



# Status and Prospects of the TAx4 Experiment

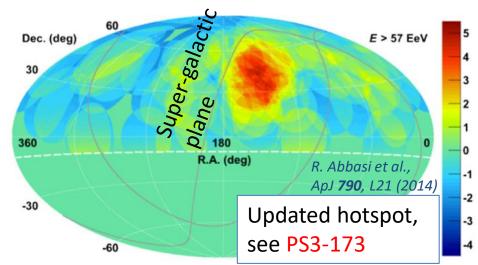
Eiji Kido for the Telescope Array Collaboration ICRR, University of Tokyo



#### Outline

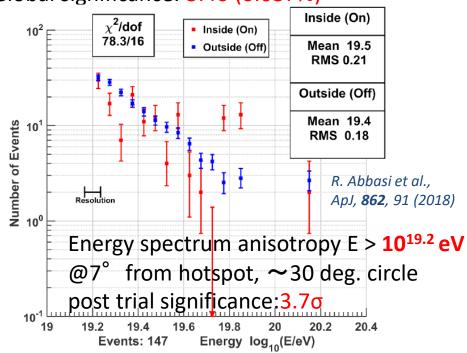
- Motivation
- Construction of Surface Detectors (SDs)
- Construction of Fluorescence Detectors (FDs)
- Prospects
- Summary

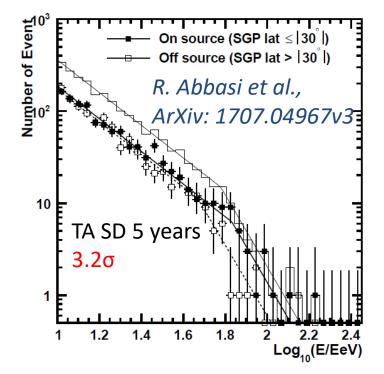
Significance map from isotropy

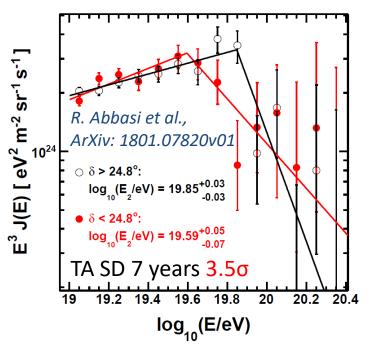


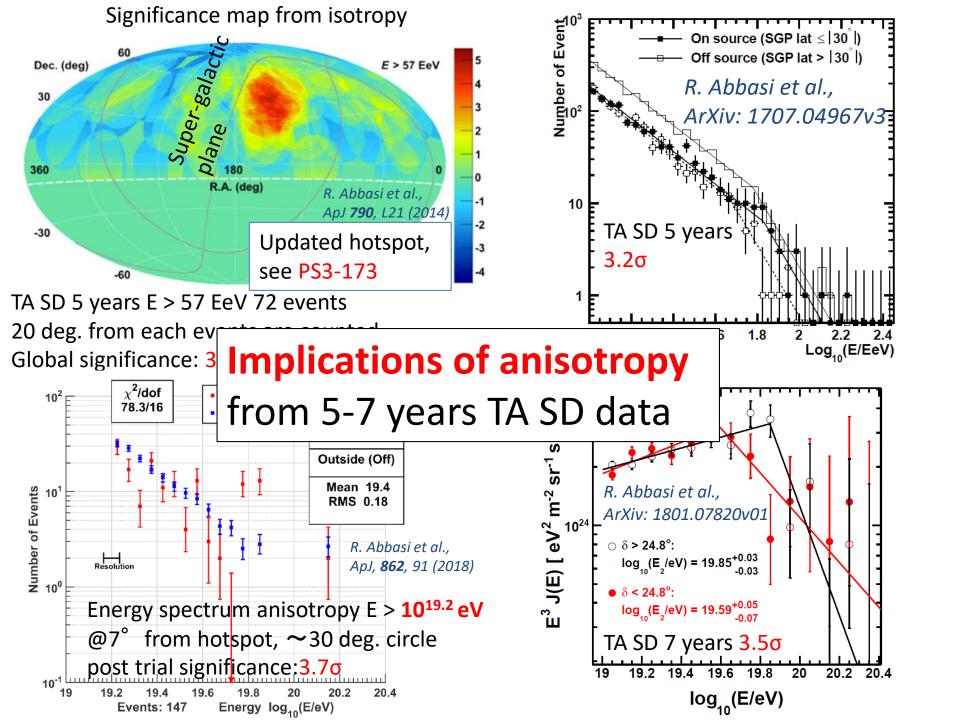
TA SD 5 years E > 57 EeV 72 events 20 deg. from each events are counted.

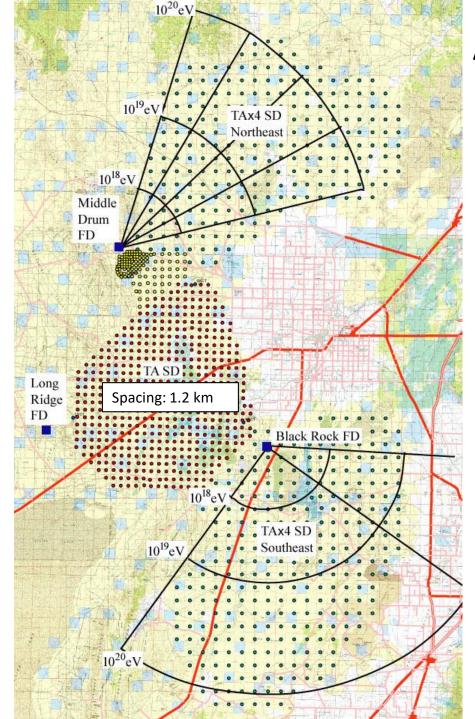
Global significance: 3.4o (0.037%)











# Arrangement of Detectors

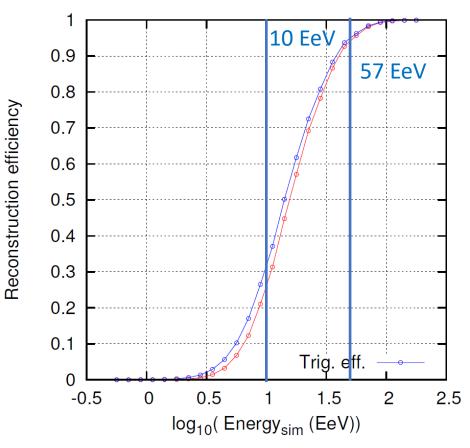
To study more about the highest energies and examine the implications obtained by TA

500 new SDs with 2.08 km spacing and TA SDs cover

4 × TA SD detection area (~3000 km²)

2 new Fluorescence Detector (FD) stations (4+8 HiRes Telescopes)

#### Expectation of the Performance of New SD Array



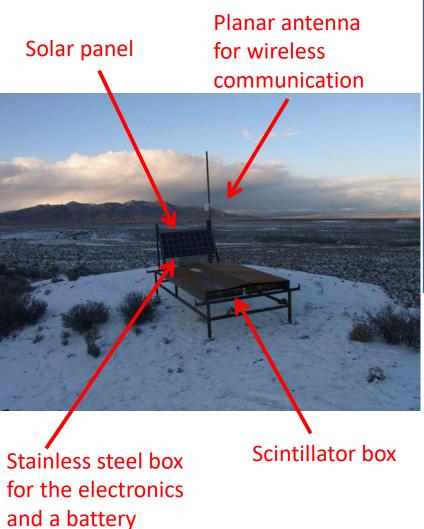
SD array: square grid with 2.08 km spacing

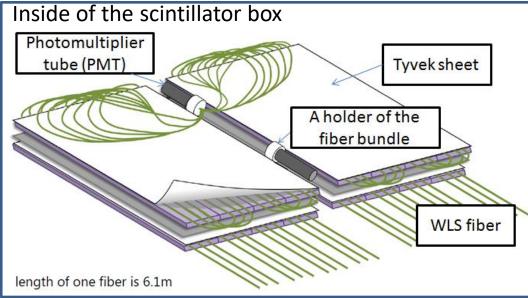
Trigger condition: adjacent 3 SDs within 14 usec

E > 57 EeV:

- Reconstruction efficiency > 95%
- Angular resolution: 2.2°
- Energy resolution:  $\sim$ 25%

### Design of SDs





- 2 layers 3 m<sup>2</sup> 1.2 cm thick plastic scintillators
- → Calibration of signals using single muons
- DAQ with 2.4 GHz wireless communication

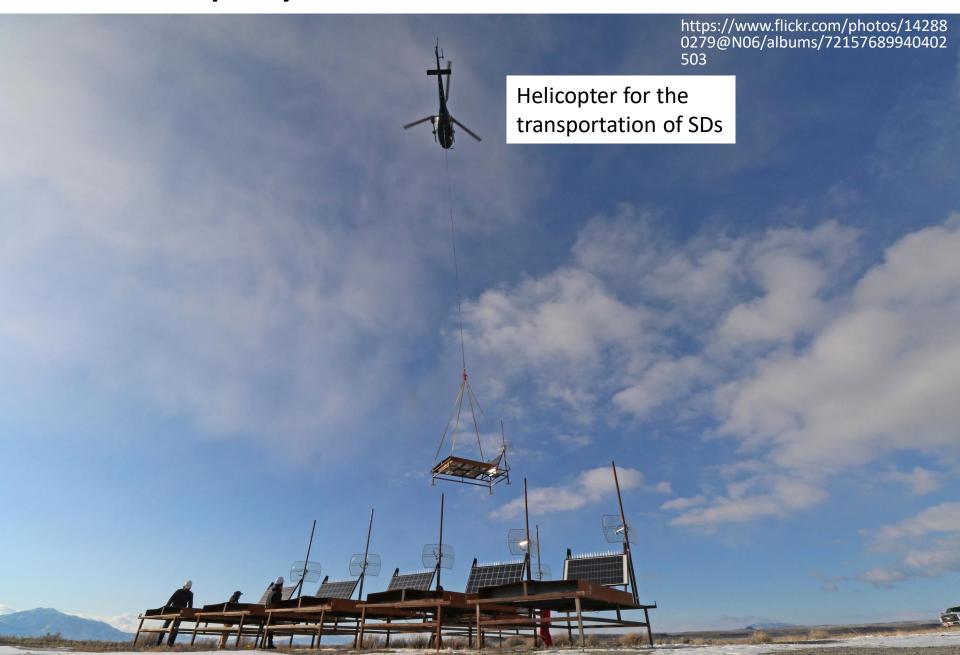
PMT and arrangement of WLF fibers was changed from TA SD for the cost reduction

Single peak: 21 p.e. in average

Non-uniformity: < 15 %

Pulse linearity:  $50 \text{ mA} (\sim 2 \text{ x TA SDs})$ 

## Deployment of Assembled SDs



#### Deployment of SDs, Team on the Ground



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# TA/TALE/TAx4 Array TAx4 North TALE Delta, Utah 10 km TAx4 South 30 km 50 Kilometers 22 Apr 2019 S. Thomas

# Deployed SDs and Communication Towers

- 257 SDs were deployed
- 6 communication towers were constructed.
- SD array is divided into 6 sub-arrays for each communication towers.
- Locations of SDs were decided to optimize hybrid events above 10 EeV and consider practical conditions of wireless communications

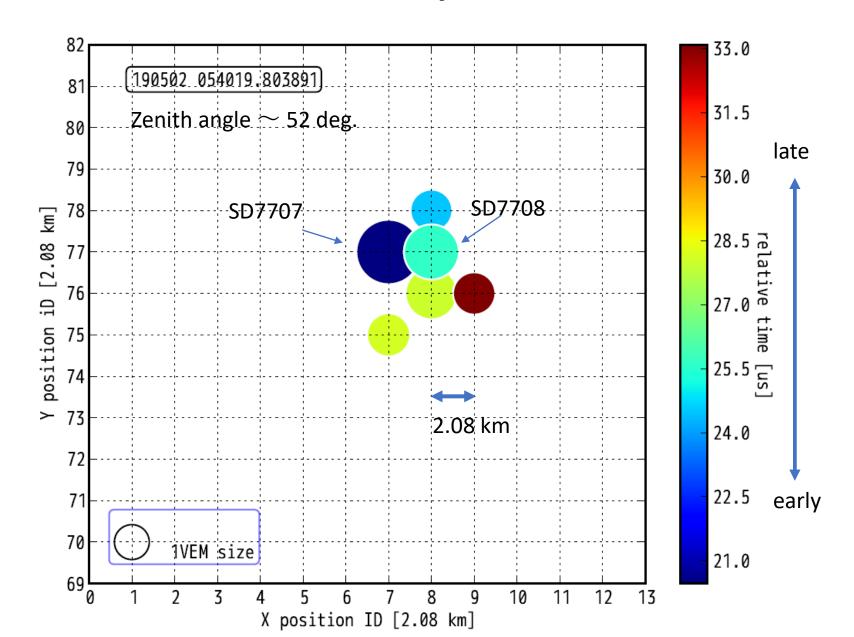
Detailed set up of communication towers, see PS1-204 poster presentation by T. Nonaka

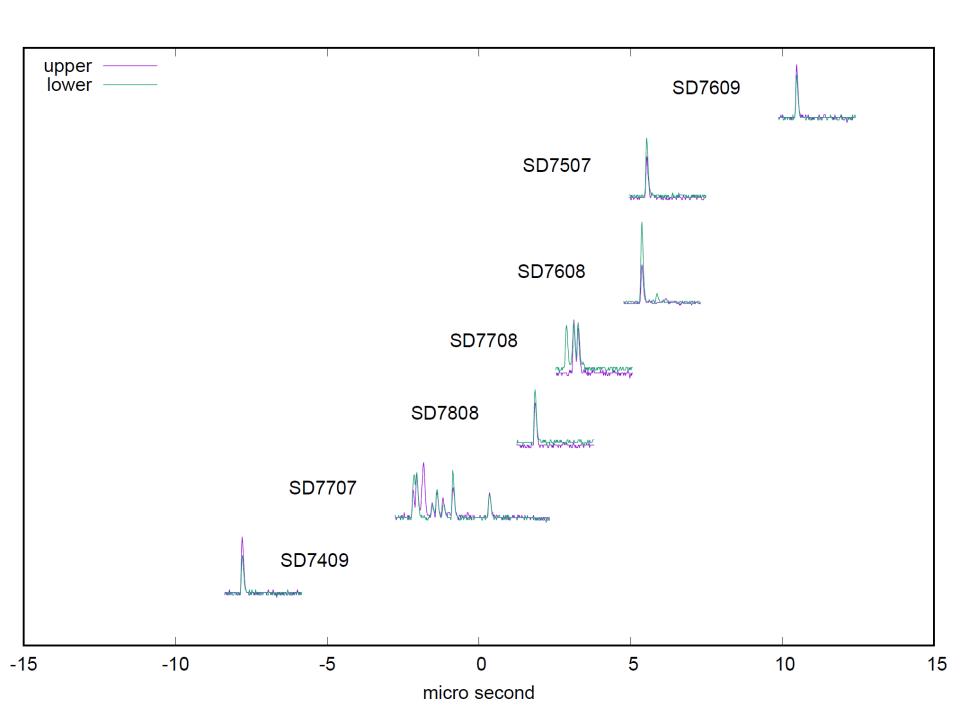
#### **Tuning Antennas**



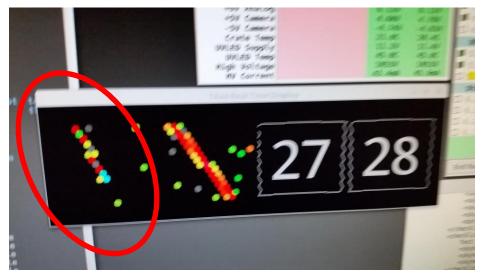
- March and June 2019
- Antennas of each SD were aligned to each communication tower.
- → Test run of DAQ of each sub-array was started from the end of Apr when the sub-array is ready.
- Cosmic ray events are being collected. (next slide)

### Cosmic Ray Event





#### Construction of North FD Station



16<sup>th</sup> Feb. 2018

First light was observed.

(camera 28: Xe Flasher)

Stable operation was started from 8<sup>th</sup> June 2018.

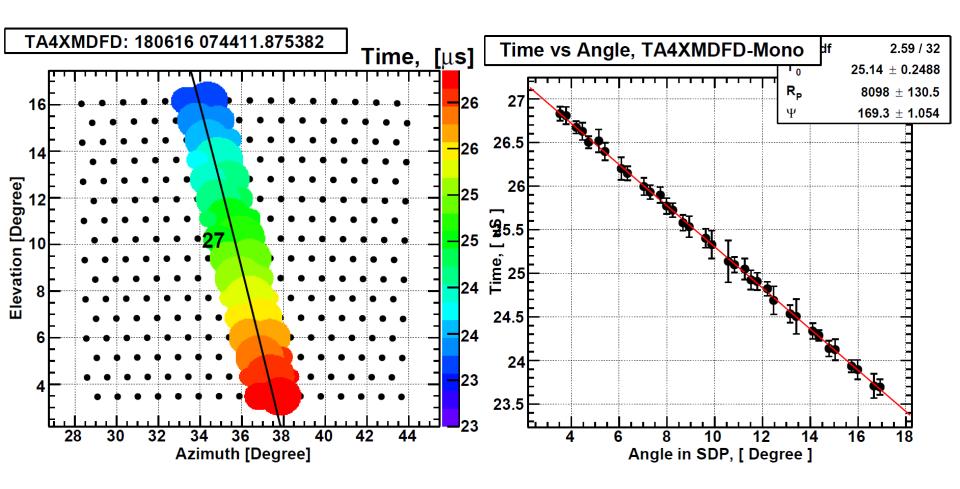


#### Construction of South FD Station



- Construction is ongoing.
- Mirrors were installed. The directions of mirrors were aligned.
- FADC electronics racks were installed.
- PMT cluster, other electronics will be installed.

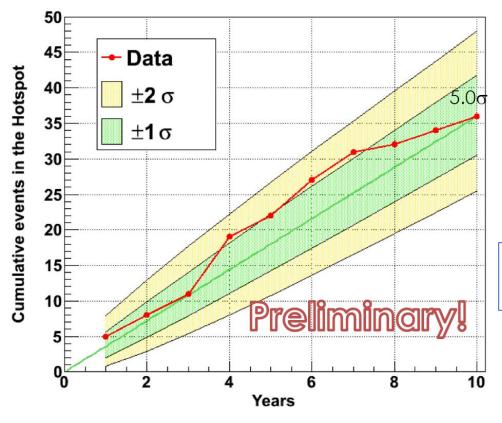
### Cosmic Ray Event



- Data analysis is ongoing.
- MC simulations will be generated, and quality cuts will be studied.

#### **Expectation of Hotspot**

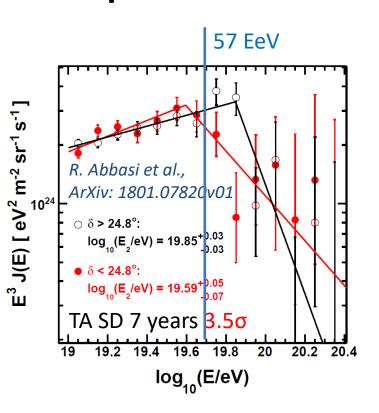
Hotspot position	Search radius	ON	α*OFF (OFF)	ON/OFF ratio (α)	σ
RA:144.3° Dec: 40.3°	25°	36	12.6 (121)	0.10435	5.0

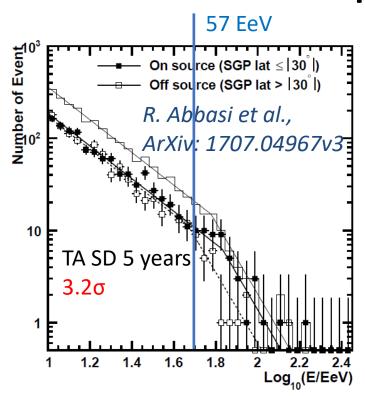


Independent  $\sim$  40 events with E > 57 EeV are expected at the hotspot from  $\sim$  2.5 years of full operation of TAx4 SD

For detail, see PS3-173 by K. Kawata

#### **Expectation of Spectrum Anisotropy**





- 7 years TA SD equivalent energy spectrum with E > 57 EeV will be obtained from ~2 years of full operation of TAx4 SD.
- Xmax at the highest energies will be also measured.
- $\sim$ 3 × TA SDFD hybrid equivalent data will be obtained by the full TAx4.

#### Summary

- Implications on anisotropy were obtained by the TA experiment.
- Arrangement of TAx4 detectors:
  - 500 new SDs with 2.08 km spacing + TA SDs → Coverage of 4 × TA SDs ~3000 km<sup>2</sup>
  - 2 new Fluorescence Detector (FD) stations (4+8 HiRes Telescopes)
- More than half of TAx4 SDs were deployed.
- North TAx4 FD was constructed.
- Data acquisition was started.
  - SD: from Apr. 2019, FD: from Jun. 2018
- Cosmic ray events are being collected.
- Prospects
  - Hotspot and spectrum anisotropy with E > 57 EeV will be studied with  $\sim$ 4 × TA SD equivalent statistics if the full operation is started.
  - Xmax:  $\sim$  3 × TA SDFD equivalent events will be expected at the highest energies if the full operation is started.

## Backup