

# A Consistent Model of the Interstellar Gamma-Ray Emission to Interpret Fermi LAT Observations of Diffuse Emissions

From E.O. (2019) Phy.Rev.D & E.O. (2018) MNRAS

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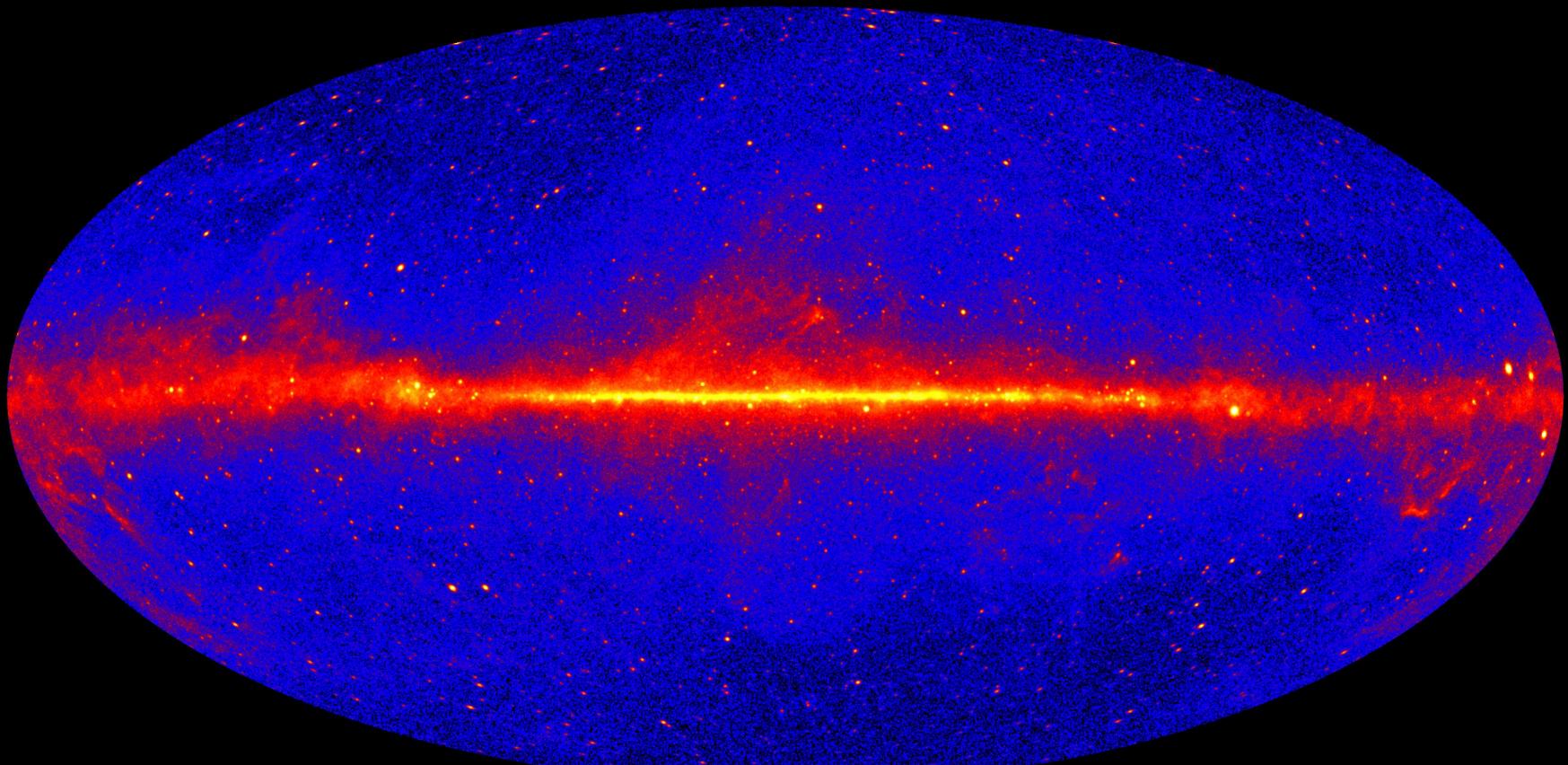


ICRC2019 – Madison, WI



# Interstellar Emission

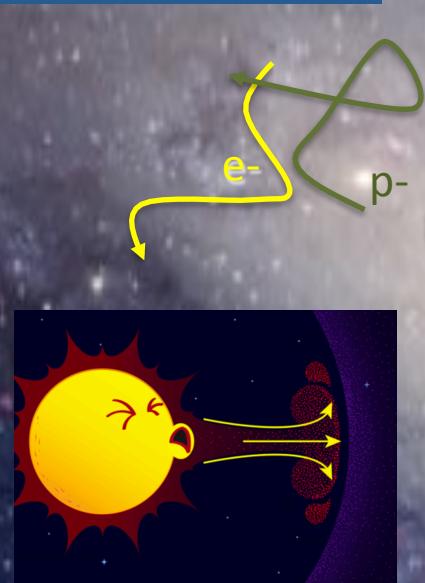
Fermi LAT > 1 GeV



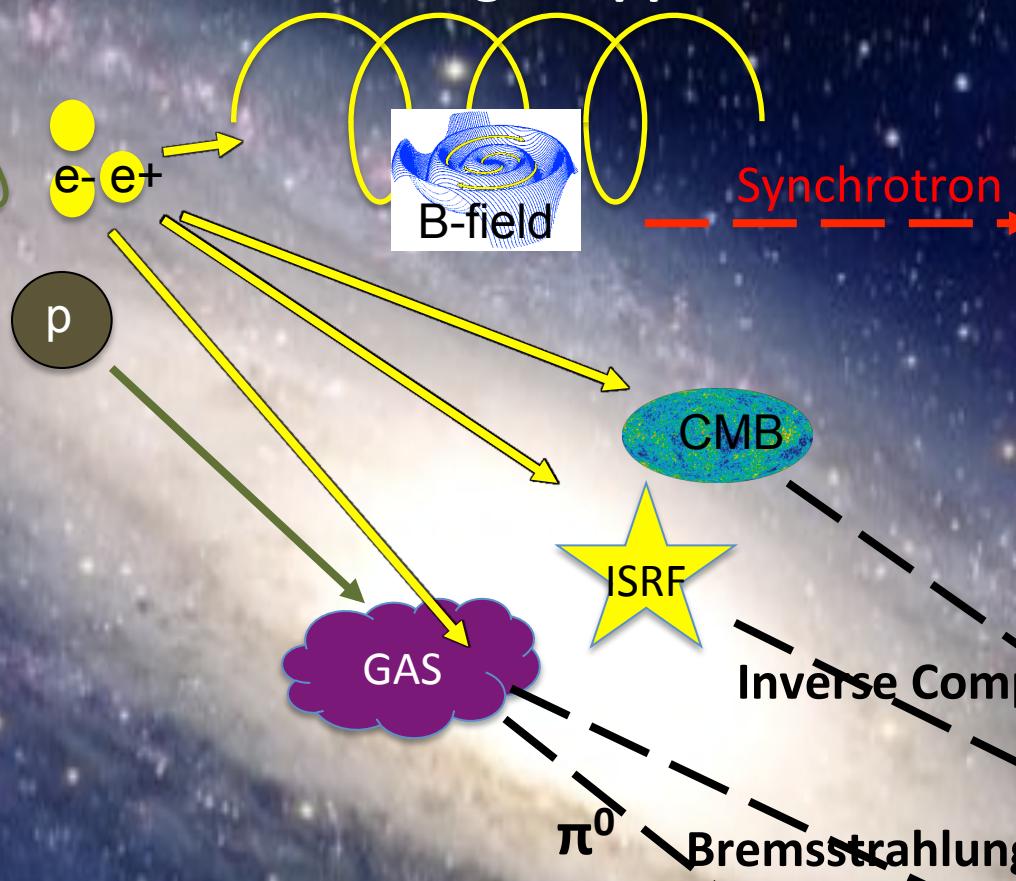
Credit: NASA/DOE/Fermi LAT Collaboration

# Multimessenger Approach

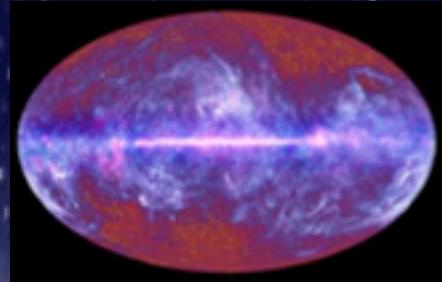
## PRIMARY CRs



Diffusion,  
energy losses,  
reacceleration,  
secondaries ...



Radio/microwave



CMB

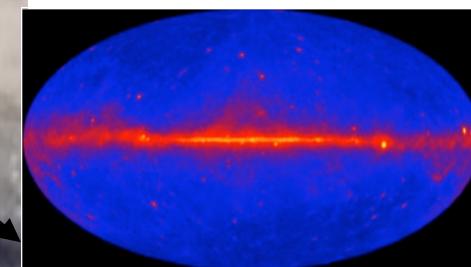


Inverse Compton

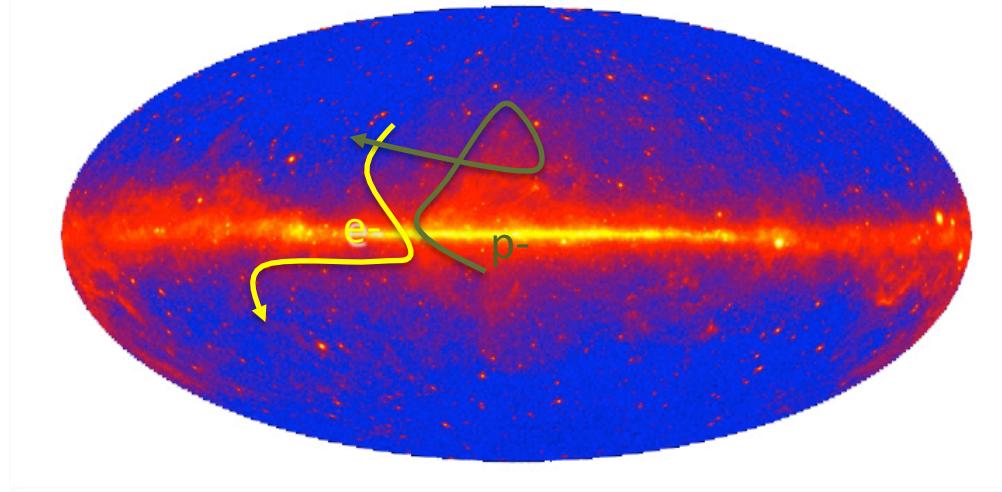
$\pi^0$

Bremsstrahlung

Gamma rays



# Present Fermi LAT Interstellar Models



From *Abdo et al (2012) ApJ*: Extensive study; Great physically based models! Many models! Based on GALPROP!

**But:**

- only standard reacceleration models for CR propagation
  - not consistent with radio and microwave data
  - not updated with new CR measurements

# This talk will show

- 1) Consistent local interstellar spectrum and CR propagation models
- 2) Effects on gamma-ray interstellar spectral and spatial templates accounting for synchrotron constraints
- 3) Predictions at MeV energies

by using a variety of multimessenger data

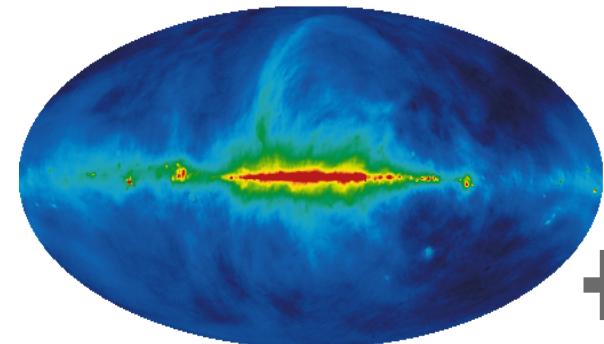
# Improvements in GALPROP: synchrotron modeling

*Strong, A., Orlando, E., Jaffe, T., 2011 A&A, 534, 54*

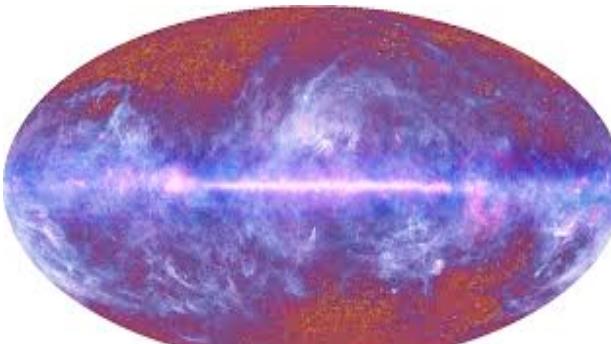
*Orlando & Strong 2013 MNRAS 436, 2127*

- 3D B-field configuration: random + regular + anisotropic random components
- total and polarized synchrotron emission
- free-free emission model
- free-free absorption

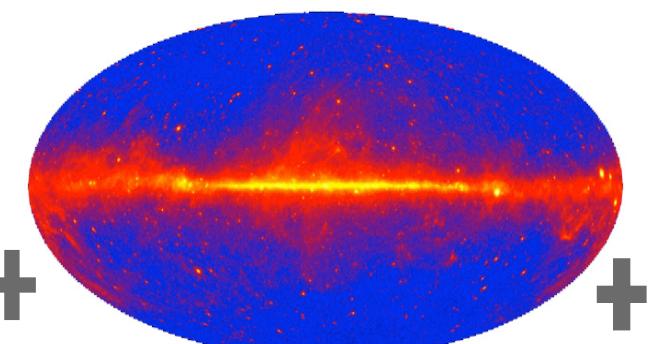
# Multimessenger Data



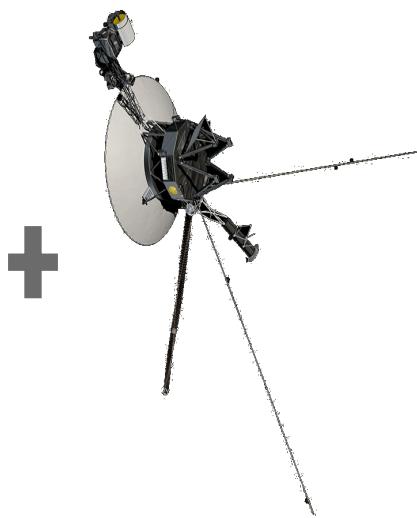
Radio at 408 MHz



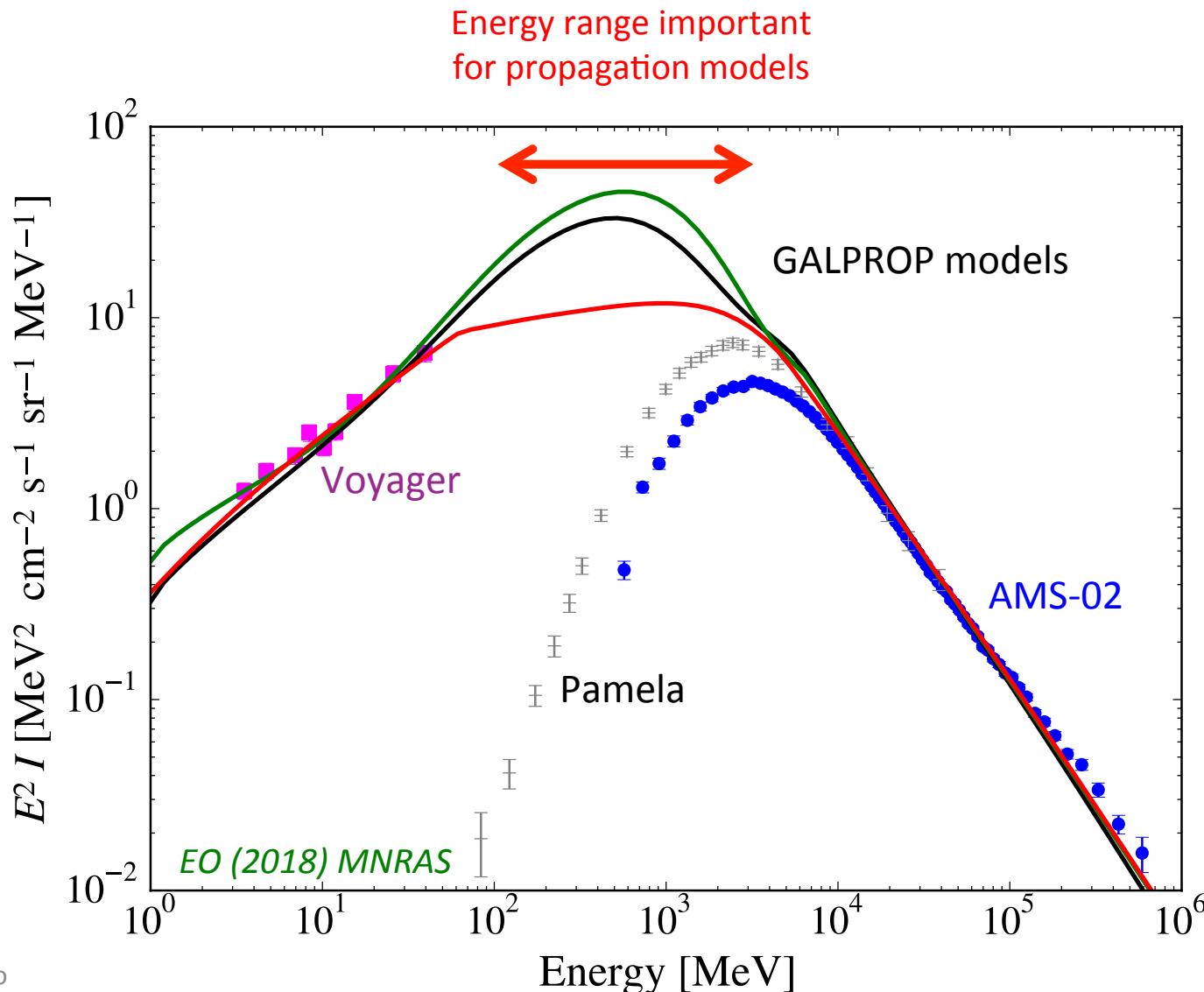
Microwaves at 30 GHz



Gamma rays at 1 GeV



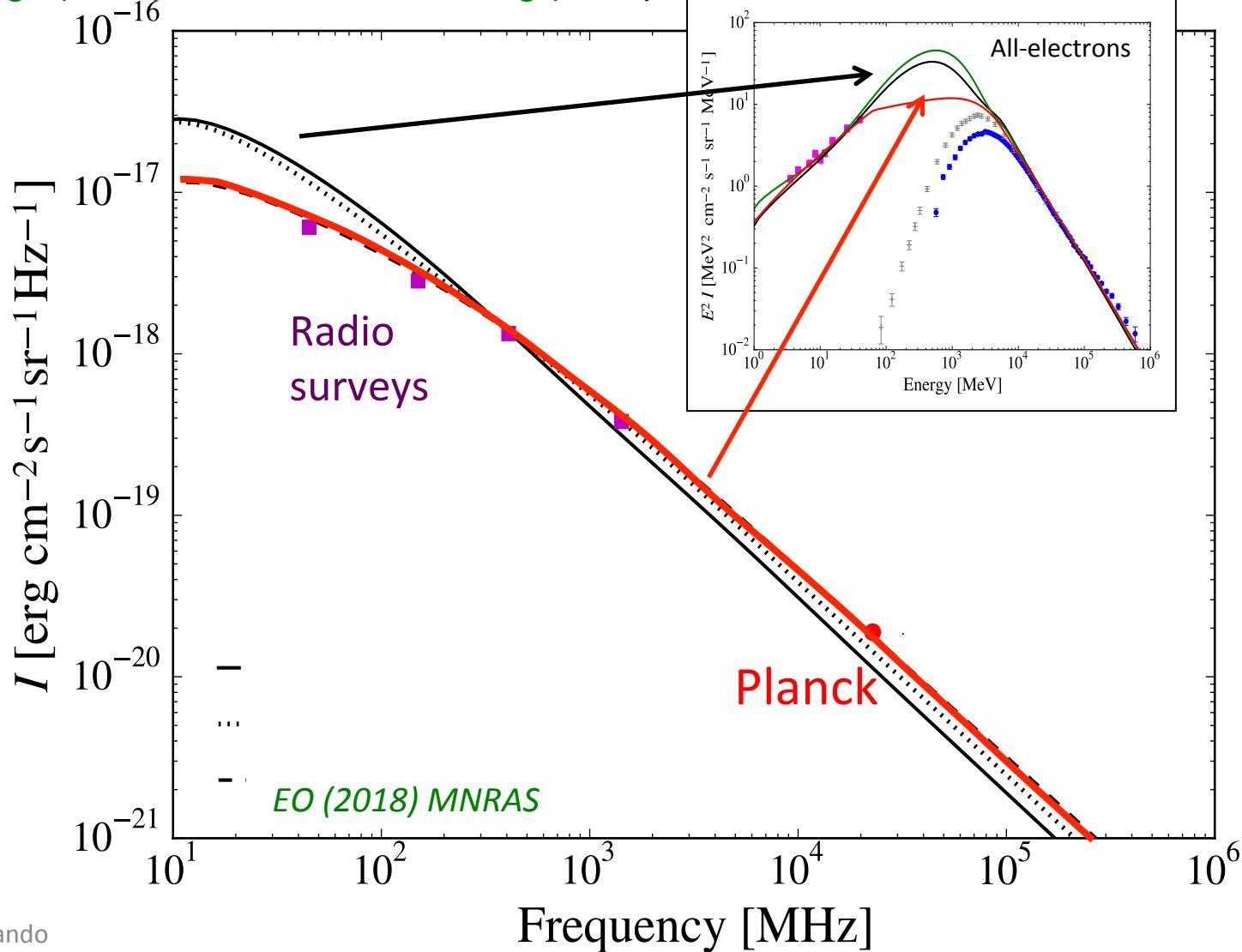
# All-Electron Spectrum



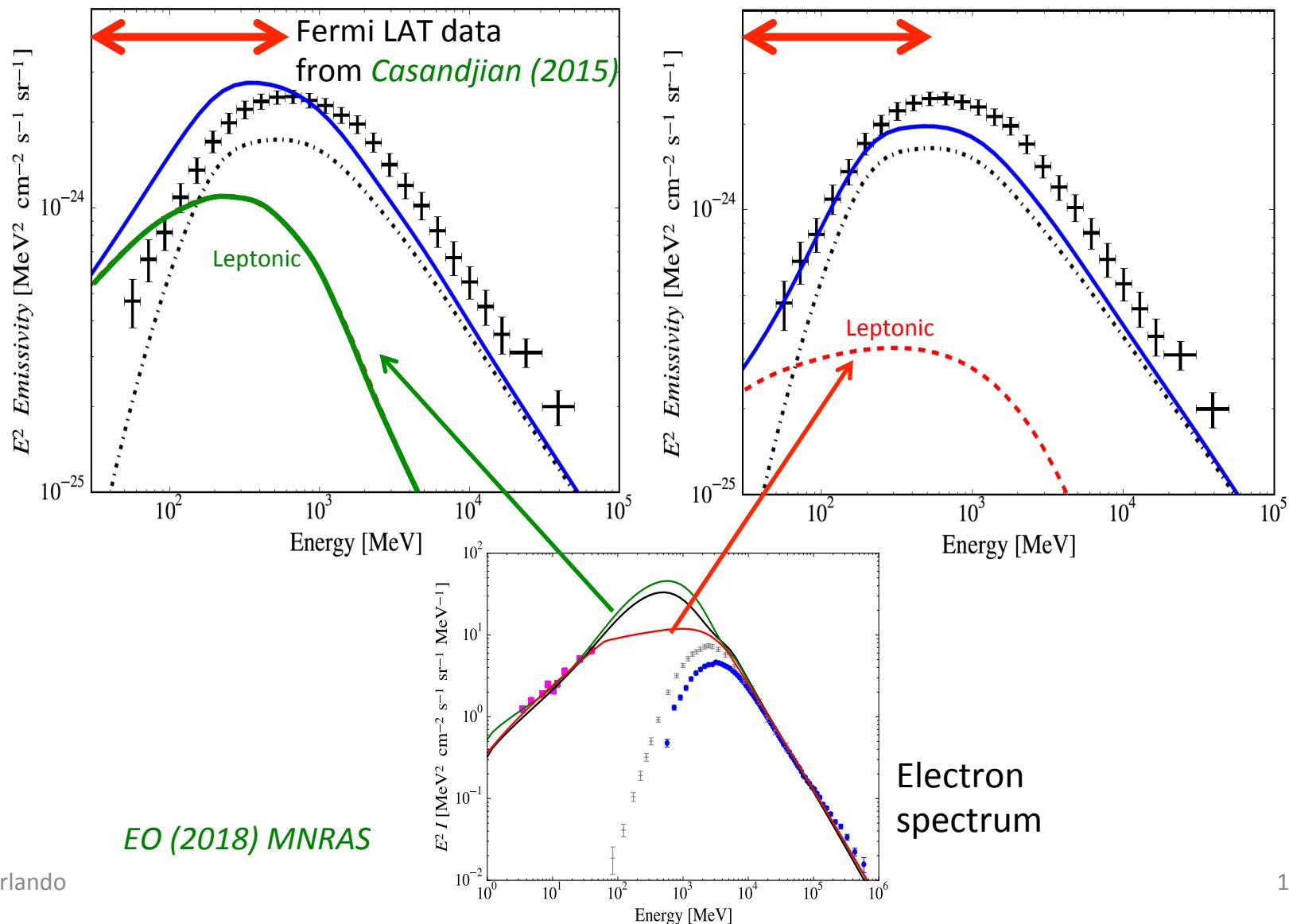
# Synchrotron Spectrum

CR & B-field intensity and data updated with respect to previous works

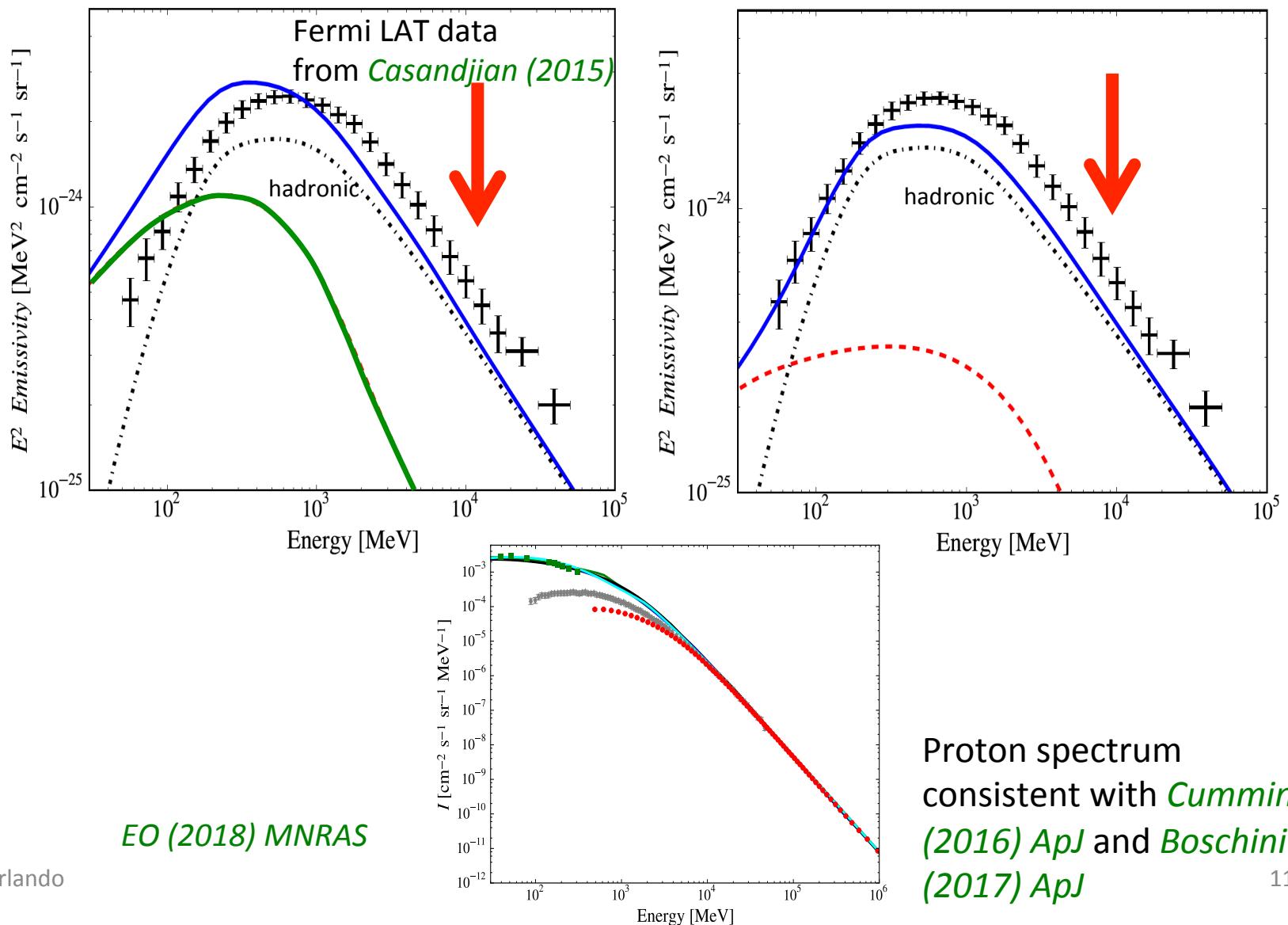
by *Strong+ (2011)* and *Orlando & Strong (2013)*



# Local HI Gamma-Ray Emissivity

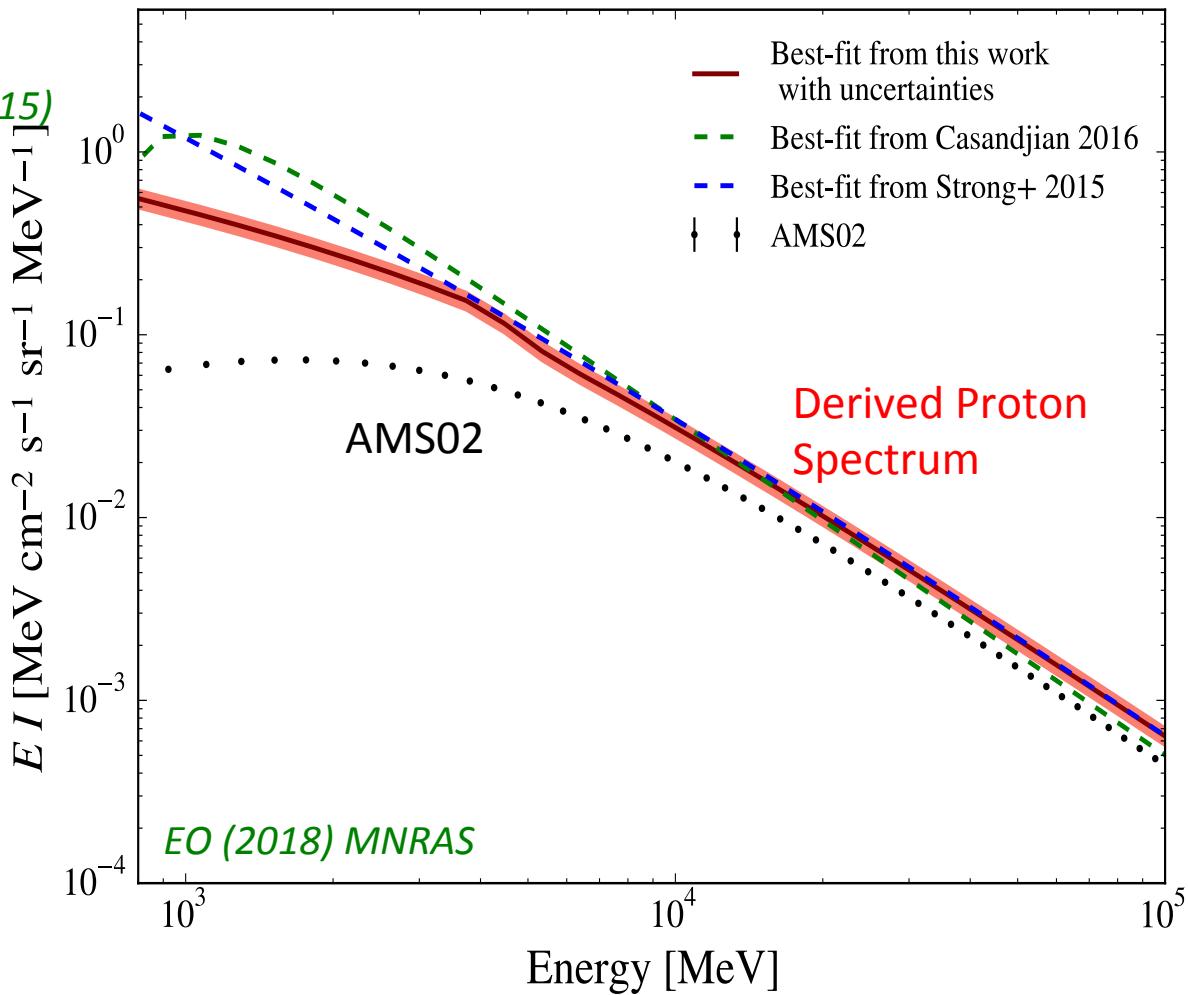
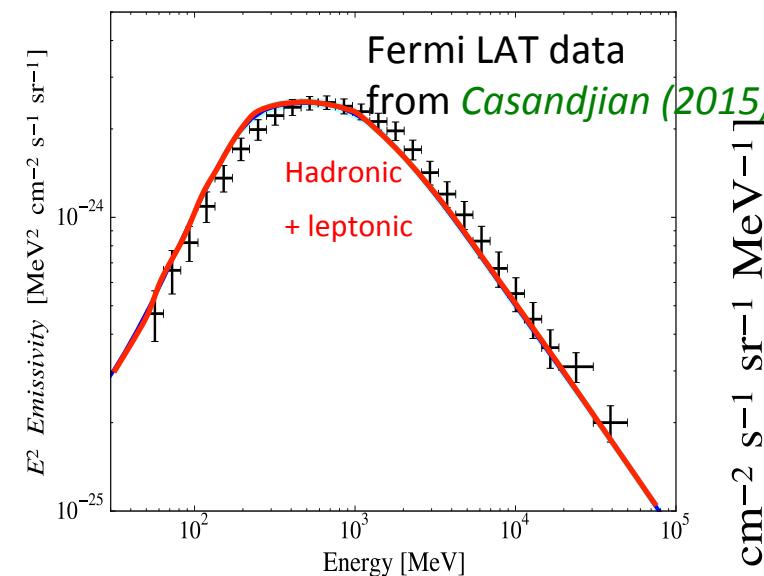


# Local HI Gamma-Ray Emissivity



# Derived Proton Spectrum

Fit to local HI Gamma-Ray Emissivity



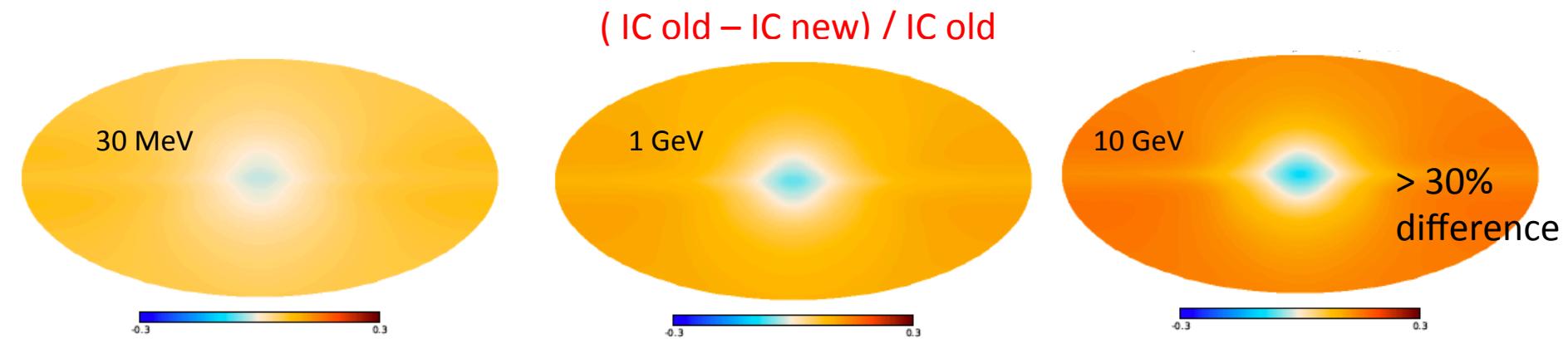
The direct measured proton spectrum may not resemble the local interstellar one

# Effects on Inverse Compton (IC) spatial templates of the gamma-ray interstellar emission

Important for the interpretation of Fermi LAT  
data and also for the generation of the  
interstellar template used for any analysis of  
the Fermi LAT data.

# Effects on Inverse Compton (IC) spatial templates of the gamma-ray interstellar emission

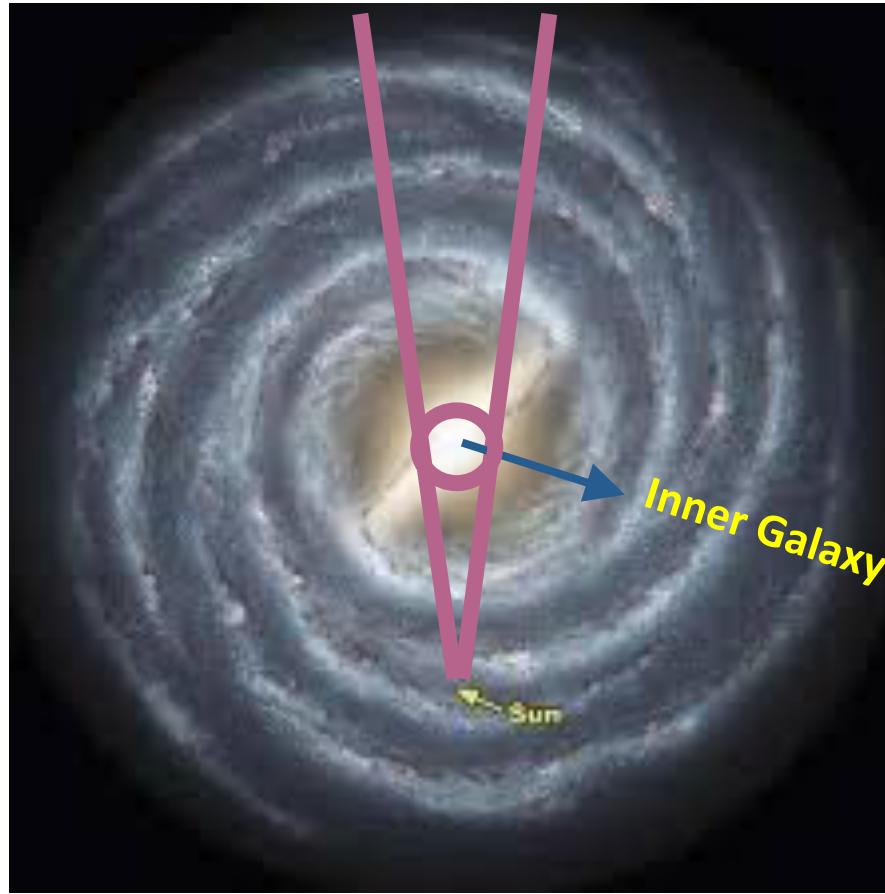
**Updated B-fields** produces brighter IC in the inner Galaxy than predicted by previous models for any photon field model



*EO (2019) Phys.Rev.D*

# Flashback: CRs toward the Galactic center

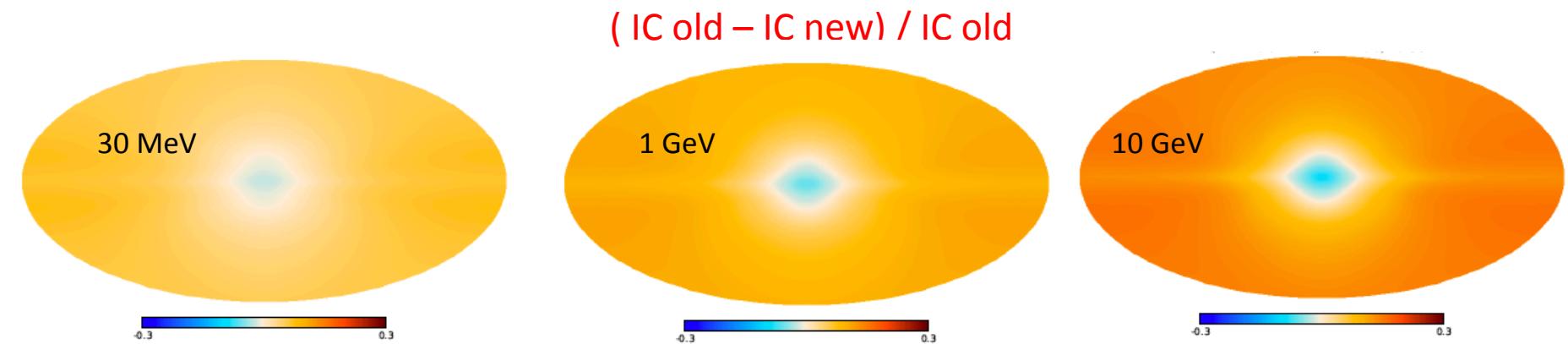
*Ajello et al. (2016) ApJ*



- 1) IC dominant and brighter than predicted by standard models
  - 2) Hardening in the inner Galaxy with respect to standard models
- CRs? – ISRF? - Unresolved sources? - Dark Matter? -

# Effects on Inverse Compton (IC) spatial templates of the gamma-ray interstellar emission

**Updated B-fields** produces brighter IC in the inner Galaxy than predicted by previous models for any photon field model



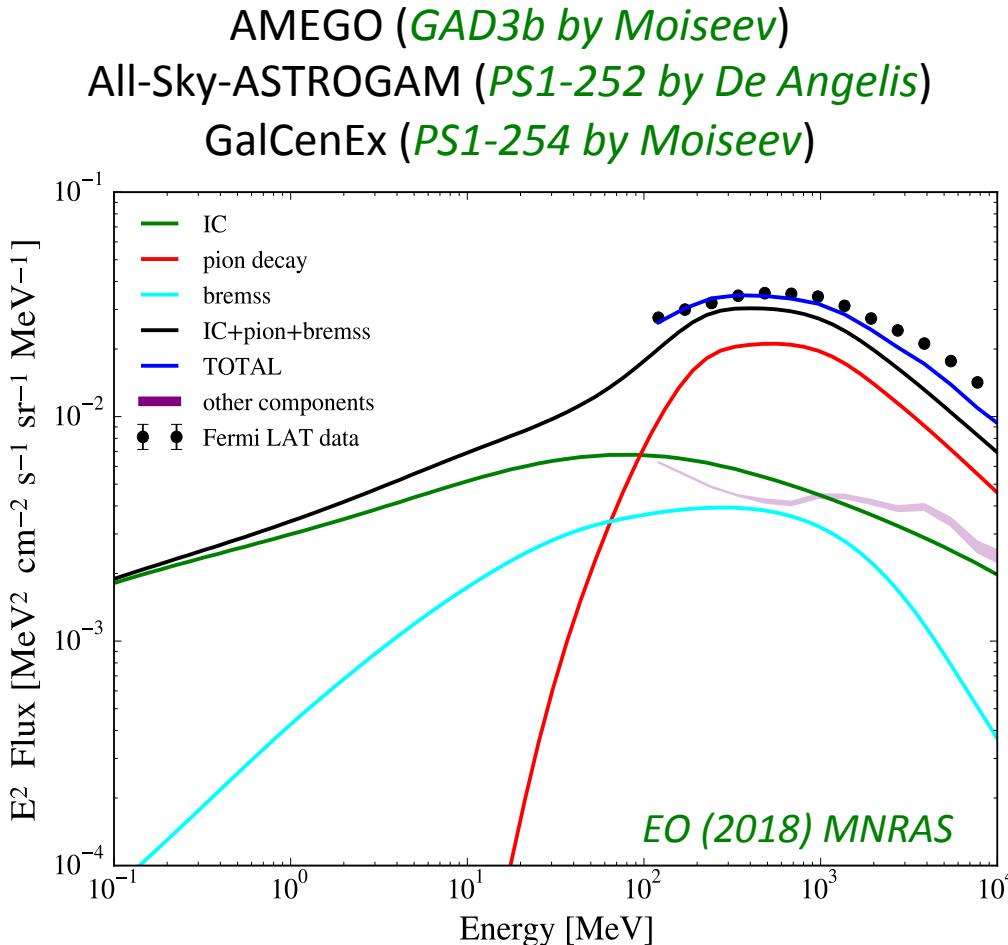
*EO (2019) Phys.Rev.D*

With respect to the standard models, these models:

- provide alternative propagation scenarios
- are consistent with radio and microwave data
  - are updated to recent CR measurements

Next step: to be tested with Fermi LAT data

# Predictions – MeV Gammas



See also *Orlando et al 2019 “Cosmic Rays and Interstellar Medium with Gamma-Ray Observations at MeV Energies” ASTRO2020 Decadal Survey on Astronomy and Astrophysics, Science white paper 151 (also Poster PS3-247)*

# Summary

