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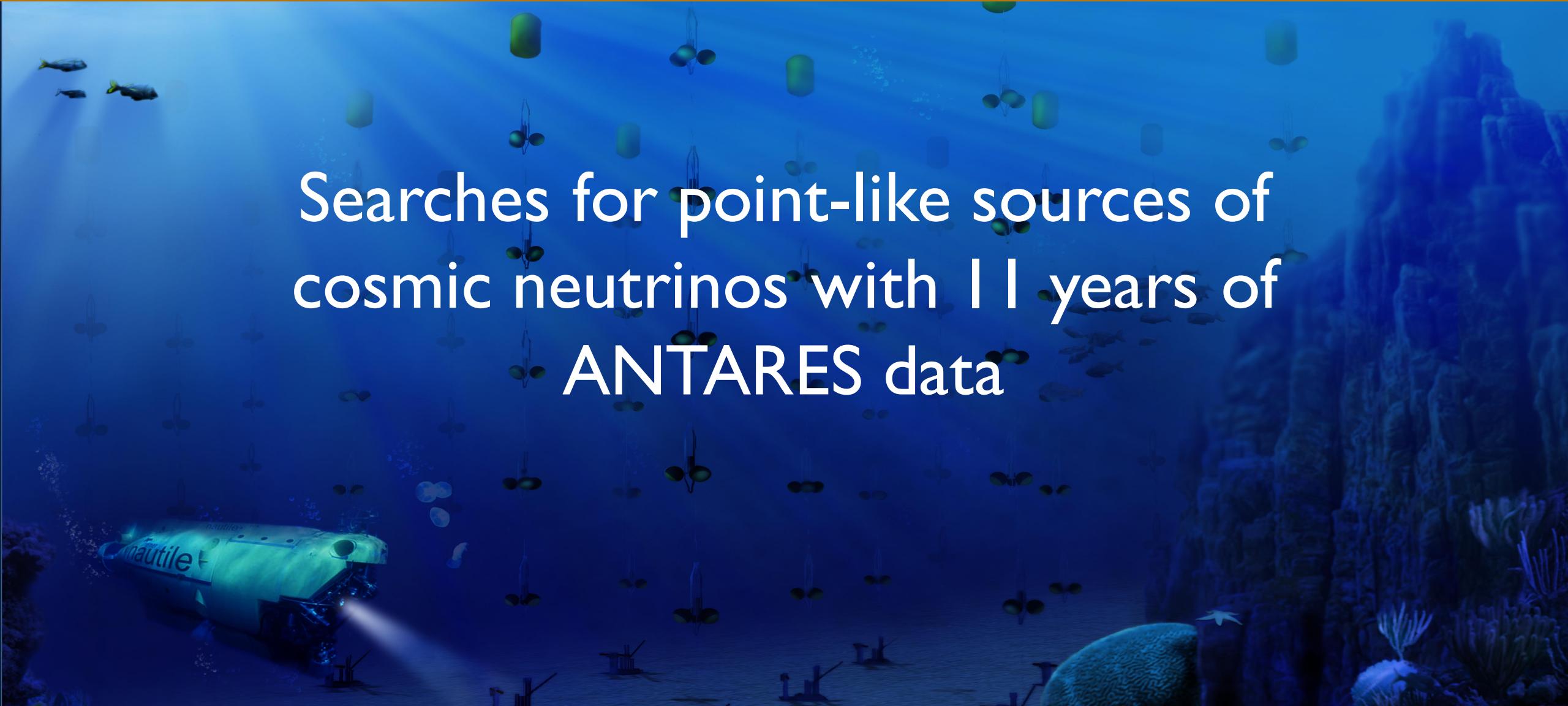
UNIVERSITAT  
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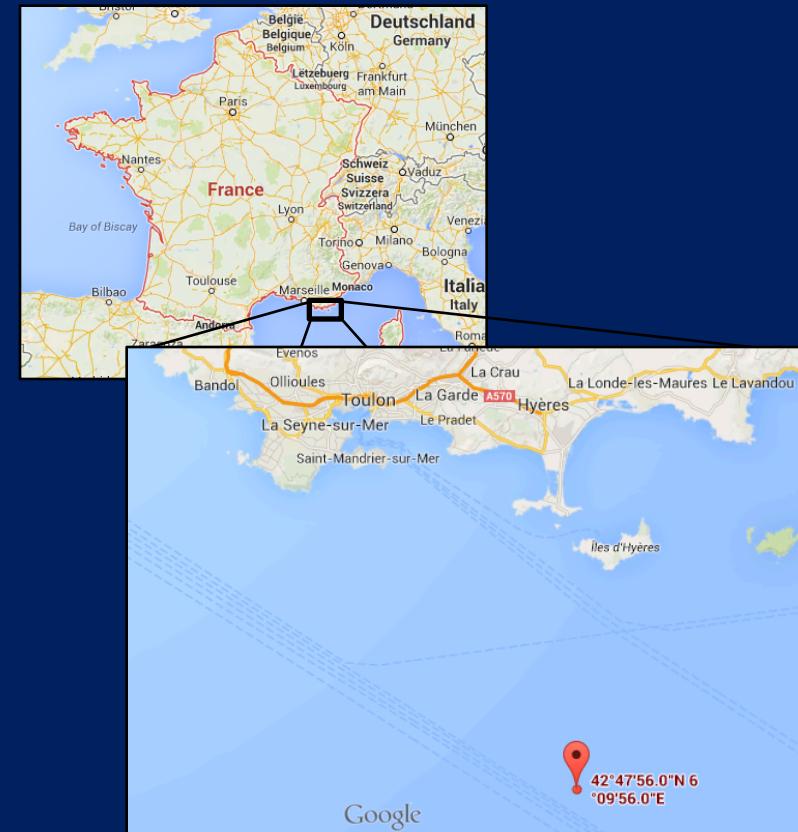
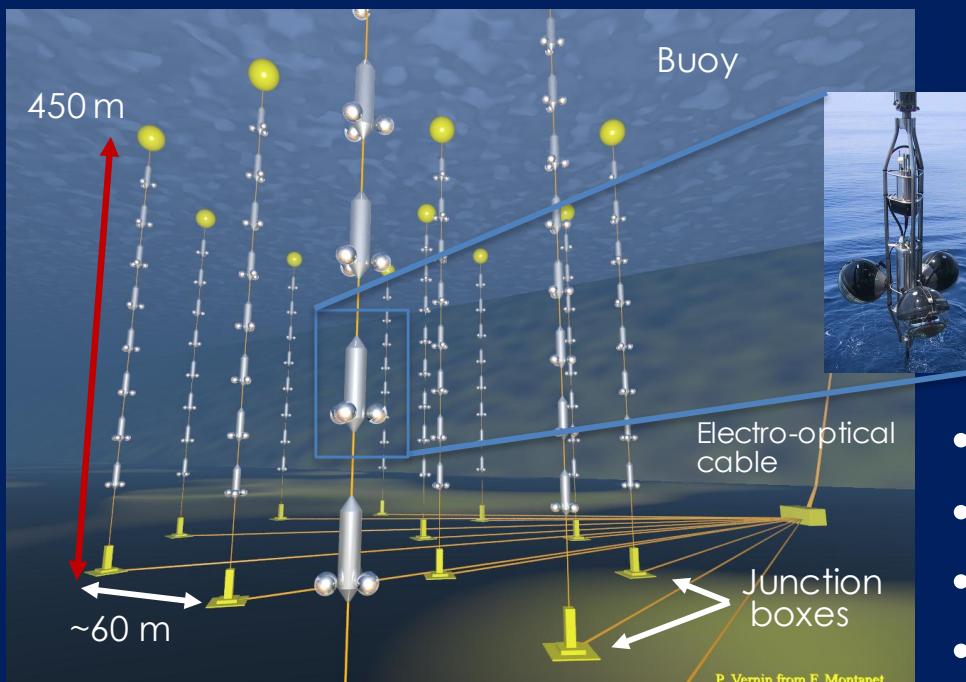


# Searches for point-like sources of cosmic neutrinos with 11 years of ANTARES data



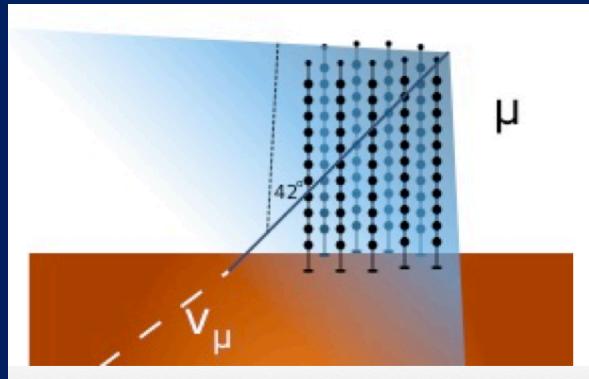
# The ANTARES neutrino telescope

- First detection line installed in early 2006
- Completed in 2008
- 2475 m depth in the Mediterranean Sea
- 40 km offshore from Toulon

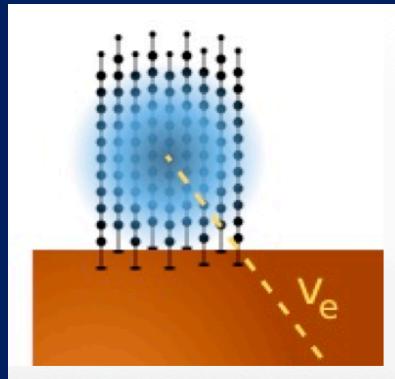


- Three-dimensional array of 885 PMTs
- 12 vertical lines, 25 storeys
- 3 PMTs per storey
- PMT facing 45° downwards

# Data sample

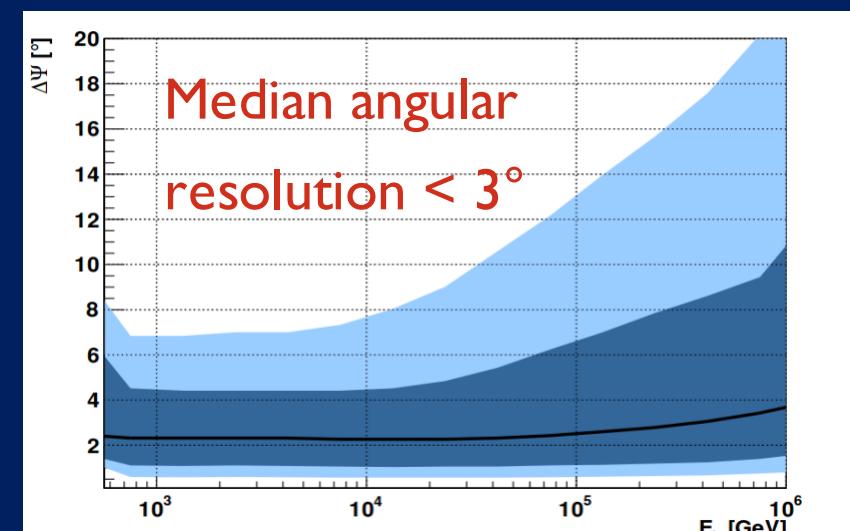
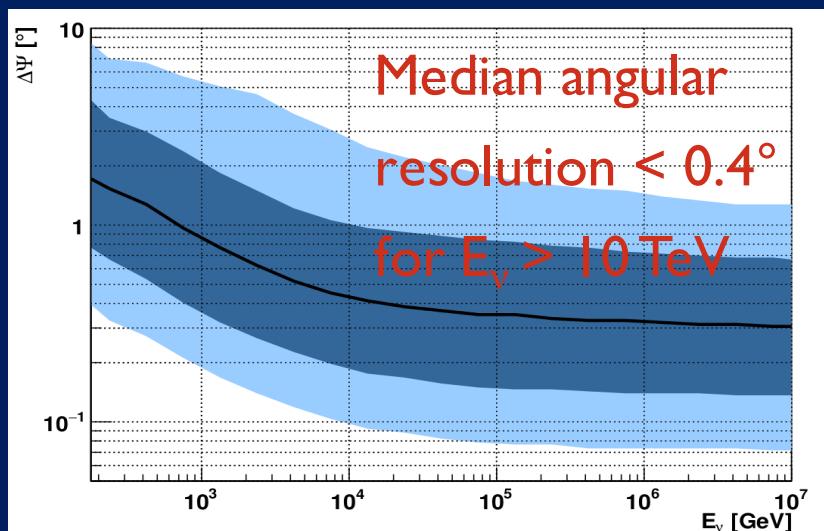


Track-like events:  
 $\nu_\mu$  ( $\nu_\tau$ ) neutrino  
CC interaction near the  
detector



Shower-like events:  
all neutrinos NC,  $\nu_e, \nu_\tau$  CC  
interaction inside or very close to  
the detector

Period: Jan 29, 2007 to Dec 31, 2017  
Livetime: 3136 days  
Events: 8754 tracks and 195 showers  
Same event selection as in the  
9 years ANTARES point-source analysis  
Phys. Rev. D96 (2017) 082001



# Method: Unbinned Likelihood

Time  
independent  
approach

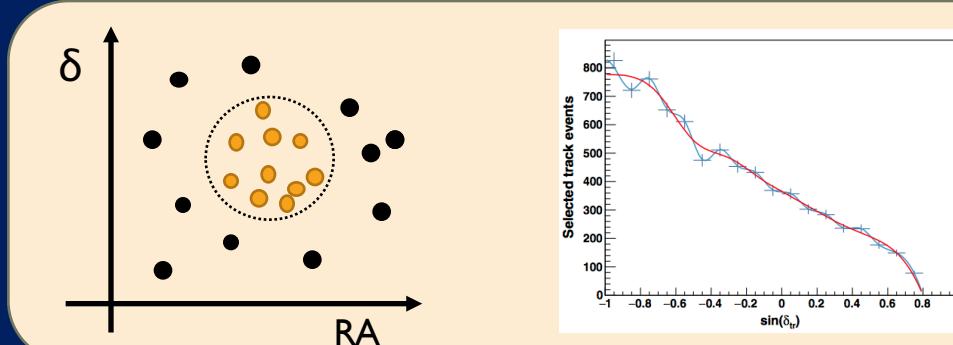
$$\log L_{s+b} = \sum_j \sum_{i \in j} \log \left[ \mu_{sig}^j S_i^j + N^j B_i^j \right] - \mu_{sig}$$

$J$ : sample (tracks or showers)  
 $i$ : event in sample  $J$

$N$ : total # of events  
 $\mu_{sig}$ : # of fitted signal events

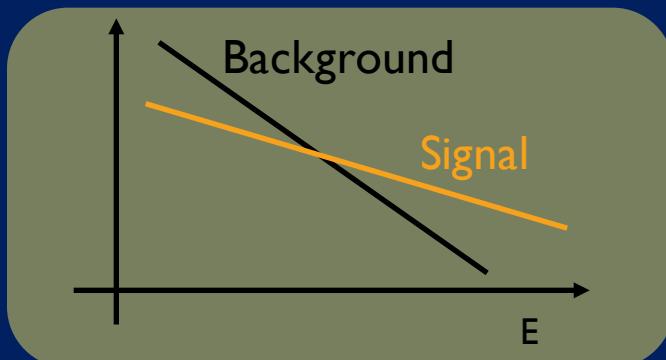
Signal PDFs:

$$S_i = S^{space} \cdot S^{energy}$$



Background PDFs:

$$B_i = B^{space} \cdot B^{energy}$$



Source spectrum  
hypothesis:  $E^{-2}$

# Method: Unbinned Likelihood

Time  
dependent  
approach

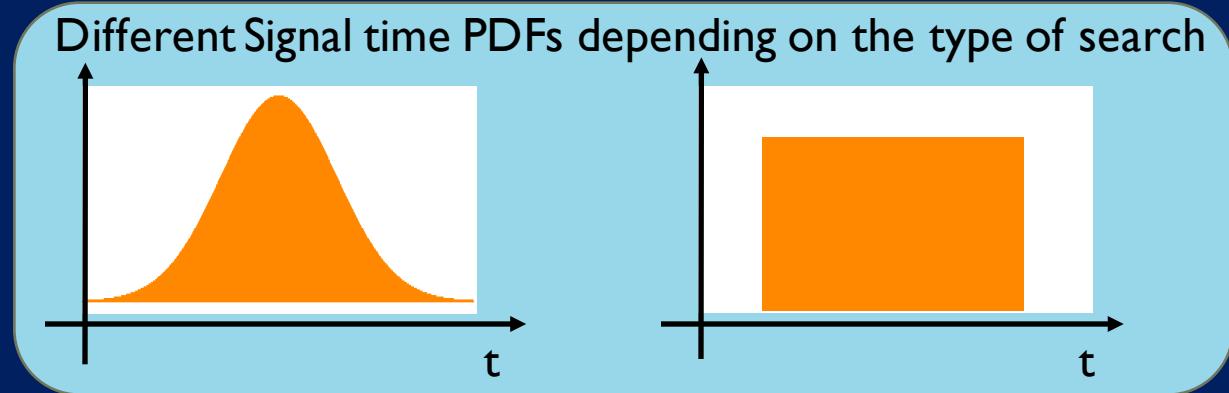
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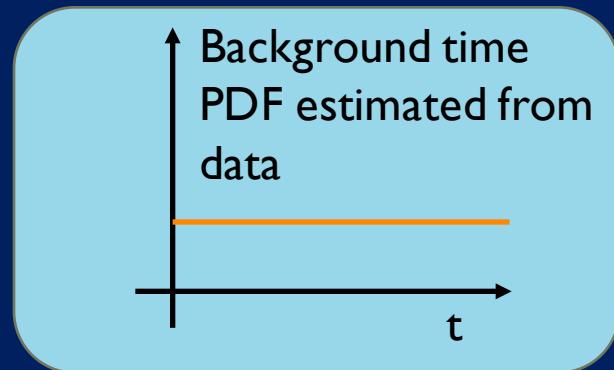
Signal PDFs:

$$S_i = S^{space} \cdot S^{energy} \cdot S^{time}$$



Background PDFs:

$$B_i = B^{space} \cdot B^{energy} \cdot B^{time}$$



# Searches

## All-sky search

- Scan of the ANTARES visible sky in squares of  $1^\circ \times 1^\circ$
- Most significant cluster: cluster with lowest p-value
- Source location free to vary in the likelihood maximization within the  $1^\circ \times 1^\circ$  boundaries

**Eta Carinae**  
 $(\alpha, \delta) = (161.27^\circ, -59.68^\circ)$

- Dedicated search for cosmic neutrinos from the stellar binary Eta Carinae assuming three different predicted neutrino spectra

## Candidate list searches

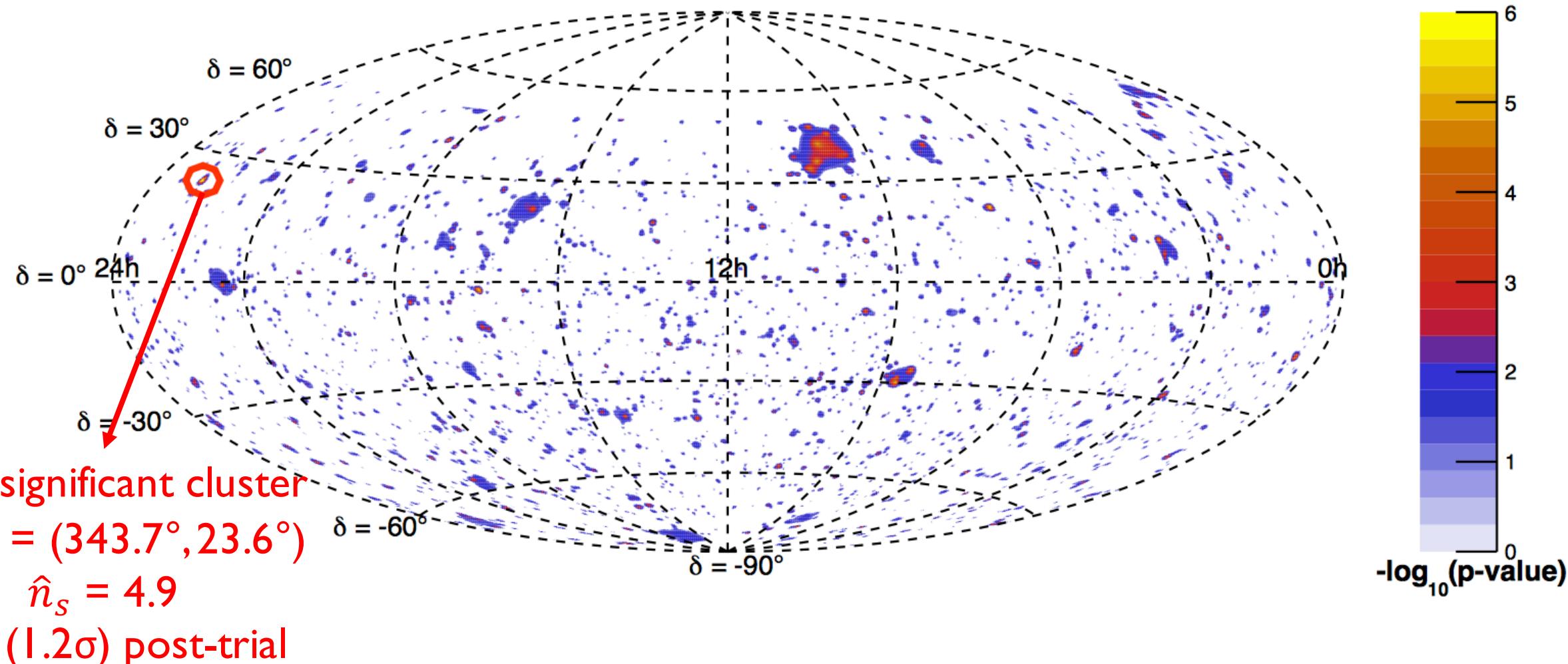
- Search for cosmic neutrinos from the location of preselected candidates
- Three candidate searches:
  - 112 known astrophysical objects (TeVCat sources)
  - 75 IceCube tracks (HESE, EHE, alerts)
  - 54 IceCube tracks (HESE, EHE): space and time correlation

**TXS 0506+056**  
 $(\alpha, \delta) = (77.36^\circ, 5.69^\circ)$

- Dedicated search for
  - steady emission
  - transient emissionfrom the direction of the blazar

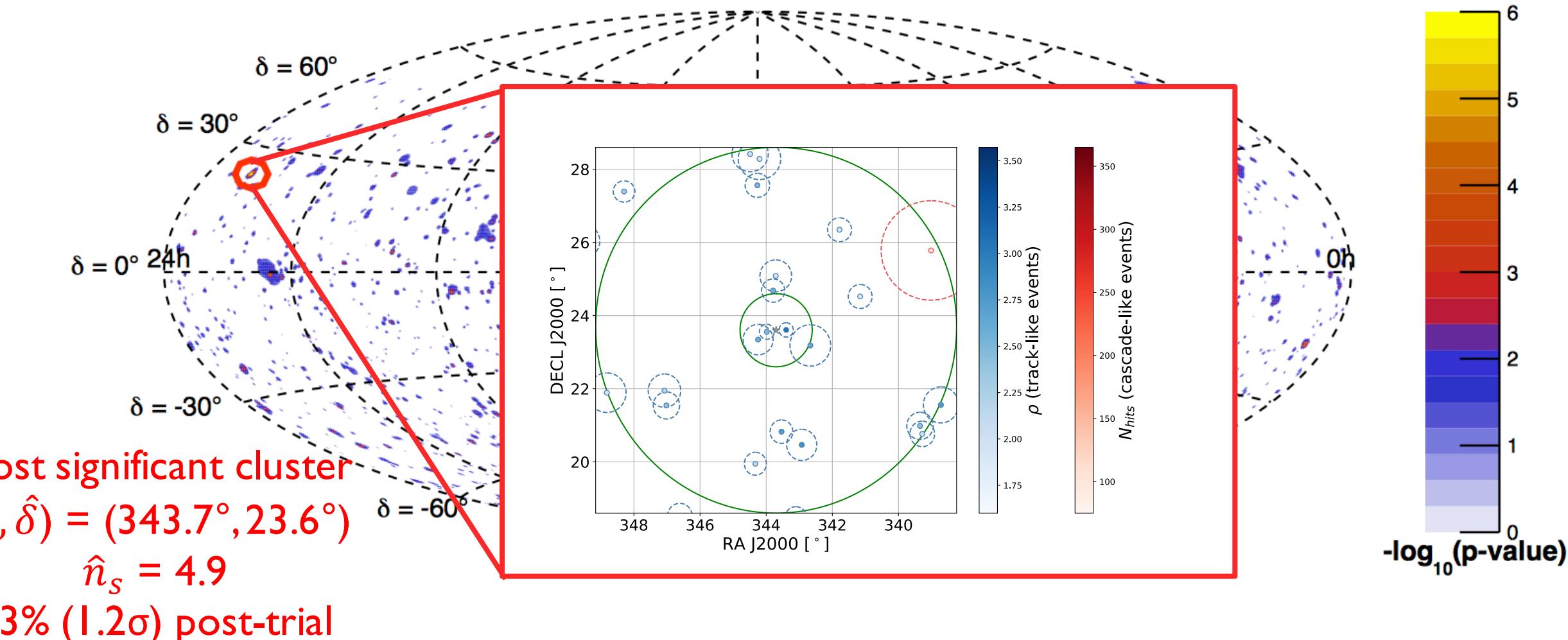
# Results: All-sky search

PRELIMINARY



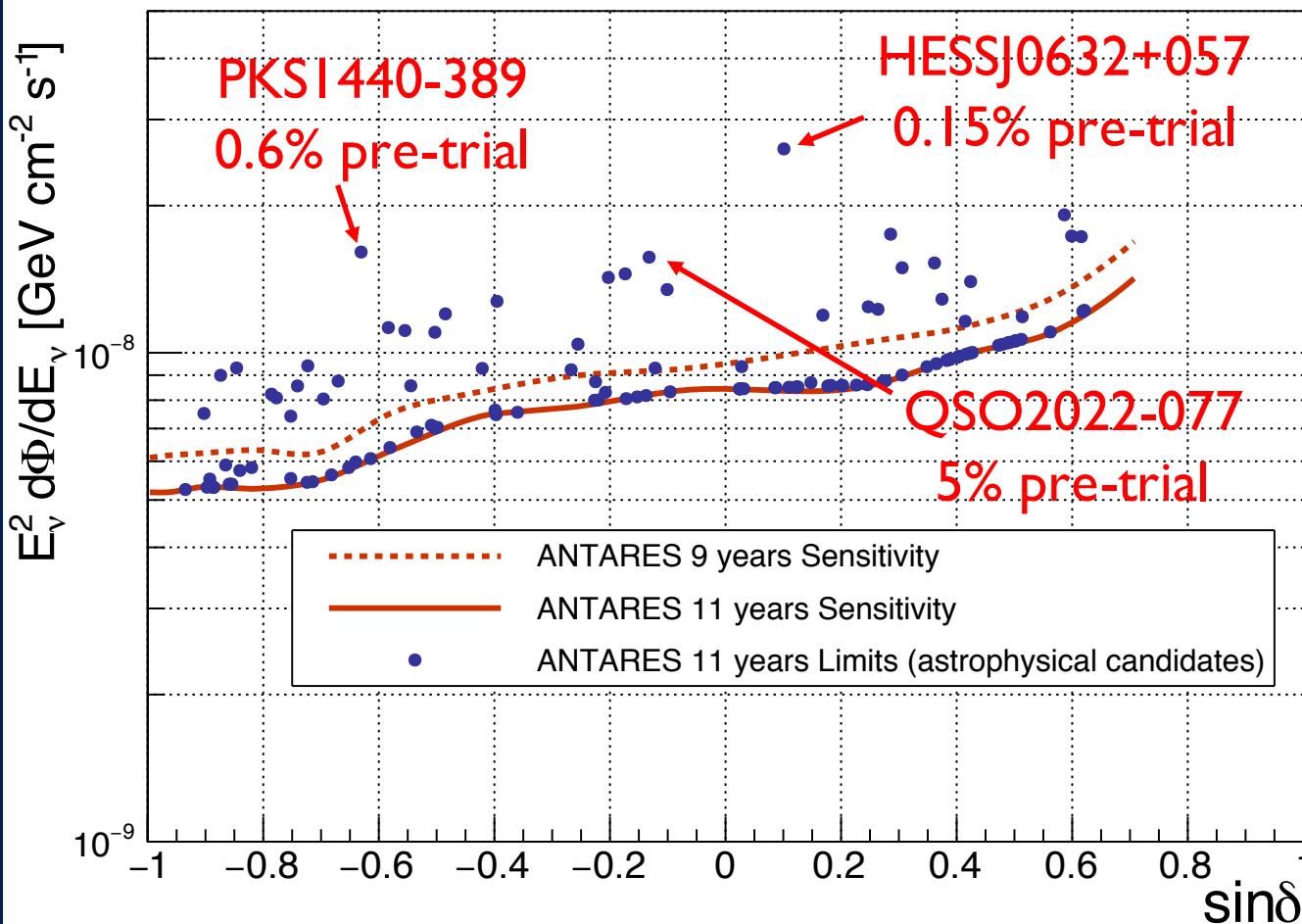
# Results: All-sky search

PRELIMINARY



# Results: Candidate list search

PRELIMINARY

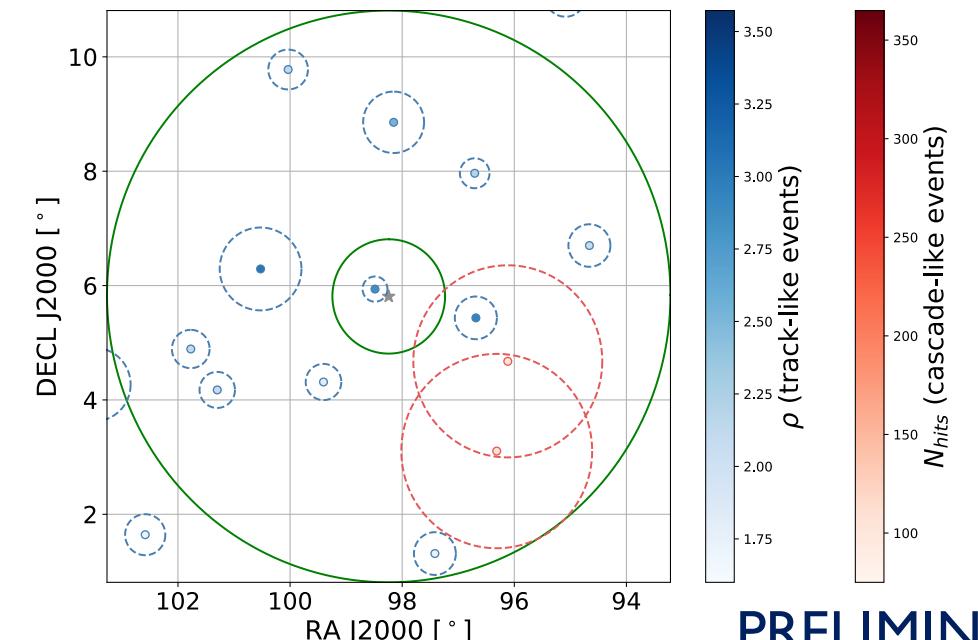


Most significant source  
HESSJ0632+057

$$\hat{n}_s = 2.7$$

16% ( $1.4\sigma$ ) post-trial

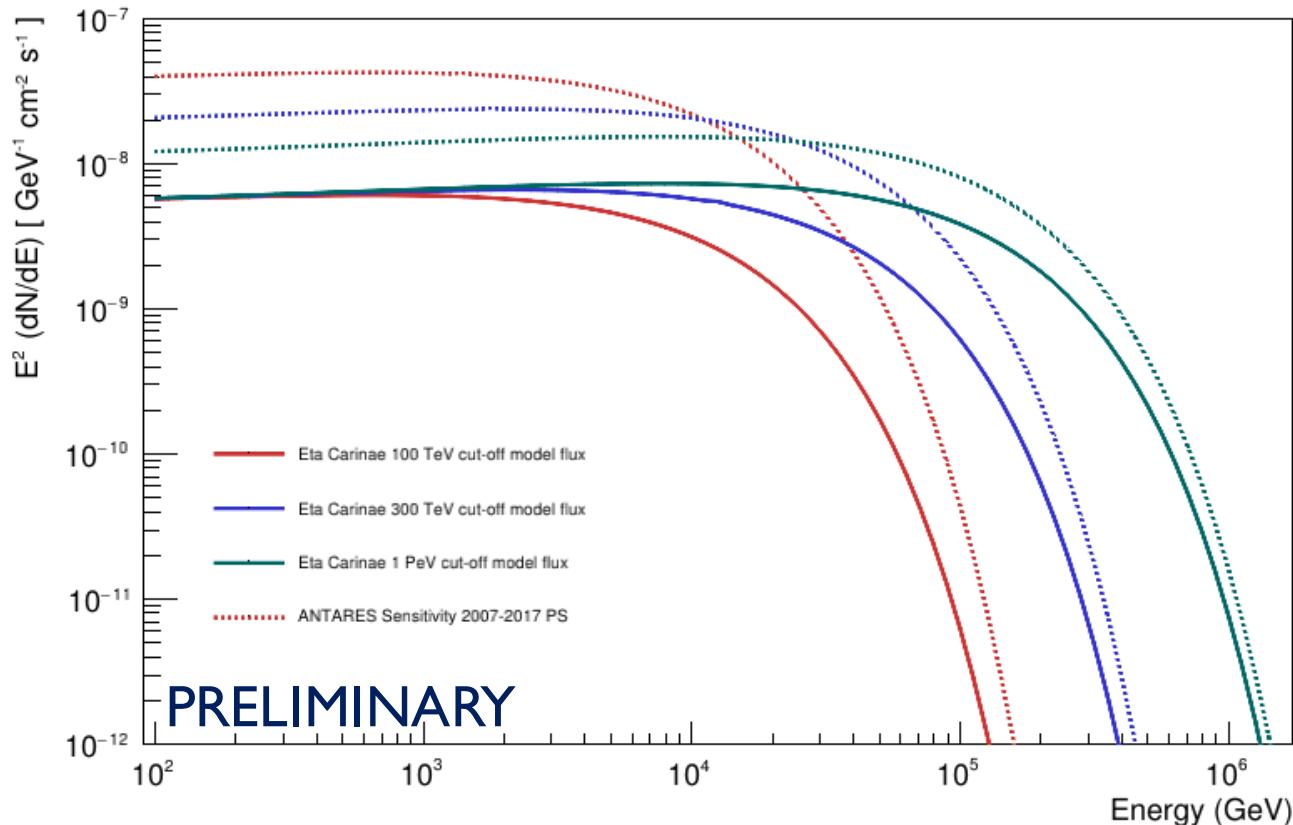
Events close to HESSJ0632+057 in 2007-2017 ANTARES data



PRELIMINARY

# Results: Eta Carinae

Spectrum models according to [Phys.Rev. D96 \(2017\) no.12, 123017](#)



- Dedicated search at the location of Eta Carinae
- Neutrino spectrum of the form:

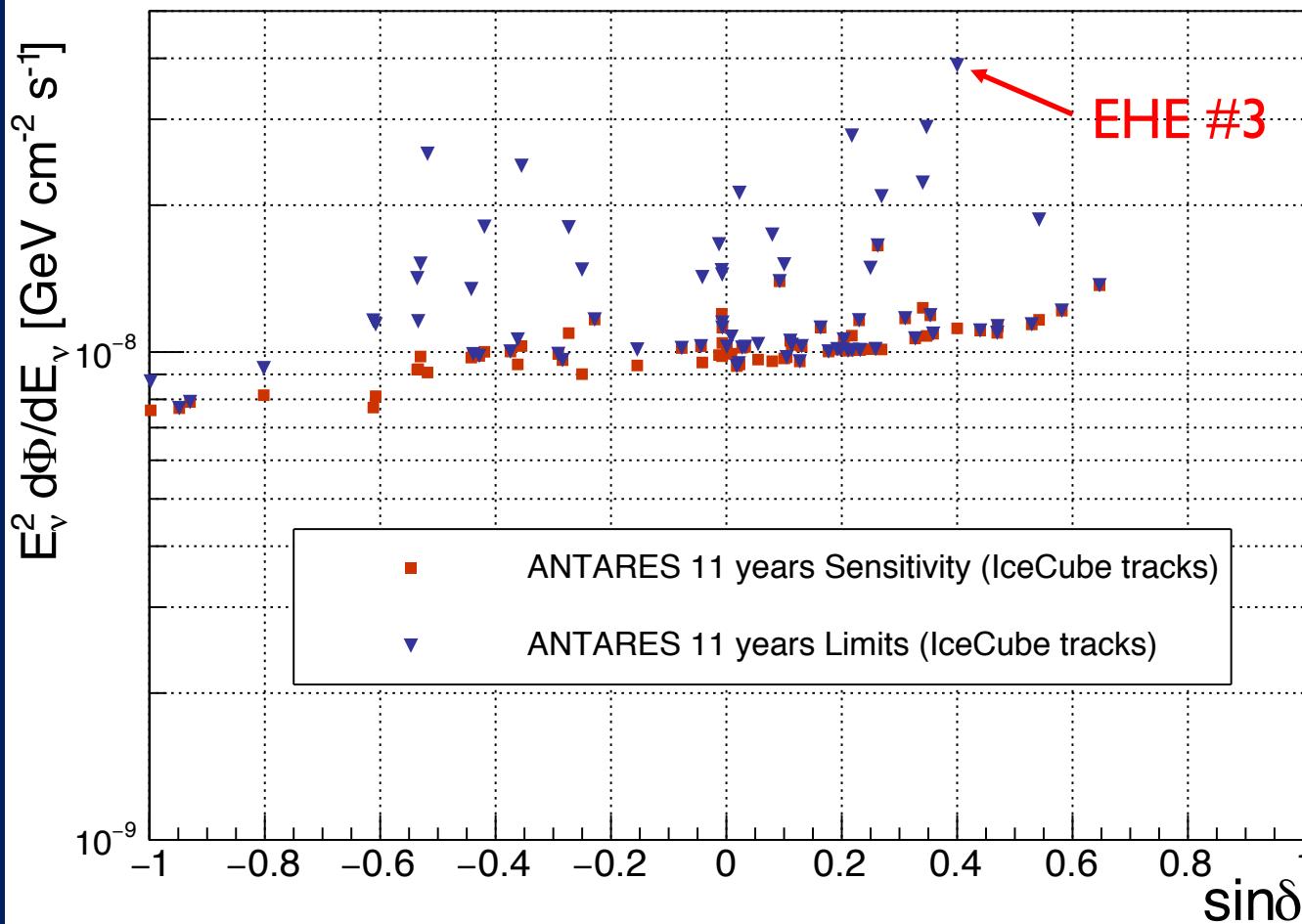
$$\frac{dN}{dE_\nu} = \Phi_0(E_\nu)^{-2} \exp\left(-\frac{E_\nu}{E_{cut}}\right)$$

- Three values of the energy cut-off
- 90% C.L. Upper limits on the neutrino flux:

$E_{cut}$	$\Phi^{90\% C.L.}/\Phi_0$
100 TeV	6.9
300 TeV	3.6
1 PeV	2.1

# Results: IceCube tracks search

PRELIMINARY



Most significant IceCube candidate

EHE ID3

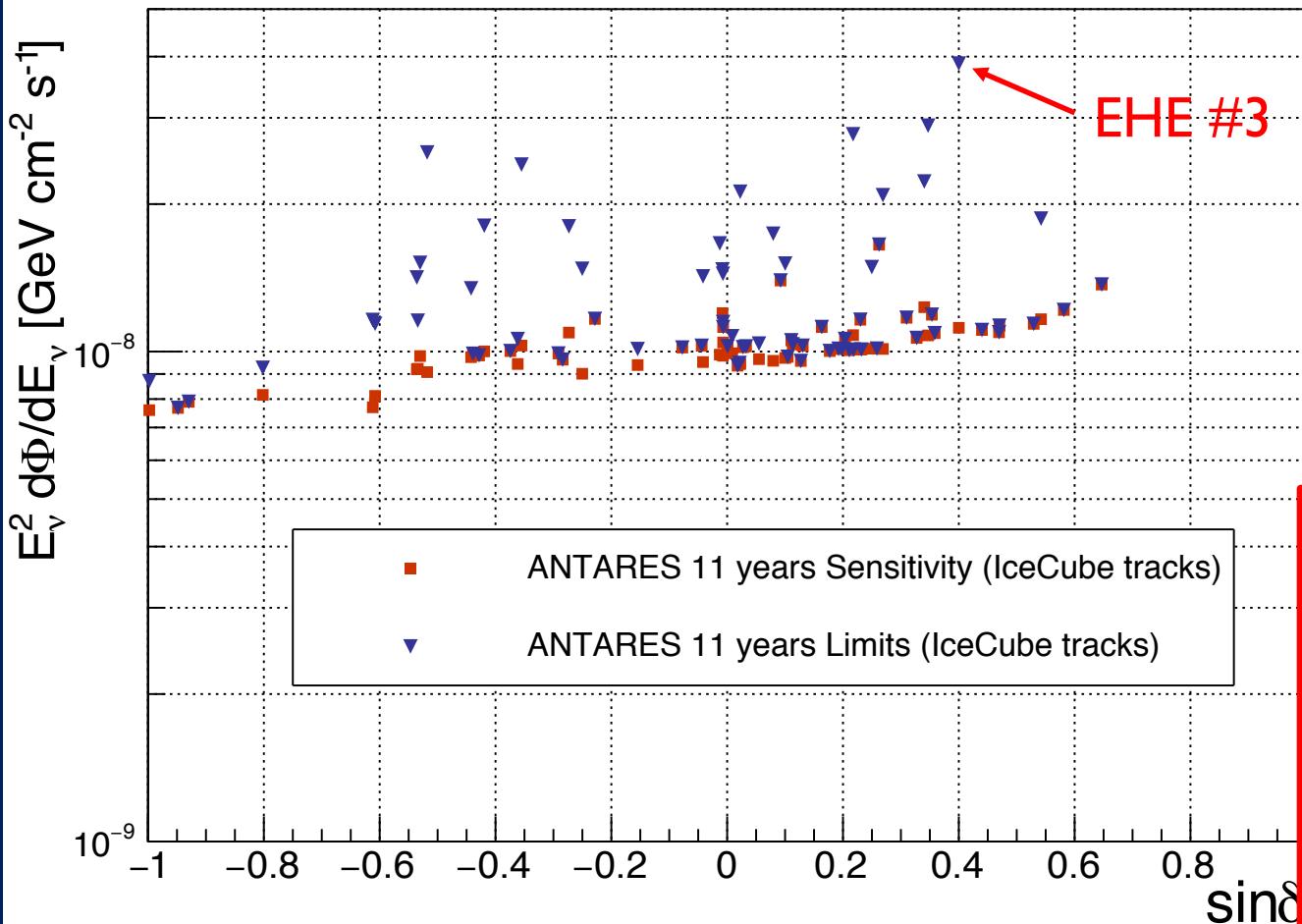
$$\hat{n}_s = 4.9$$

1.5% (2.4 $\sigma$ ) post-trial

Best-fit coordinates:  
 $(\hat{\alpha}, \hat{\delta}) = (343.7^\circ, 23.6^\circ)$

# Results: IceCube tracks search

PRELIMINARY



Most significant IceCube candidate

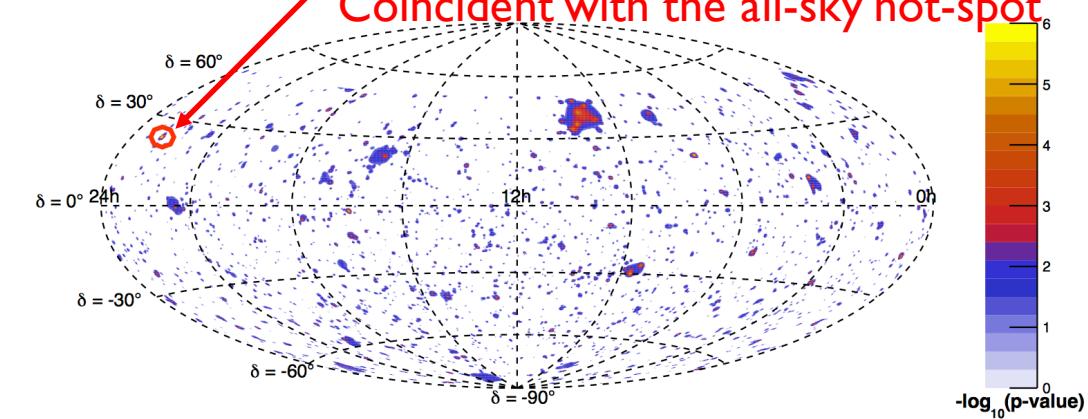
EHE ID3

$$\hat{n}_s = 4.9$$

1.5% (2.4 $\sigma$ ) post-trial

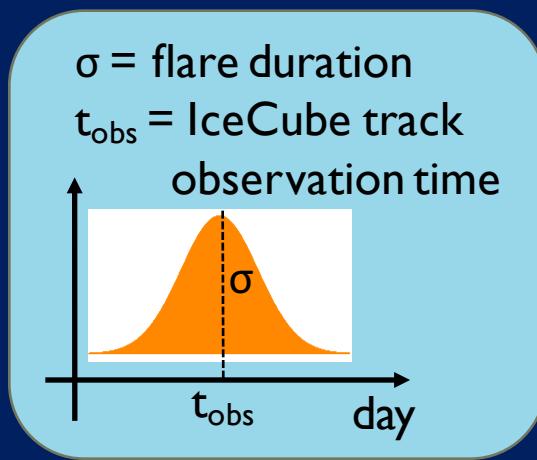
Best-fit coordinates:  
 $(\hat{\alpha}, \hat{\delta}) = (343.7^\circ, 23.6^\circ)$

Coincident with the all-sky hot-spot

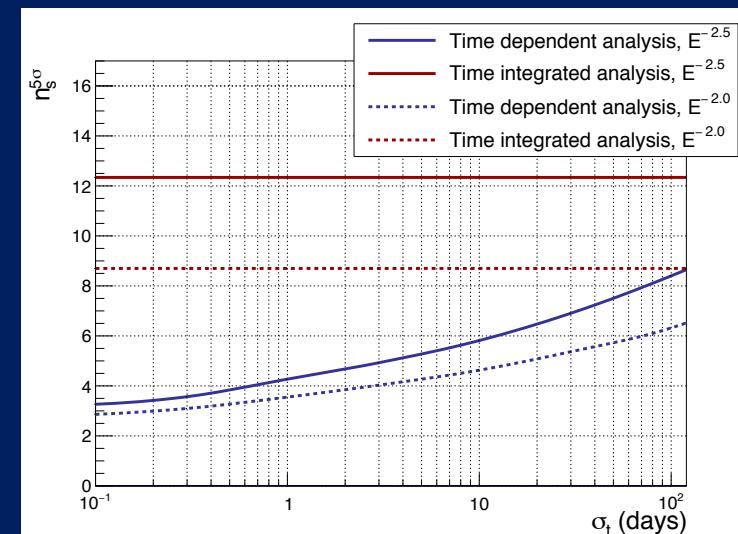


# Time/Space correlation with IceCube tracks

$$\text{Likelihood PDF} = P^{\text{space}} \cdot P^{\text{energy}} \cdot P^{\text{time}}$$



Flare duration additional free parameter  
 $0.1 \text{ days} \leq \sigma \leq 120 \text{ days}$



90% C.L. upper limits on the neutrino fluence:

Most significant candidate

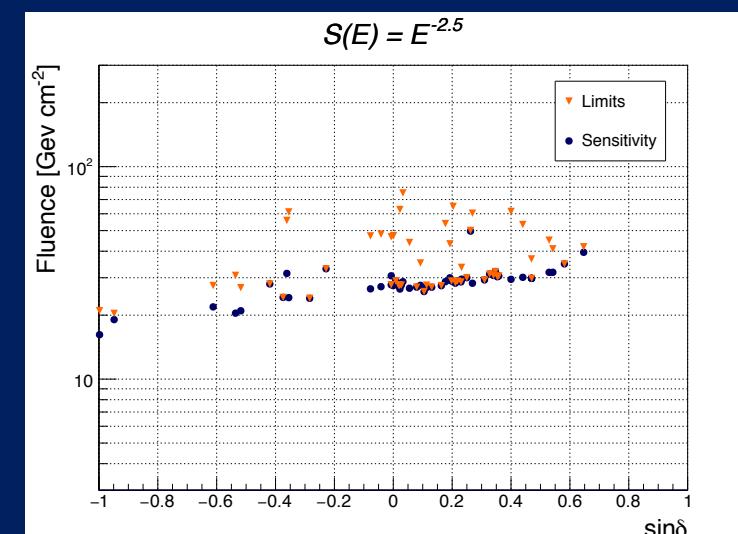
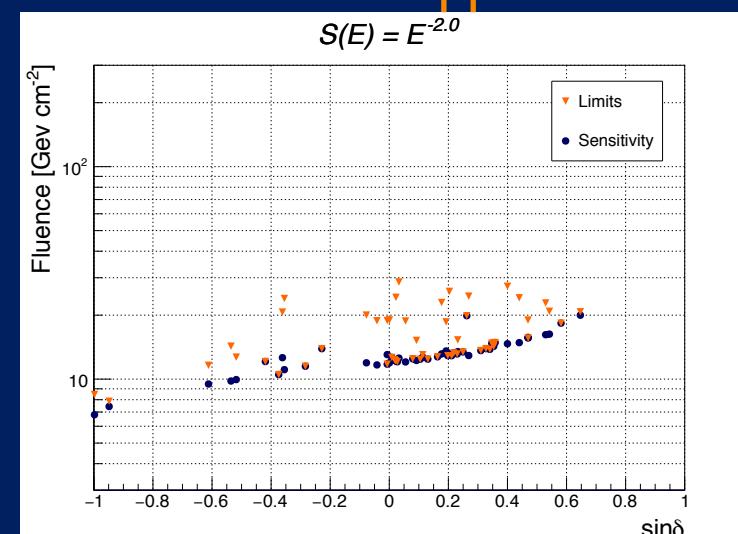
EHE ID15

$$\hat{n}_s = 2.2$$

$$\hat{\sigma}_s = 120 \text{ days}$$

3.7% pre-trial

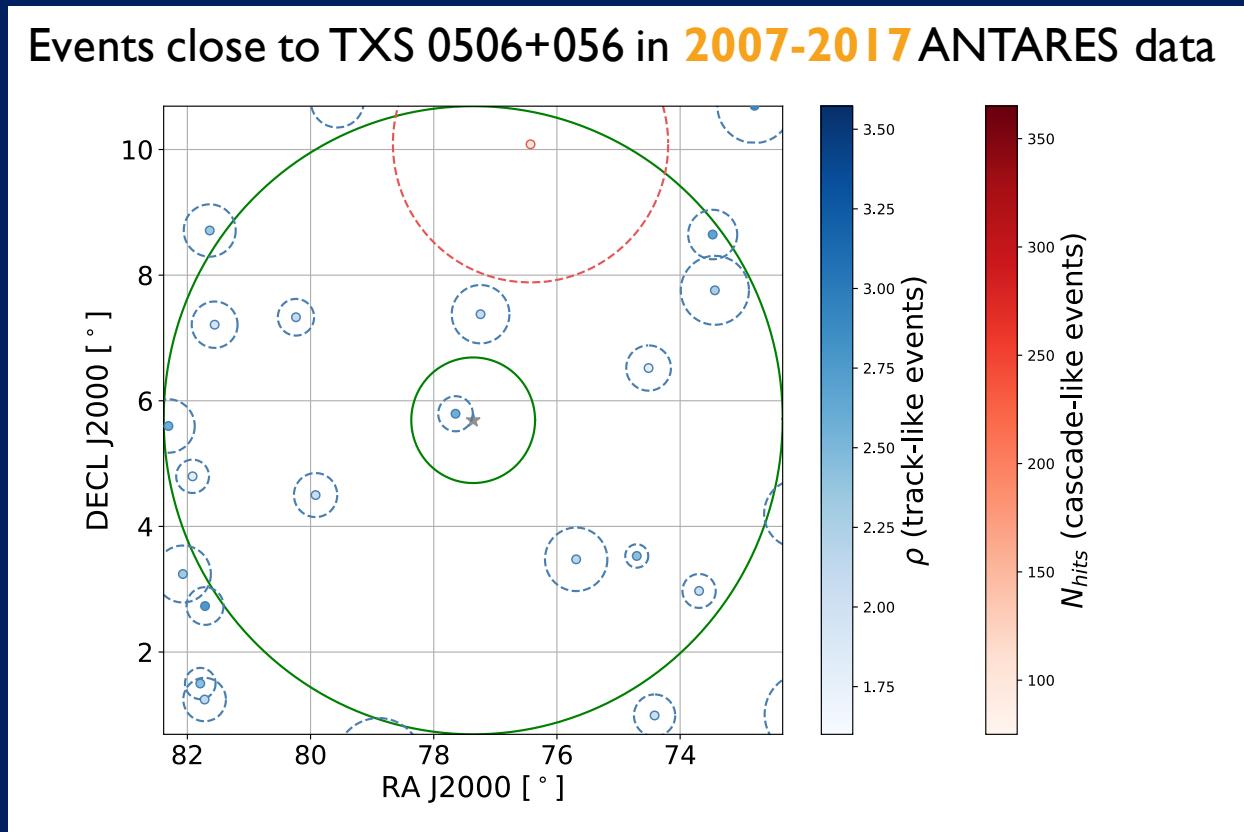
90% post-trial



# Search for steady emission from TXS 0506+056

- TXS 0506+056 added to the list of 106 neutrino source candidates analysed in the latest ANTARES point-source search (Phys. Rev. D **96**, 082001 (2017))
- TXS 0506+056 third most significant source
- Best fitted # of signal events  $\mu_{sig} = 1.03$
- 3.4% pre-trial p-value
- 87% post-trial p-value
- 90% C.L. flux upper limits on the flux:

Spectrum	$\Phi_{100\text{TeV}}^{90\%} [10^{-18} \text{ GeV}^{-1} \text{cm}^{-2}\text{s}^{-1}]$
$E^{-2.0}$	1.6
$E^{-2.3}$	1.4
$E^{-2.5}$	1.0

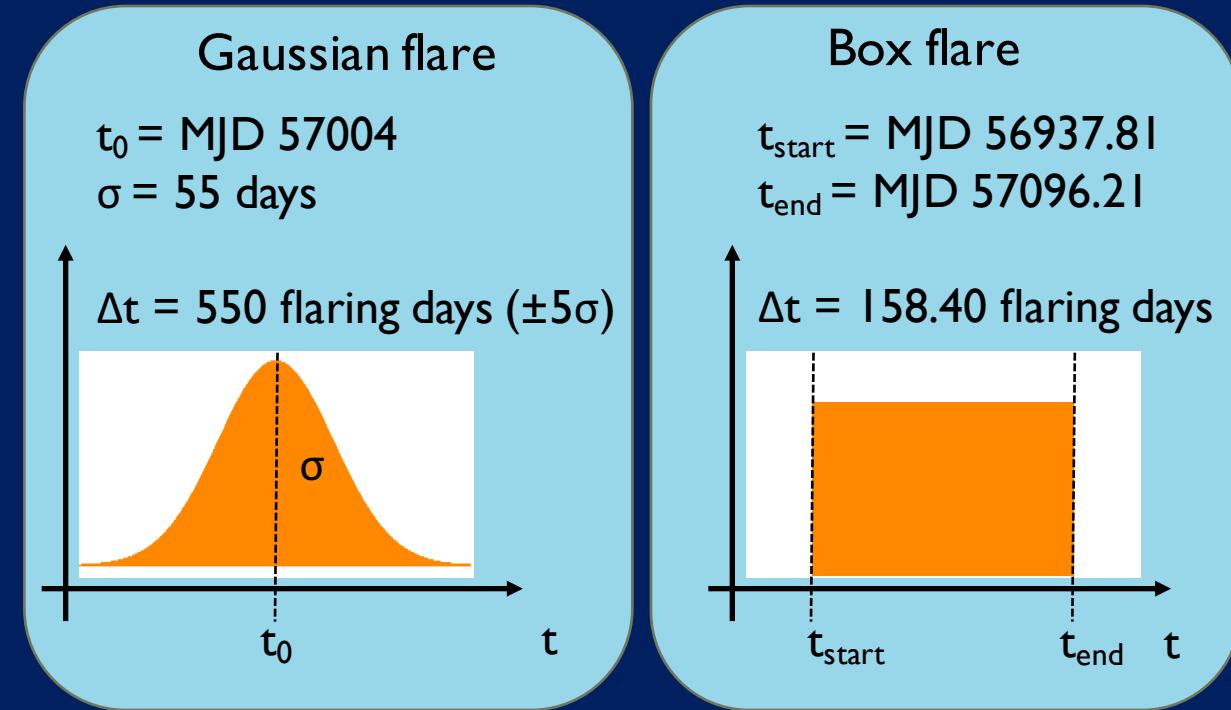


- Closest event at  $0.3^\circ$  from the source position
- Recorded on December 12, 2013
- Only 9% of  $\nu$  events have higher estimated energy

# Search for transient emission from TXS 0506+056

$$PDF = P_{space} \cdot P_{energy} \cdot P_{time}$$

- Bursting period defined by the two profiles provided by the IceCube Collaboration:
  - Gaussian flare (centered on December 13, 2014)
  - Box flare (centered on December 26, 2014)
- No signal found during either of the considered flares
- 90% C.L. upper limits on the neutrino flux derived for the Gaussian-shaped period:



Spectrum	$\Phi_{100\text{TeV}}^{90\%} [10^{-18} \text{ GeV}^{-1} \text{cm}^{-2}\text{s}^{-1}]$	5%-95% energy range
$E^{-2.0}$	4.6	2.0 TeV – 3.2 PeV
$E^{-2.1}$	4.4	1.3 TeV – 1.6 PeV
$E^{-2.2}$	4.2	1.0 TeV – 1.0 PeV

For the box-shaped period the flux normalization factors are a factor 3.3 higher

# Summary

- Various searches for steady and transient point-like neutrino sources presented
- No significant point-like emission found, upper limits set on neutrino flux and fluence
- All-sky search: largest excess with  $1.2\sigma$  post-trial at  $(\alpha, \delta) = (343.7^\circ, 23.6^\circ)$
- Candidate list over astrophysical objects: largest excess with  $1.4\sigma$  post-trial found for HESS J0632+057
- Candidate list over IceCube tracks: largest excess with  $2.4\sigma$  post-trial found for EHE ID3
- Candidate list over IceCube tracks (time-dependent): largest excess with  $90\%$  post-trial found for EHE ID15
- Upper limits on the neutrino flux and fluence from TXS 0506+056
- Upper limits on the neutrino flux for three proposed neutrino emission models for Eta Carinae



**NEW PUBLIC DATA**

## **Data set for the 2007-2017 ANTARES search for cosmic neutrino point sources**

Available here:

<http://antares.in2p3.fr/publicdata2017.html>