

Gamma/Proton discrimination for LHAASO-WCDA



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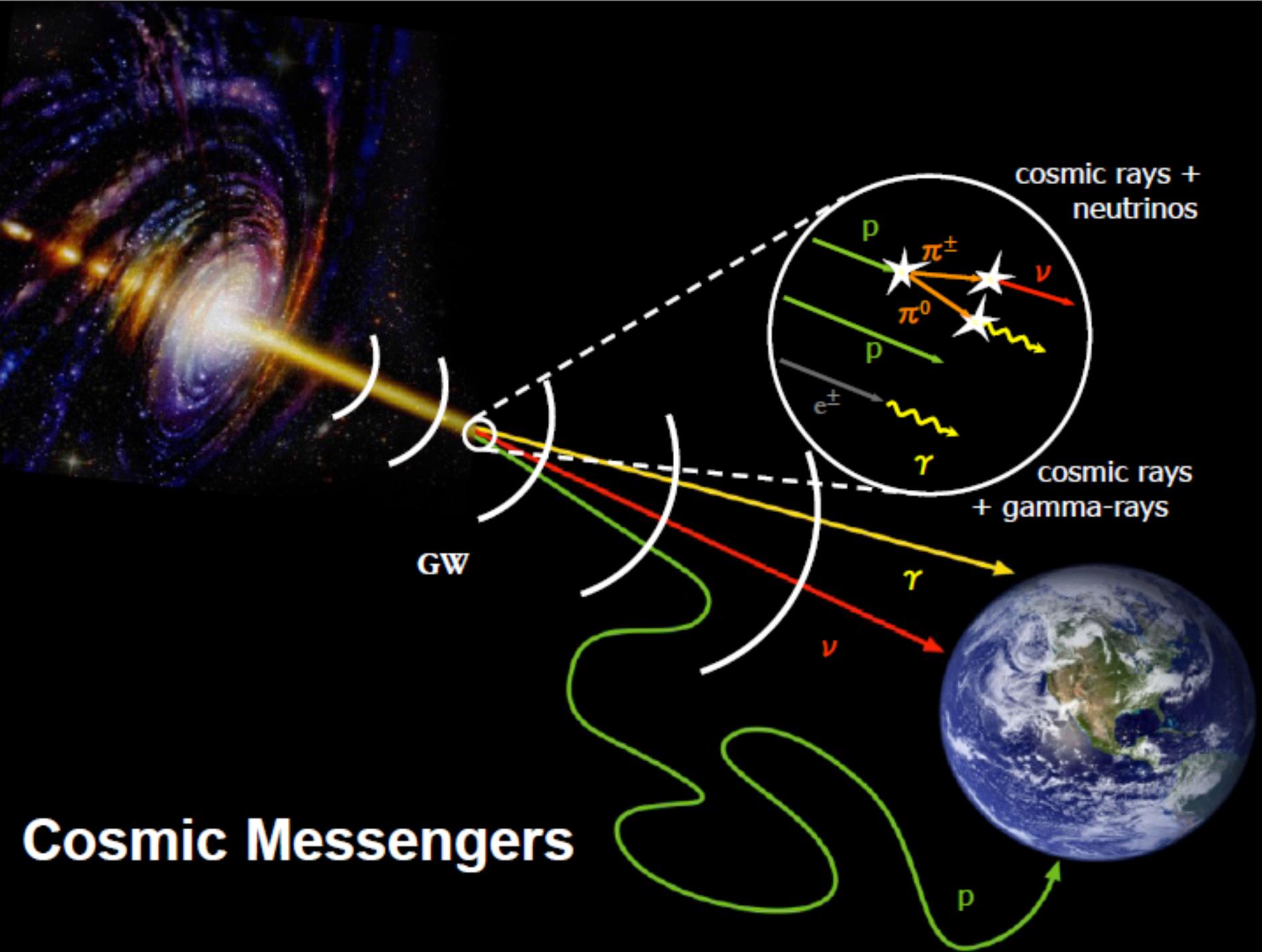
For the LHAASO collaboration

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Madison WI , 26/07/2019

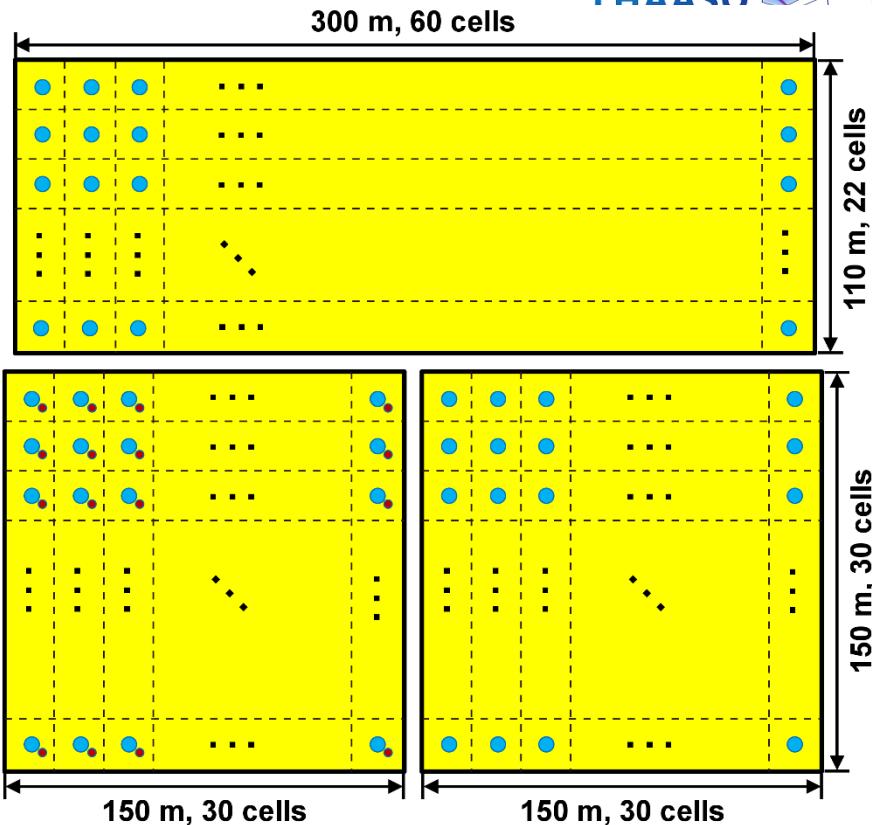
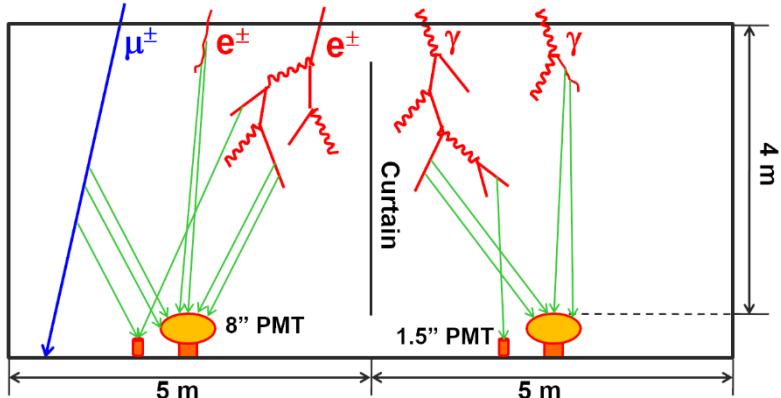
Outline

- ◆ Background introduction
- ◆ Methods
- ◆ Parameters & results & discussions
- ◆ Summary & outlook



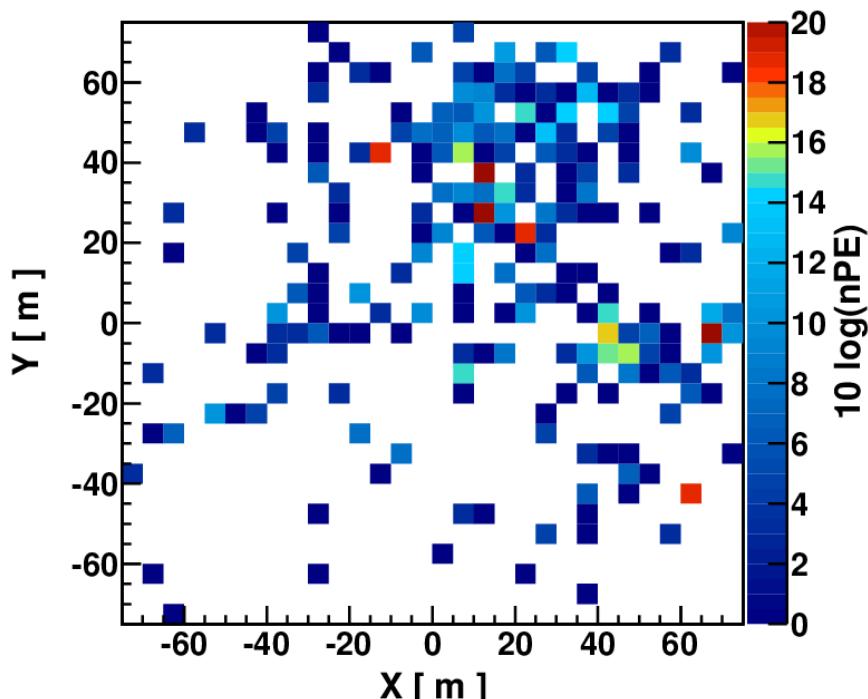
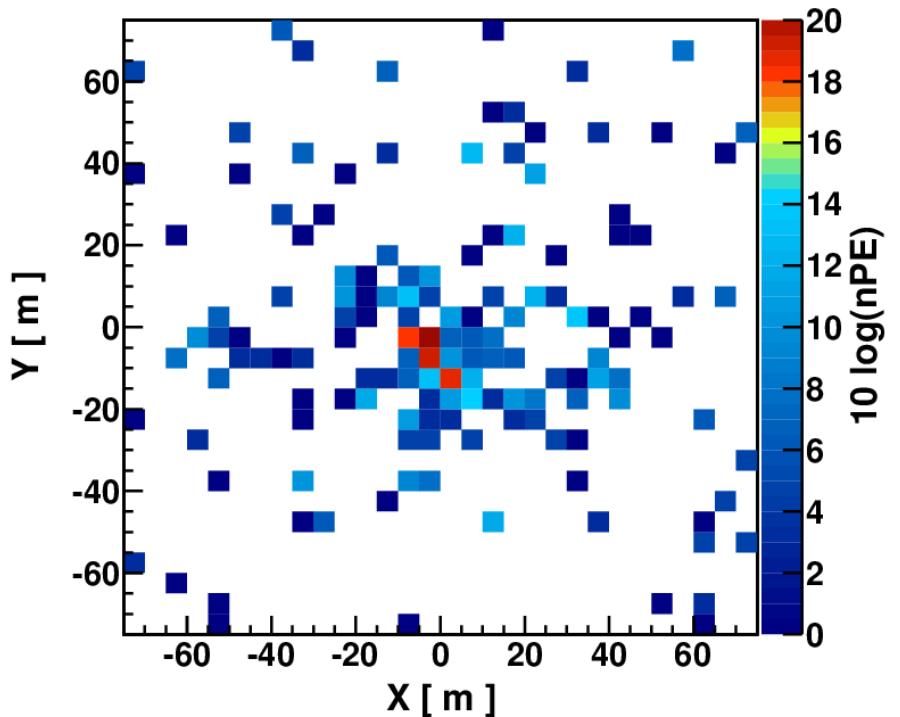
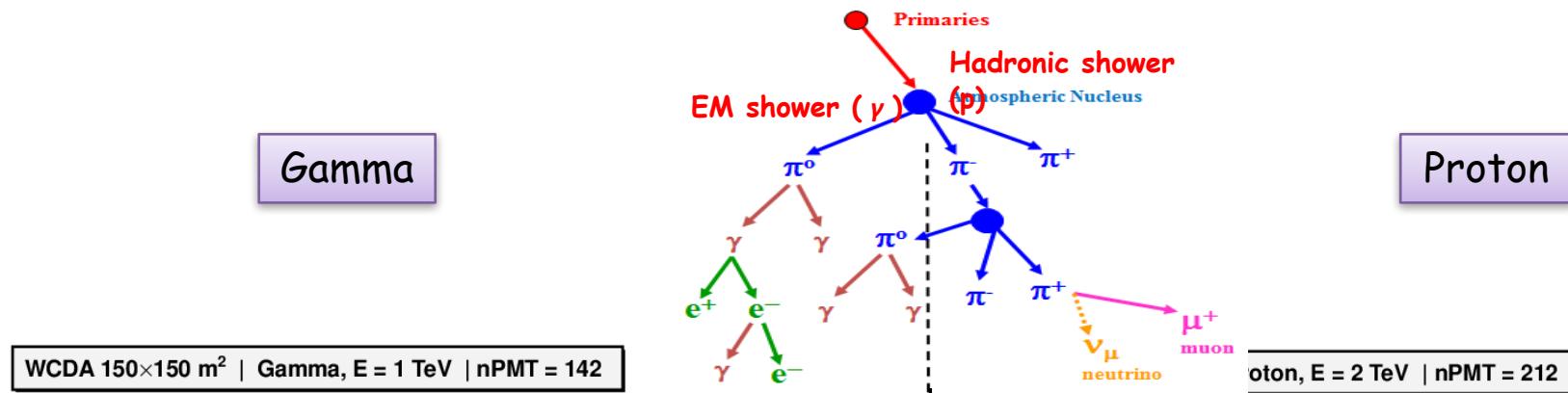
LHAASO-WCDA

- 3 pools
- 3120 cells
 - ✓ 2 PMTs/cell
- 1st pool
 - ✓ 900 cells



Background rejection
 (*gamma/proton discrimination*)
 is very important

Mechanism & methods



Discrimination variables

Take advantages of muons & sub-core

$$1. \text{ Compactness}(C) = \frac{nFit}{cxPE_{45}}$$

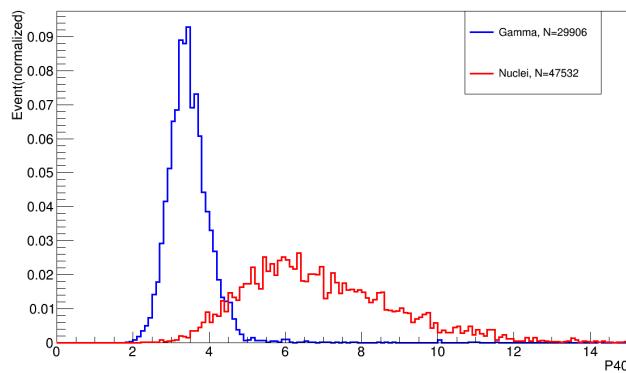
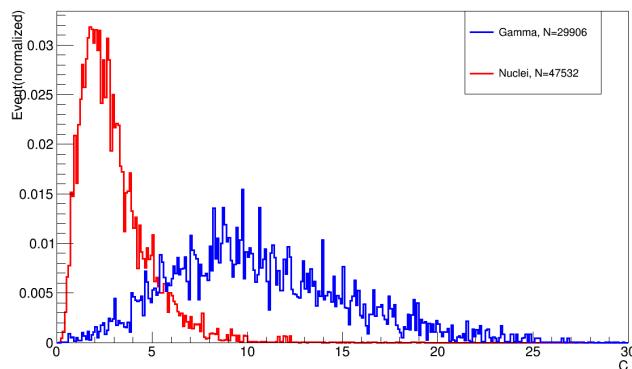
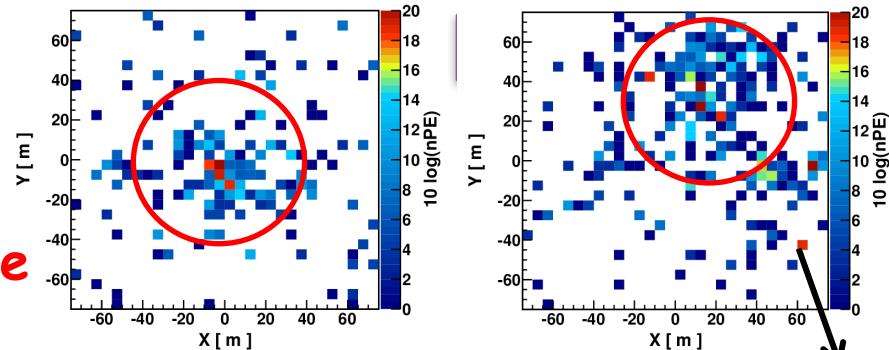
nFit: number of PMTs involved in reconstruction

cxPE₄₅: maximum PE count outside the reconstruction core with a distance of 45m ;

$$2. \text{ Density out}(\rho_{40}) = \frac{\sum PE_{40}}{\sum PMT_{40}}$$

$\sum PE_{40}$: sum number of PEs outside core within 40m

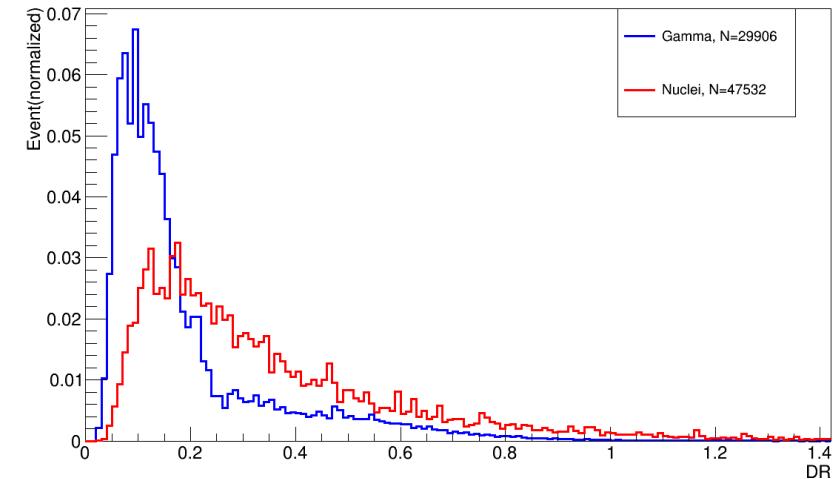
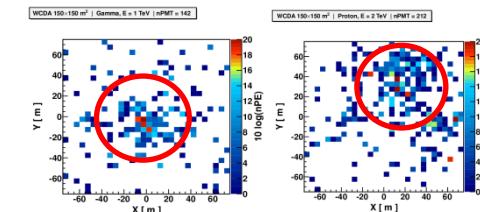
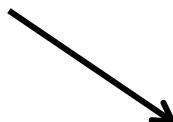
$\sum PMT_{40}$: sum number of fired PMTs outside shower core with distance of 40m



Discrimination variables

3. Density ratio(DR)

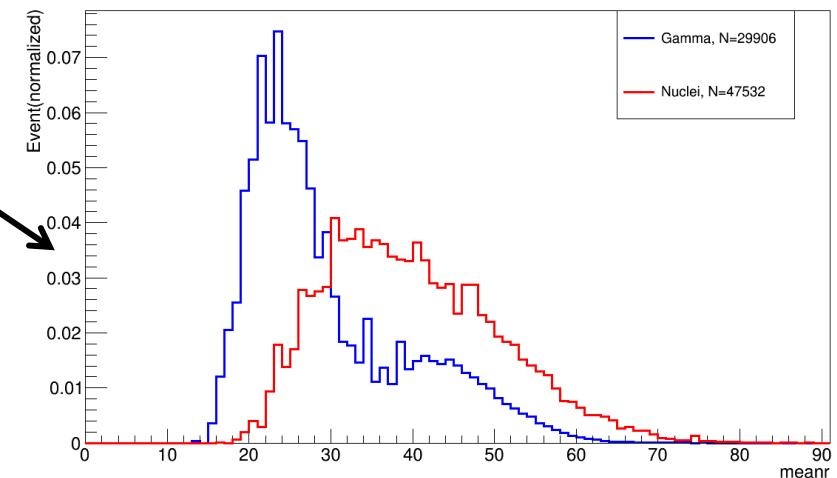
$$DR = \frac{\sum PE_{50} / \sum PMT_{50}}{\sum PE_{10} / \sum PMT_{10}}$$



Use the lateral distribution of shower

$$4. \langle R \rangle = \frac{\sum PE_i R_i}{\sum PE_i}$$

PE_i : number of nPE of the ith fired PMT;
 R_i : distance between the ith fired PMT to shower core

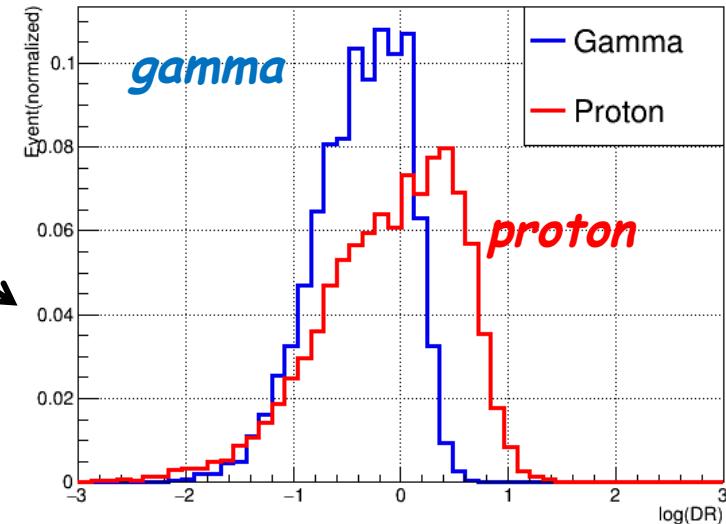
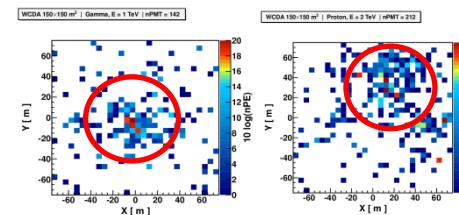


V. Grabski et al 32nd ICRC

Discrimination variables

5. r70

Minimum radius which contained 70% PE charge of the shower



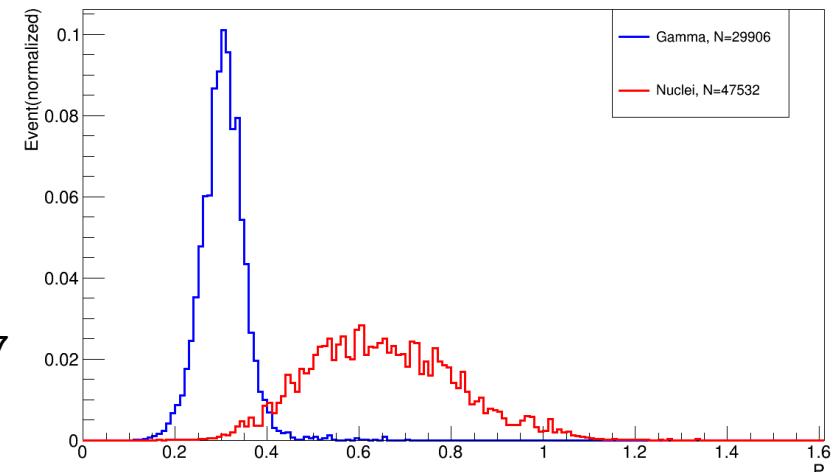
$$6. P = \frac{1}{N} \sum_{i=0}^N \frac{(\zeta_i - \langle \zeta_i \rangle)^2}{\sigma_{\zeta_i}^2}$$

$$\zeta_i = \log_{10}(Q_i)$$

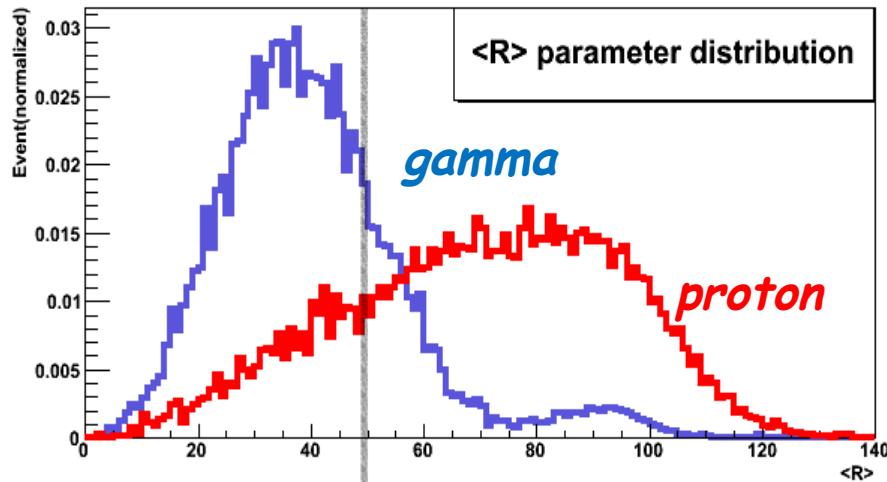
$\langle \zeta_i \rangle$: average ζ_i in an annulus with width

Of 5m

A. U. Abeysekara et al APJ 2017

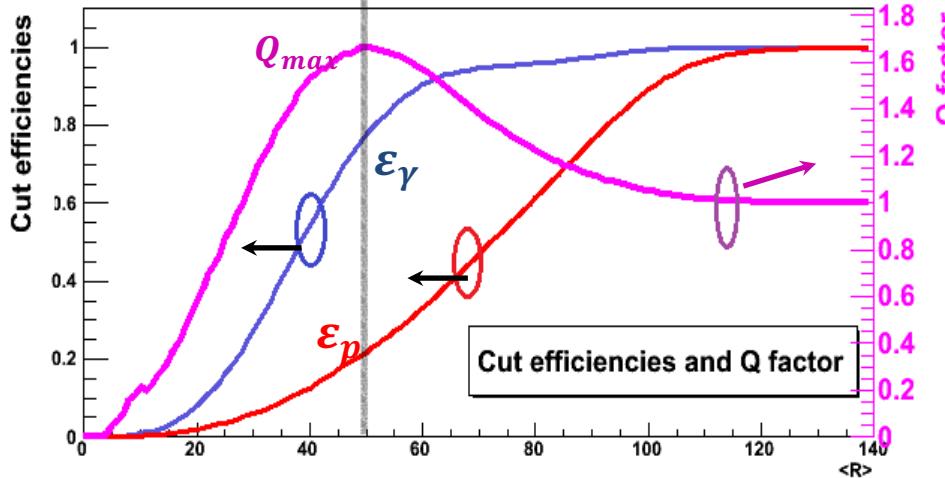


Judgment parameter——Q factor



$$Q = \frac{\varepsilon_\gamma}{\sqrt{\varepsilon_p}}$$

ε_γ : efficiency
 ε_p : misidentification rate



For example,

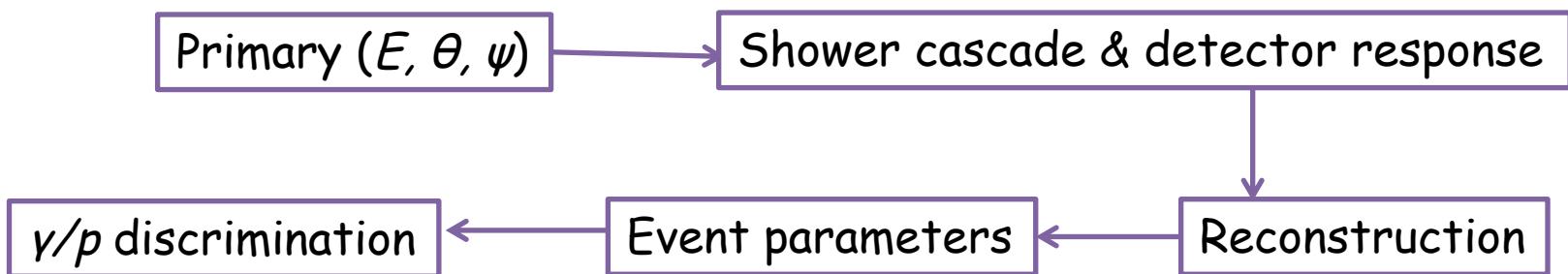
$$Q_{max} = 1.68$$

$$\varepsilon_\gamma = 78\%$$

$$\varepsilon_p = 20\%$$

Background rejection ratio is 80% , keeping 78% of signal

G/P separation procedure



MC simulation

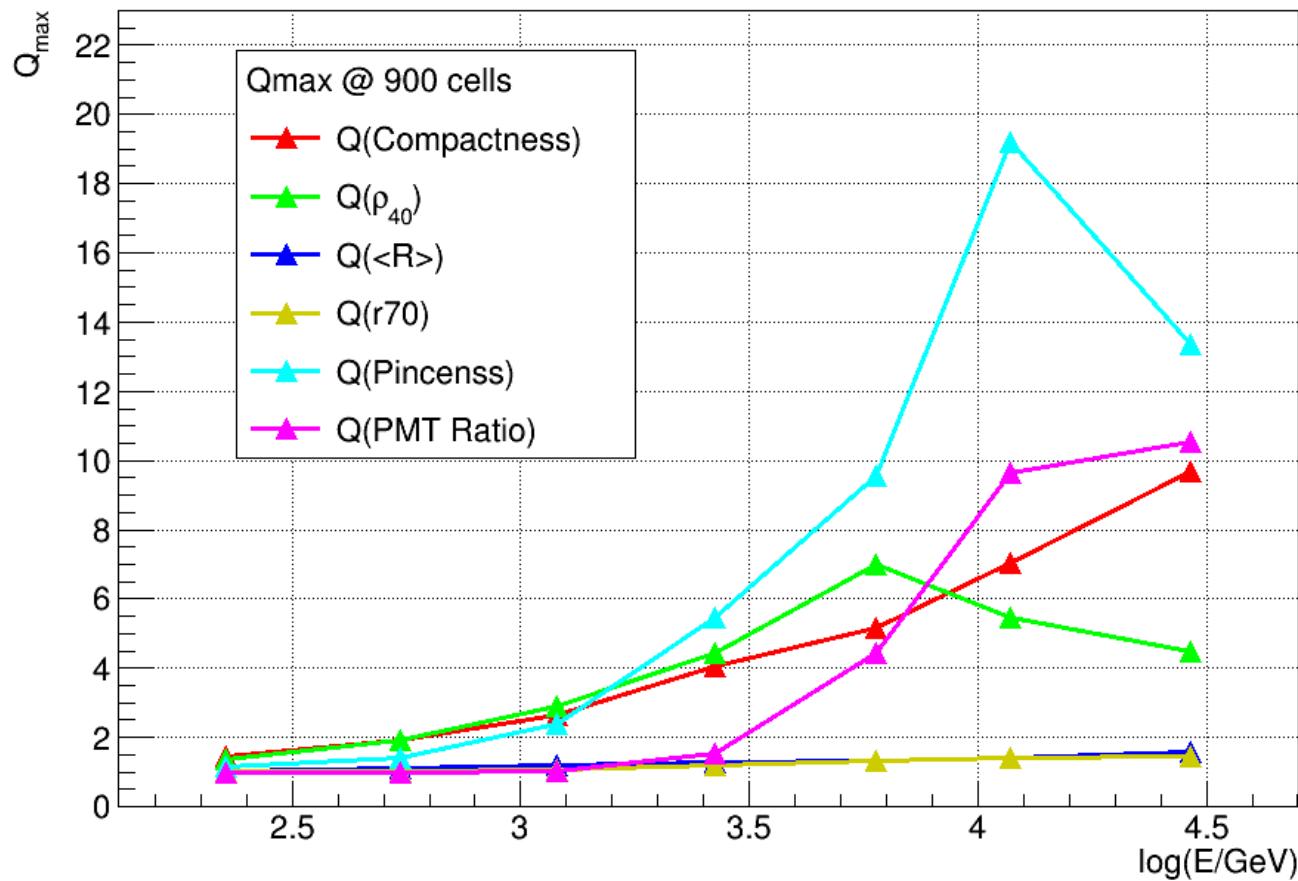
- ◆ Site: Daocheng @ 4410m a.s.l.
- ◆ Code: Corsika 6720 + **QGSJET-II** (was EPOS) + GHEISHA
- ◆ Primary: point source (γ)
Spectrum & Flux: Crab measured by **HEGRA (astro-ph/0407118)** $\propto 2.05 \times 10^{-6} (E/\text{GeV})^{-2.62} \text{ cm}^{-2}\text{s}^{-1}\text{GeV}^{-1}$.
Energy:
4 segments: 100 GeV-1-10-100 TeV
- ◆ Primary: background(p)
Spectrum & flux: **J.R. Hoerandel, Astroparticle Physics 19 (2003) 193-220**
Energy: same, but min energy = $\min(10, 1.1 \times A)$.
- ◆ Energy cuts:
50 (hadron), 50 (muon), 0.3 (electron), 0.3 (photon / pion)
MeV

Events selection

Cuts :

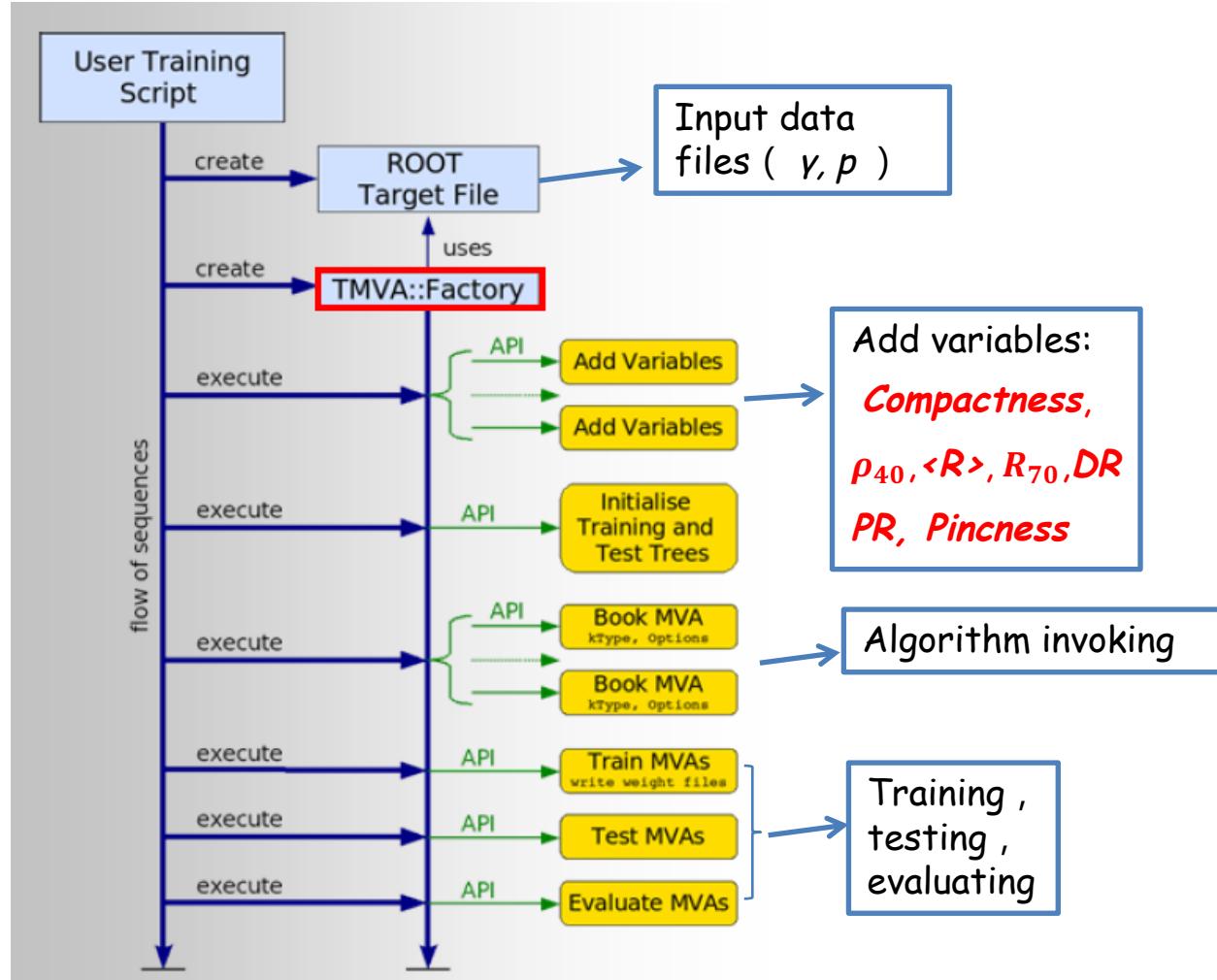
- ◆ reconstruction core fall in $150 \times 150 \text{ m}^2$
- ◆ zenith : $0 \sim 45^\circ$
- ◆ Gamma efficiency $> 50\%$

Single parameter result



TMVA

Analysis steps :



TMVA:

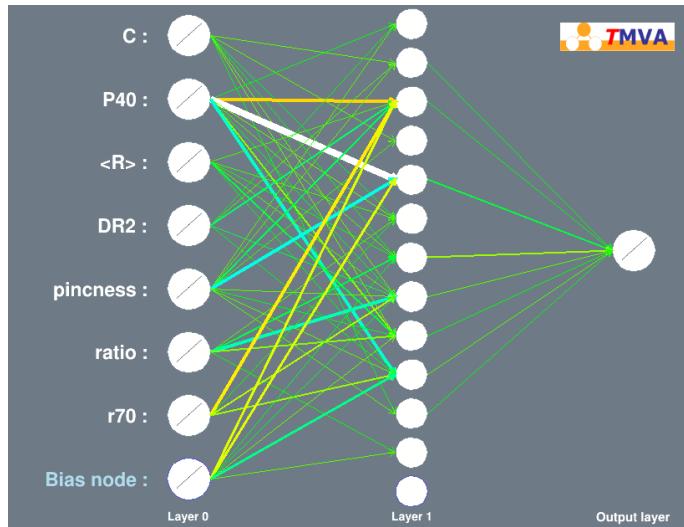
Toolkit for Multivariate Data Analysis with ROOT

Algorithms :

- Artificial Neural Network (ANN)
- Boosting decision trees (BDTG)

Artificial Neural Network (ANN)

1. Input variable & neural network layout



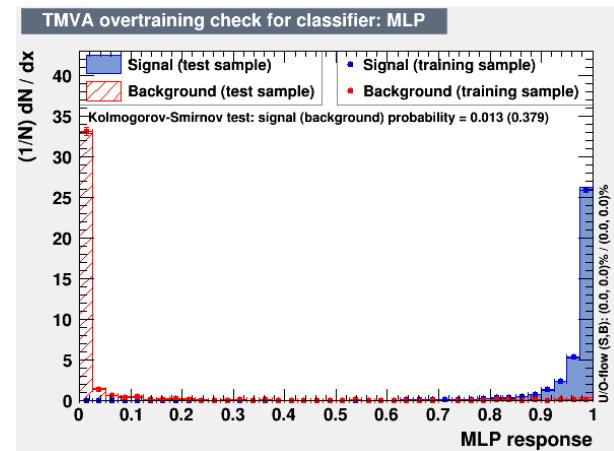
2. Input data

Signal: 28856

Background: 27040

Half of them are used in training,
the other half for testing

3. ANN result ($200 < n_{fit} < 300$)



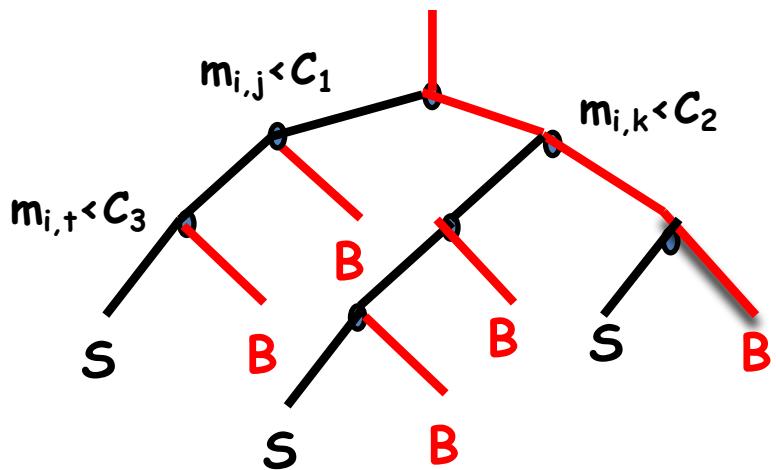
	P	C	MLP
Q max	9.55	5.18	7.78
ϵ_γ	62.9%	57.2%	54.7%
ϵ_p	0.4%	1.2%	0.5%

1.5 times of C !

Gradient Boost Dicision Tree(BDTG)

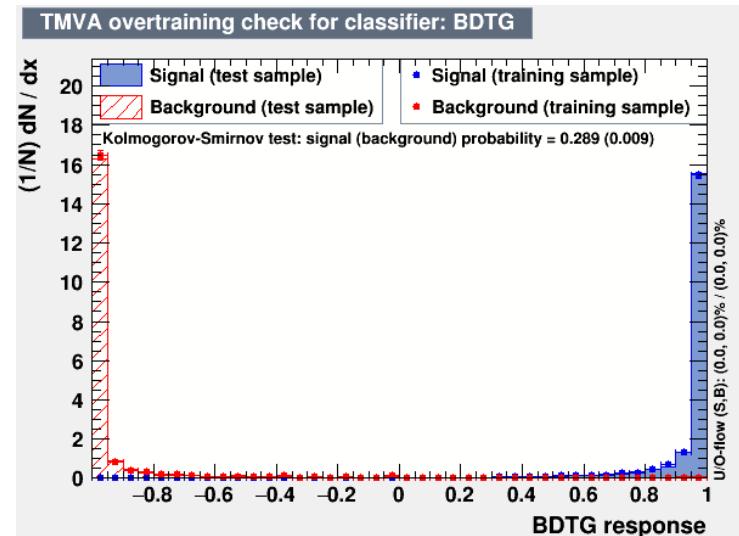
BDTG result ($200 < nFit < 300$)

Events(with parameter $m_{i,j,k,\dots,t}$)



Input data

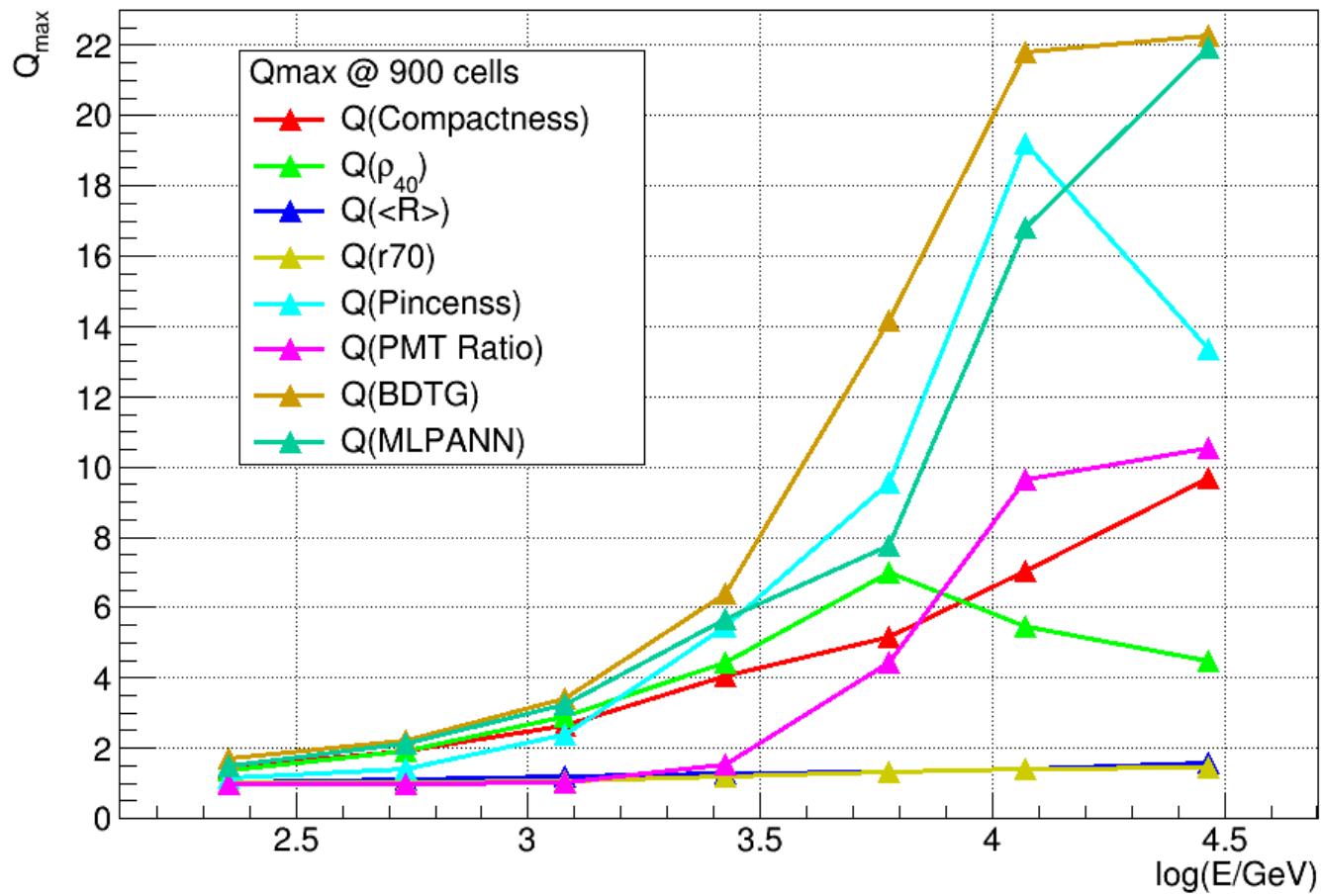
- ✓ Signal: 28856
- ✓ Background: 27040
- ✓ Half of them are used in training,
✓ the other half for testing



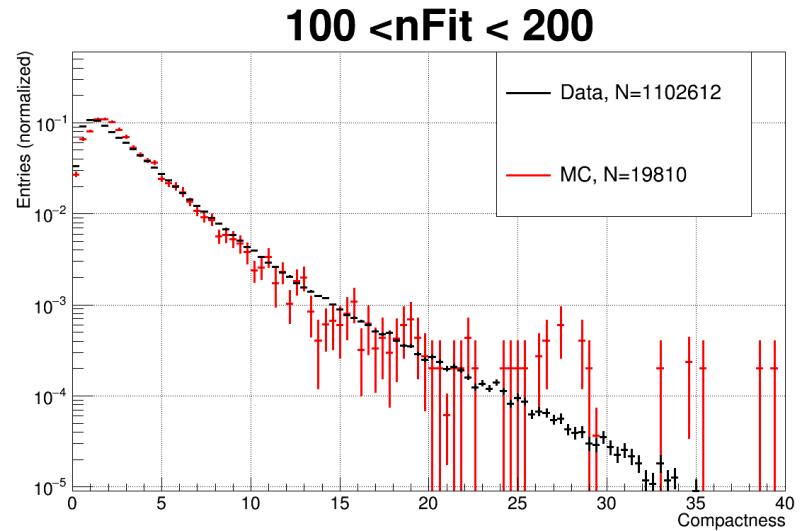
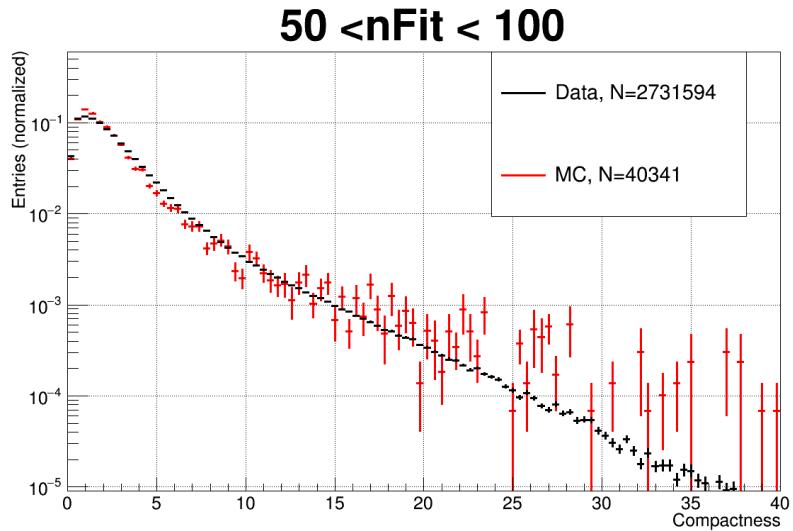
	P	C	BDTG
Q max	9.55	5.18	21.82
ϵ_γ	62.9%	57.2%	81.6%
ϵ_p	0.4%	1.2%	0.1%

4.2 times of C !

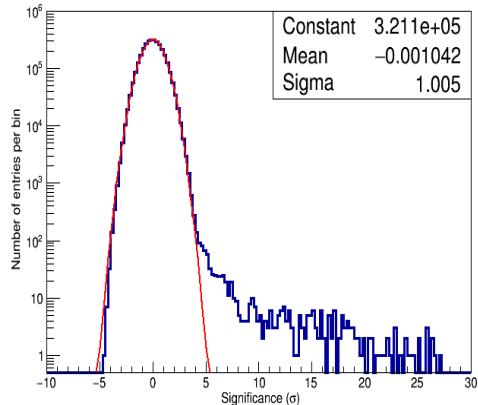
Qmax vs. Energy



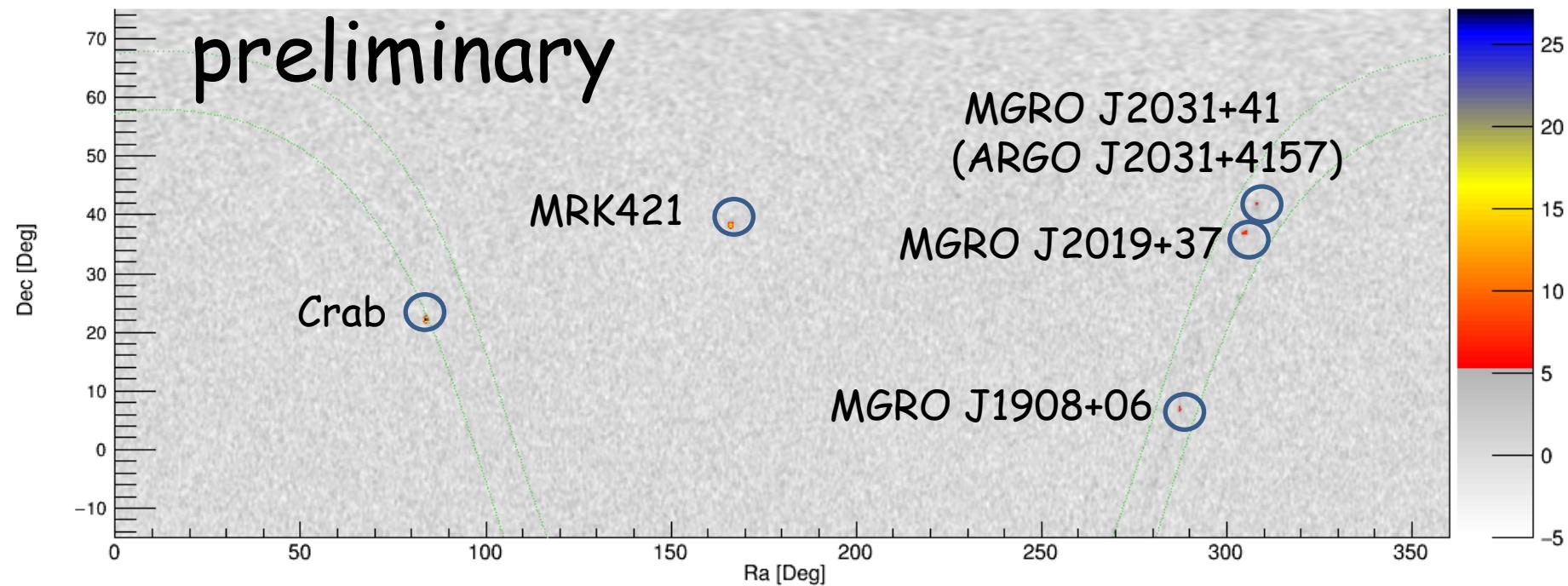
Simulation vs Data



WCDA#1 Sky Map



- Data up to 2019-06-30
- Live time: 57 days
- Crab: 27σ



Summary & Outlook

- ◆ LHAASO-WCDA can separate gamma/proton well
- ◆ Energy $>3\text{TeV}$ reject 99% of the background events
- ◆ Multivariate analysis can improve Q factor value obviously.
- ◆ Compactness distribution of data is basically agree with MC's
- ◆ There are still some problems need to be resolved , with more data, we hope better results

THANK YOU !

Back up

Training & testing sample

nfit bin	Number of signal	Number of background
10-20	6172	52322
20-50	69964	140948
50-100	141316	128340
100-200	113468	78656
200-300	28856	27040
300-500	12520	14468
500-900	1424	1612

input variable correlations

