





Extended Measurement of Cosmic-Ray Electron and Positron Spectrum from CALET on the ISS

Shoji Torii Waseda University & Kanagawa University for the CALET collaboration CRD2c 25 July, 2019

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CALET Payload





- Mass: 612.8 kg
- JEM Standard Payload Size: 1850mm(L) × 800mm(W) × 1000mm(H)
- Power Consumption: 507 W (max) ٠
- **Telemetry:**

Medium 600 kbps (6.5GB/day) / Low 50 kbps





JEM/Port #9



CALET Instrument

٨	Plastic	Scintillator + PMT + 64anode PM	er T Scintillator(PWO) + APD/PD or PMT (X1)	CALORIMETER
				CHD-FEC CHD-FEC MC-FEC IMC-FEC
		IMC	TASC Contraction of the second	ASC-FEO TASC TASC TASC TASC
1.1				
		CHD (Charge Detector)	IMC (Imaging Calorimeter)	TASC (Total Absorption Calorimeter)
	Measure	CHD (Charge Detector) Charge (Z=1-40)	IMC (Imaging Calorimeter) Tracking , Particle ID	TASC (Total Absorption Calorimeter) Energy, e/p Separation
	Measure Geometry (Material)	CHD (Charge Detector) Charge (Z=1-40) Plastic Scintillator 14 paddles x 2 layers (X,Y): 28 paddles Paddle Size: 32 x 10 x 450 mm ³	IMC (Imaging Calorimeter) Tracking , Particle ID 448 Scifi x 16 layers (X,Y) : 7168 Scifi 7 W layers (3X ₀): 0.2X ₀ x 5 + 1X ₀ x2 Scifi size : 1 x 1 x 448 mm ³	TASC (Total Absorption Calorimeter) Energy, e/p Separation 16 PWO logs x 12 layers (x,y): 192 logs log size: 19 x 20 x 326 mm ³ Total Thickness : 27 X ₀ , ~1.2 λ ₁
	Measure Geometry (Material) Readout	CHD (Charge Detector) Charge (Z=1-40) Plastic Scintillator 14 paddles x 2 layers (X,Y): 28 paddles Paddle Size: 32 x 10 x 450 mm ³ PMT+CSA	IMC (Imaging Calorimeter) Tracking , Particle ID 448 Scifi x 16 layers (X,Y) : 7168 Scifi 7 W layers (3X ₀): 0.2X ₀ x 5 + 1X ₀ x2 Scifi size : 1 x 1 x 448 mm ³ 64-anode PMT+ ASIC	TASC (Total Absorption Calorimeter)Energy, e/p Separation16 PWO logs x 12 layers (x,y): 192 logs log size: 19 x 20 x 326 mm³ Total Thickness : 27 X ₀ , ~1.2 λ ₁ APD/PD+CSA PMT+CSA (for Trigger)@top layer



CALET Capability







Examples of Event Display



Fe, ∆E=9.3 TeV

Gamma-ray, E=44.3 GeV





Simple Two Parameter Cut

- **F**_E: Energy fraction of the bottom layer sum to the whole energy deposit sum in TASC
- **R**_E: Lateral spread of energy deposit in TASC-X1

Cut Parameter K is defined as follows:

 $K = log_{10}(F_E) + 0.5 R_E (/cm)$

Boosted Decision Trees (BDT)

In addition to the two parameters in the left, TASC and IMC shower profile fits are used as discriminating variables.





In the final electron sample, the resultant contamination ratios of protons are:

5 % up to 1 TeV ; 10% - 20% in the 1 - 4.8 TeV region , while keeping a constant high efficiency of 80 % for electrons.

Number of Events Number of Event: 475.5< E/[GeV] <598.6 Flight Data 1194.3< E/[GeV] <4754.7 Flight Data 10 χ^2 /d.o.f. = 0.94 for -0.5 < R_{BDT} < 0.5 MC Electrons χ^2 /d.o.f. = 1.21 for -0.5 < R_{BDT} < 0.5 MC Electrons MC Protons MC Protons — MC Total ----- MC Total 10 10 | -0.5 -1 -0.5 BDT Response (R_{BDT}) BDT Response (R_{BDT})

476 < E < 599 GeV

1196 < E < 4755 GeV (highest energy bin)



- Constant and high efficiency is the key point in our analysis.
- Simple two parameter cut is used in the low energy region while the difference in resultant spectrum are taken into account in the systematic uncertainty.





Stability of BDT Analysis and Energy Dependence of Systematic Uncertainties

Stability of BDT analysis with respect to independent training samples and BDT - cut efficiency in the 949 < E < 1194 GeV

- Color maps show the flux ratio dependence on efficiency, where the bin value (number of trials) increases as color changes from violet, blue, green, yellow to red.
- A projection onto the Y -axis is shown as a rotated histogram (in gray color).

Energy dependence of systematic uncertainties

- The red squares represent the systematic uncertainties stemming from the electron identification based on BDT.
- The bands defined by black lines show the sum in quadrature of all the sources of systematics, except the energy scale uncertainties.



Shoji Torii

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Extended Measurement by Observation over 780 days

Approximately doubled statistics above 500GeV by using full acceptance of CALET





Prospects for the CALET All-Electron Spectrum

Five years or more observations \Rightarrow 3 times more statistics, reduction of systematic errors





Summary and Future Prospects

CALET was successfully launched on August 19th, 2015, and is successfully carrying out observations since October 2015 with stable instrument performance.

- As of the end of May, 2019, the exposure amount, SΩT, reached more than 110 m² sr day for electron observations over 10 GeV.
- □ We have reported results of the all-electron (e^++e^-) spectrum in the energy range from 10 GeV to 4.8 TeV.
- Further observations will improve the measurement of electron spectrum by better statistics and a further reduction of the systematic errors, especially in the TeV region.

See Highlight Talk (H7) by Y.Asaoka on July 27 for comprehensive results of CALET.

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