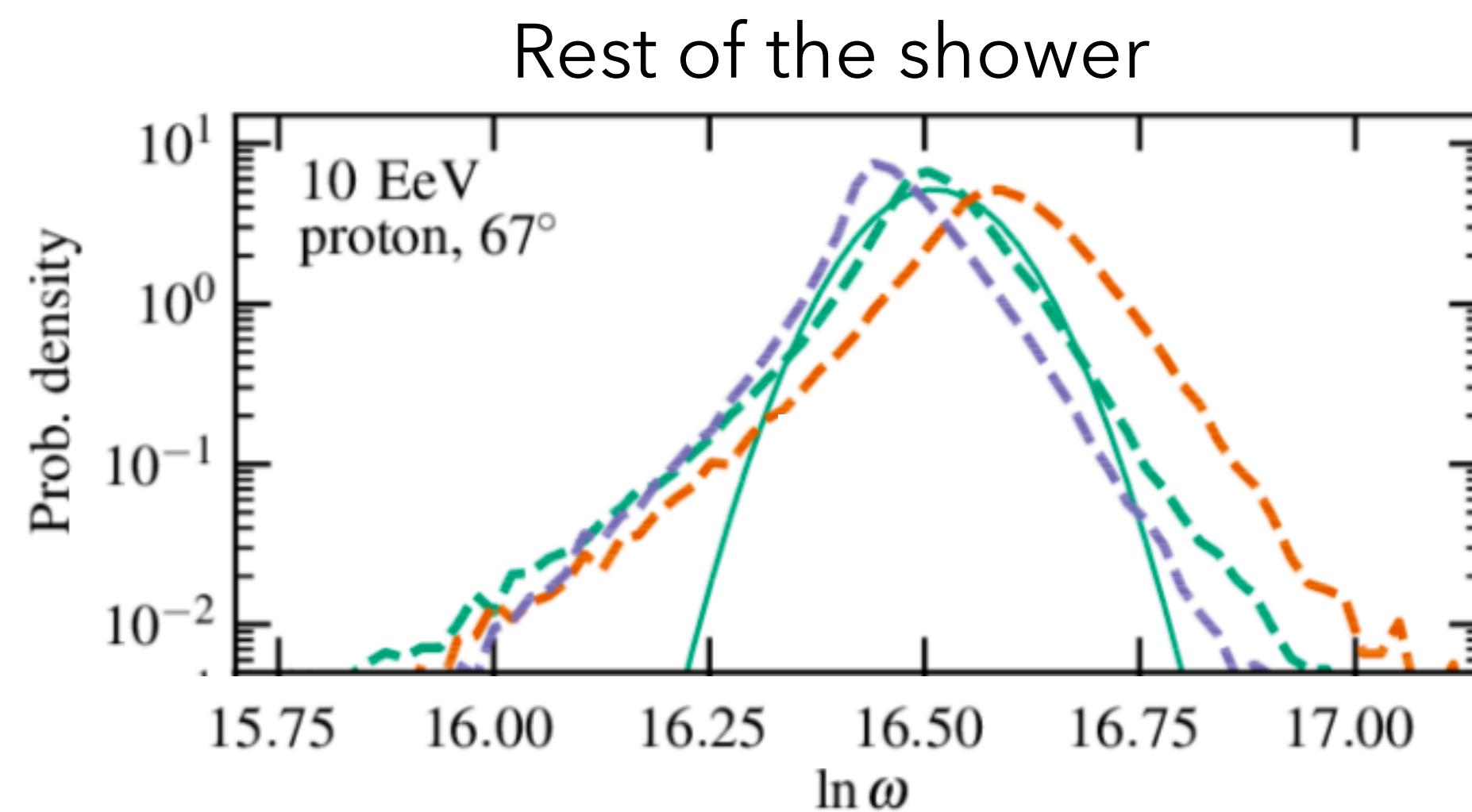
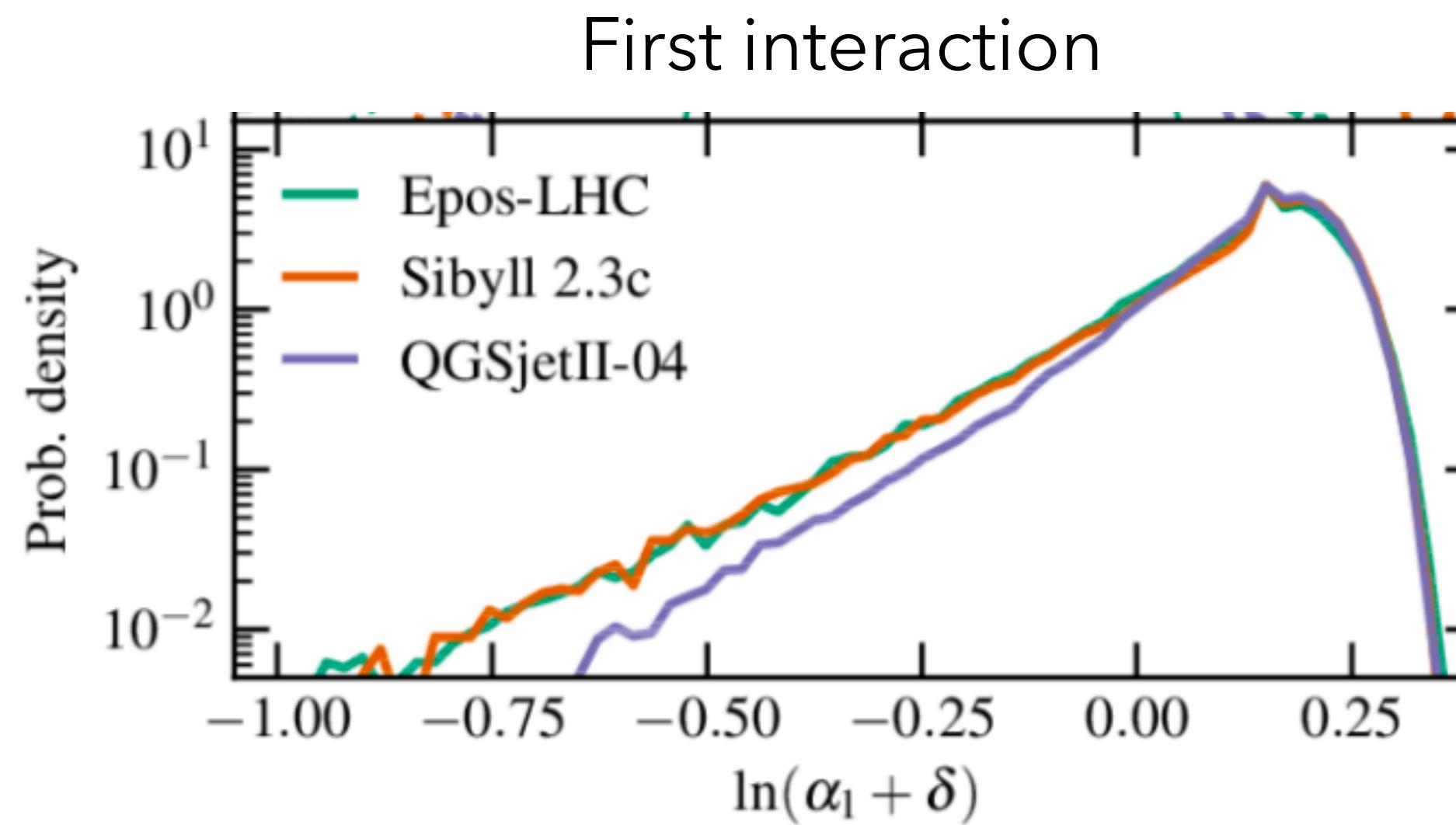
L. Cazon, RC, F. Riehn, PLB 784 (2018) 68-76



The **number of muons** strongly correlates with the fraction of **hadronic energy of the first interaction**

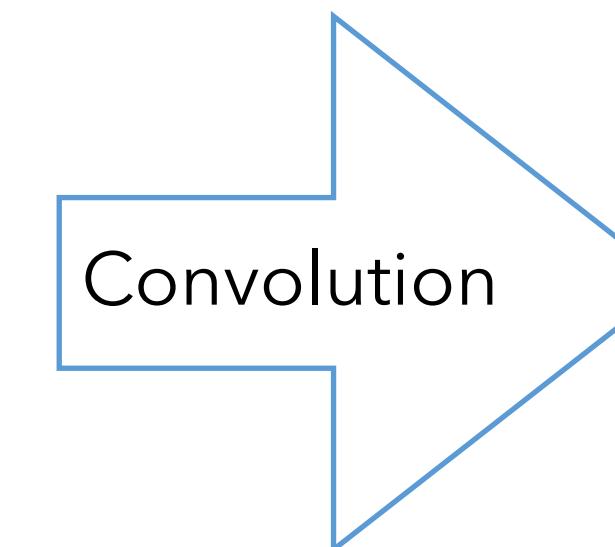
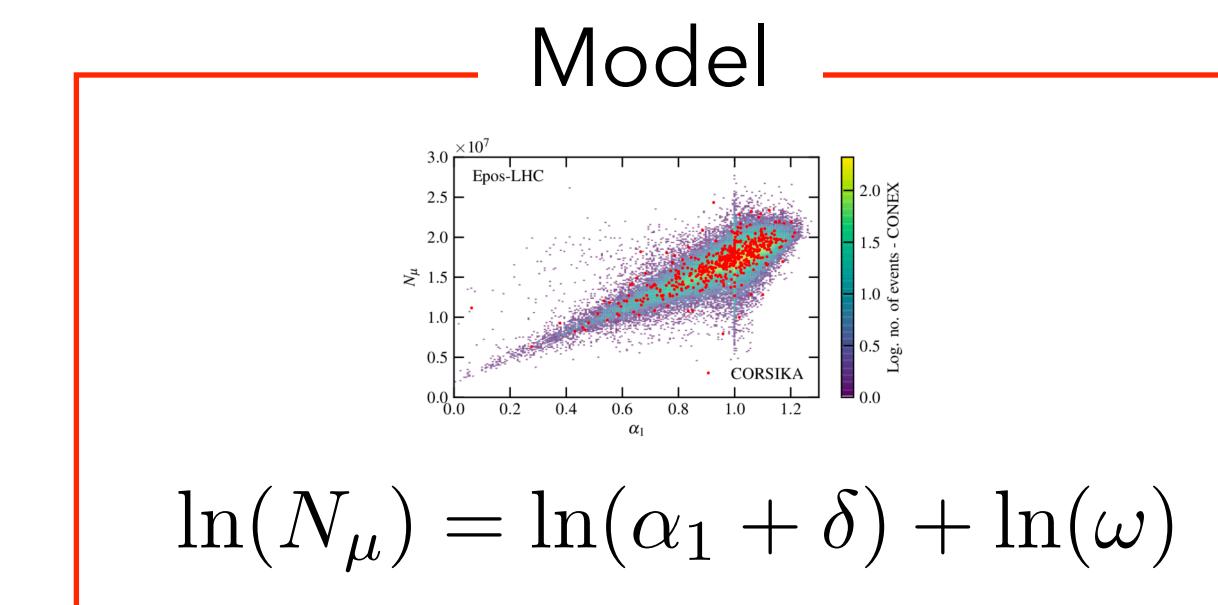
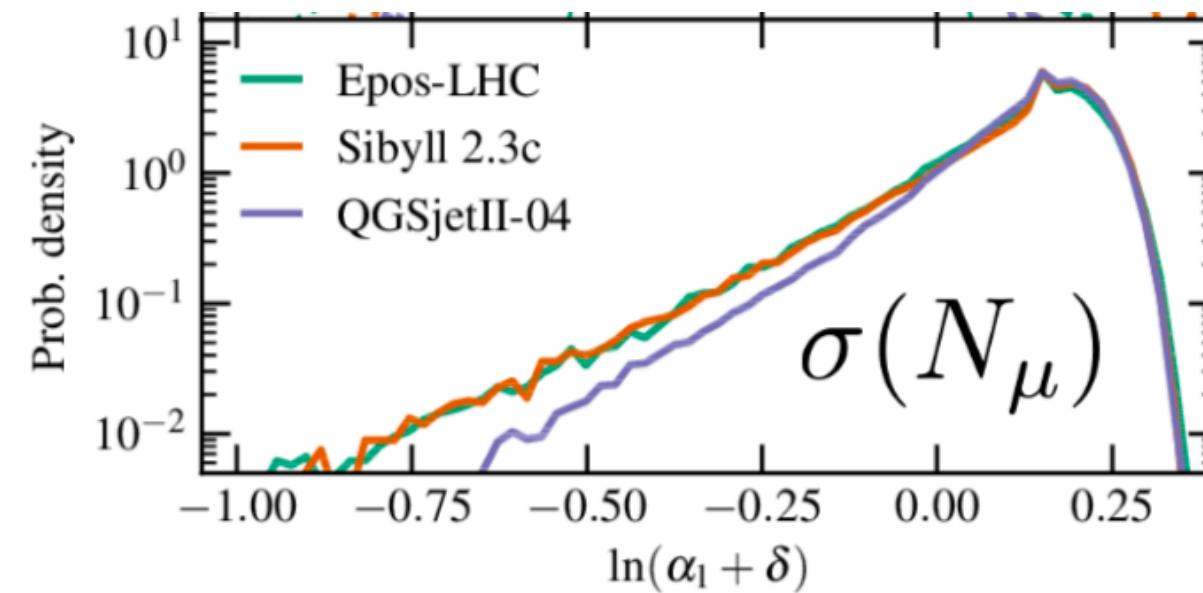
- Independently of interaction model, primary energy, primary zenith angle

Understanding the N_μ distribution

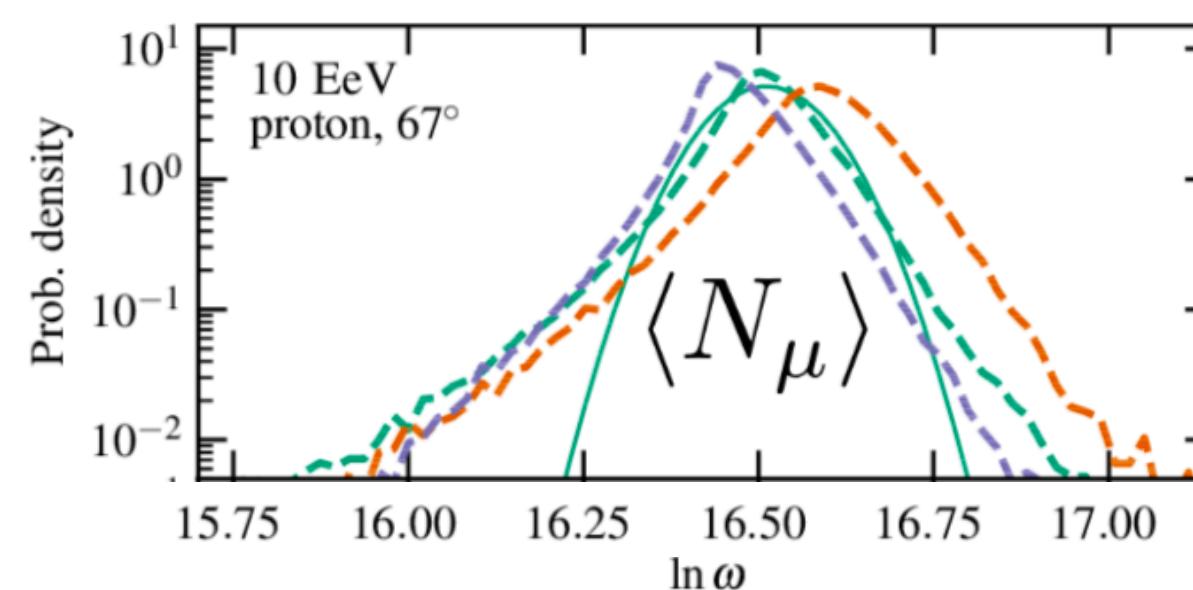


Understanding the N_μ distribution

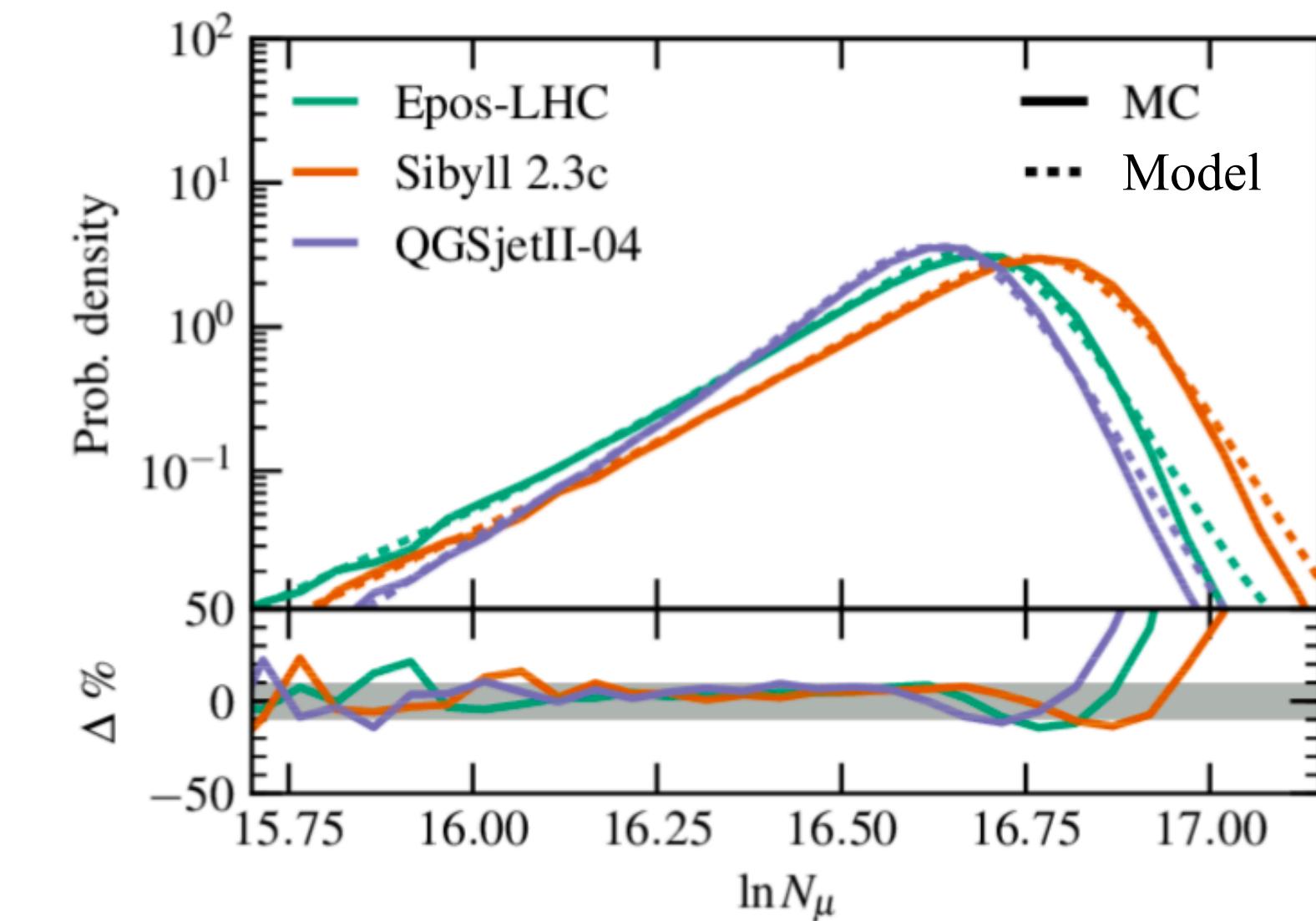
First interaction



Rest of the shower

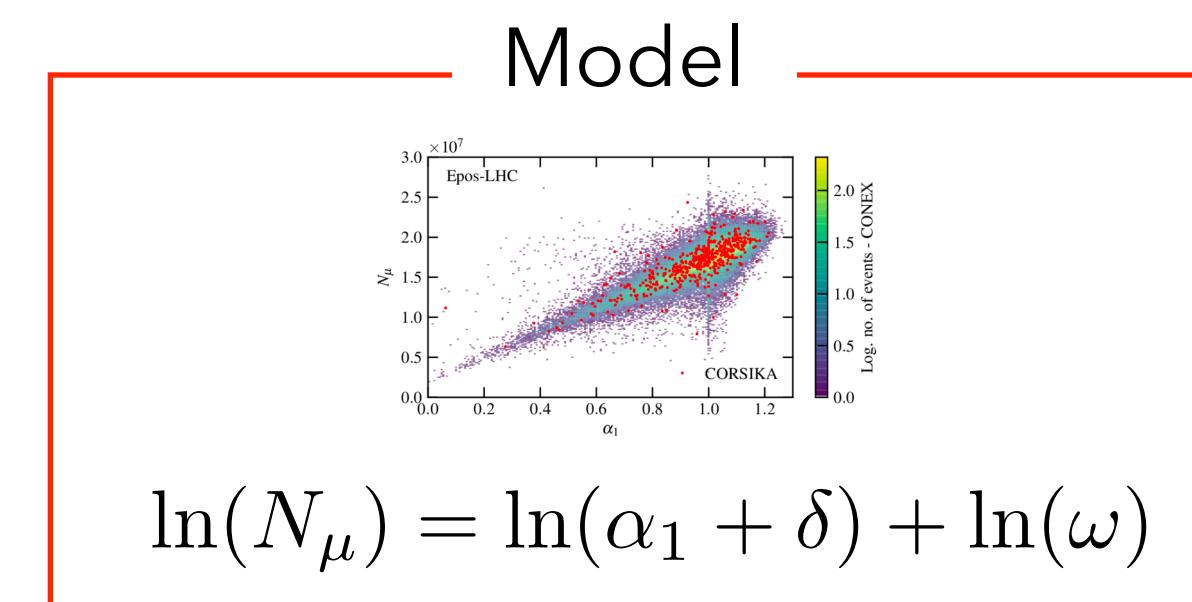
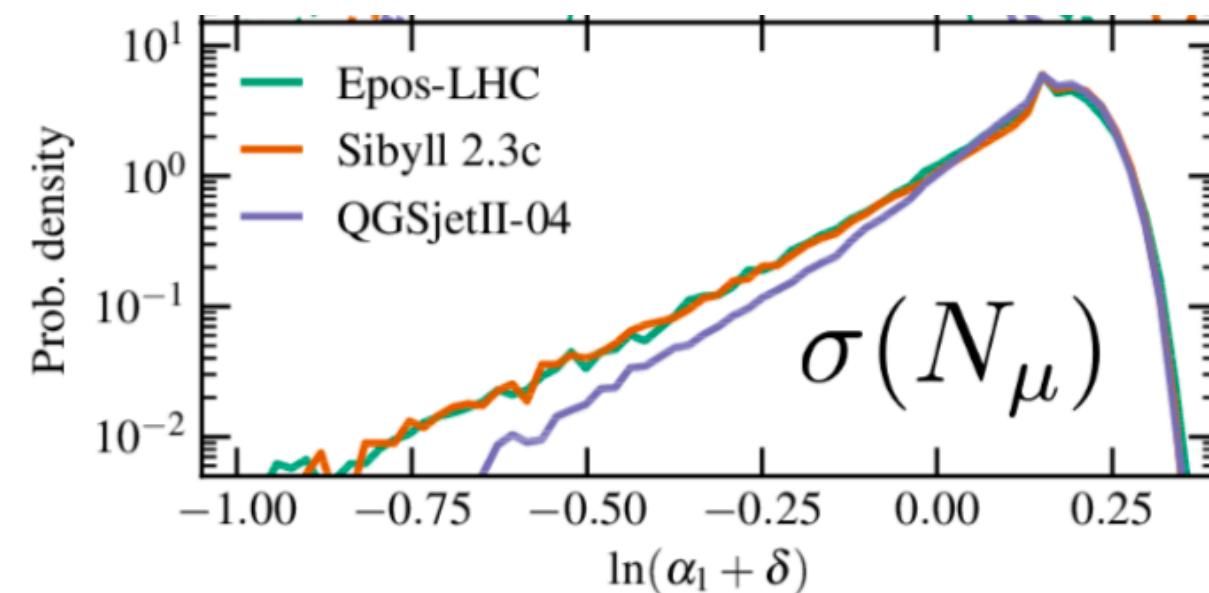


α is the main parameter that controls the shape of N_μ dist
 δ is a small contribution from photopion production, $\sim 10\%$

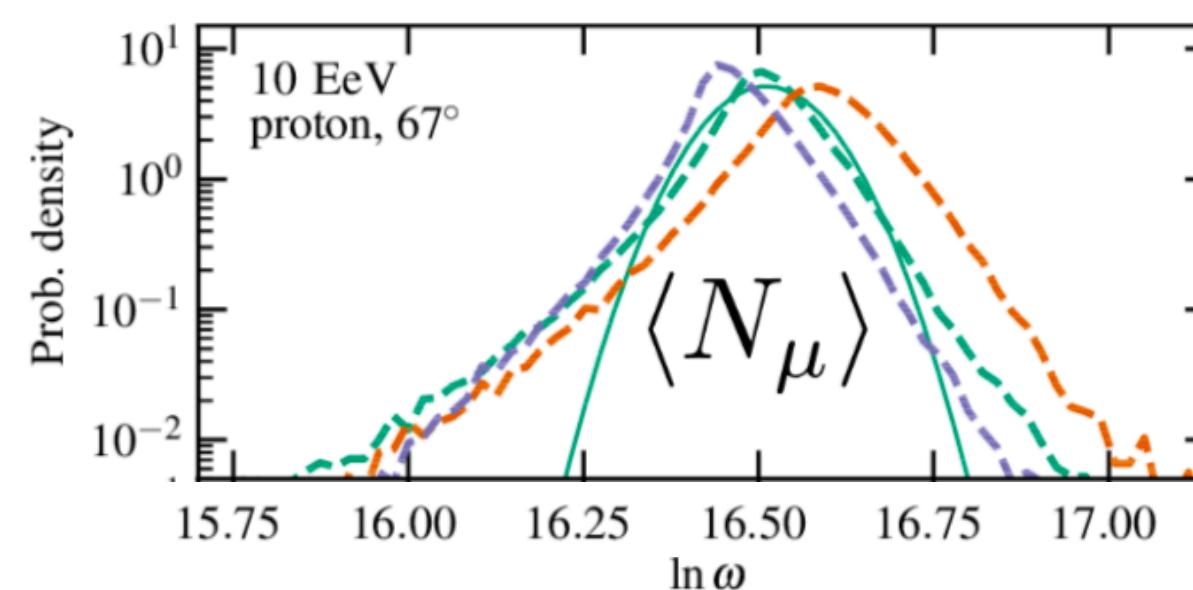


Understanding the N_μ distribution

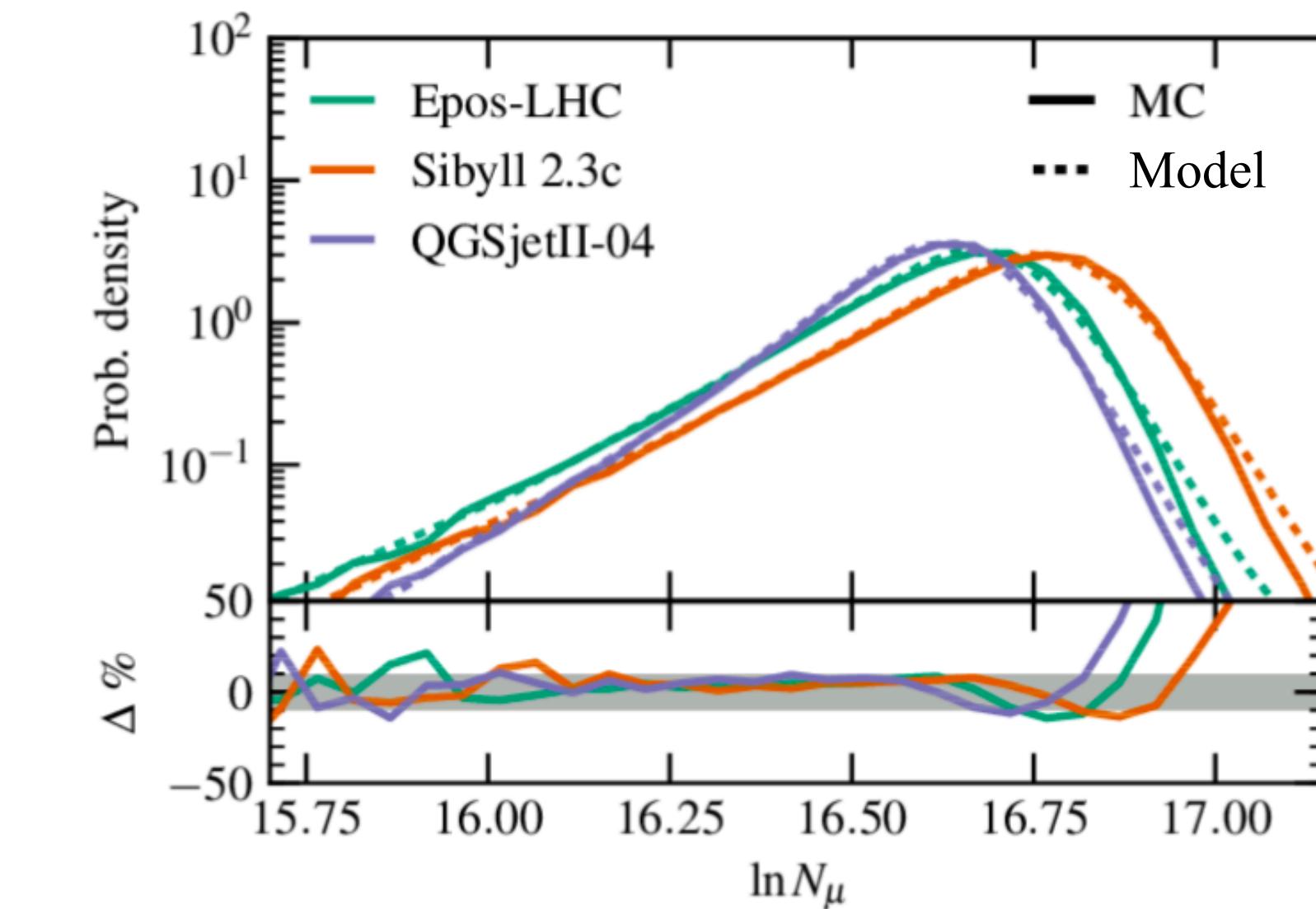
First interaction



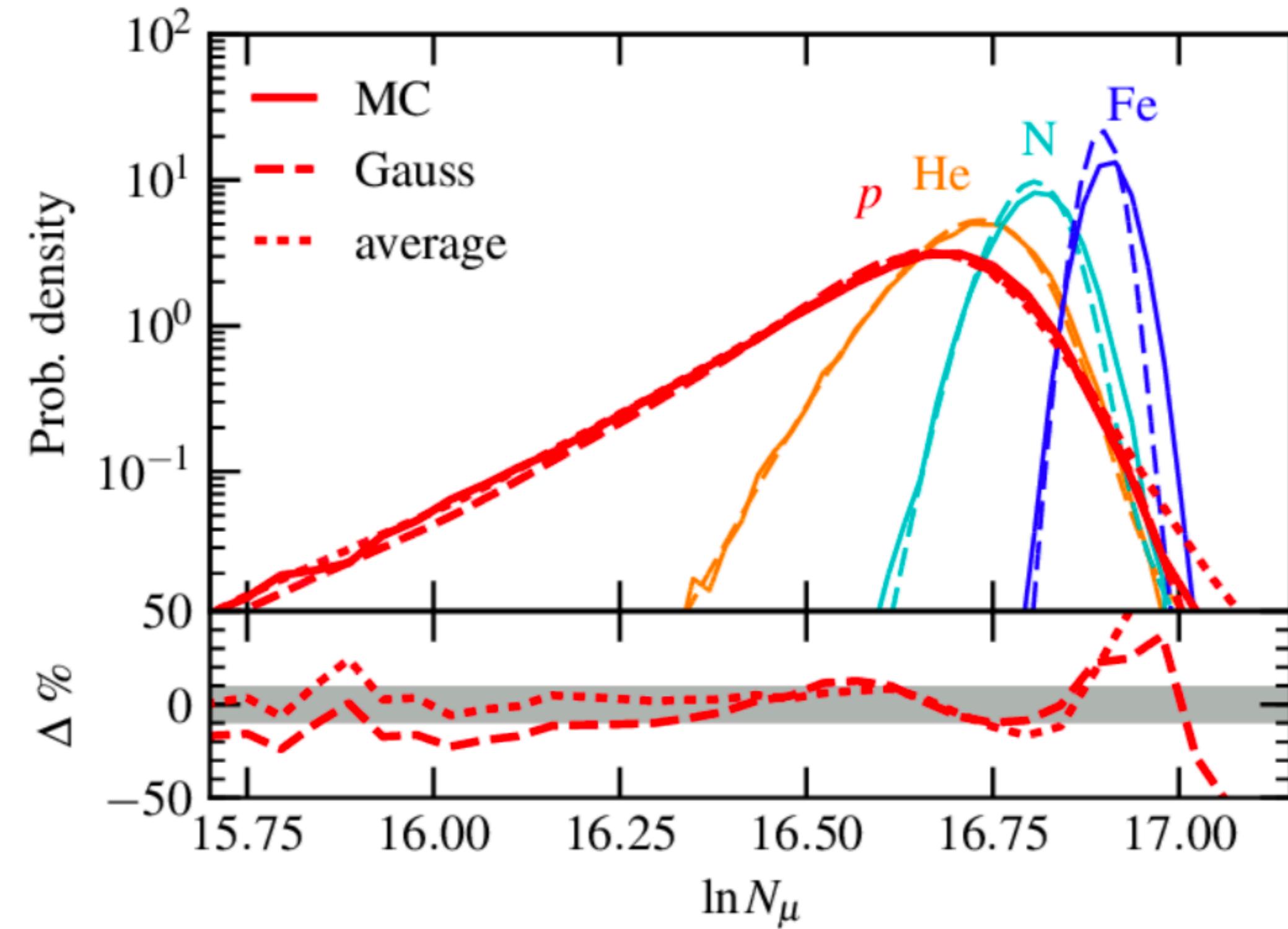
Rest of the shower



Controls N_μ fluctuations
event-by-event



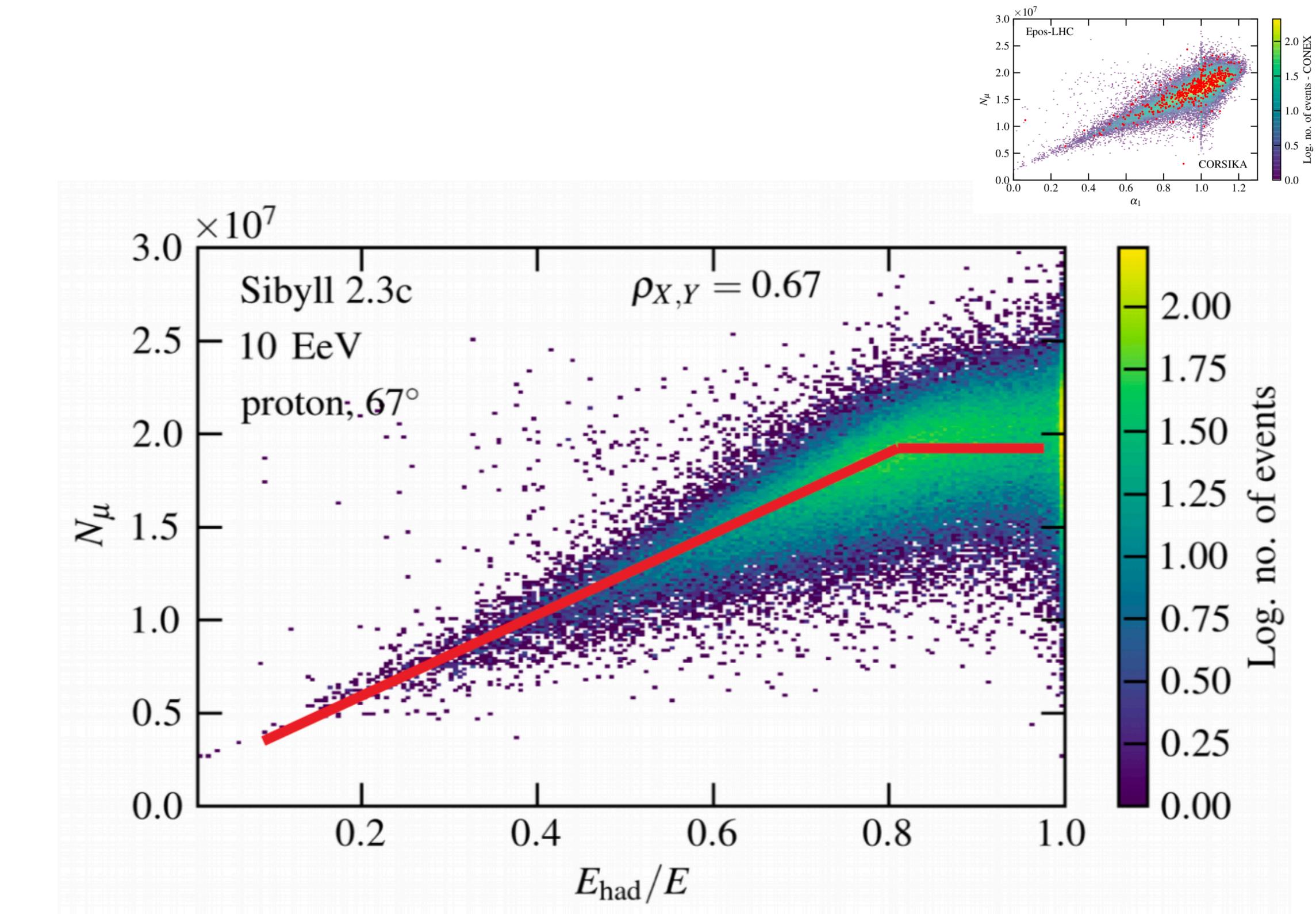
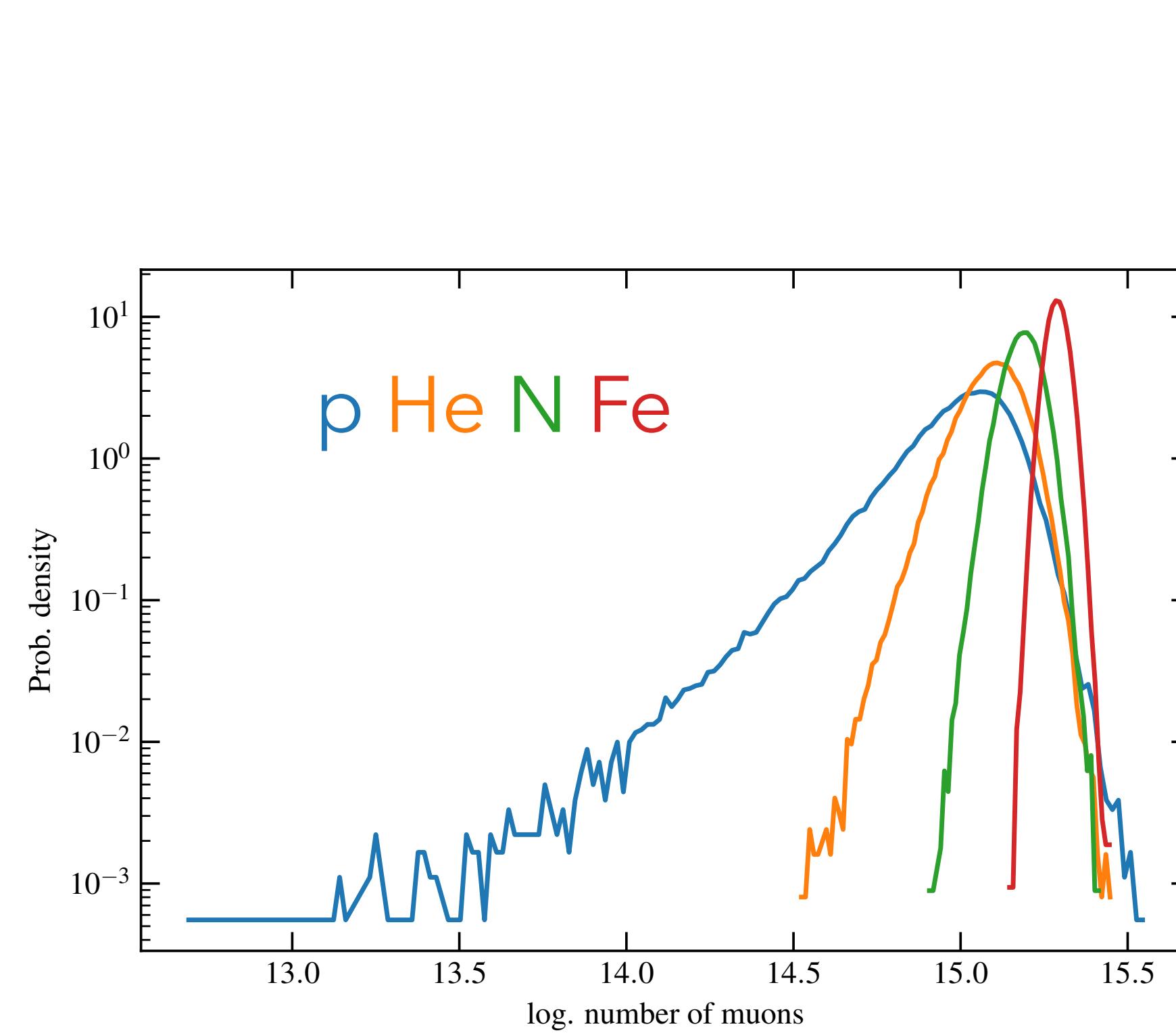
Controls overall
number of muons



What is the proton low N_μ tail?

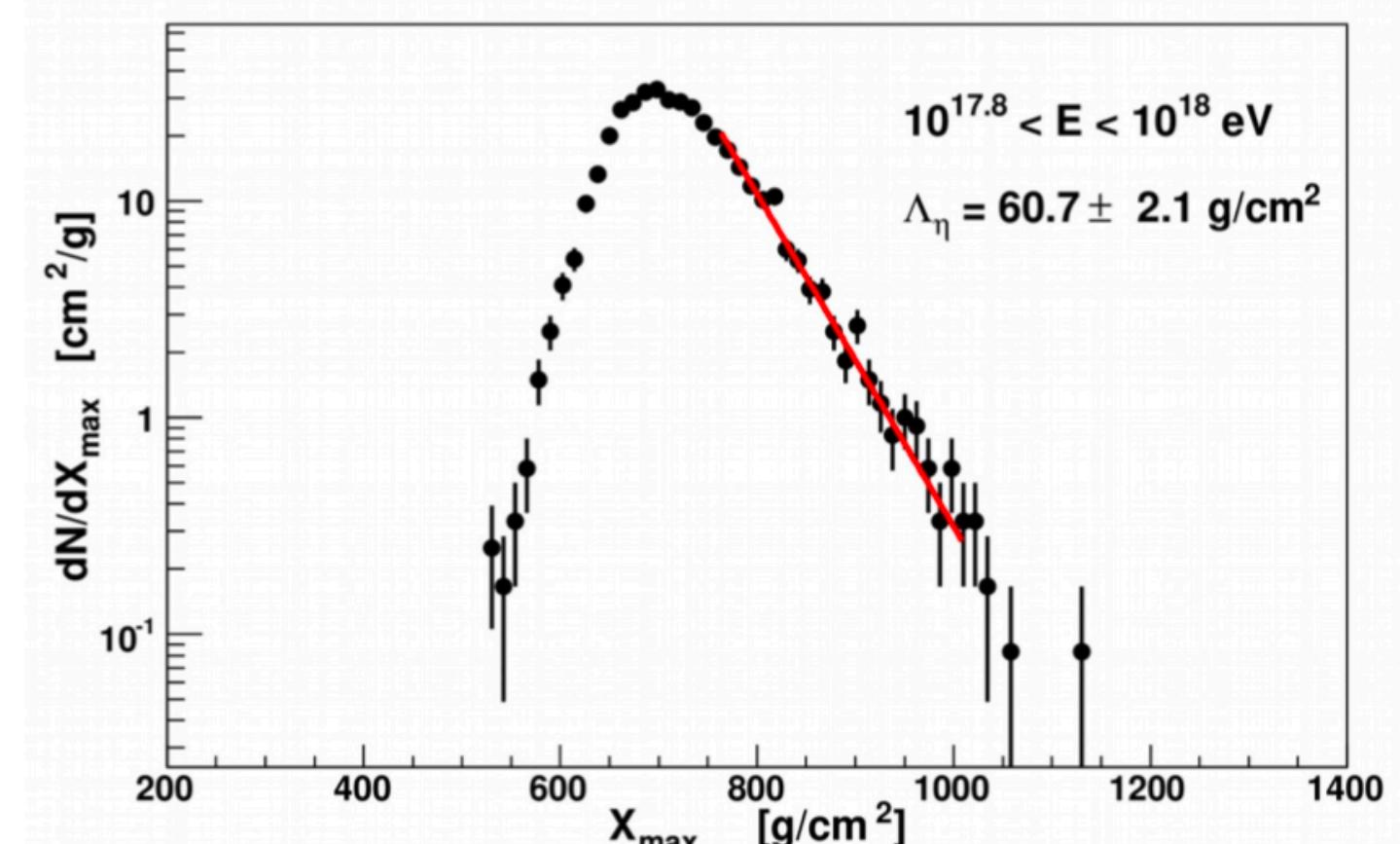
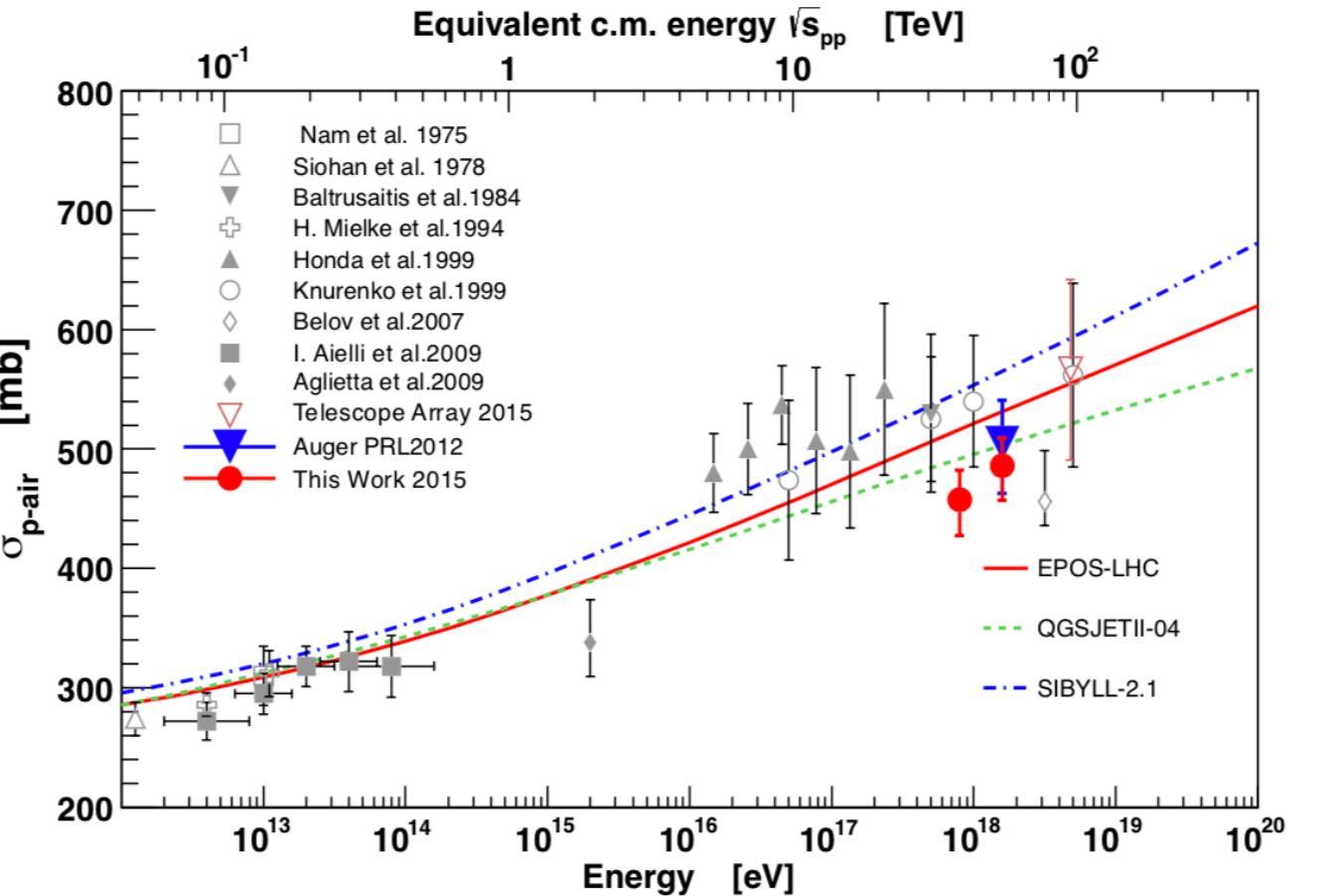
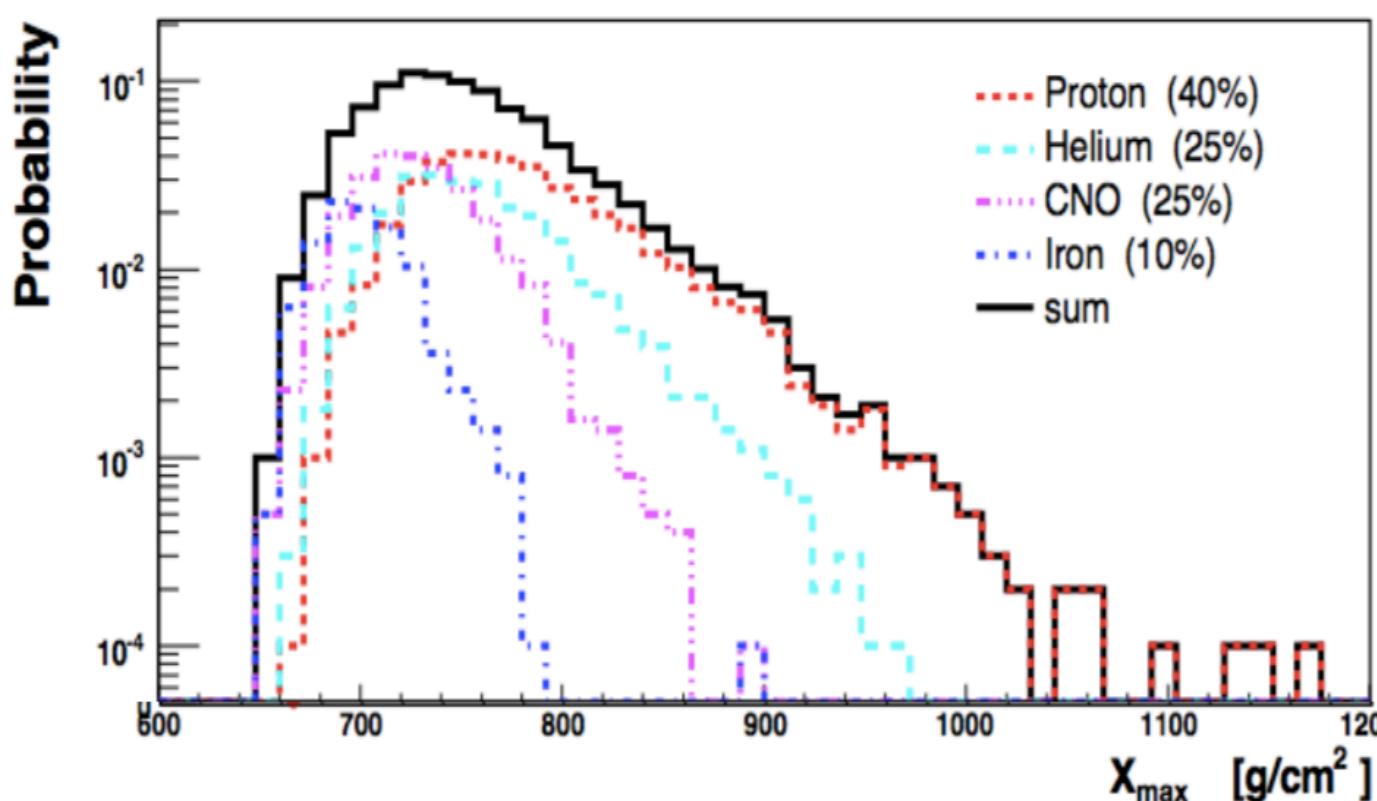
To which property of the first interaction is it related to?

On the origin of the proton N_μ distribution tail

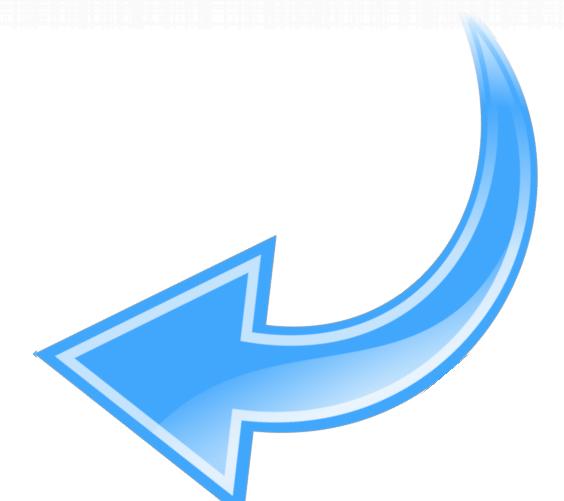
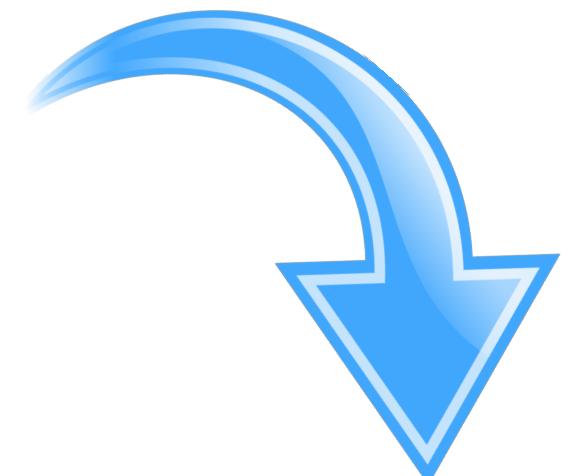
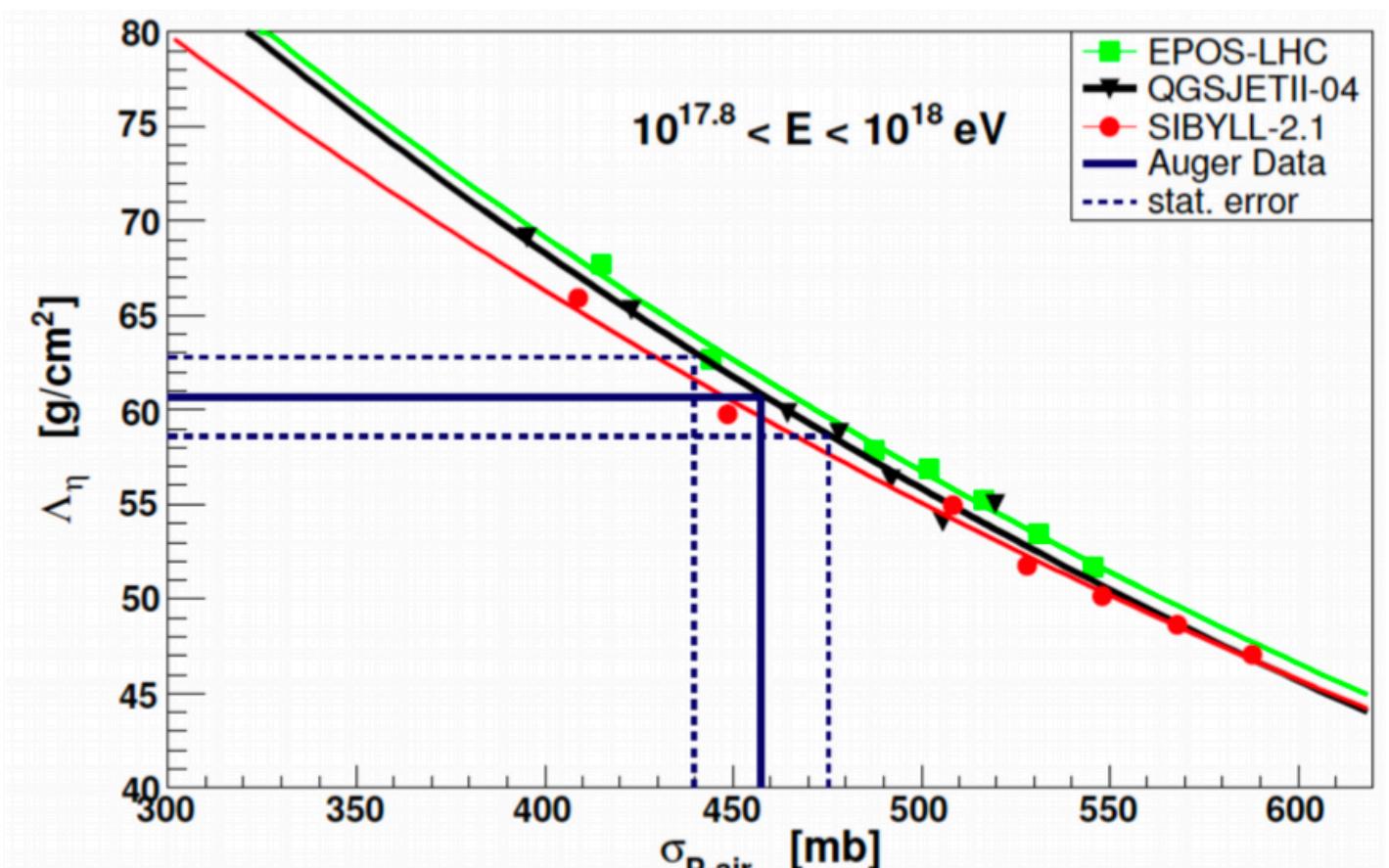


N_μ distribution tail has a very high correlation with the **hadronic energy of the first interaction**

Can it be measured?



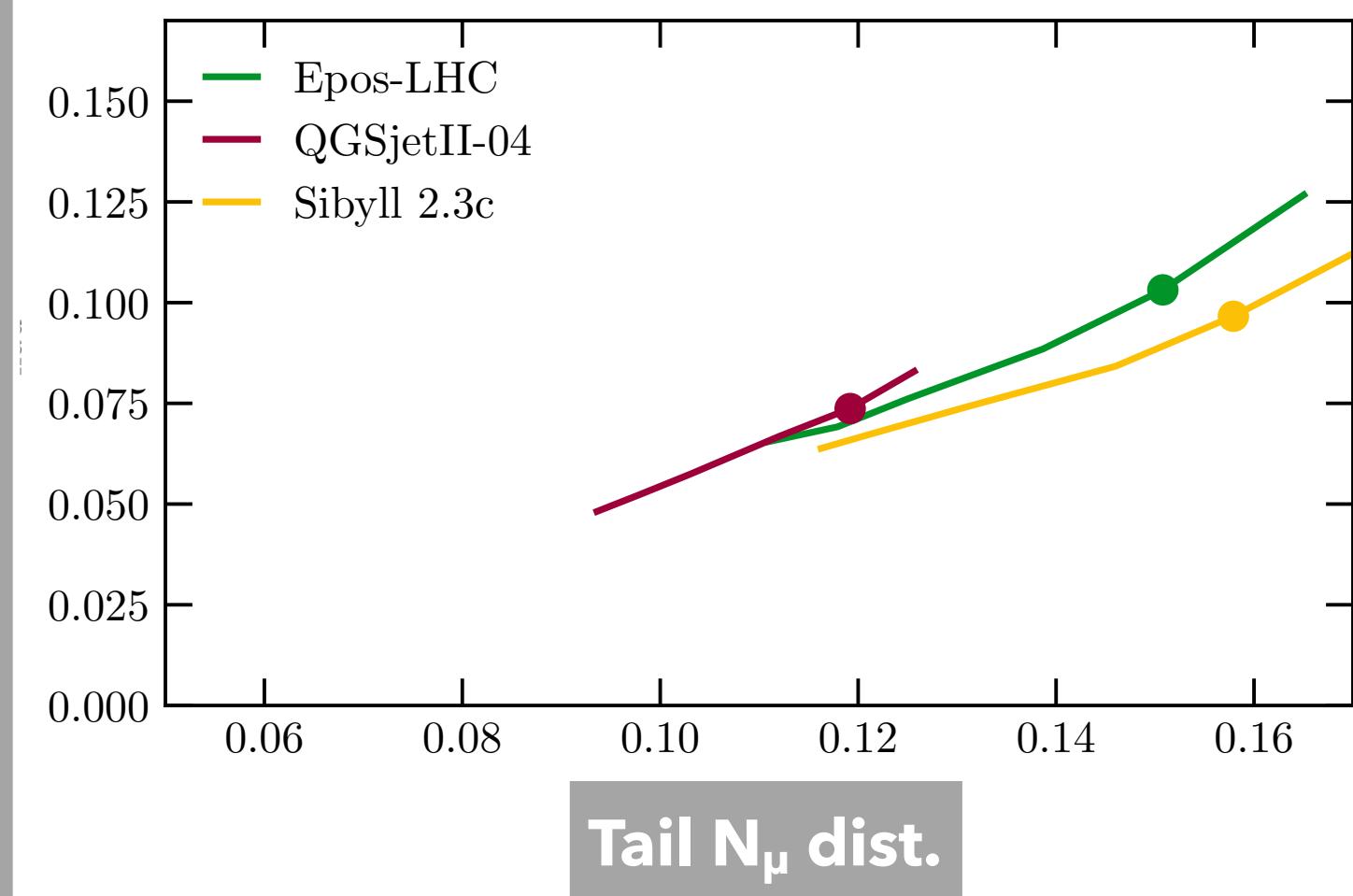
proton-Air cross-section measurement



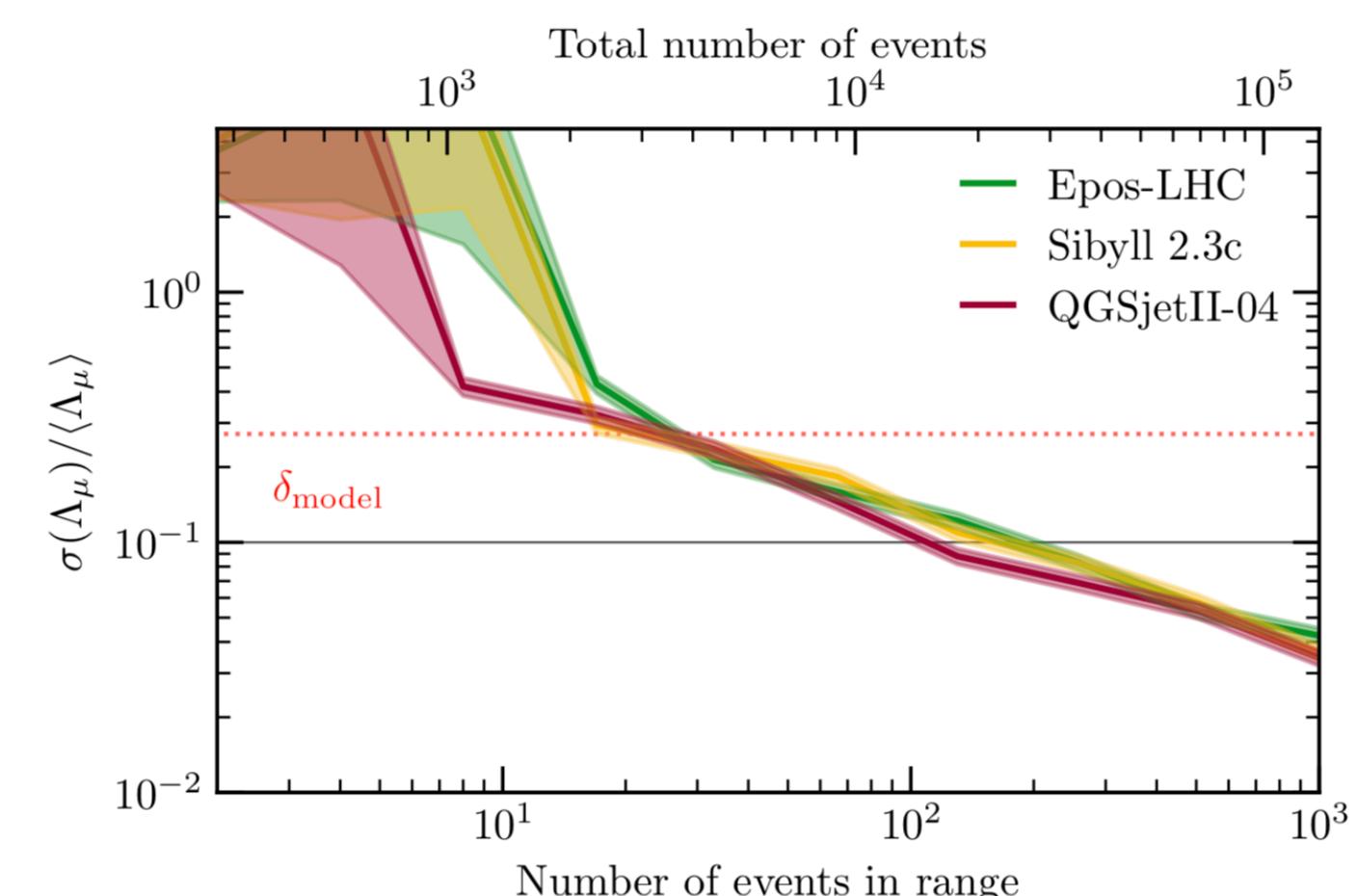
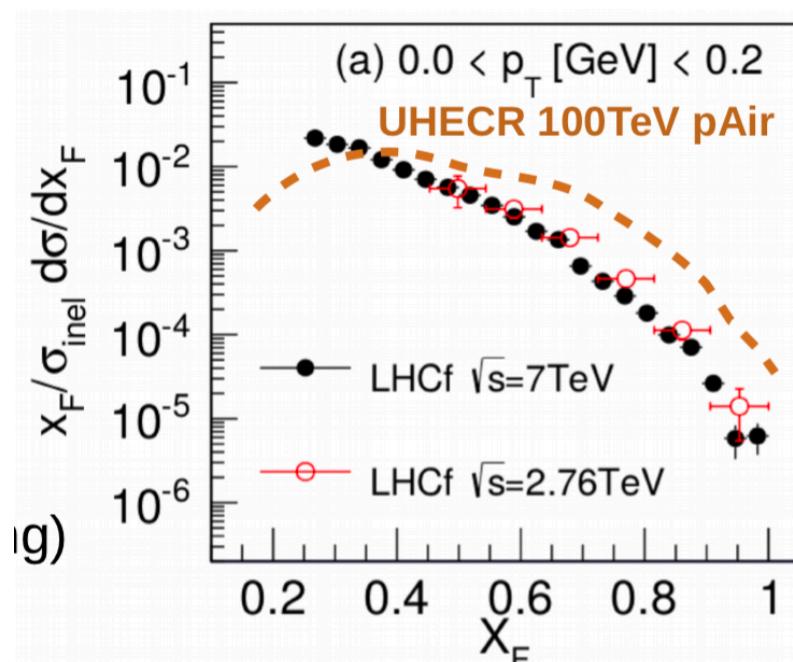
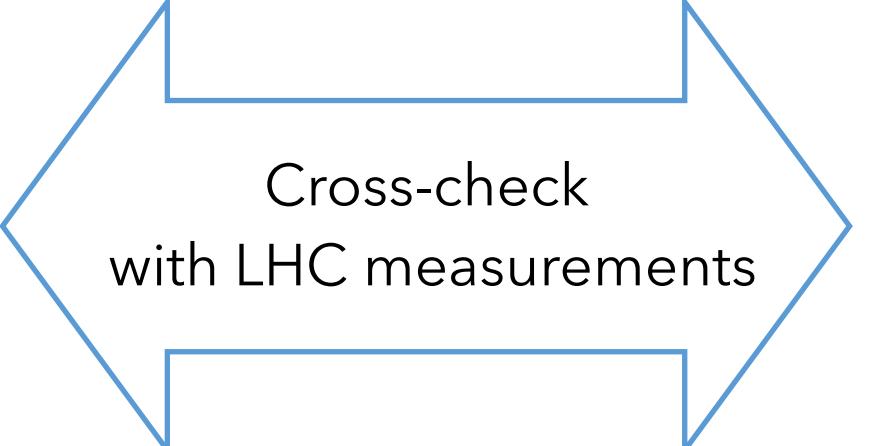
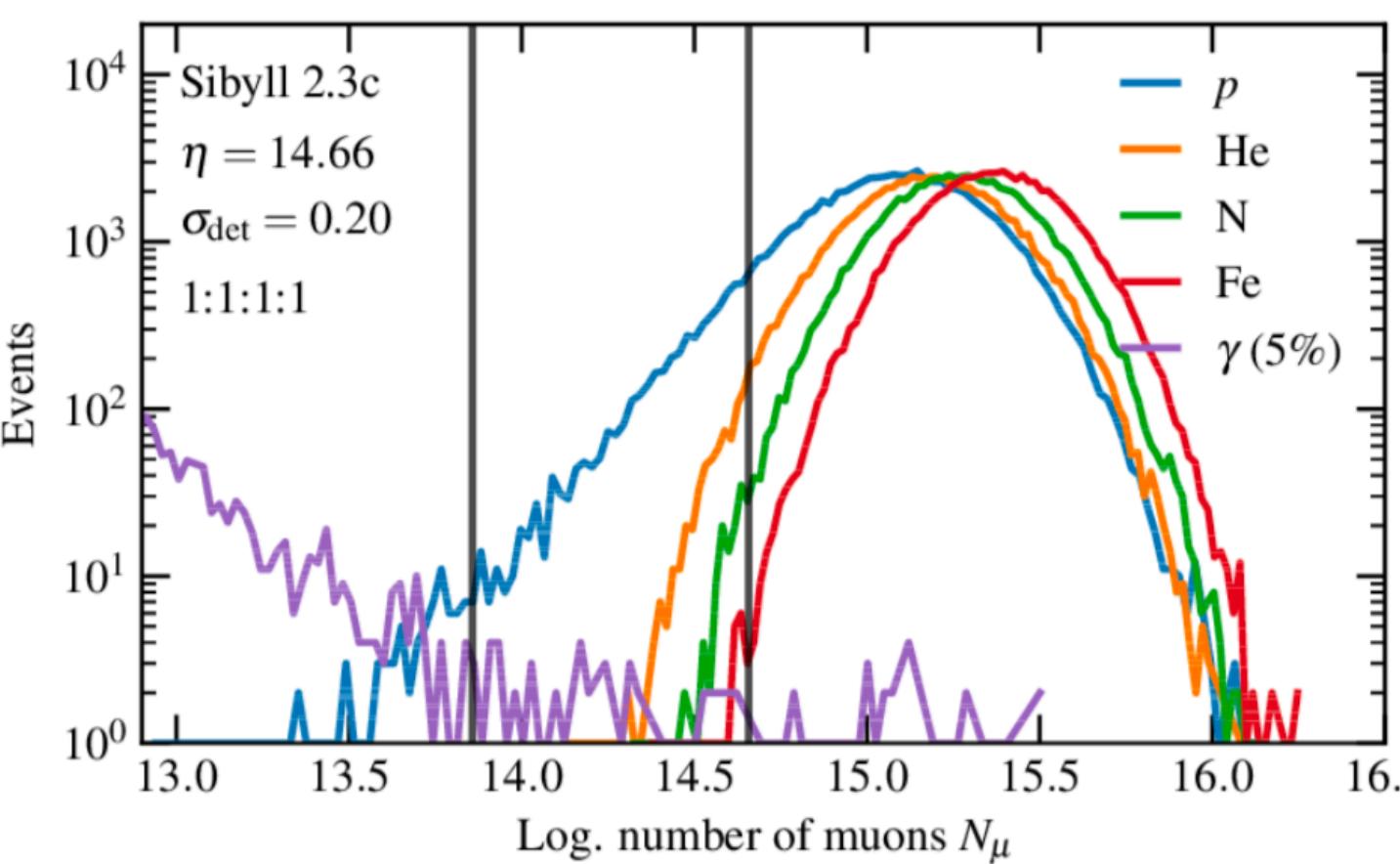
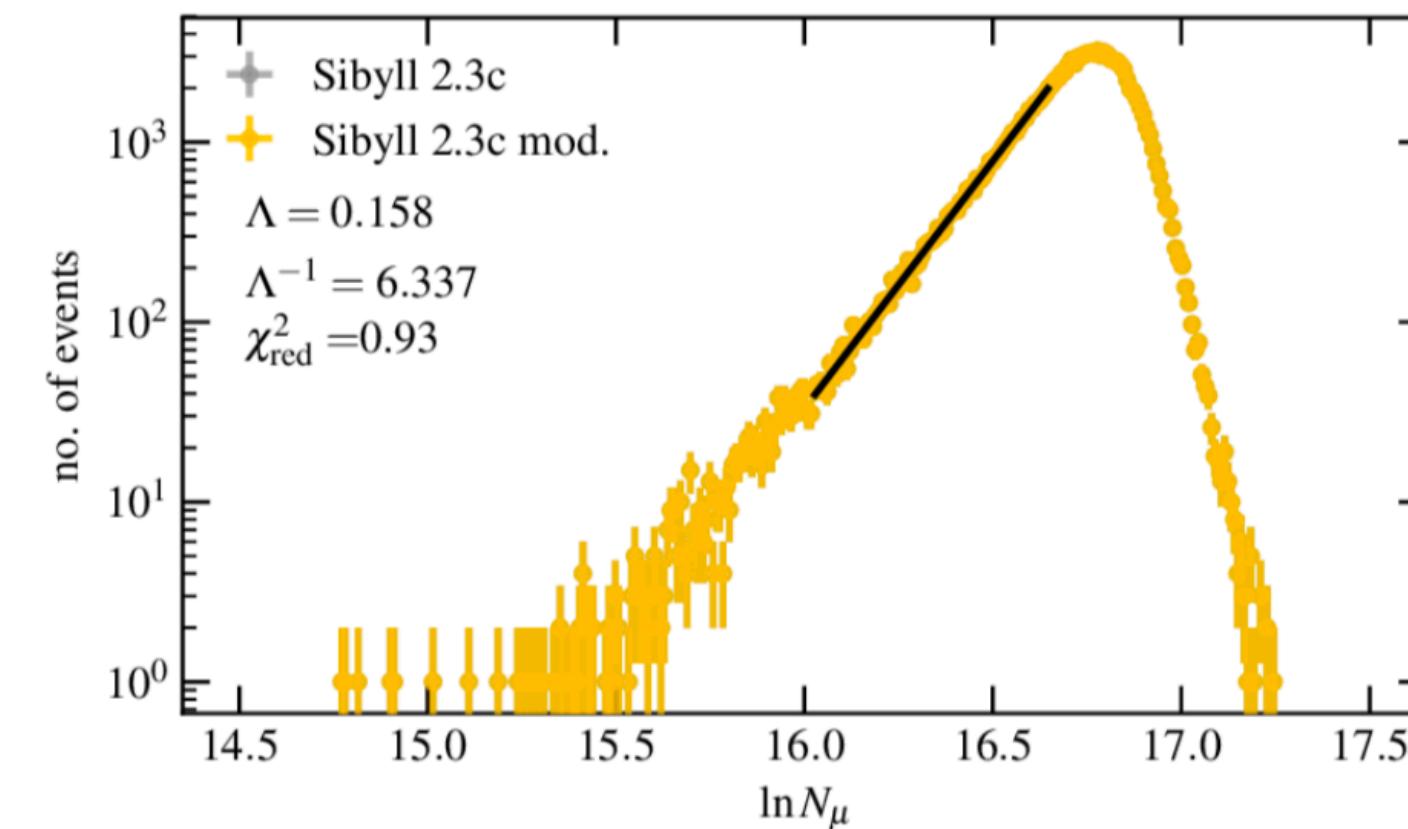
Paper to be published soon...

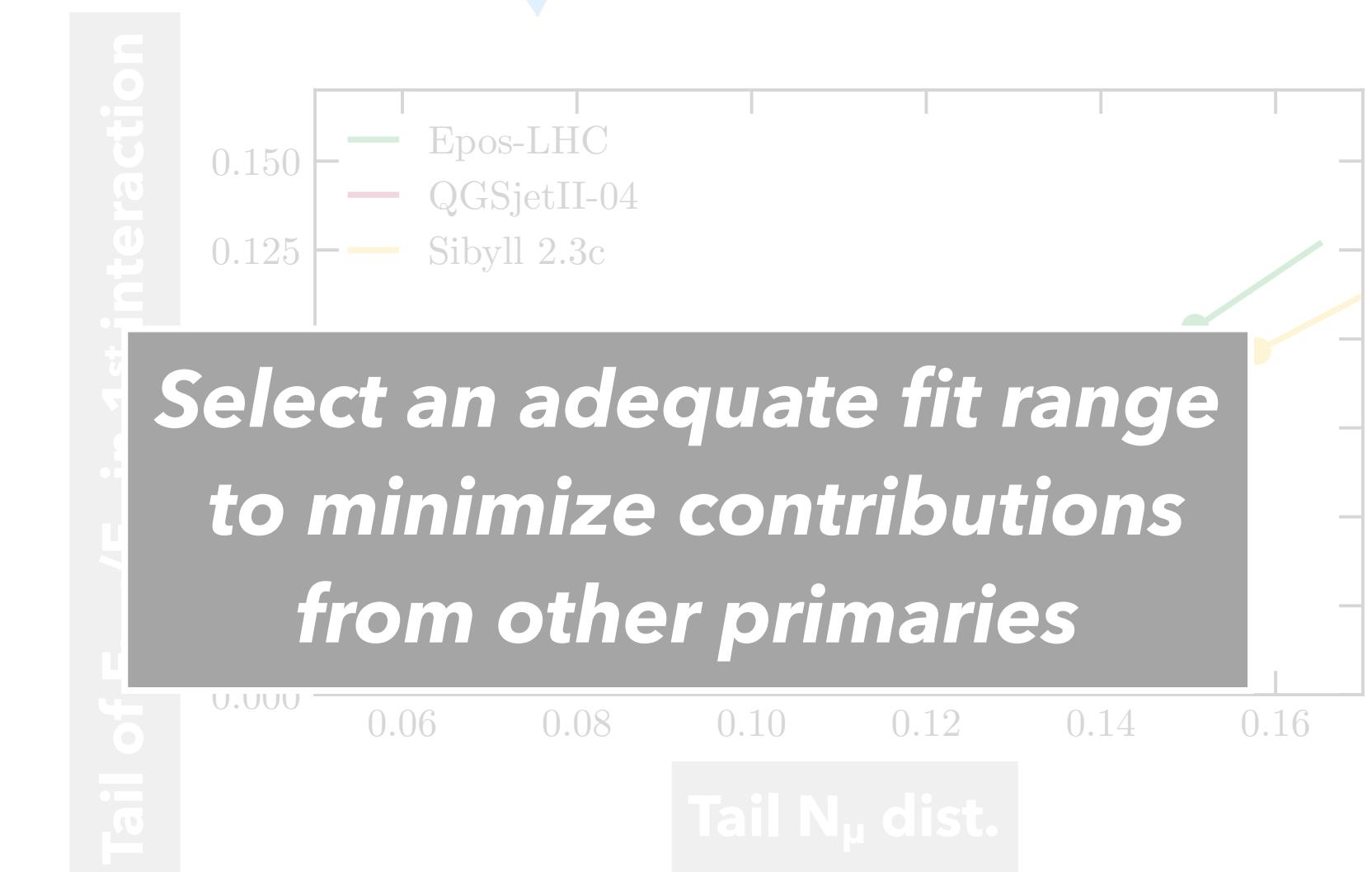
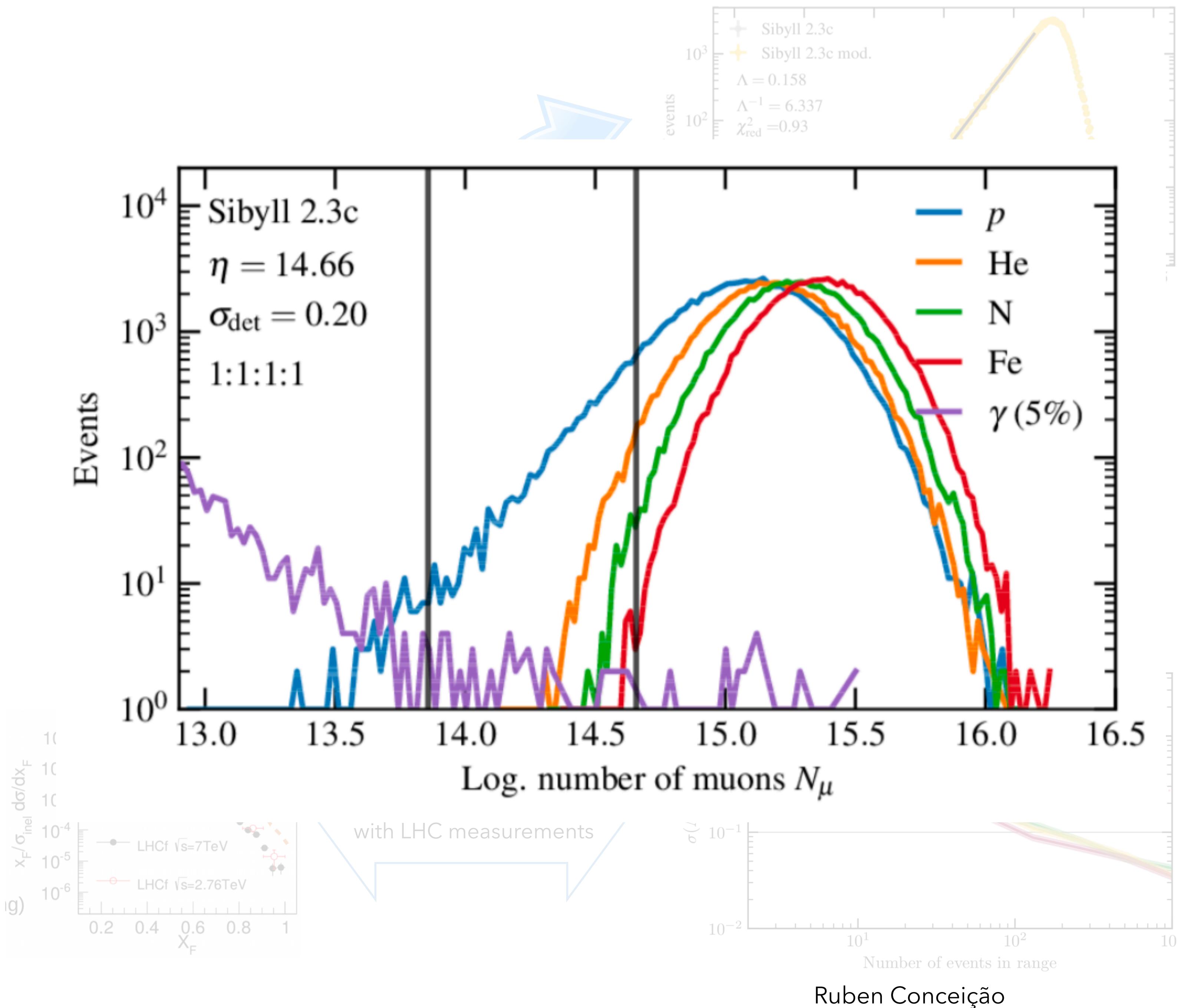


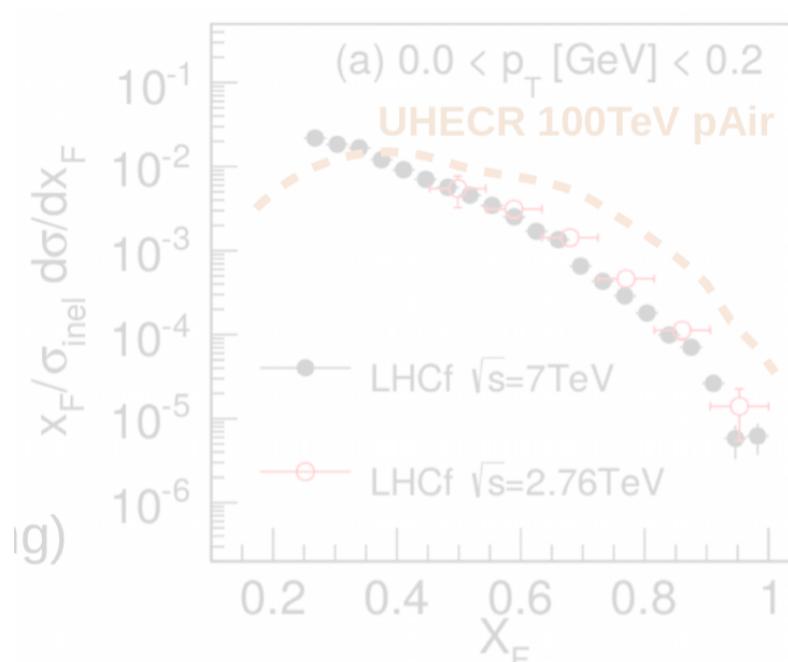
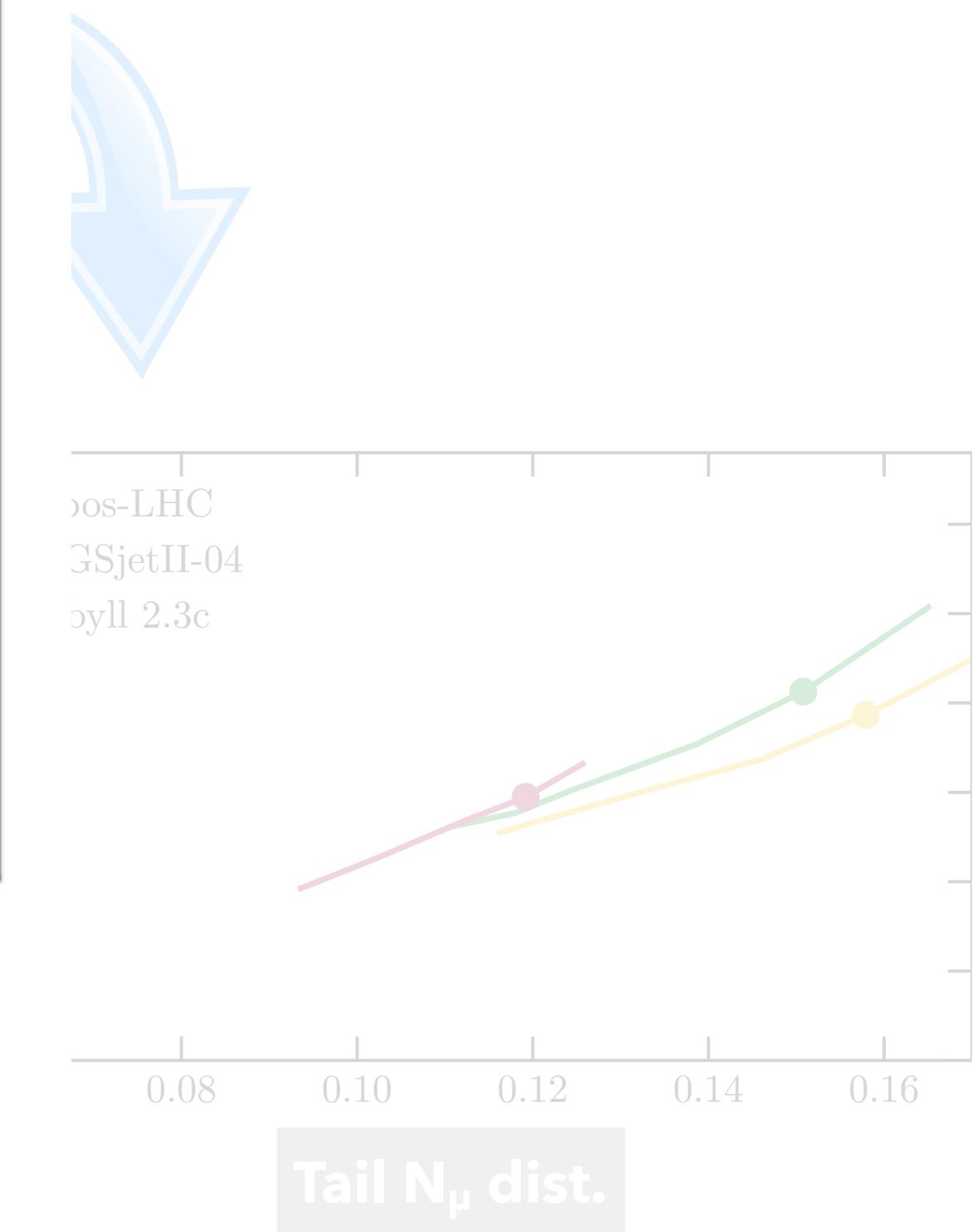
Tail of E_{had}/E_0 in 1st interaction



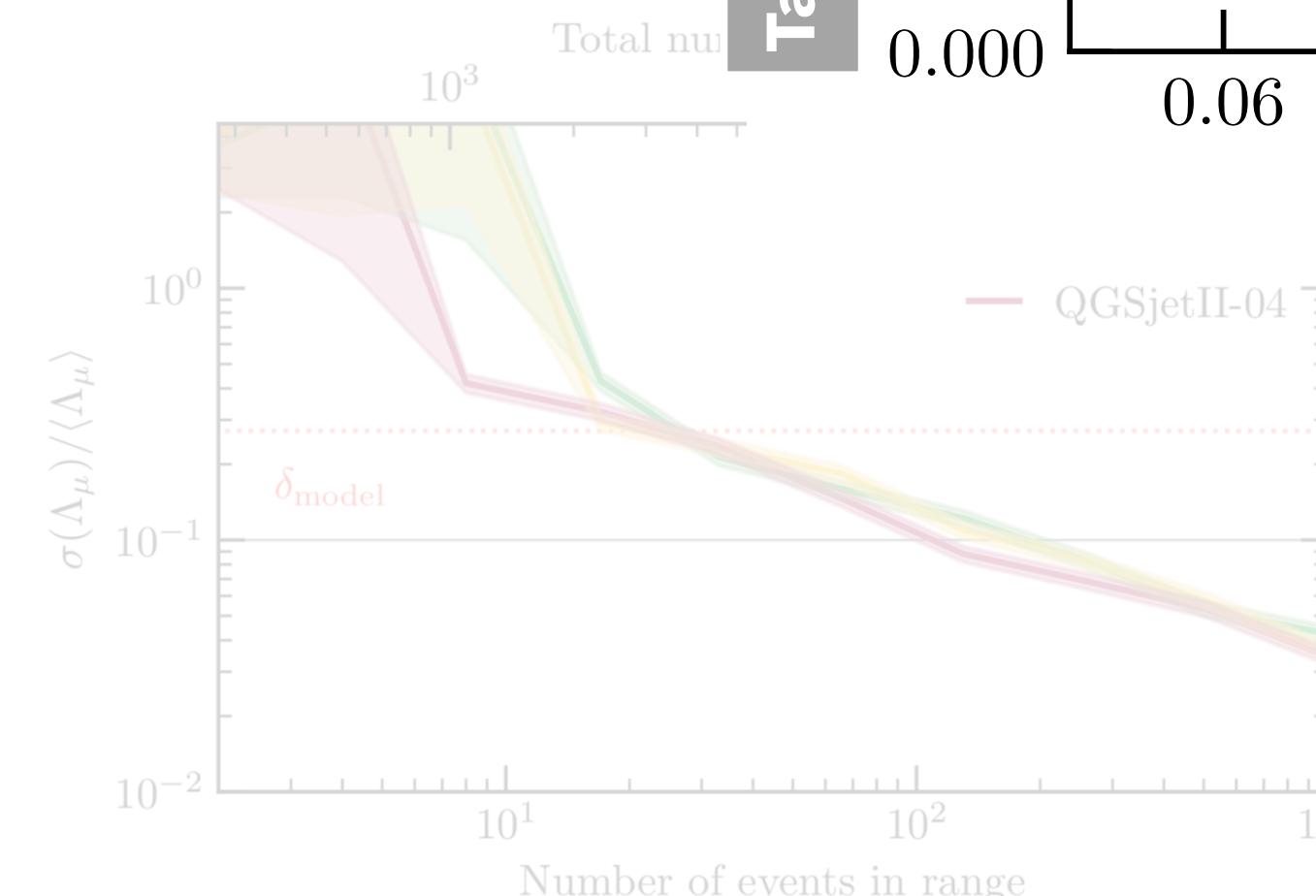
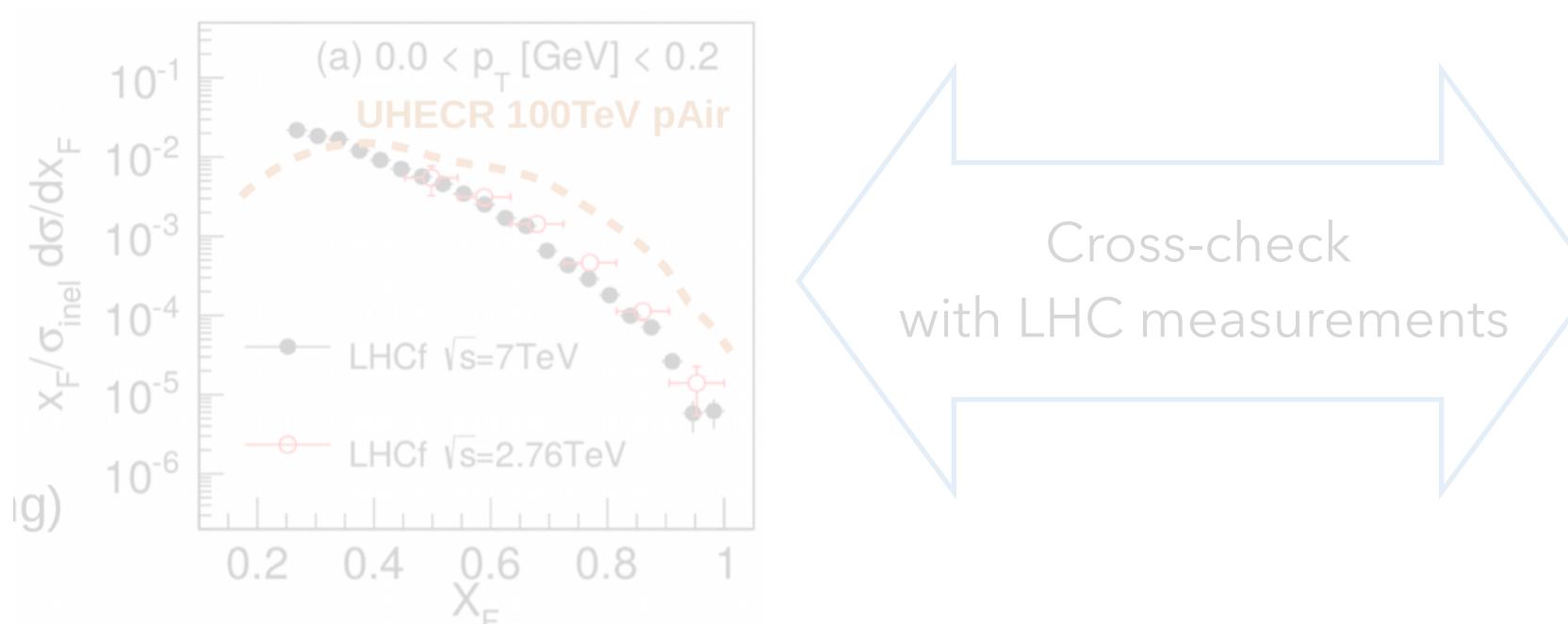
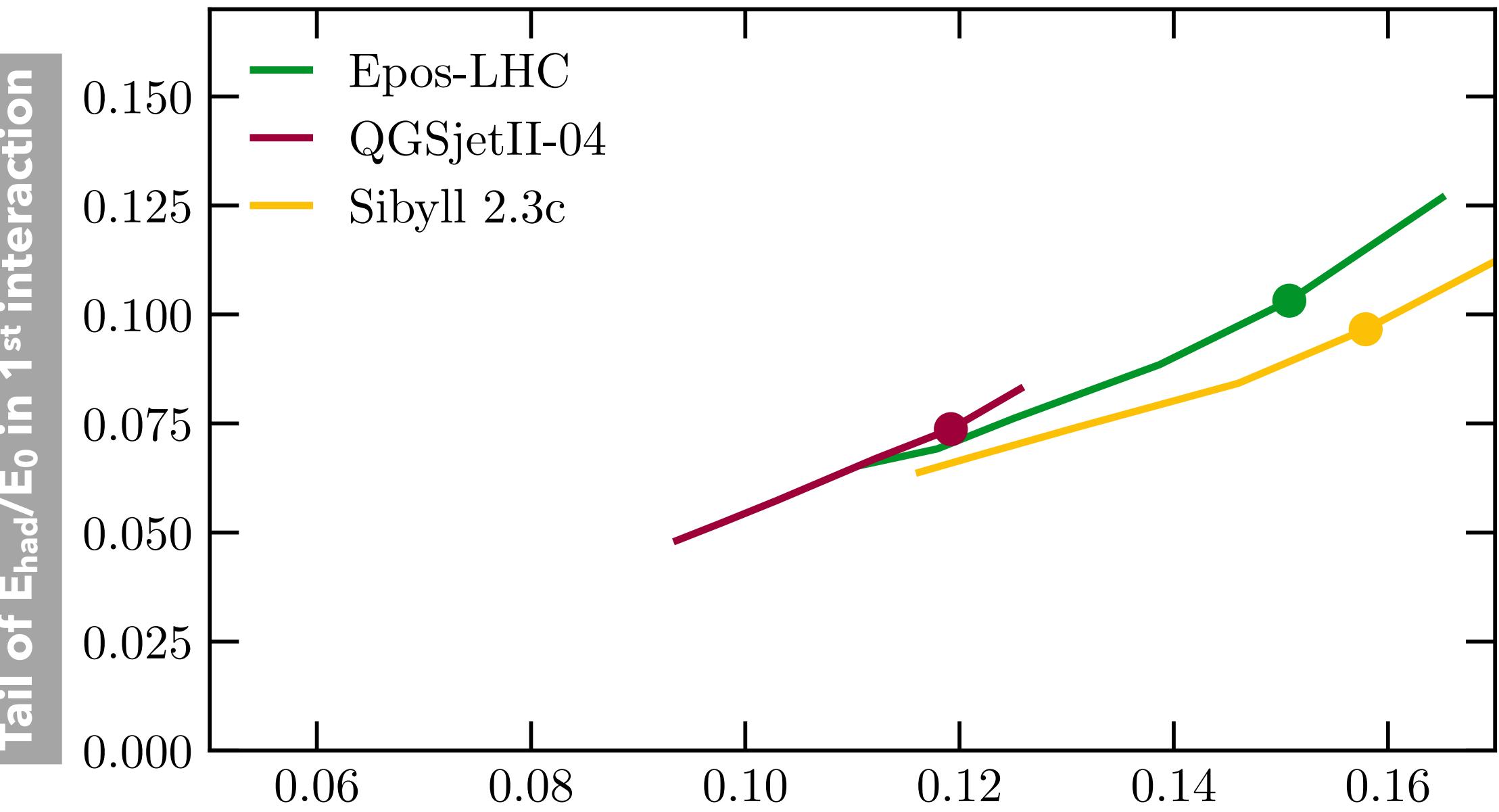
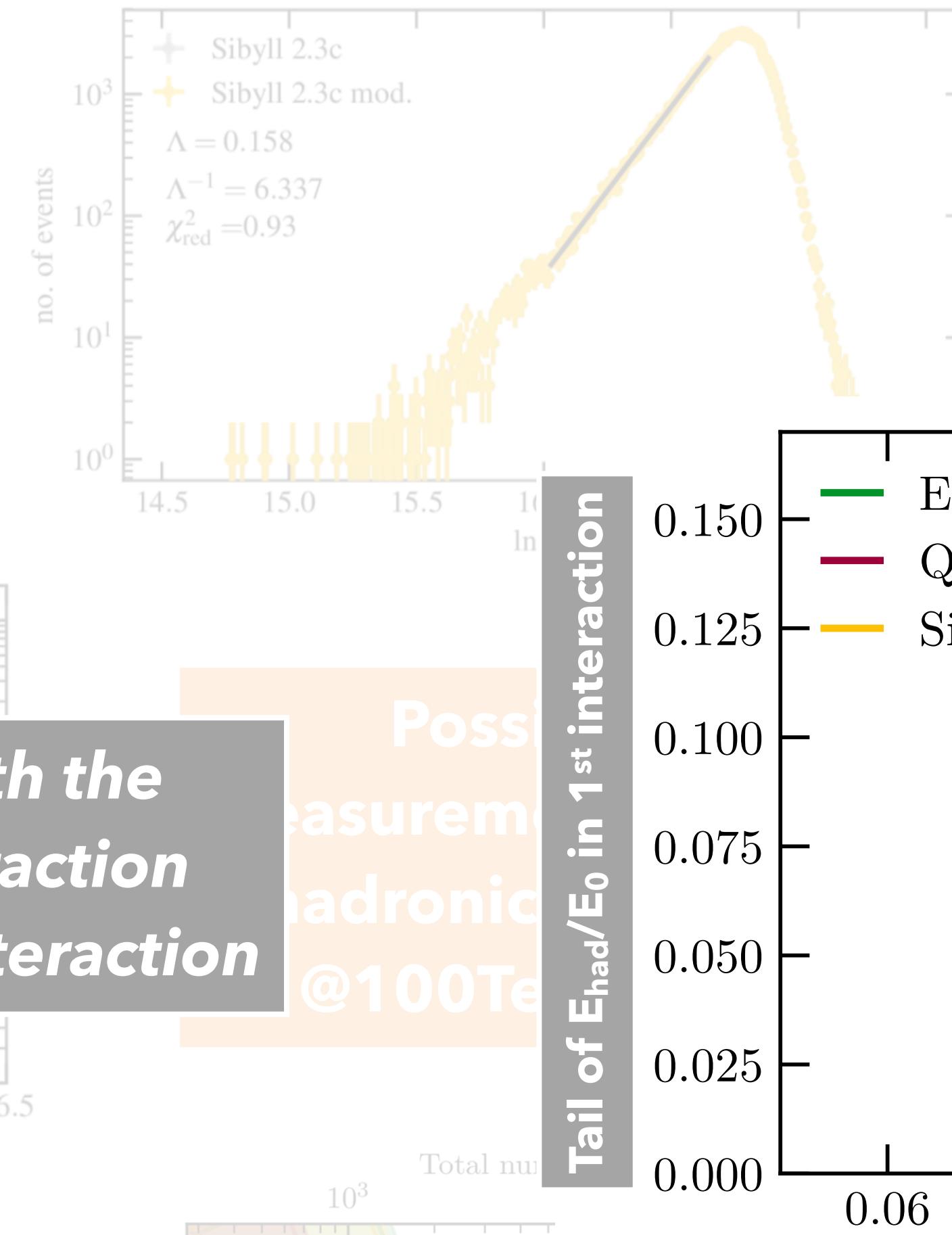
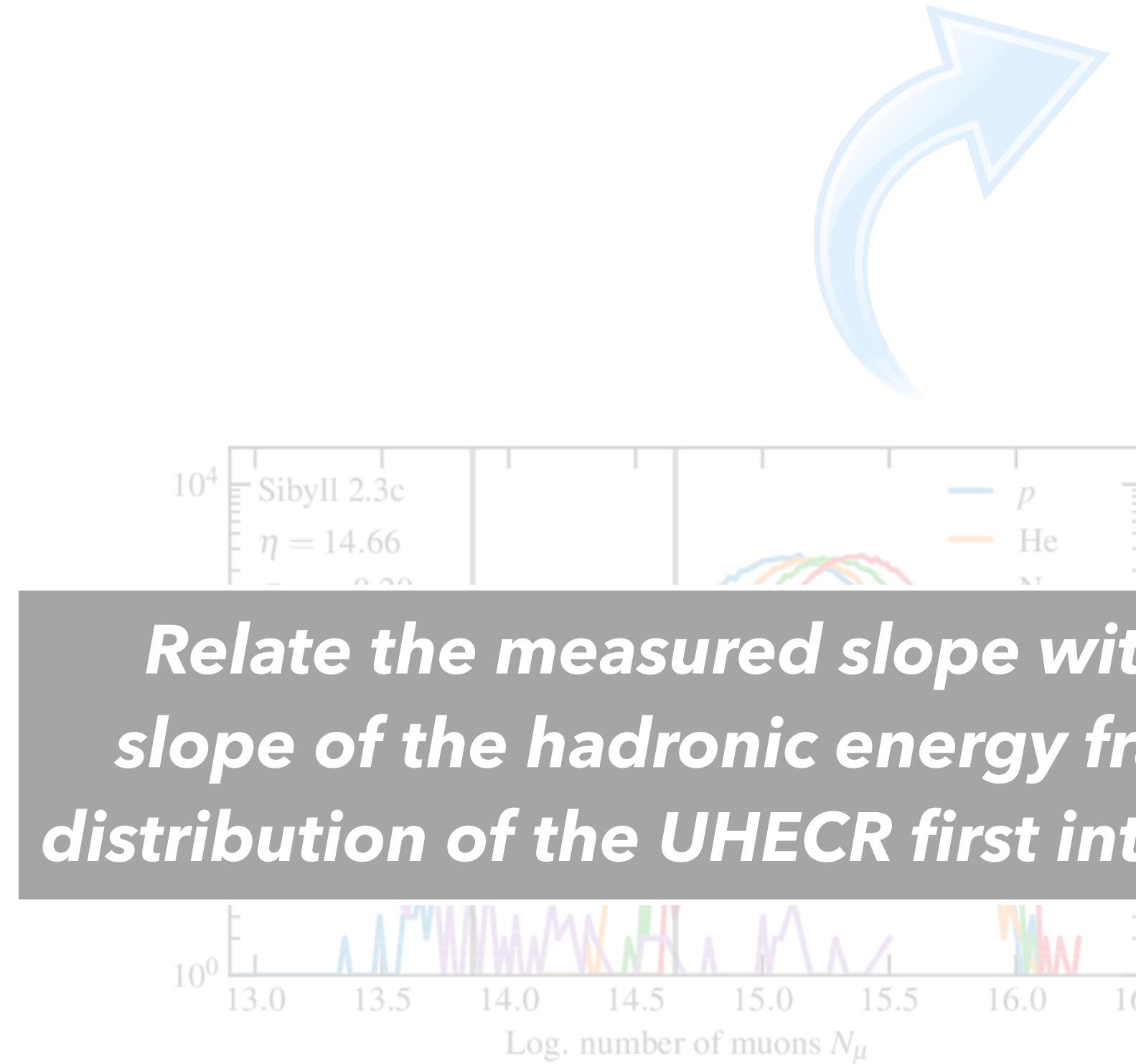
**Possible
measurement of the
hadronic energy
@100TeV c.m.**





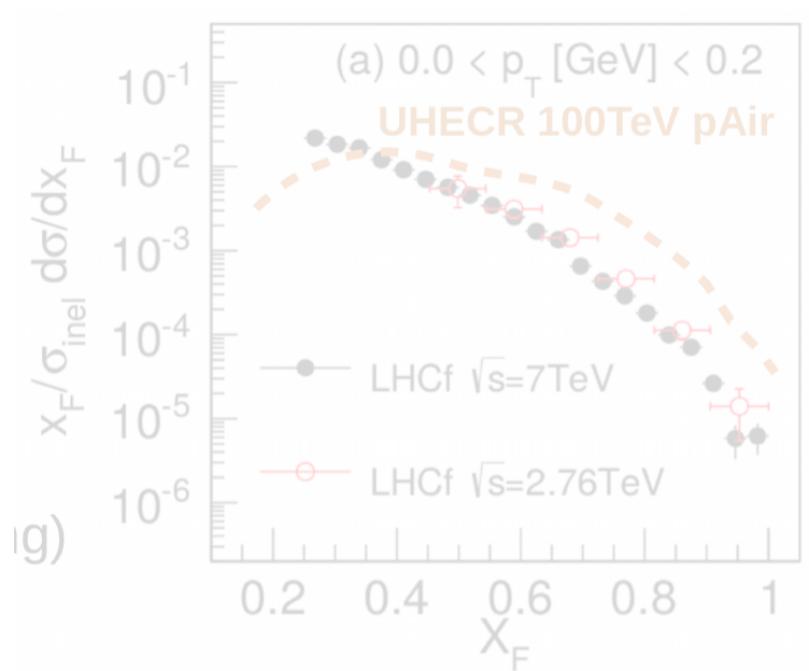
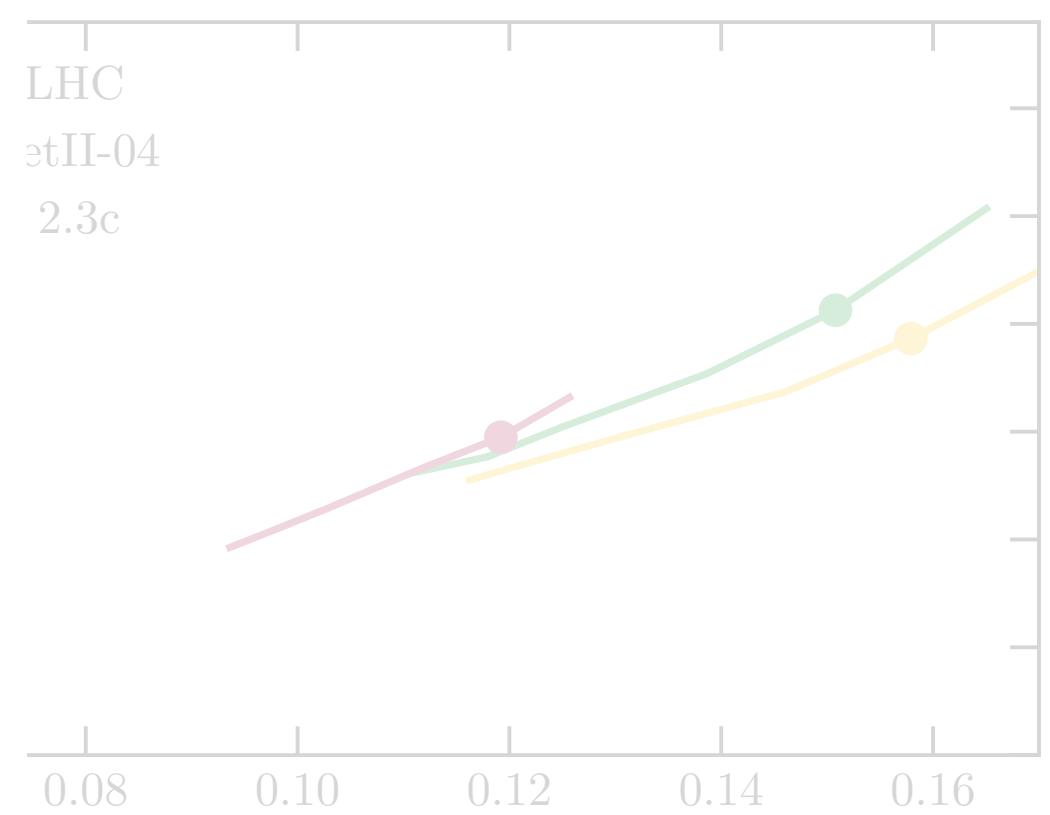
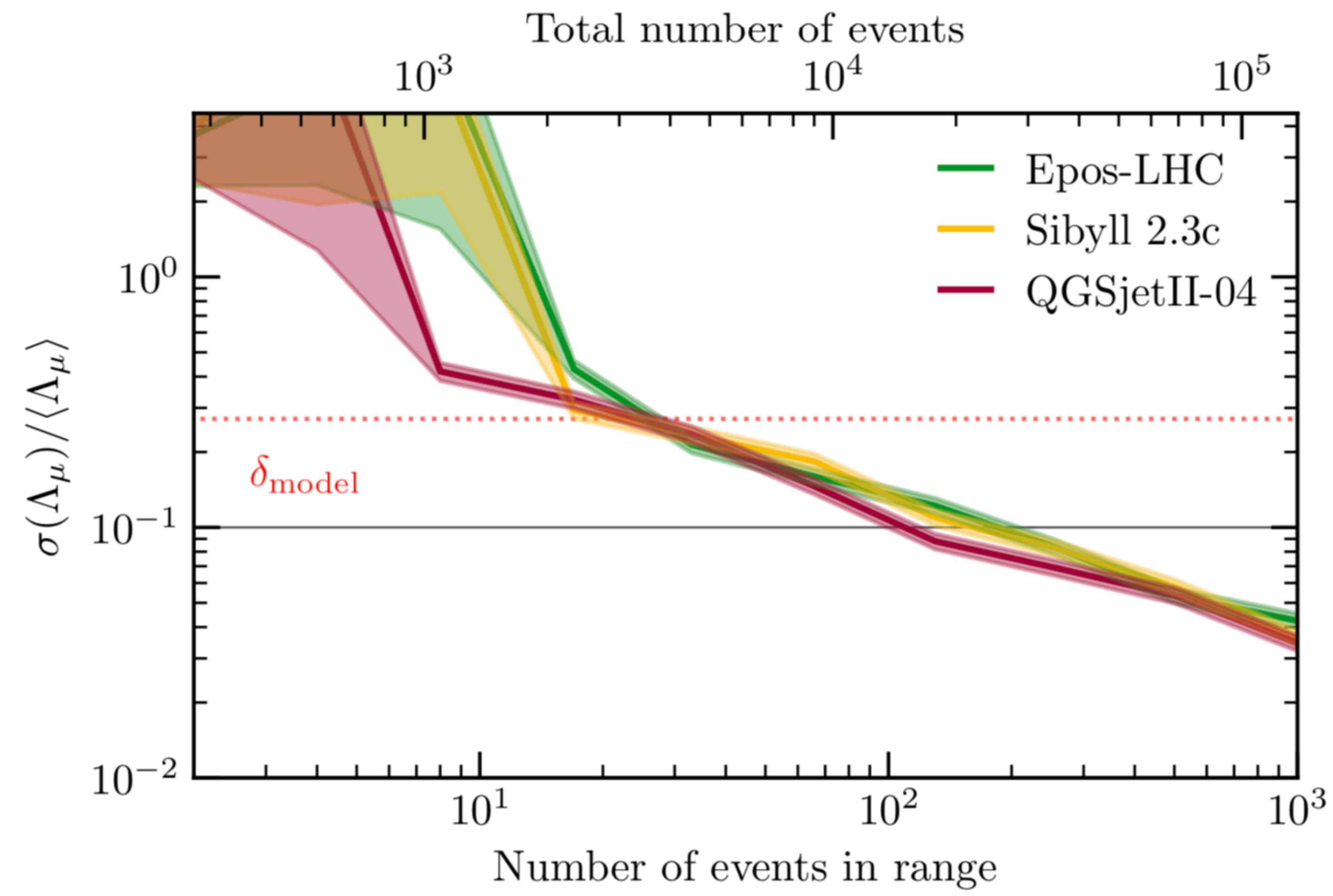
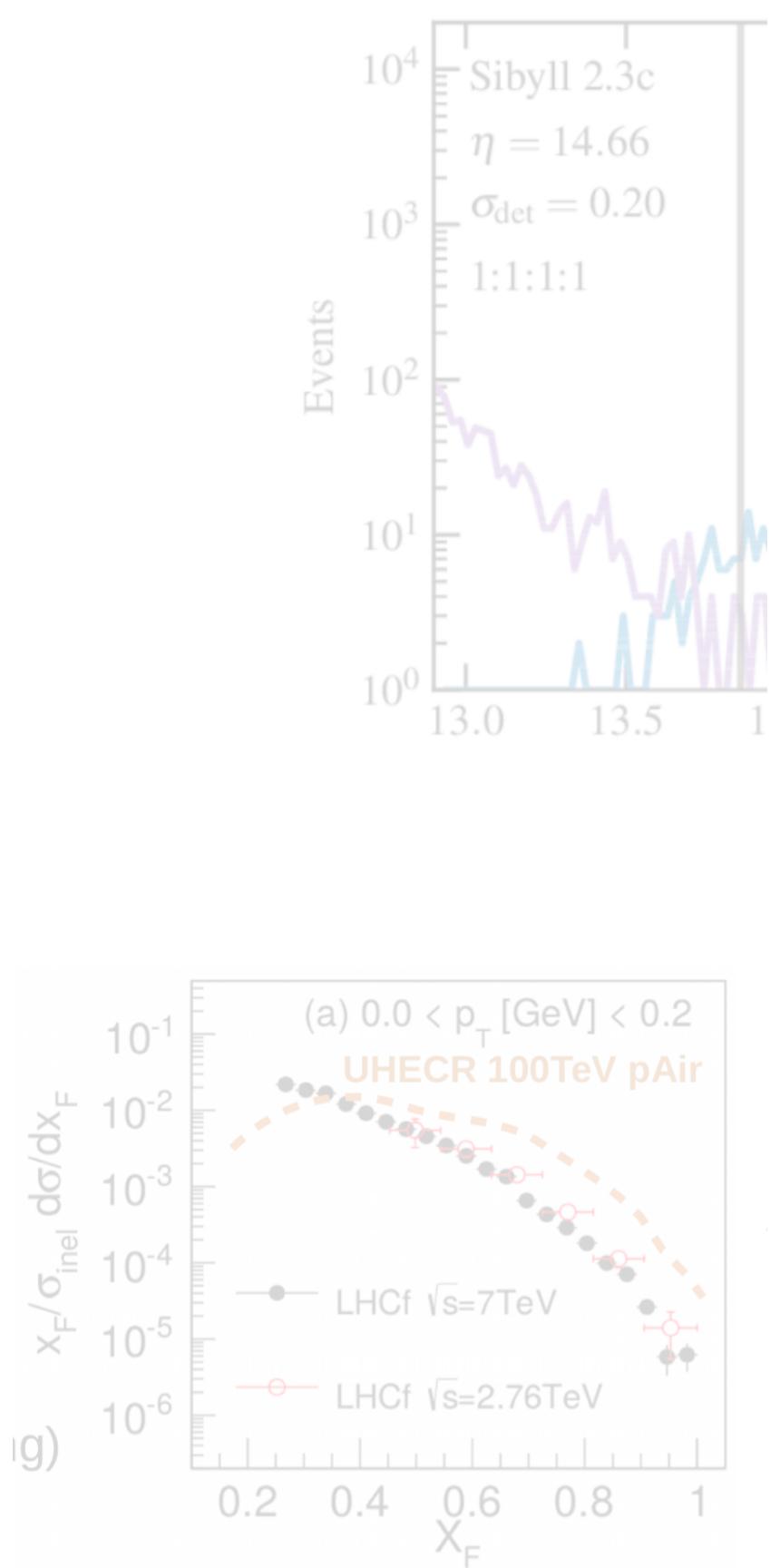


Extract the exponential slope of the muon distribution





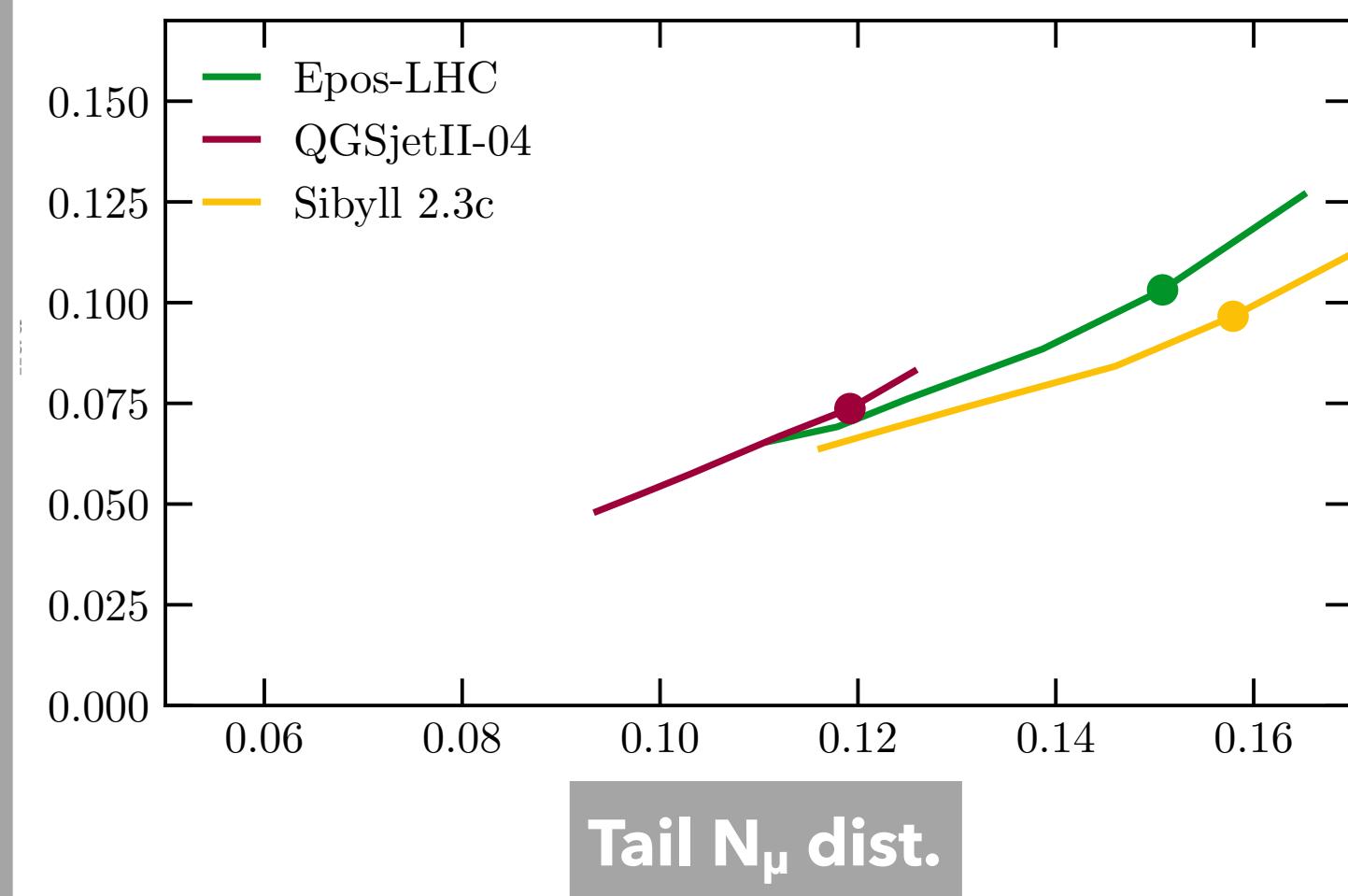
Number of shower events necessary to have a meaningful result



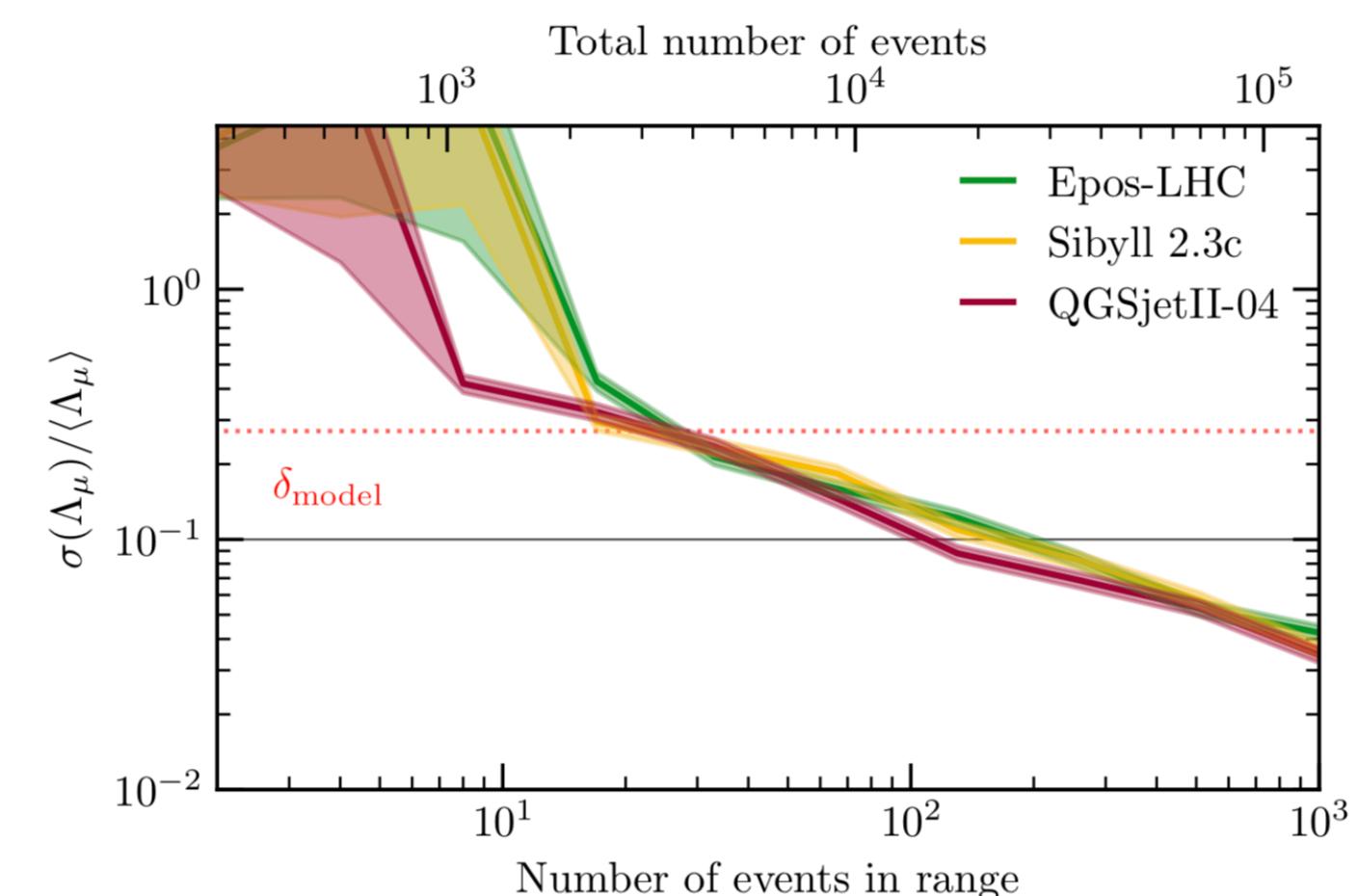
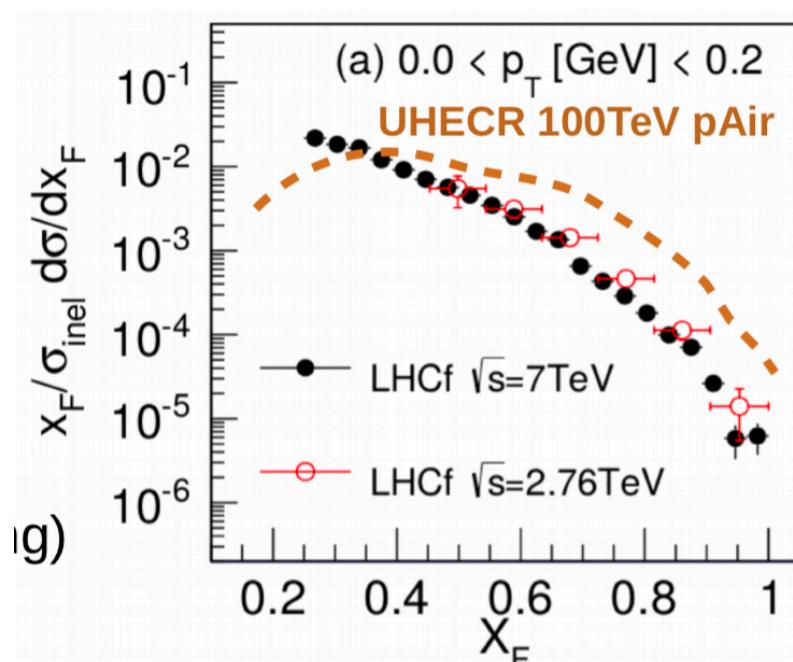
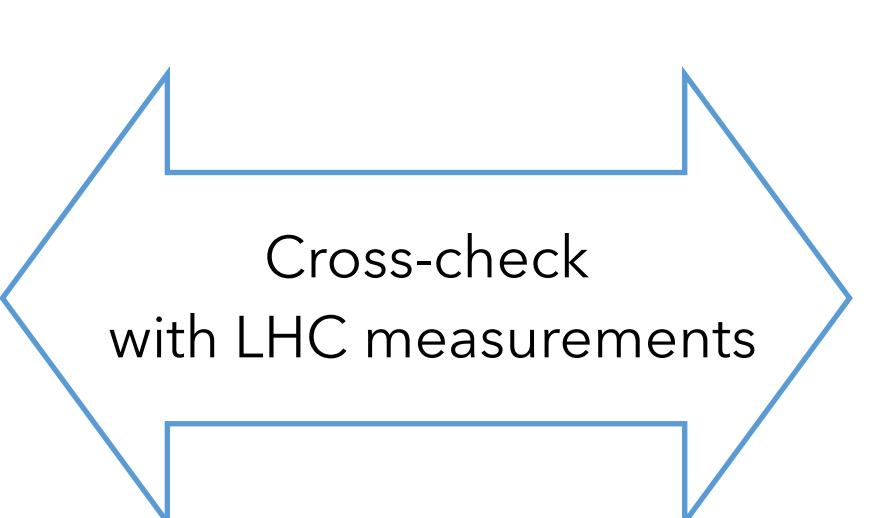
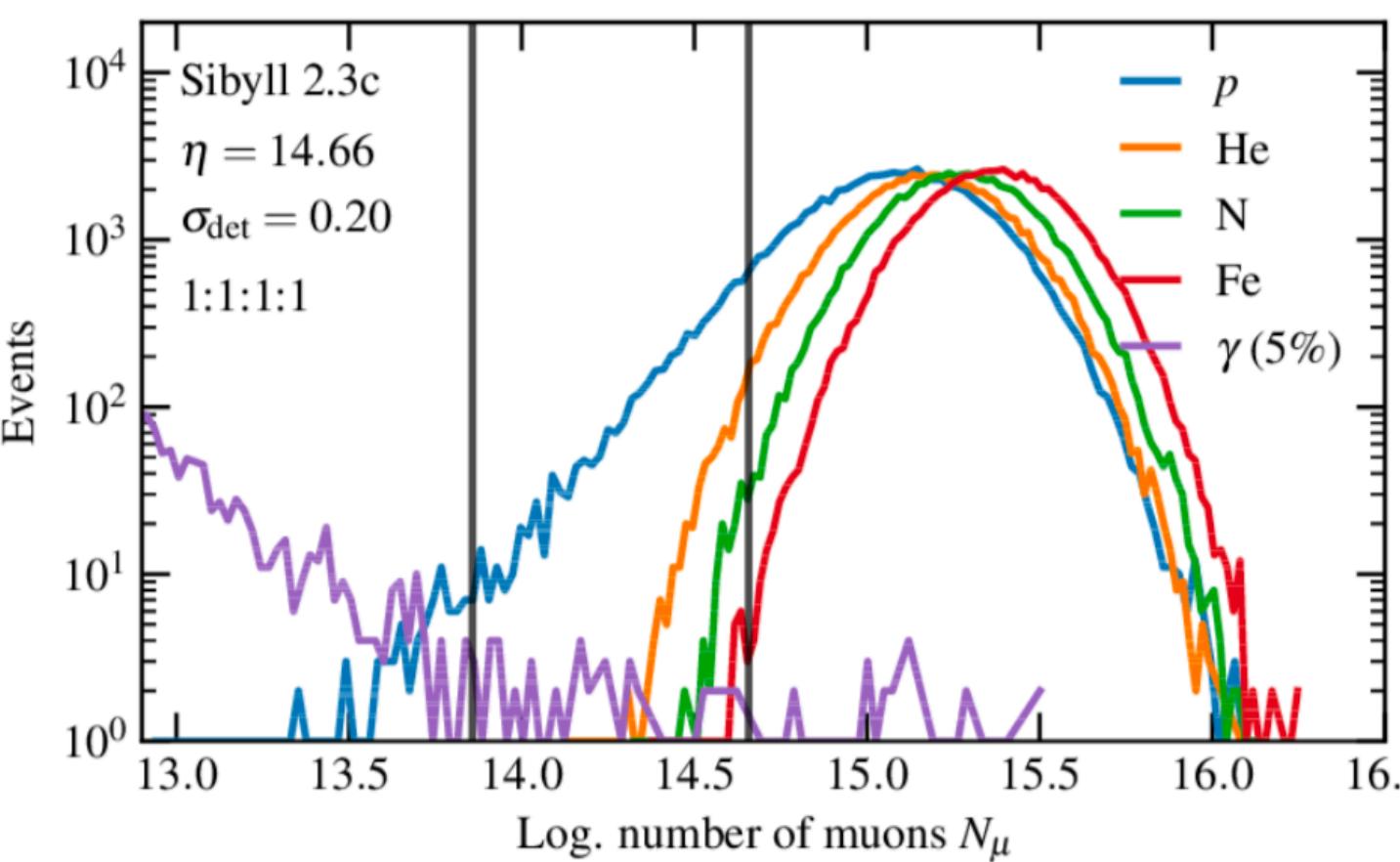
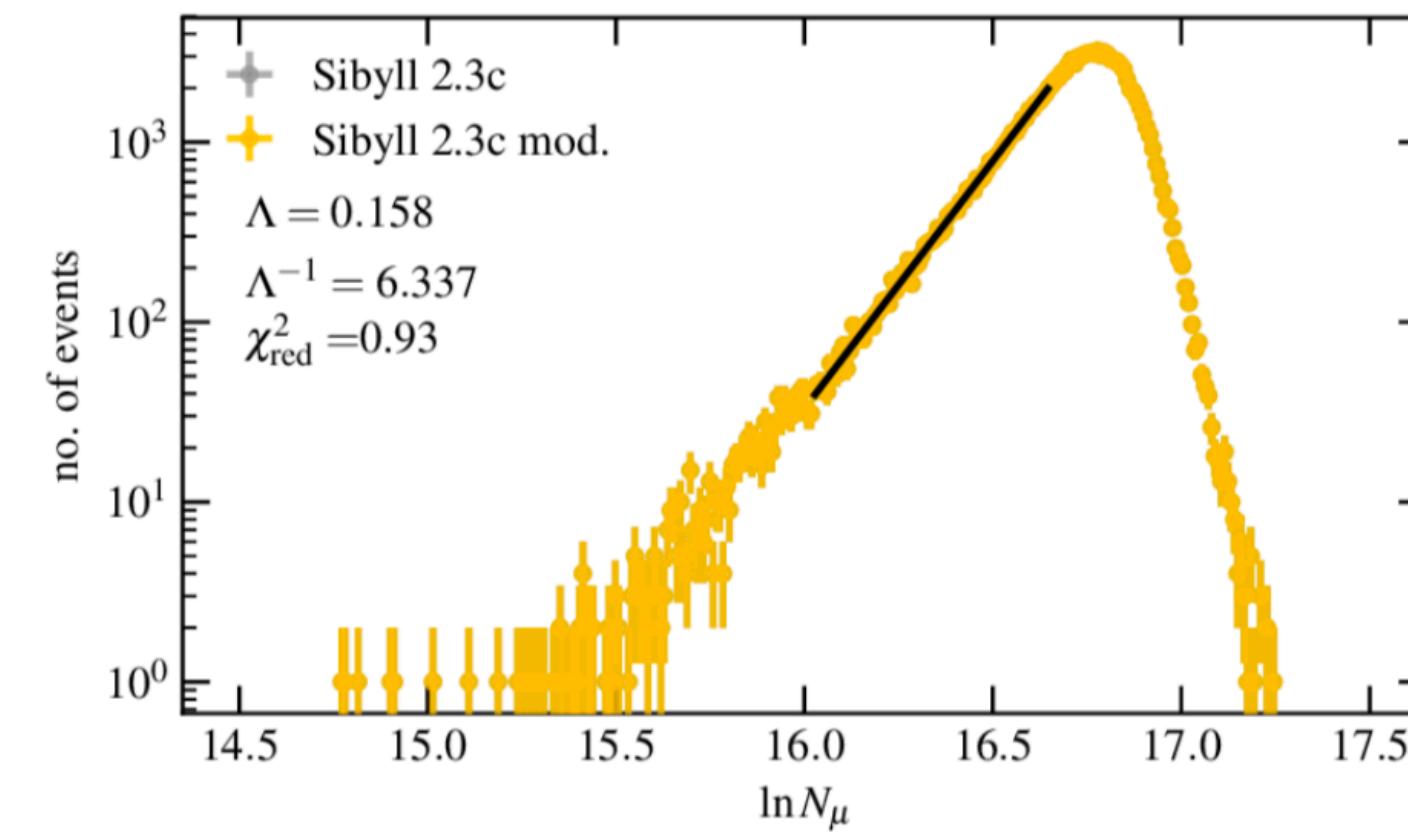
Paper to be published soon...



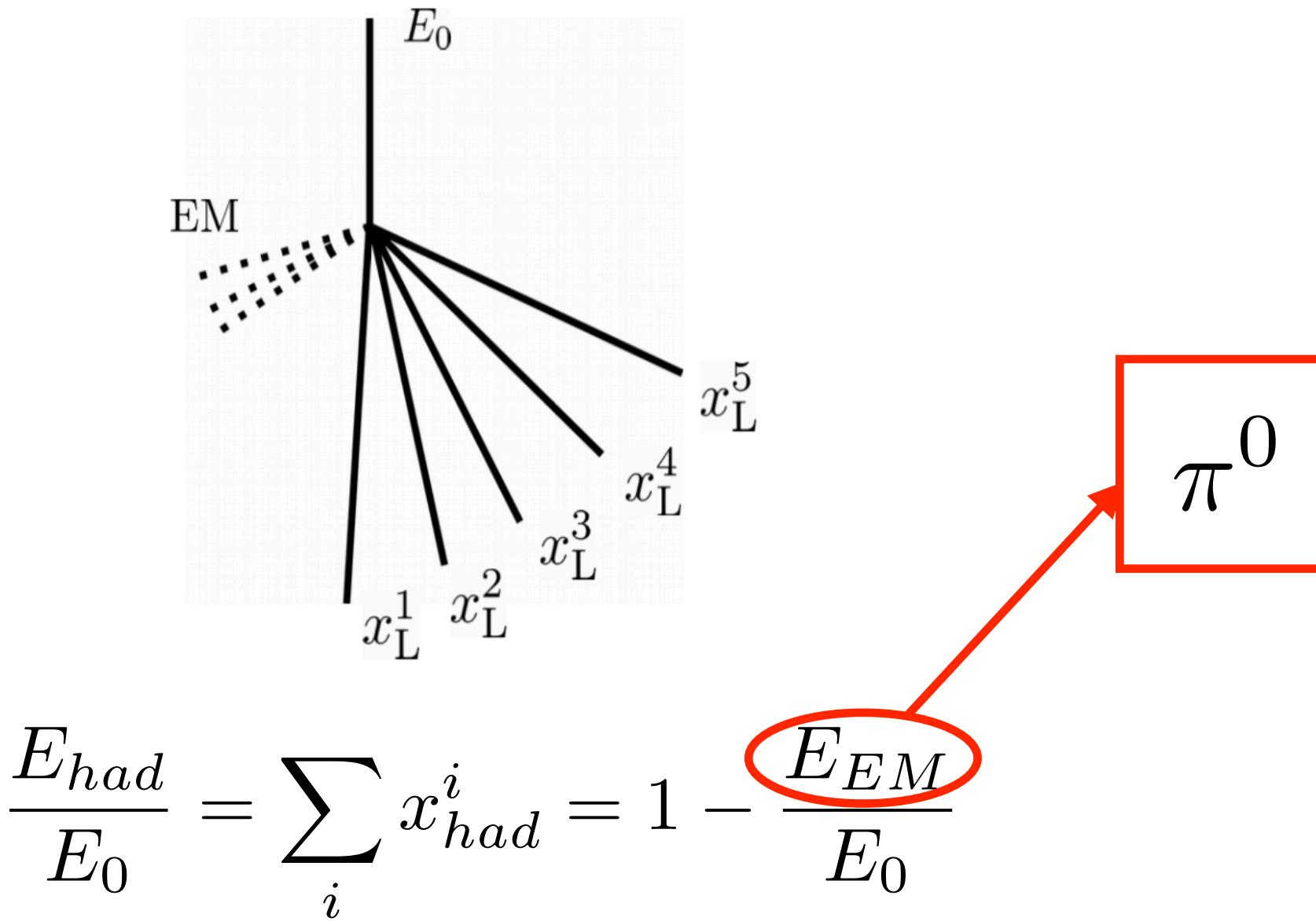
Tail of E_{had}/E_0 in 1st interaction



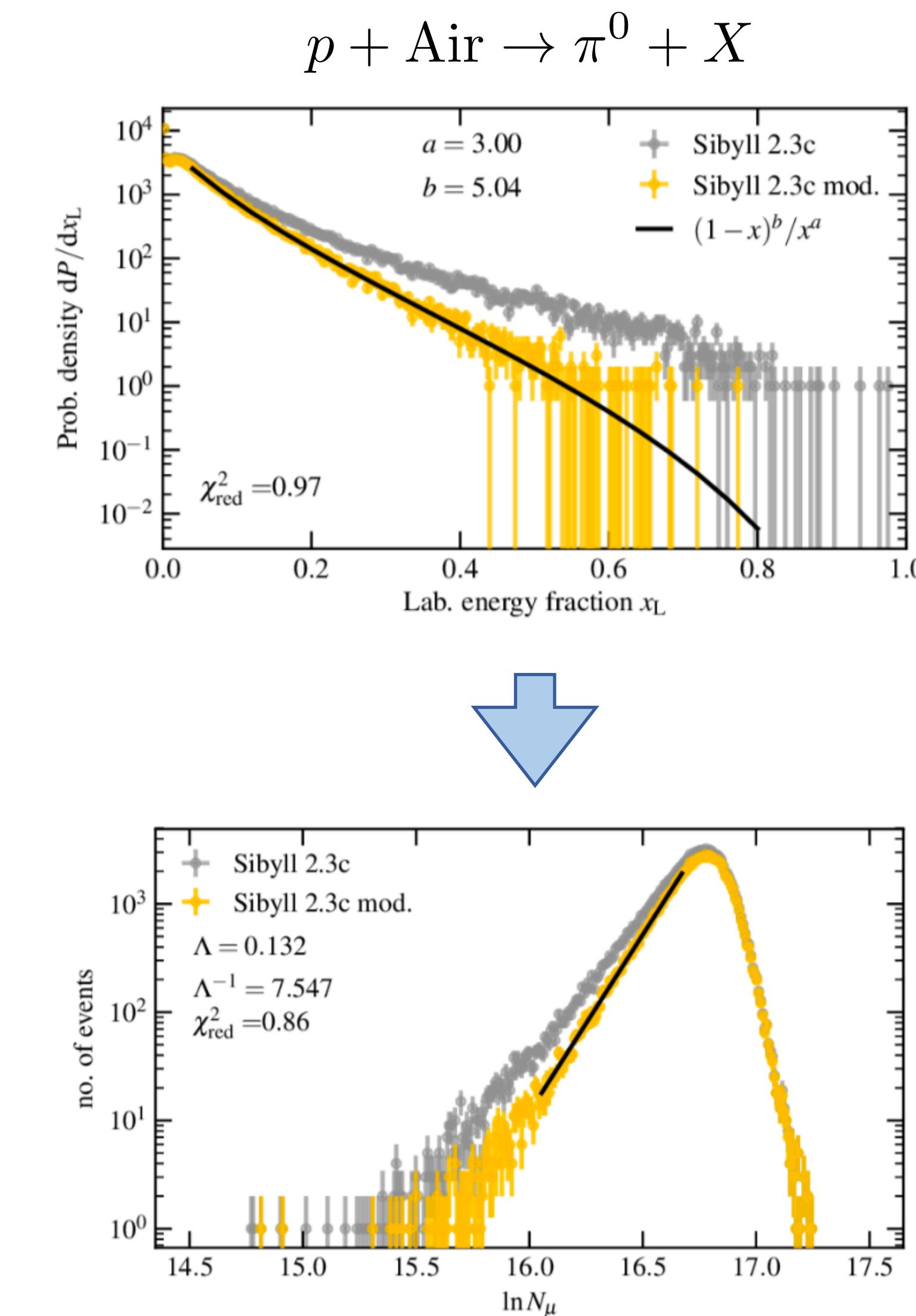
**Possible
measurement of the
hadronic energy
@100TeV c.m.**



Access the pion energy spectrum of the first interaction



- ✧ Change the high-energy tail of the inclusive production cross-section
- ✧ Check impact in the slope of muon distribution low-tail



Slope of muon distribution \Leftrightarrow Shape of inclusive production cross-section at high- x

Summary

- ❖ Fluctuations of the N_μ distribution are sensitive to the first interaction properties
- ❖ *PLB 784 (2018) 68-76*
- ❖ The slope of the low-tail of the proton N_μ distribution is sensitive to the pion energy spectrum of the first interaction
- ❖ *Paper to be published soon*

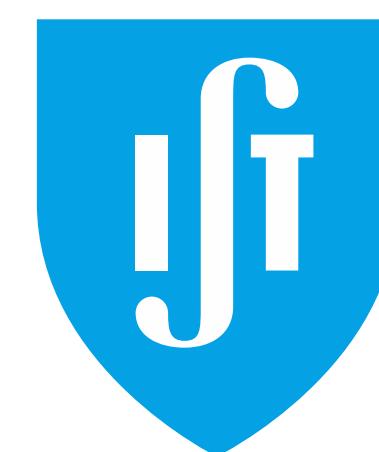
Acknowledgements



Fundação para a Ciência e a Tecnologia
MINISTÉRIO DA EDUCAÇÃO E CIÊNCIA



REPÚBLICA
PORTUGUESA

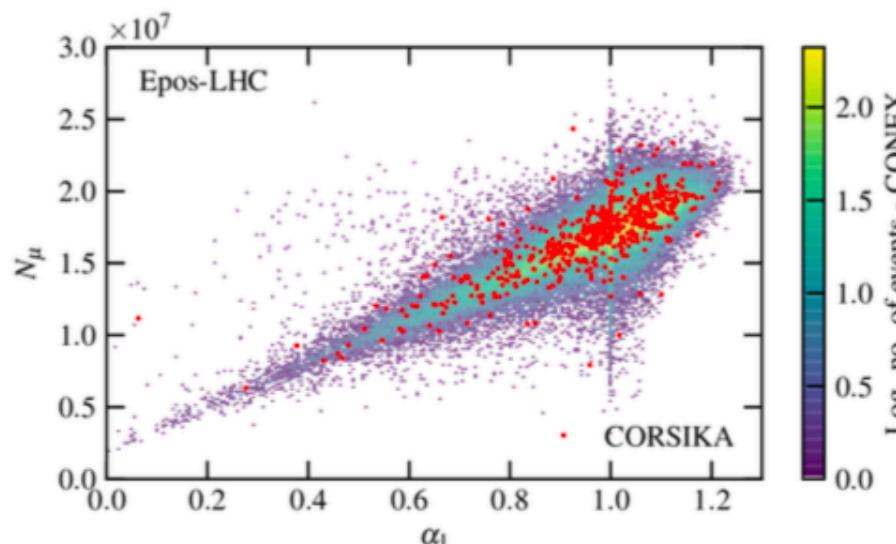


TÉCNICO
LISBOA

Backup slides

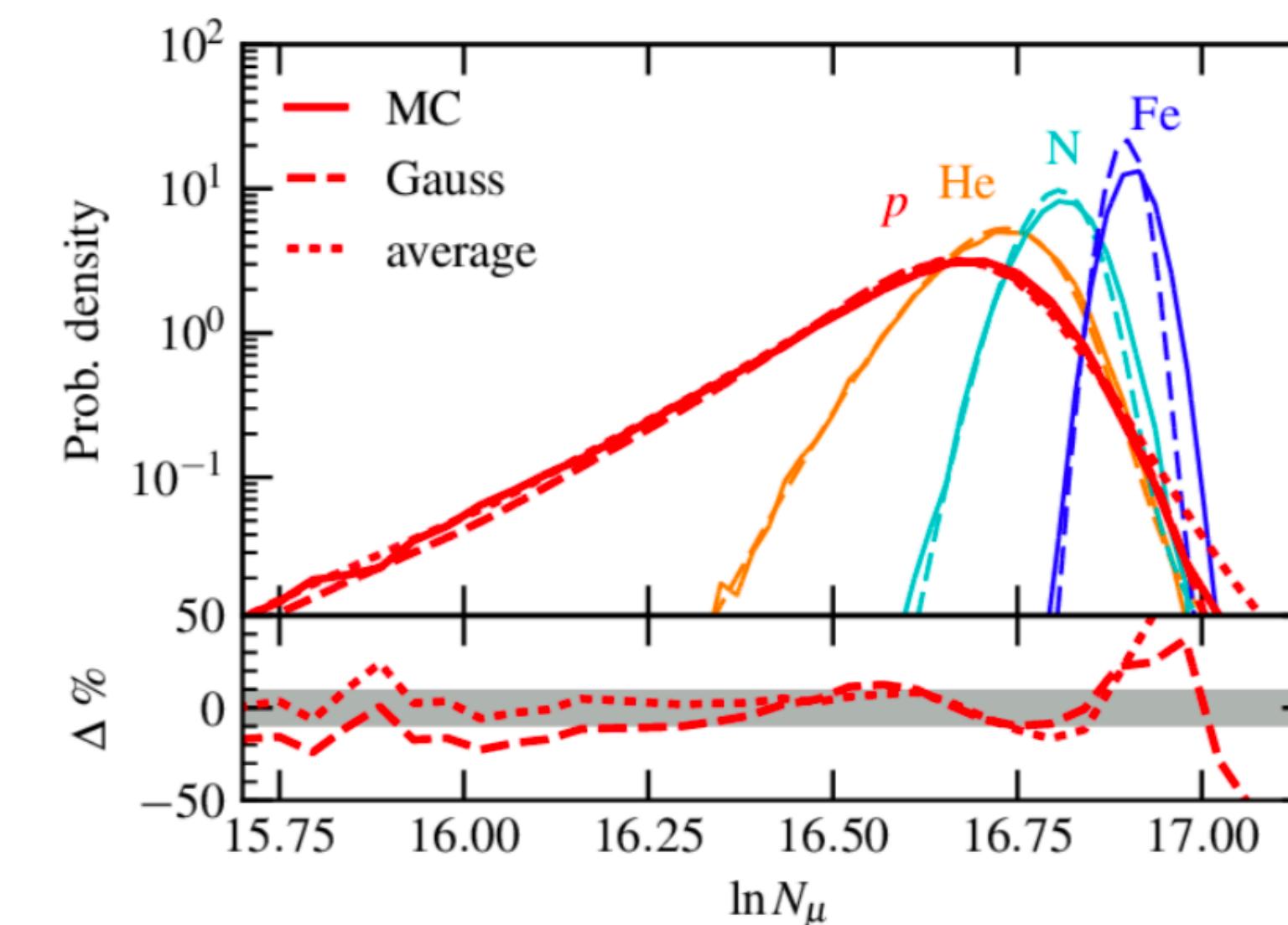
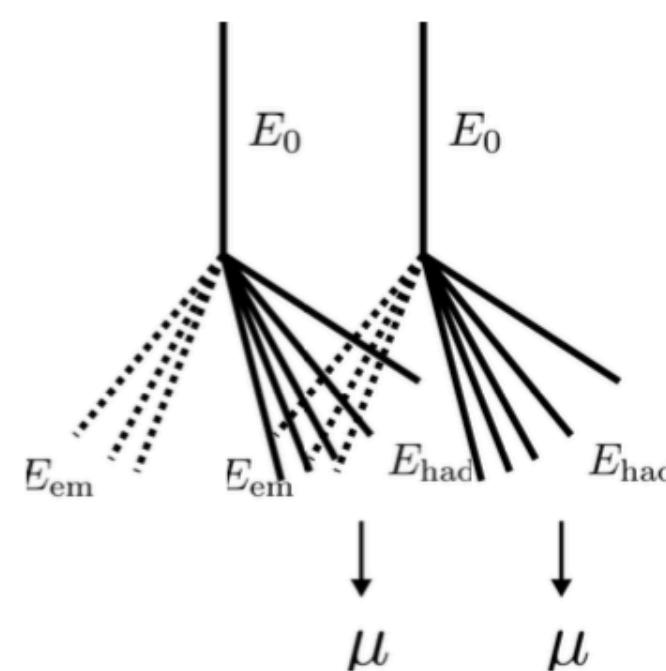
Different primaries

* Energy fluctuations p-Air

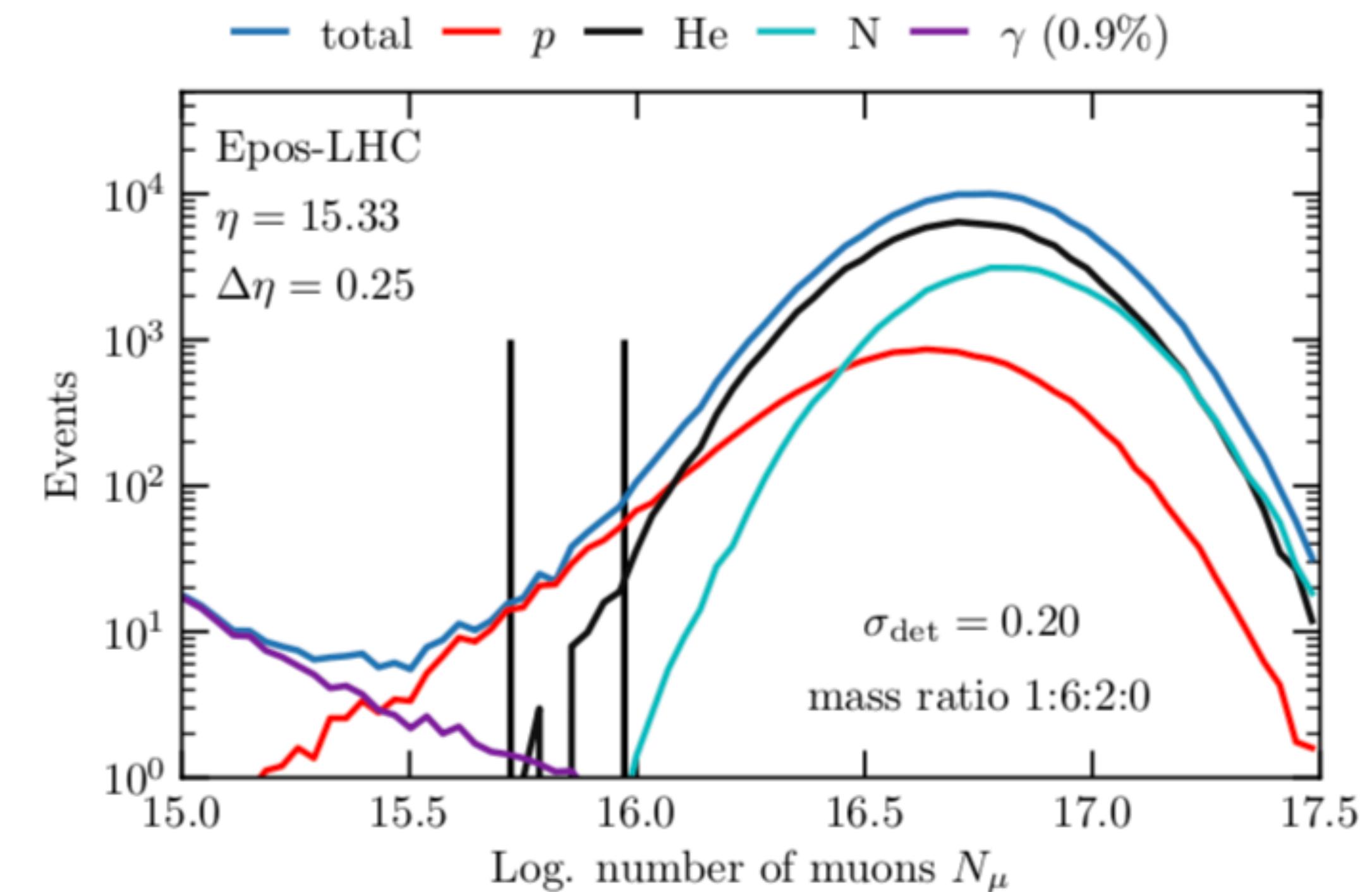
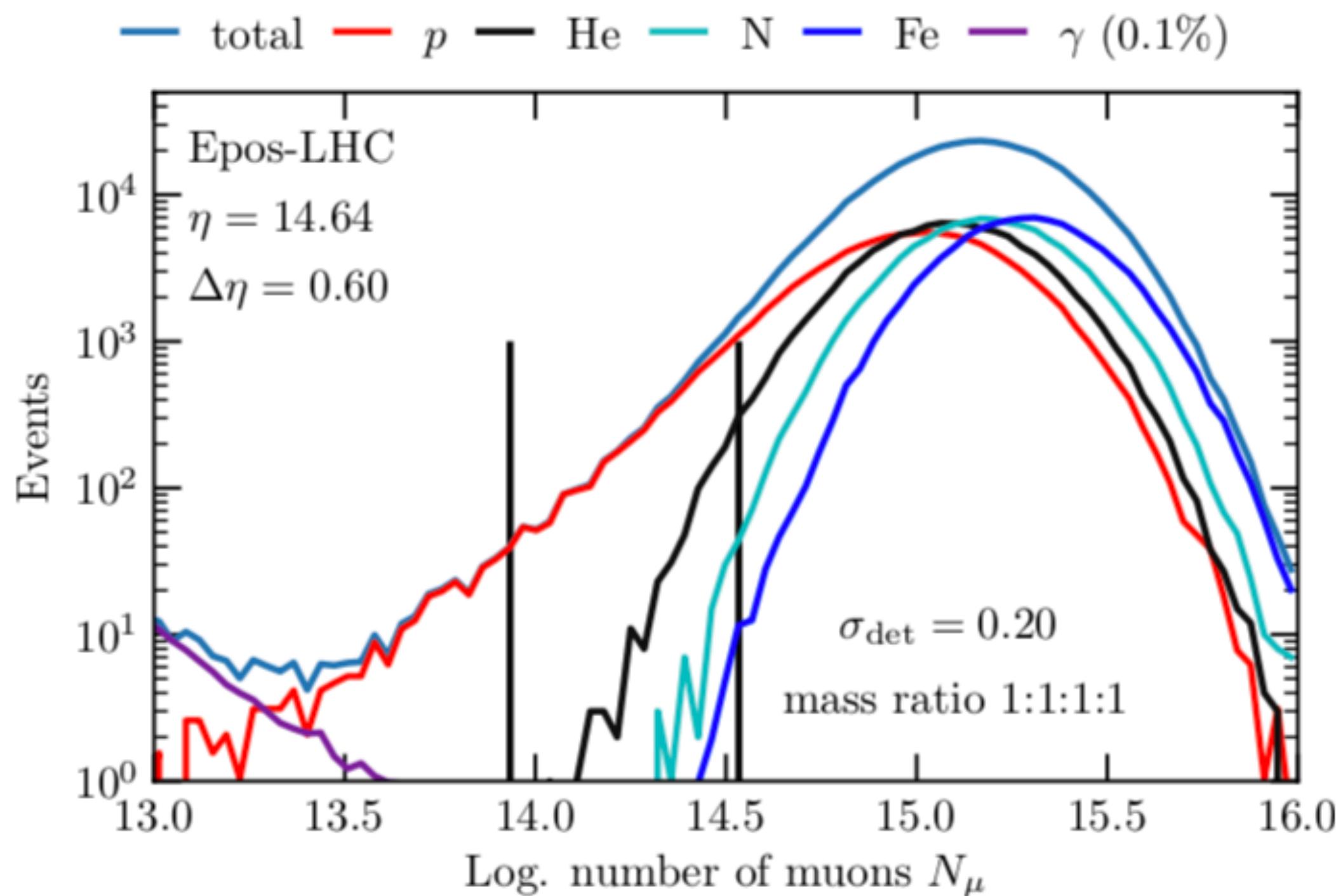


* Superposition model

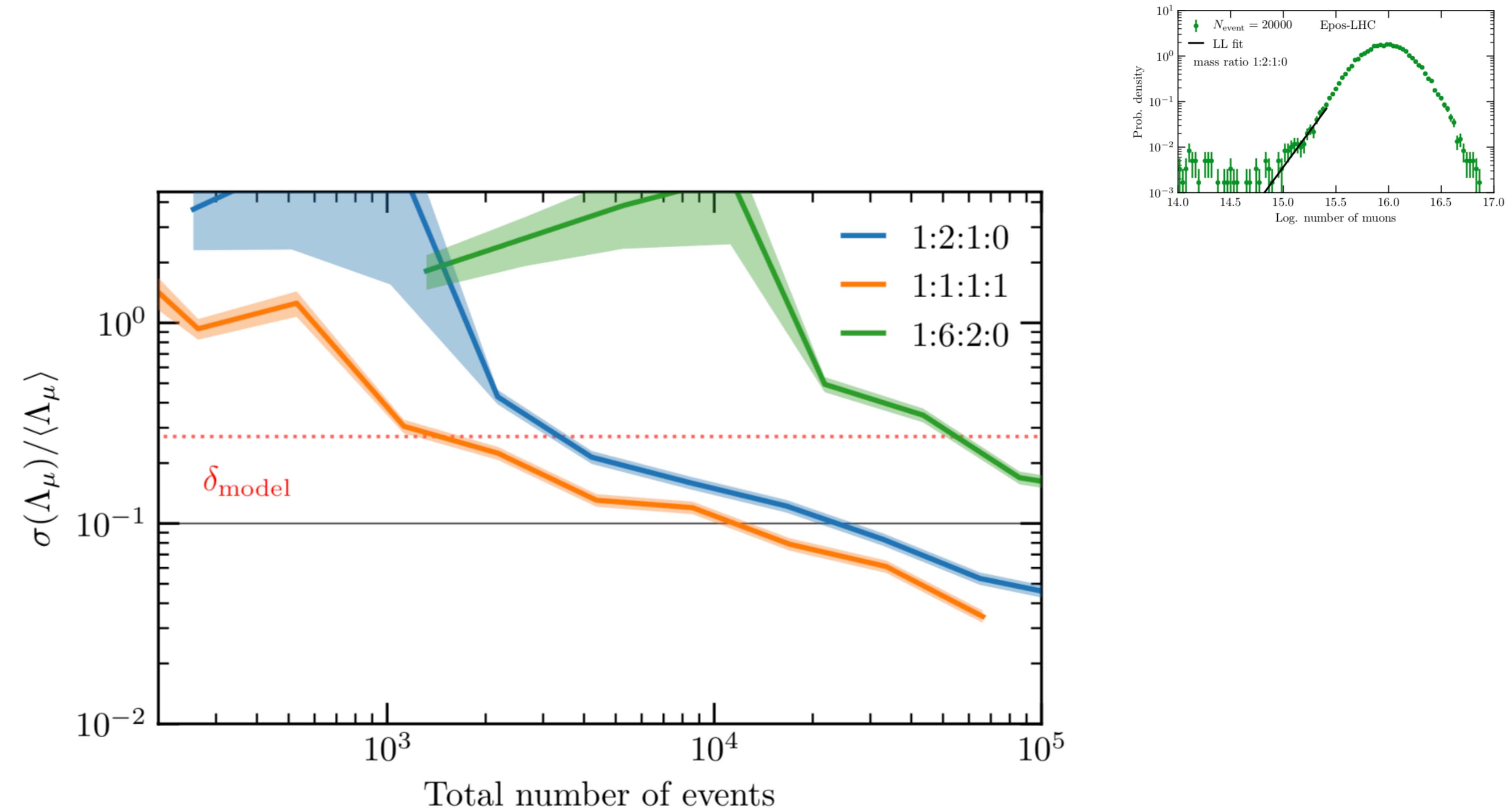
$$N_\mu(E, A) = A \cdot N_\mu(E/A)$$



Different mass composition scenarios



Different mass composition scenarios



Fitting the pion energy spectra tail

