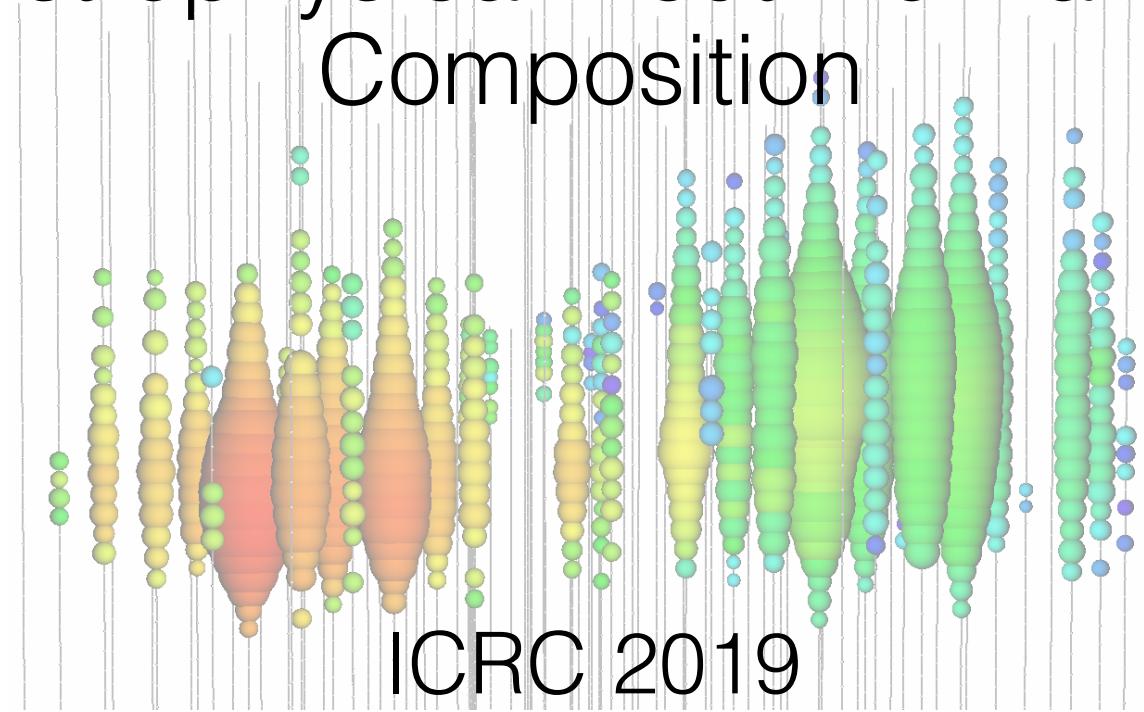


Cosmic Tau Neutrinos and the Astrophysical Neutrino Flavor Composition



July 30, 2019
Juliana Stachurska



Astrophysical Tau Neutrinos

- Measure spectrum and flavor composition → learn about environment at cosmic accelerators → constrain production mechanisms
- ν_τ needed for flavor composition measurement
- Negligible atmospheric ν_τ flux
- Guaranteed astrophysical ν_τ flux due to neutrino mixing
- Finding cosmic ν_τ :
 - Another confirmation of the astrophysical content of diffuse neutrino flux
 - Finding last SM cosmic messenger
 - Verification of neutrino mixing at largest baselines, highest energies

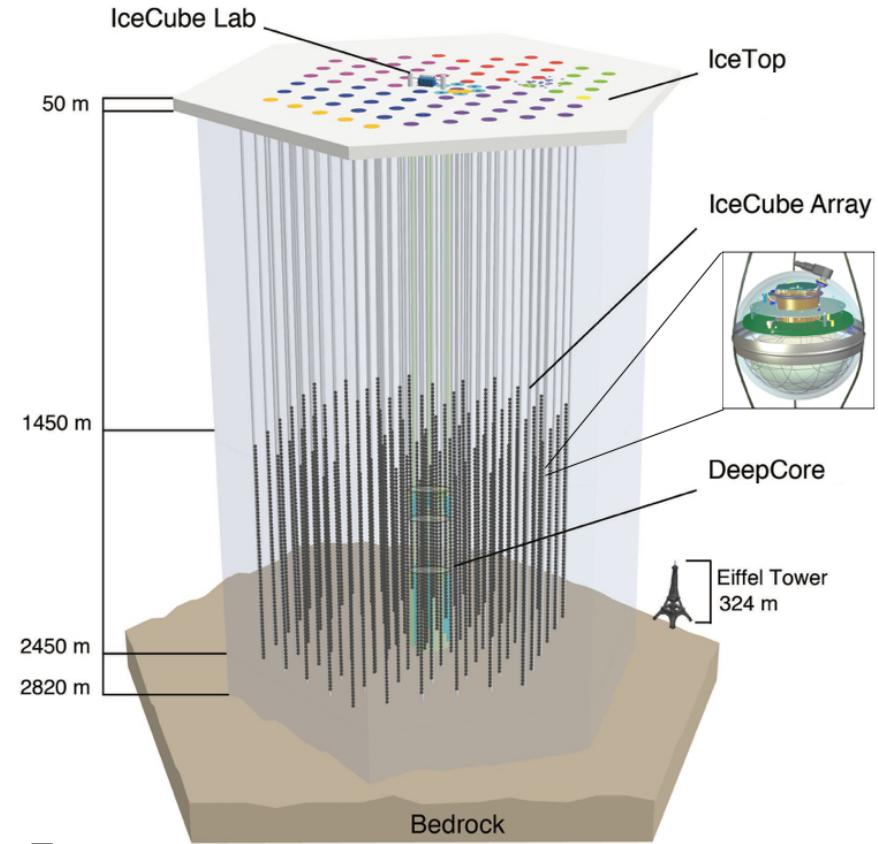
complementarity

discovery

universality &
BSM constraints

IceCube

- km³ scale Cherenkov detector
- At South Pole
- Instrumented Antarctic ice in 1.5-2.5km depth
- 5160 Digital Optical Modules (DOMs)
- ~1 ns time resolution



	Horizontal spacing [m]	Vertical spacing [m]	Energy threshold [GeV]	
IceCube	125	17	~100	Astrophysical neutrinos
DeepCore	50	7	~5	Atmospheric oscillations

Topologies in IceCube

early  late

Single Cascade

ν_e CC, NC

Double Cascade

ν_τ CC

Track

ν_μ CC

All NC interactions
 ν_e CC interactions

Good energy resolution
Bad angular resolution

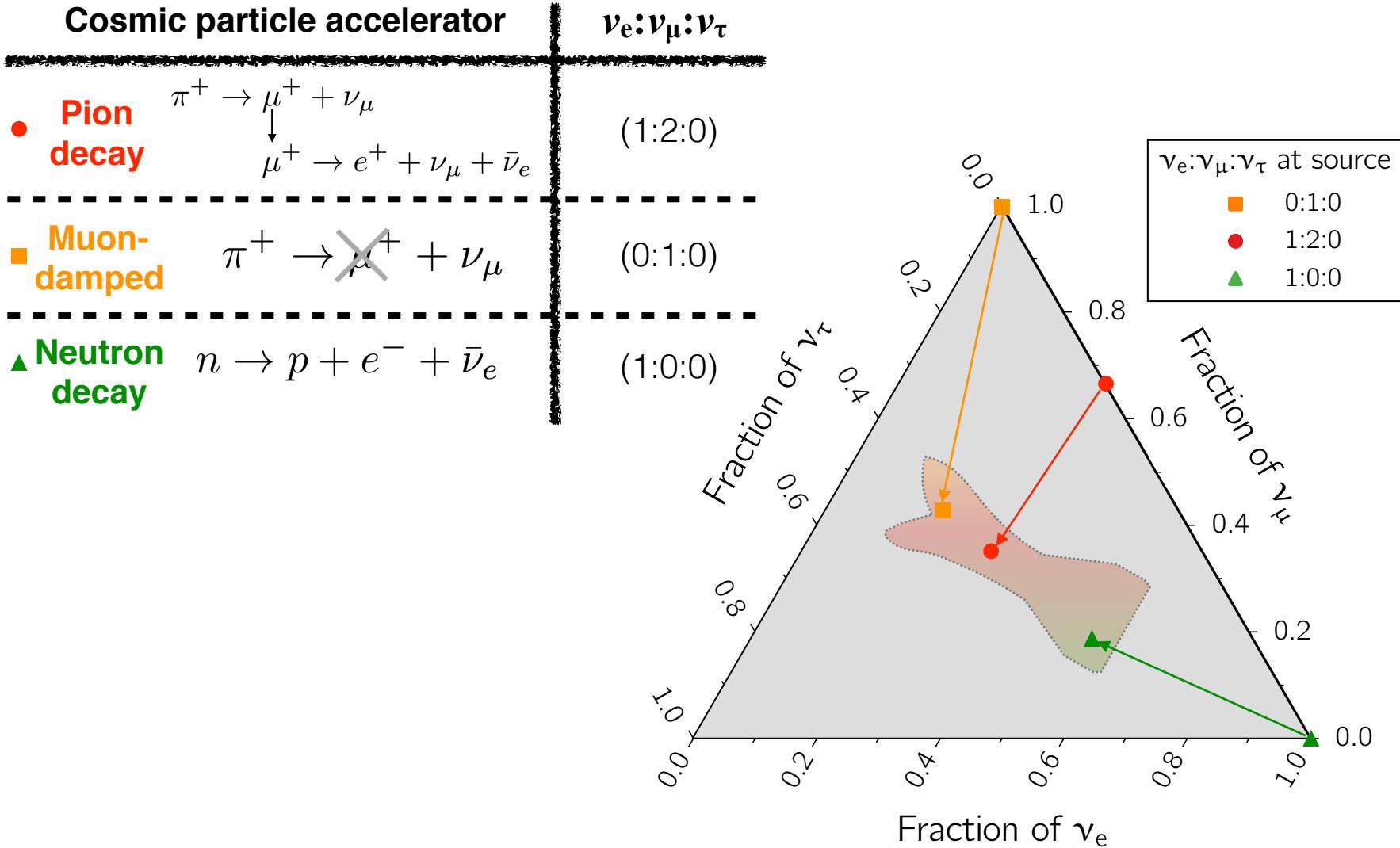
ν_τ CC interactions with hadronic / electronic tau decay

Good energy resolution
Angular resolution gets better with larger lengths

ν_μ CC interactions
Atmospheric μ
 ν_τ CC interactions with muonic tau decay

Bad energy resolution
Good angular resolution

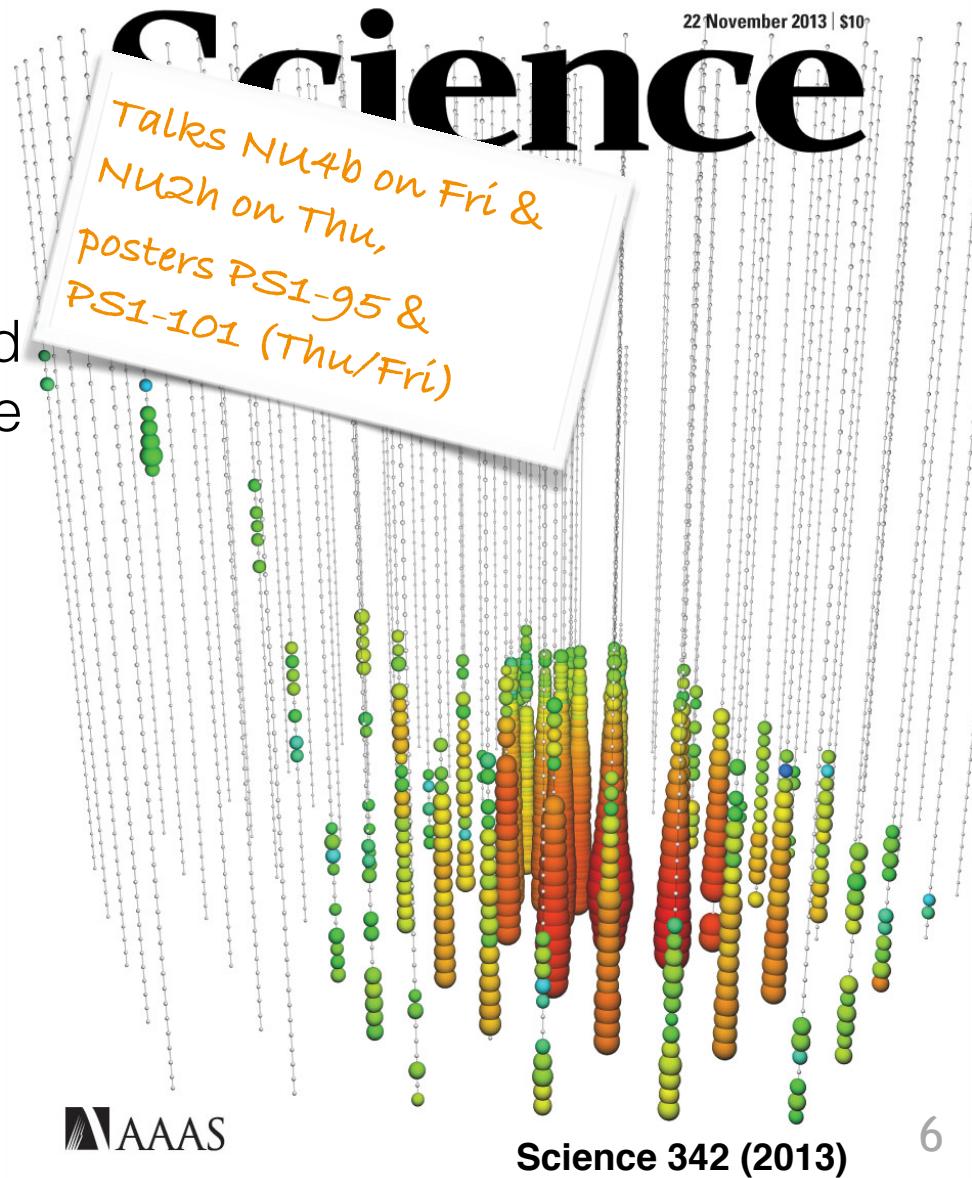
Flavor Composition





High-Energy Starting Events (HESE)

- Evidence of diffuse astrophysical neutrino flux with 2 years livetime
- All-flavor event selection
- Revisit of HESE and associated analyses with 7.5 years livetime
- Ternary topology ID algorithm added
 - Cascades
 - Tracks
 - Double Cascades

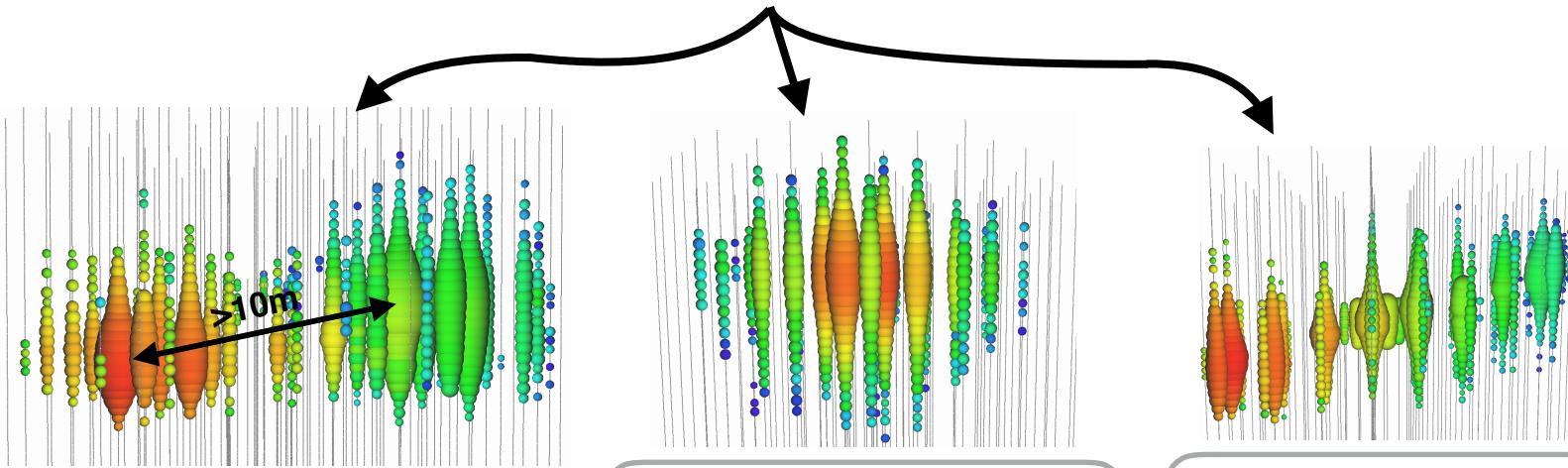


Flavor Analysis Overview

60 HESE events in 7.5 years of data above 60 TeV



Observables from direct double-cascade reconstruction



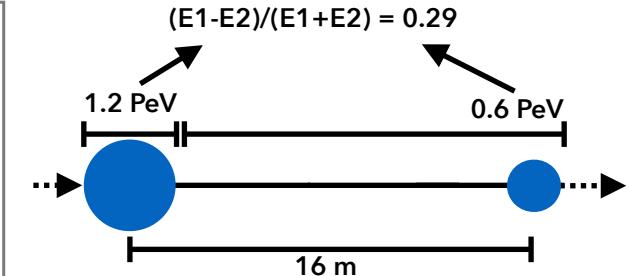
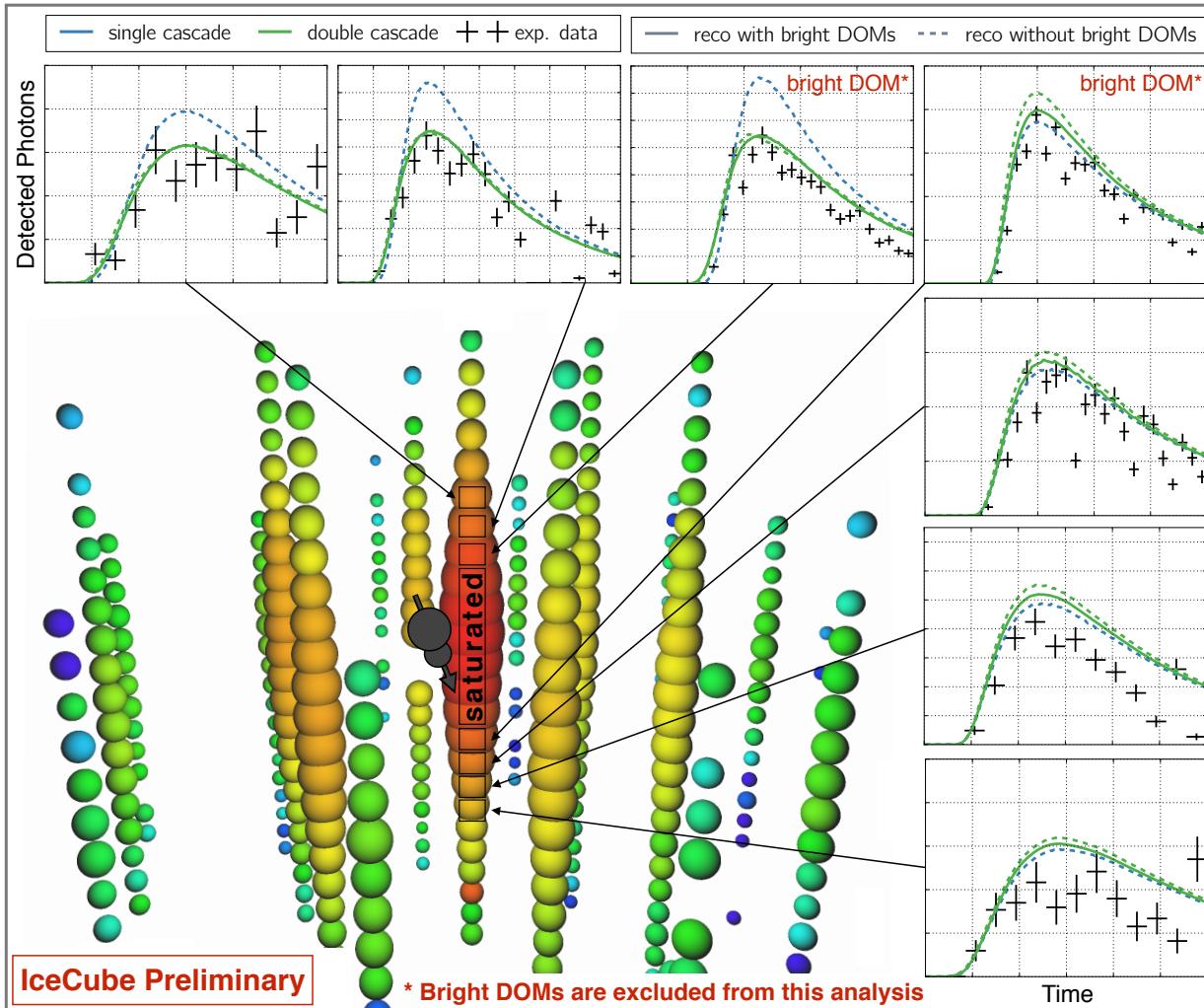
Double cascade (2):
 $\nu_\tau + N \rightarrow \tau + \text{hadrons} \xrightarrow{>10m}$
hadrons / electrons

Single cascade (42):
not well-reconstructable ν_τ
– CC interactions, all other
cascades

Track sample (16):
 ν_τ – CC interactions
creating μ , ν_μ – and
atm. μ – tracks

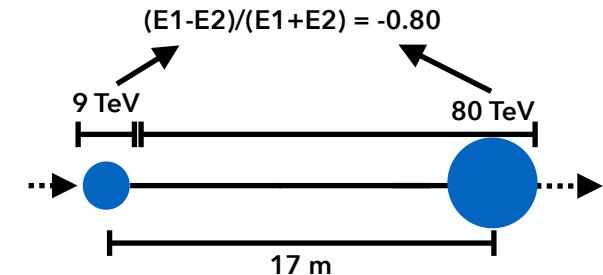
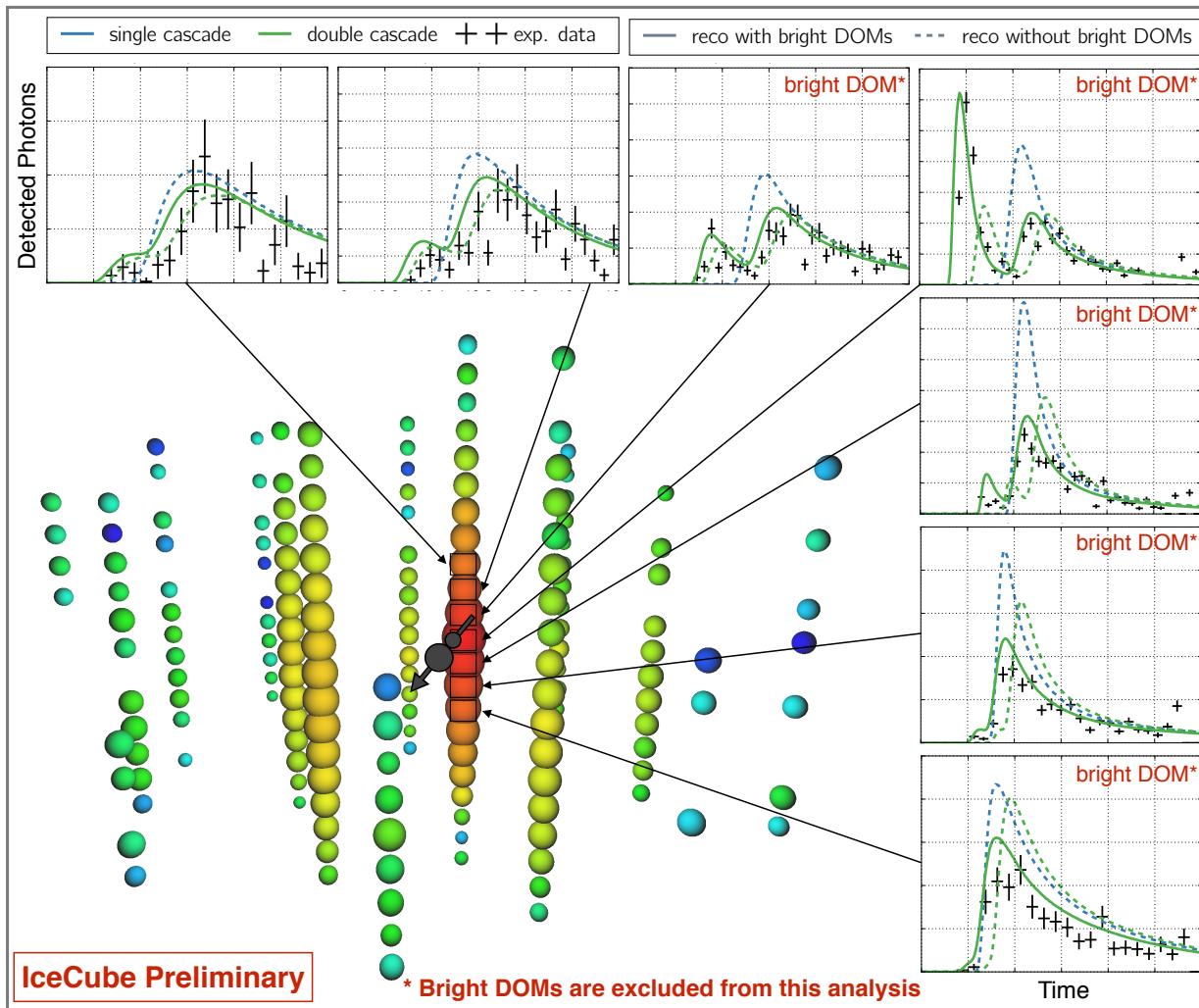
Flavor composition

Event #1 (Big Bird)



- Observed 2012
- Shows no clear preference between a single cascade and double cascade hypothesis

Event #2 (Double Double)

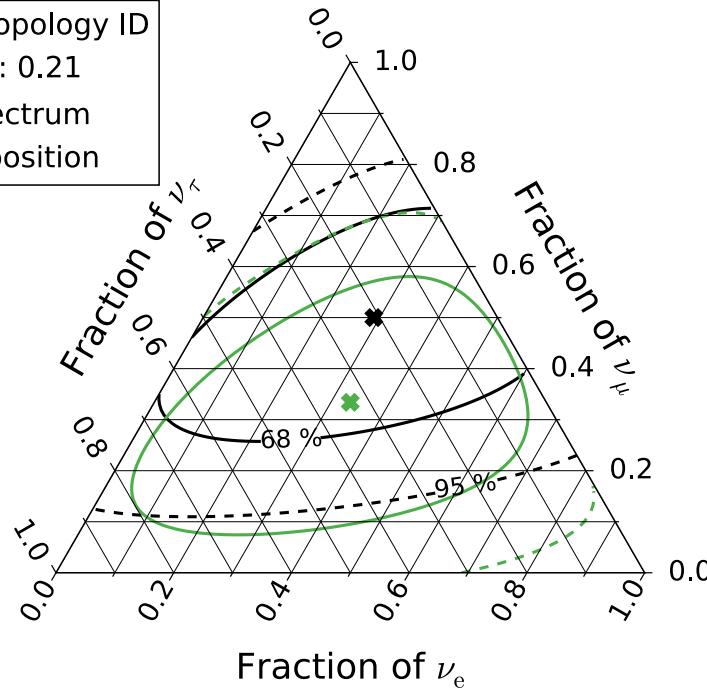


- Observed 2014
- Observed light arrival pattern clearly favors double cascade hypothesis

Flavor Composition

- HESE with ternary topology ID
- * Best fit: 0.29 : 0.50 : 0.21
- Sensitivity, $E^{-2.9}$ spectrum
- * 1 : 1 : 1 flavor composition

WORK IN PROGRESS

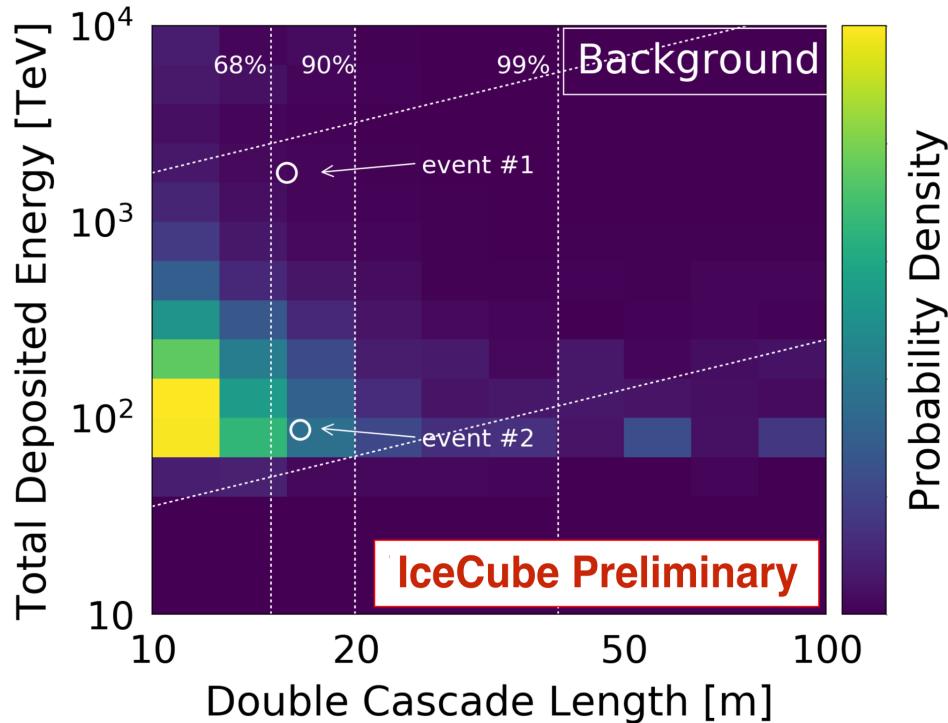
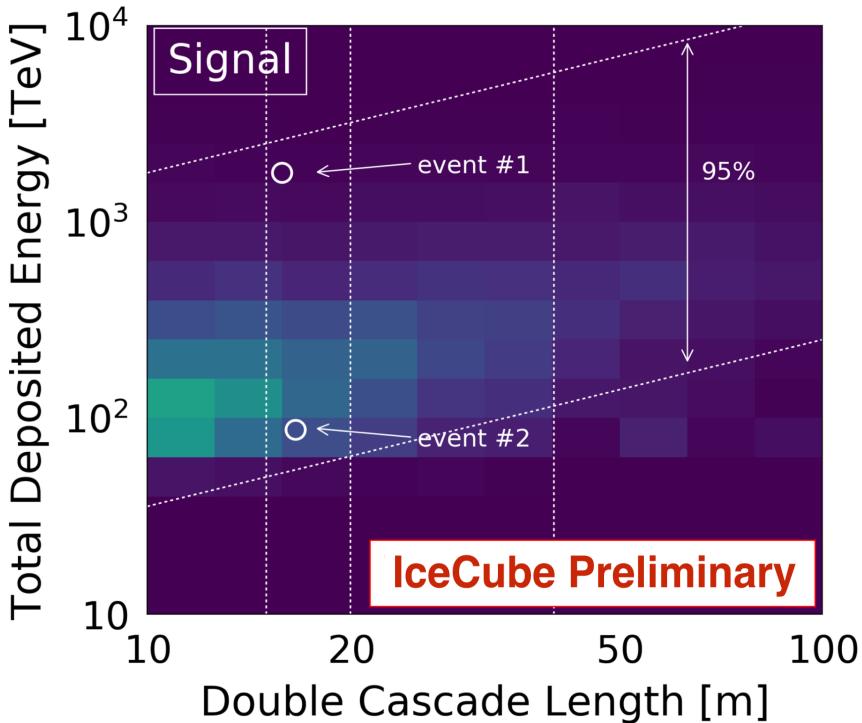


First probe of neutrino oscillations over cosmological baselines and at TeV energies!

First best fit non-zero in each flavor component!

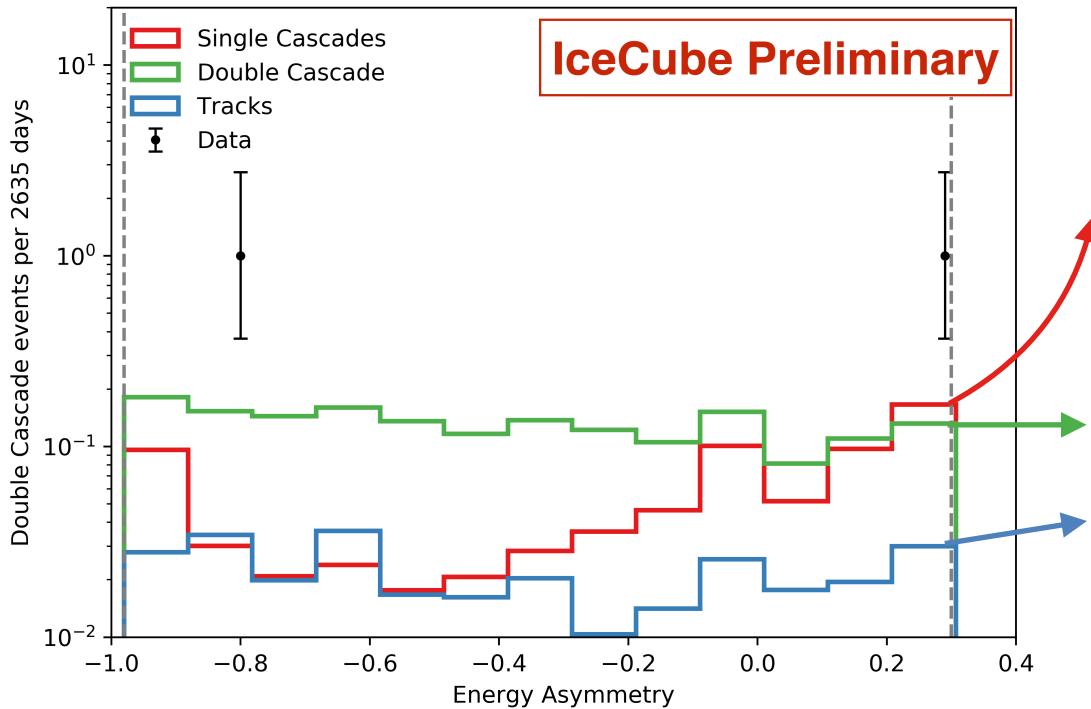
- Best-fit $\nu_e:\nu_\mu:\nu_\tau = 0.29:0.50:0.21$
- Consistent with previous measurements and expectation of $\sim 1:1:1$ for astrophysical neutrinos
- Zero ν_τ flux cannot be excluded
- Systematic errors not included

Energy & Length



- 2 events with Double Cascade topology
- Soft spectral index: 2.9 → expect ~2.1 events (~1.4 signal + ~0.7 background)

Energy Asymmetry

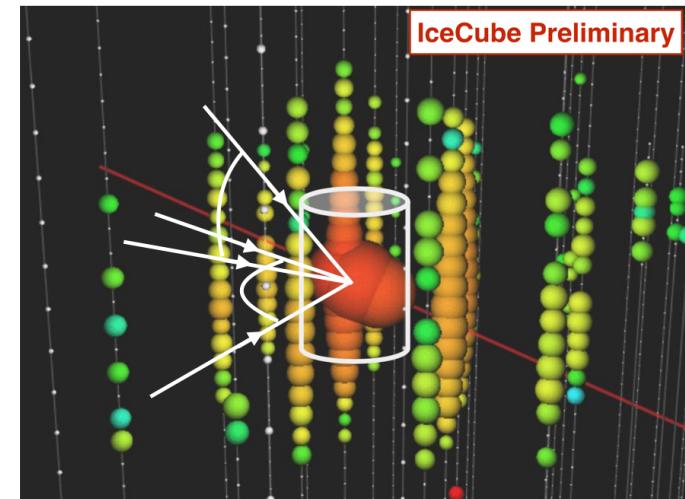


- $A_E = (E_1 - E_2) / (E_1 + E_2)$
- Only a straight cut was used
- Mainly due to computational issues
- Want to use all event properties

Energy asymmetry for best-fit spectrum,
and a $\nu_e : \nu_\mu : \nu_\tau = 1:1:1$ composition

Follow-Up

- $P(\nu_\tau | \mathbf{x})$ for Big Bird / Double Double?
→ Resimulate!
- Save computational power
→ Restrict parameter space
- Simulate ν_e, ν_μ, ν_τ and atmospheric μ (Double Double)
- 13M Double-Double-like events & 1M Big-Bird like events pass HESE selection
- Evaluate! → Tauness



Tauness:

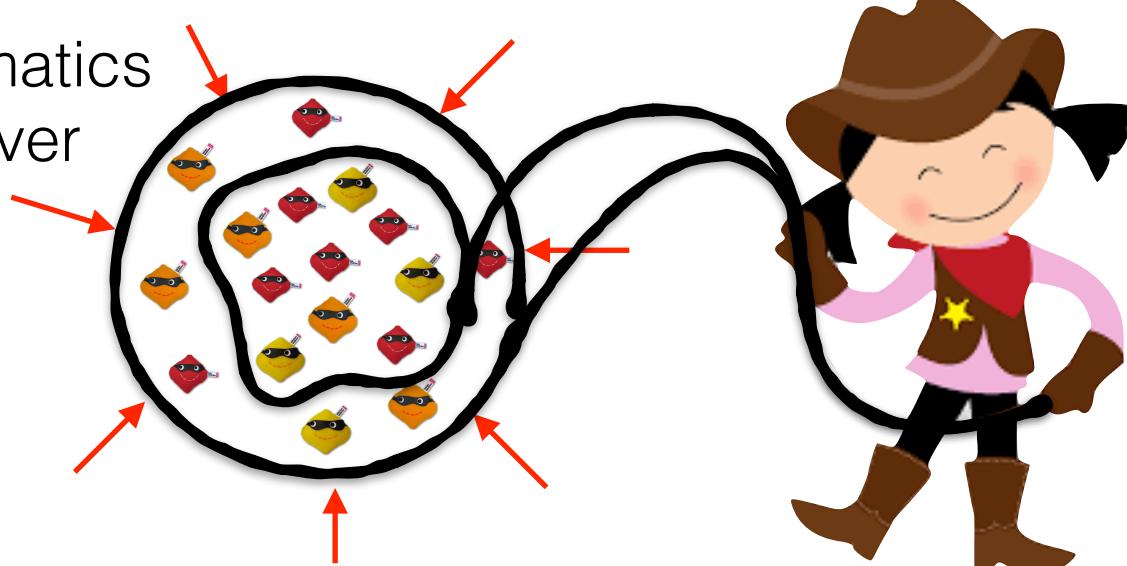
$$P(\nu_\tau | \mathbf{x}) \approx \frac{dN_S^{DC}/d\mathbf{x}}{dN_S^{DC}/d\mathbf{x} + dN_B^{DC}/d\mathbf{x}} \equiv \tau$$

differential signal event rates
↓
 $dN_S^{DC}/d\mathbf{x}$
↑
differential background event rates

with $\mathbf{x} = (x, y, z, \theta, \phi, E_{tot}, L, A_E)$

Follow-Up Results

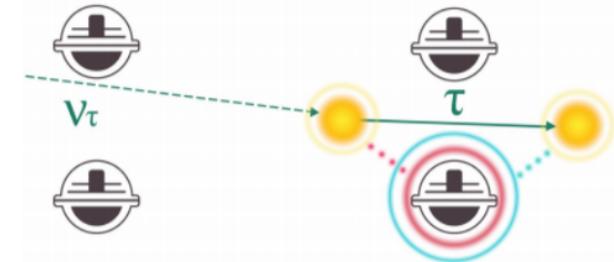
- Use 6D KDE with Rodeo* algorithm & ROI in A_E , L
- Assume best-fit $E^{-2.89}$ spectrum, 1:1:1 flavor composition (see talk NU4b on Fri)
- Uncertainties & systematics not yet marginalized over



- **$\geq 97\%$ of Double-Double-like events are ν_τ -induced**
- **$\sim 75\%$ of Big-Bird-like events are ν_τ -induced**
- Statistics still limiting factor

Double Pulse

- Another feature of ν_τ CC interactions
- Analyze waveform on brightest DOMs
- No full reconstruction of event
- Information used complementary to Double Cascade method presented here
- 2 analyses presented here:
 - Results using straight cuts
 - Results using machine learning

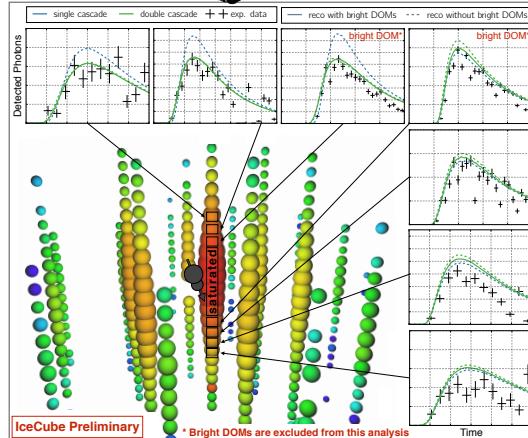


Talk NU110 on Wed

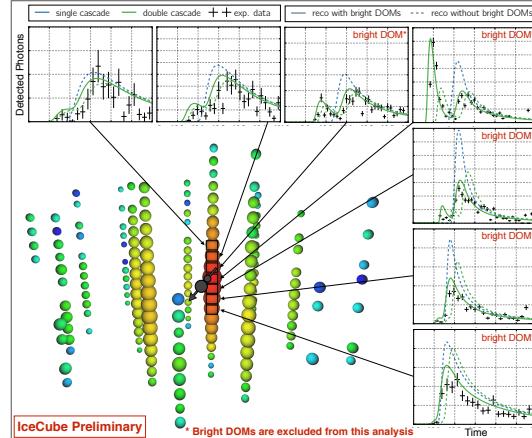
Poster PS1-117 (Thu/Fri)

Summary

Big Bird

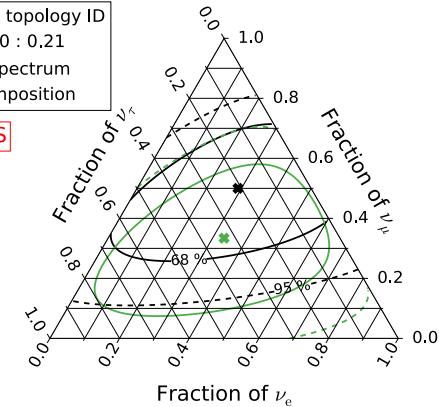


Double Double



— HESE with ternary topology ID
 * Best fit: 0.29 : 0.50 : 0.21
 — Sensitivity, $E^{-2.9}$ spectrum
 ✕ 1 : 1 : 1 flavor composition

WORK IN PROGRESS



Best fit $\neq 0$ in all flavors!



($\approx 75\% \nu_\tau$)



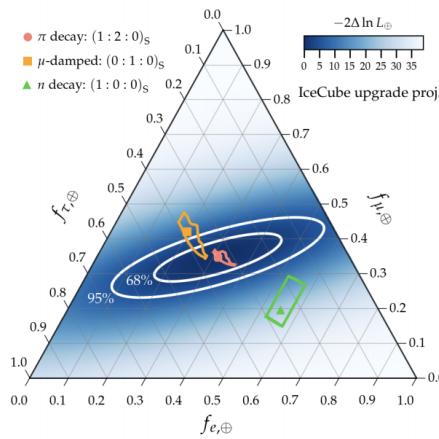
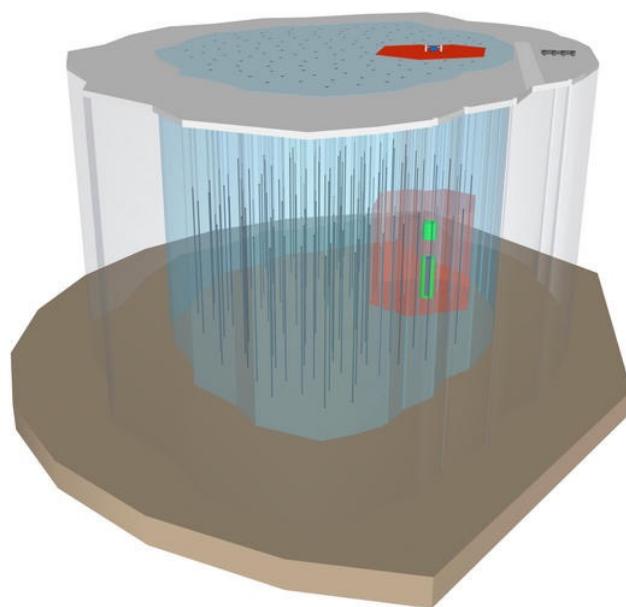
($\geq 97\% \nu_\tau$)

SM Cosmic messengers:

e, p, A	✓	ν_e	✓
γ	✓	ν_μ	✓
GW	✓	ν_τ	✓

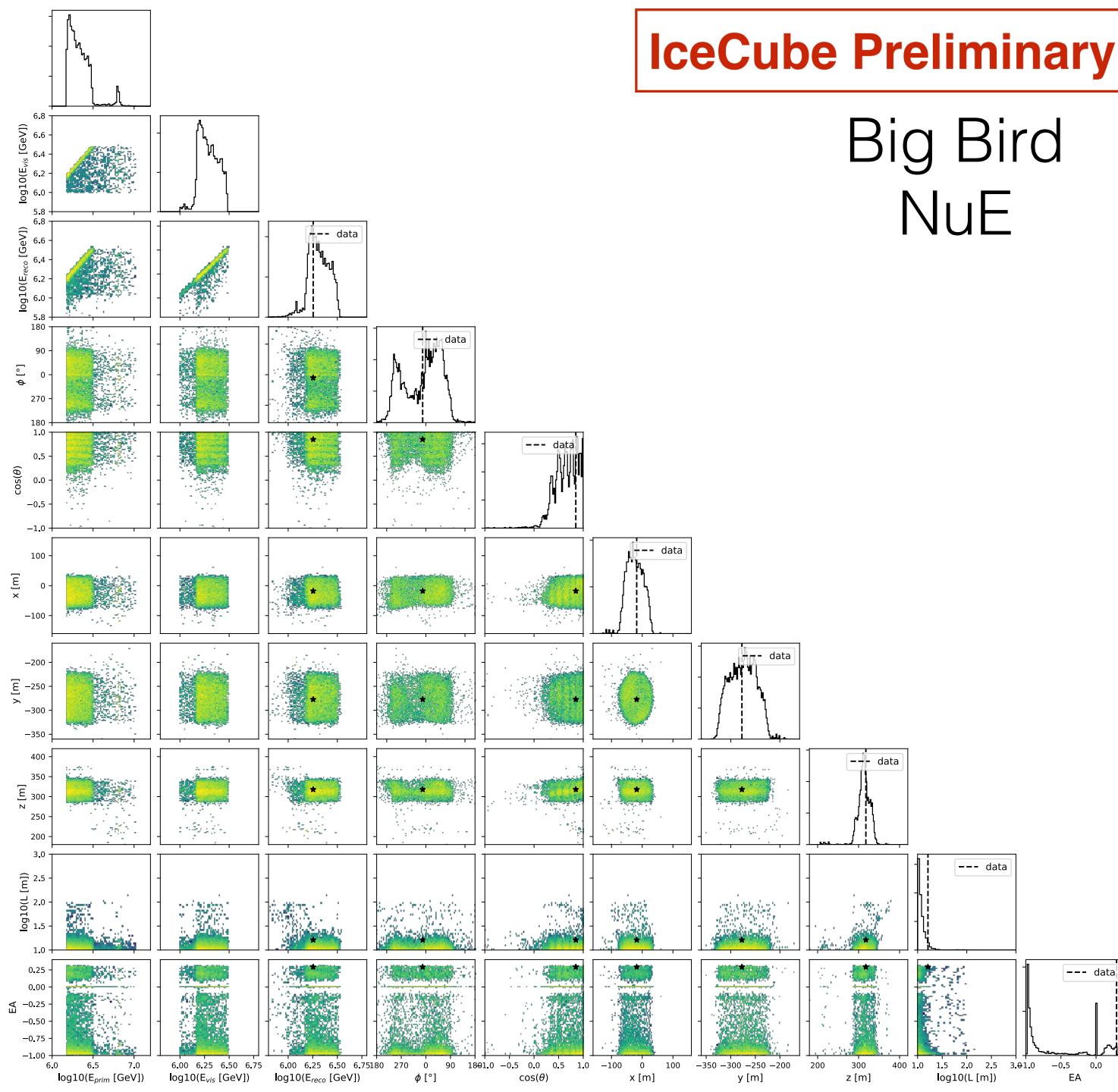
Back up

The Future of Astrophysical ν_τ

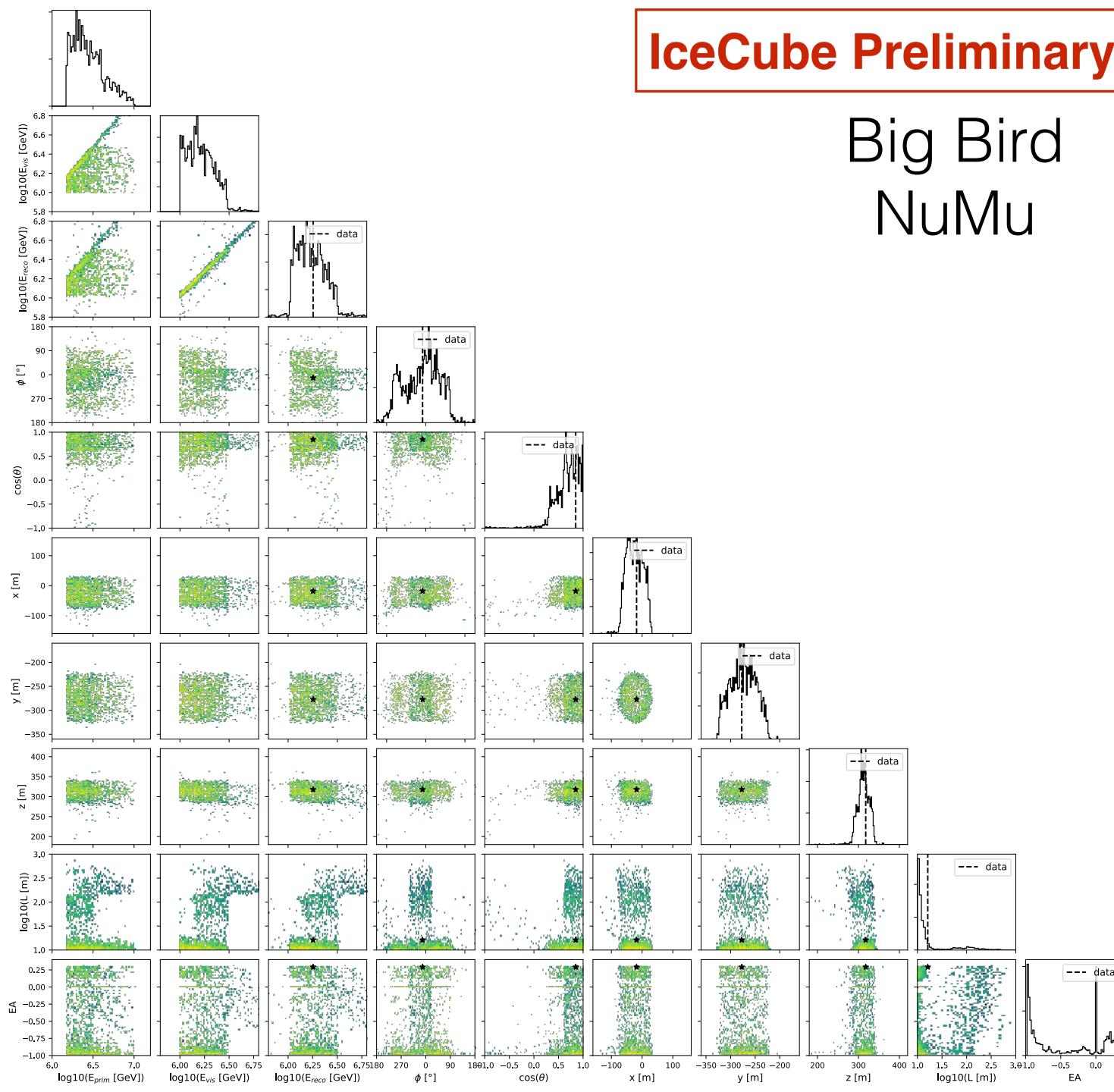


- IceCube Upgrade
 - Precision measurements of ice optical properties
 - In-situ verification of ν_τ reconstruction at ~ 20 m baselines
- IceCube Gen2
 - More volume
 - More identifiable ν_τ expected
- Higher precision for flavor composition measurement

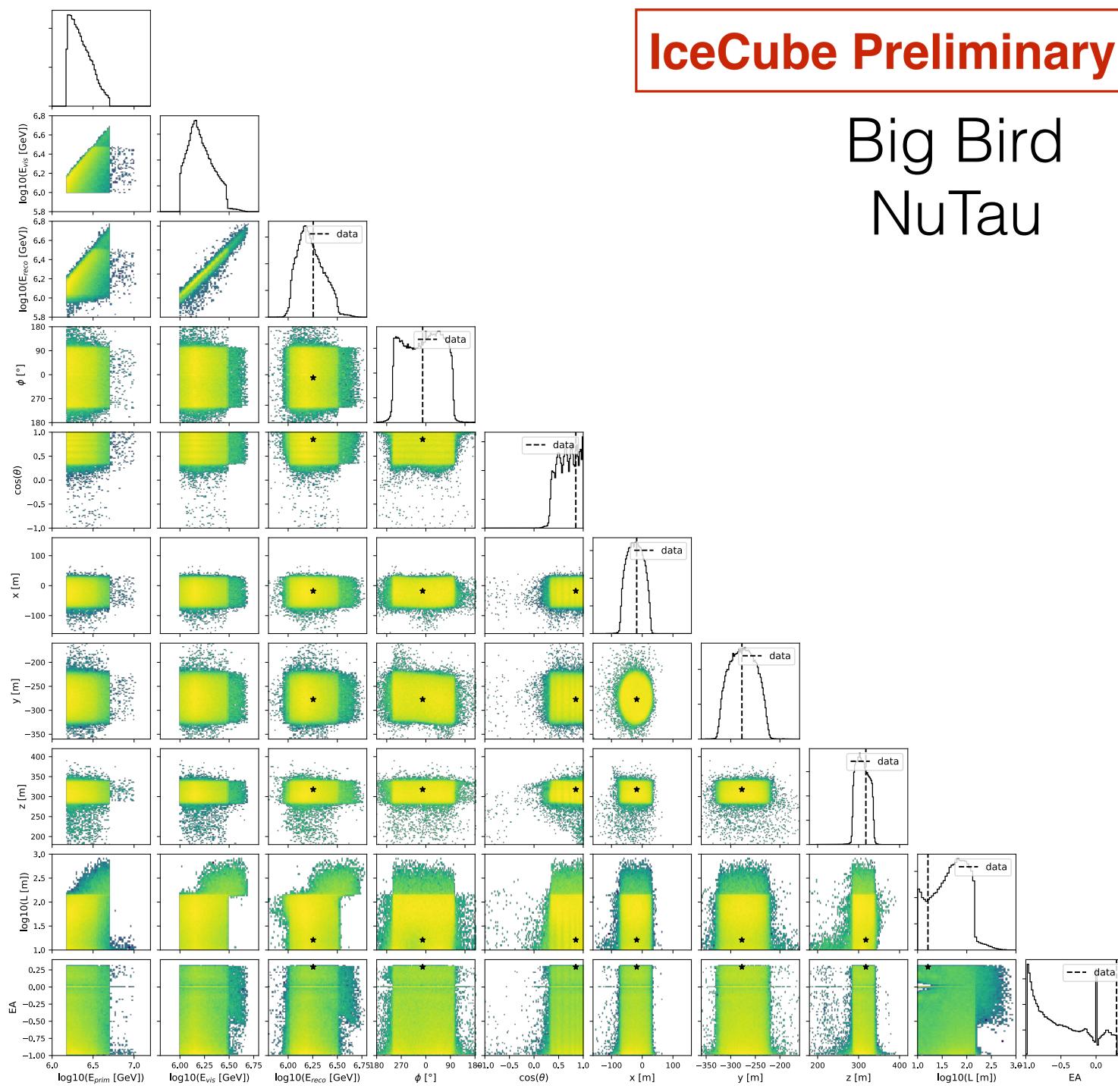
Big Bird NuE

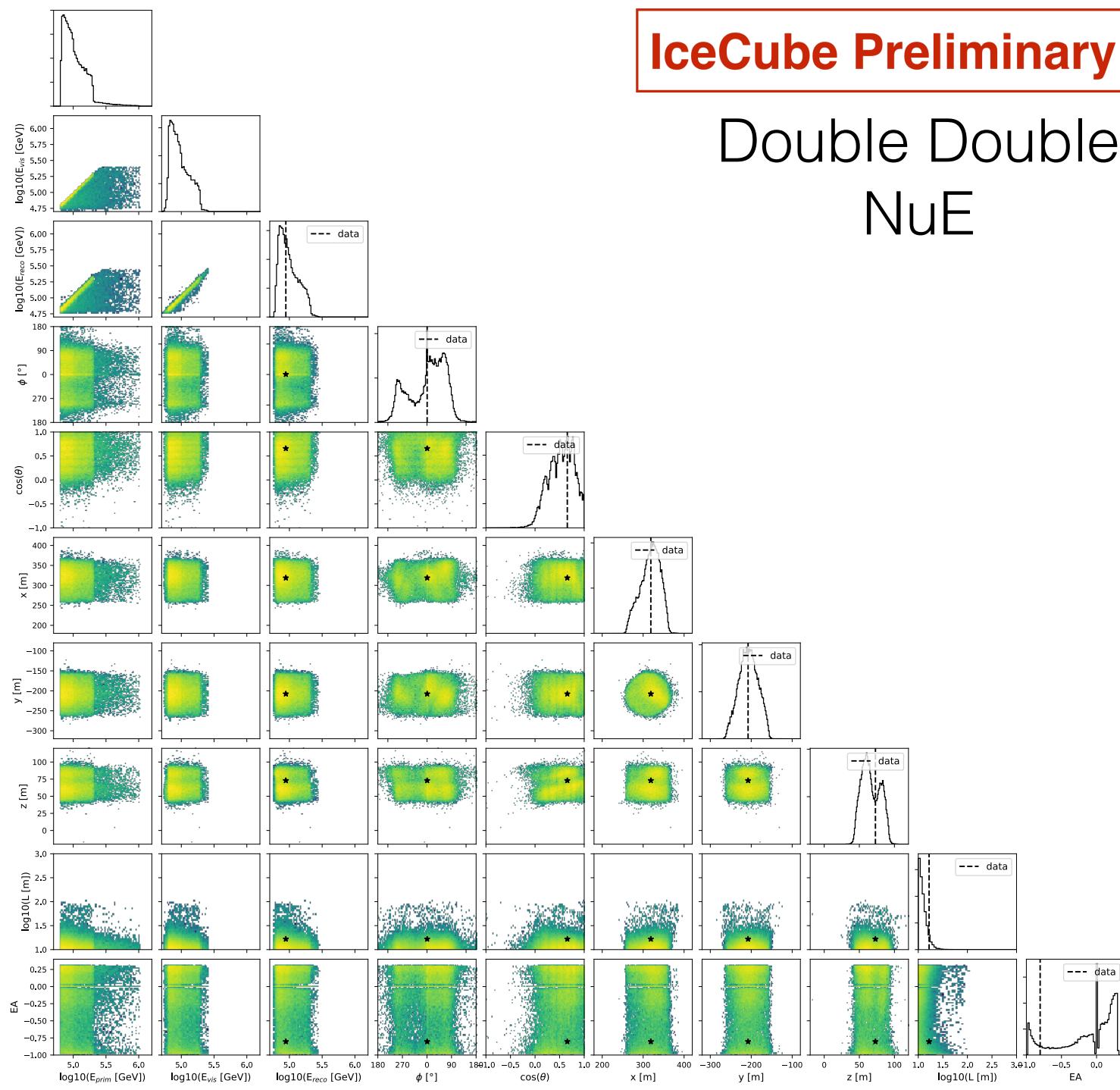


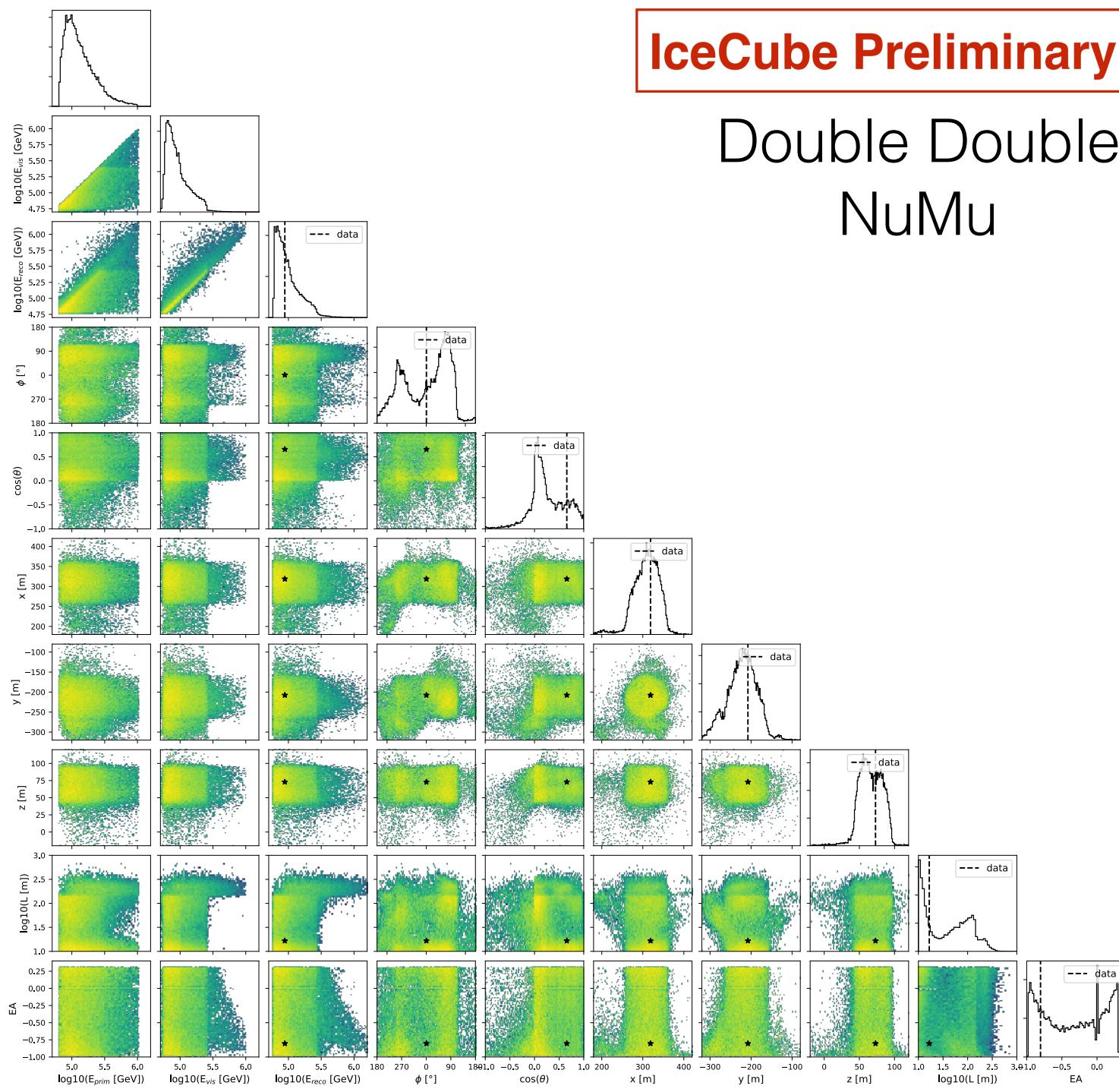
Big Bird NuMu

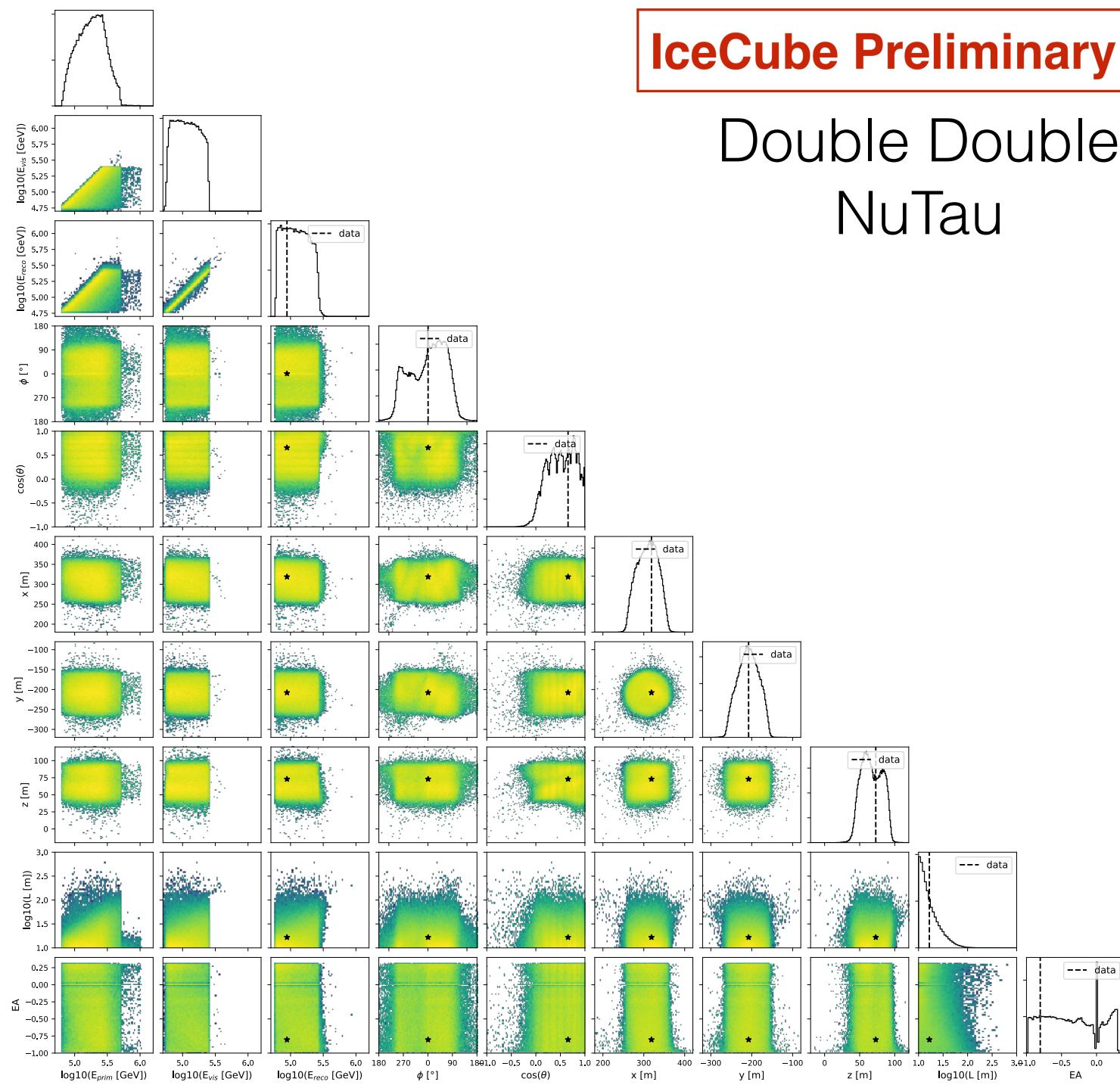


Big Bird NuTau

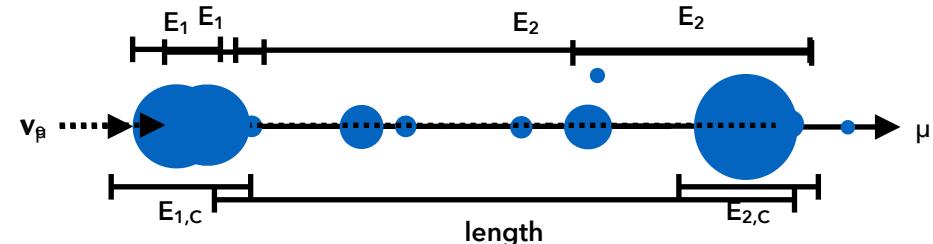
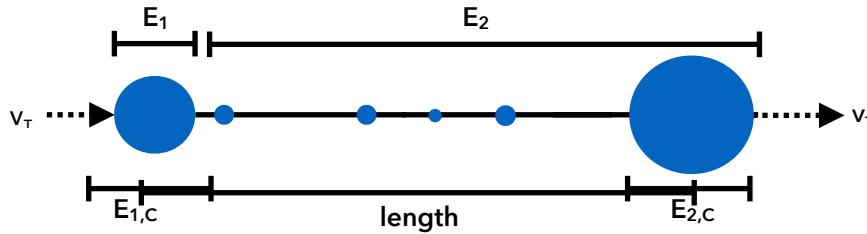
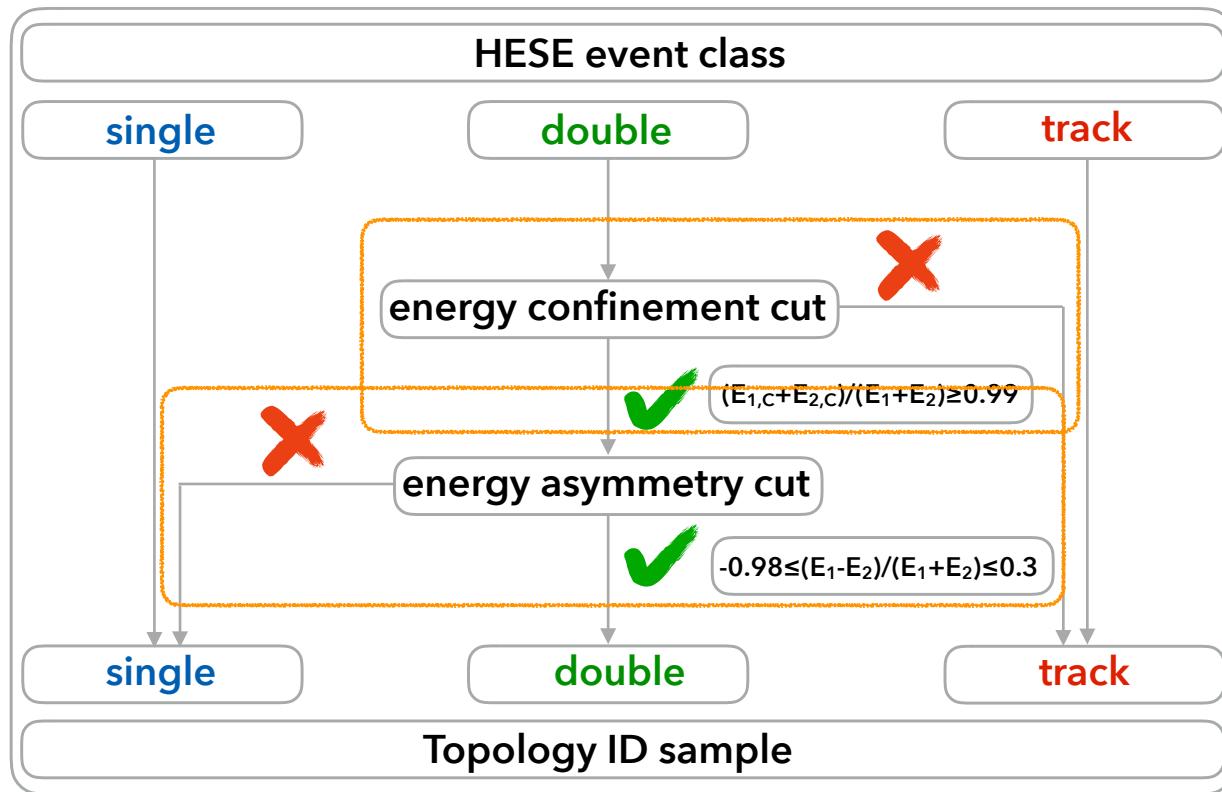






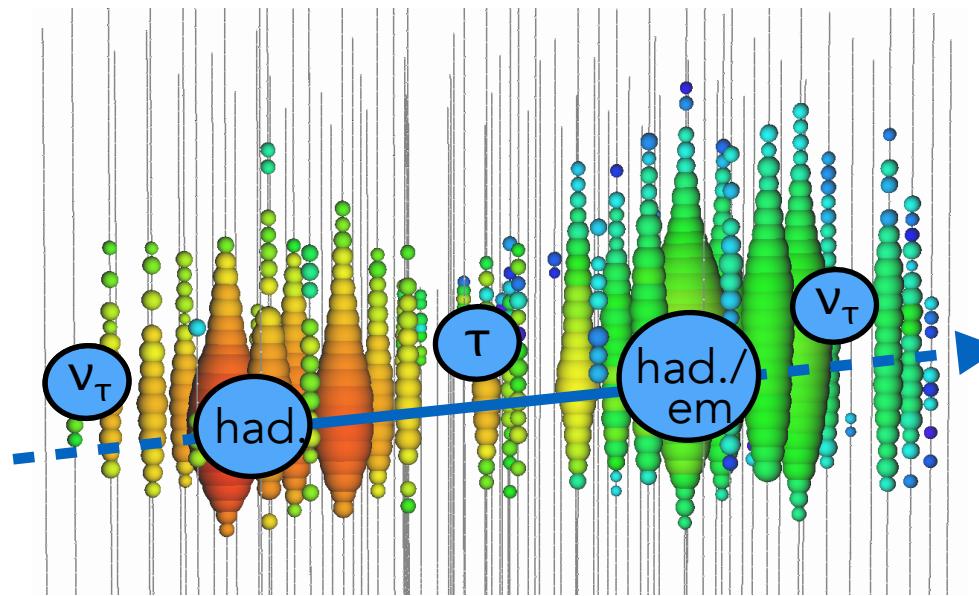


Selection



Double Cascade Signal

- ν_τ interaction
- Charged current (71%)
- Tau decays into hadrons / electrons (83%)
- Mean length: 50m x energy/1PeV



simulated 10PeV Double Cascade event

Astrophysical Tau Neutrinos

- Tau neutrinos needed for flavor composition measurement
- Negligible atmospheric ν_τ flux
- Guaranteed astrophysical ν_τ flux due to neutrino mixing
- Finding cosmic ν_τ :
 - Finding the last SM cosmic messenger
 - Yet another confirmation of the diffuse neutrino flux being astrophysical

- Discovery
- Complementarity
- Universality
- BSM constraints

on of neutrino mixing at baselines, highest energies

