

# Trinity

## An Air-Shower Imaging System for the Detection of Ultrahigh Energy Neutrinos

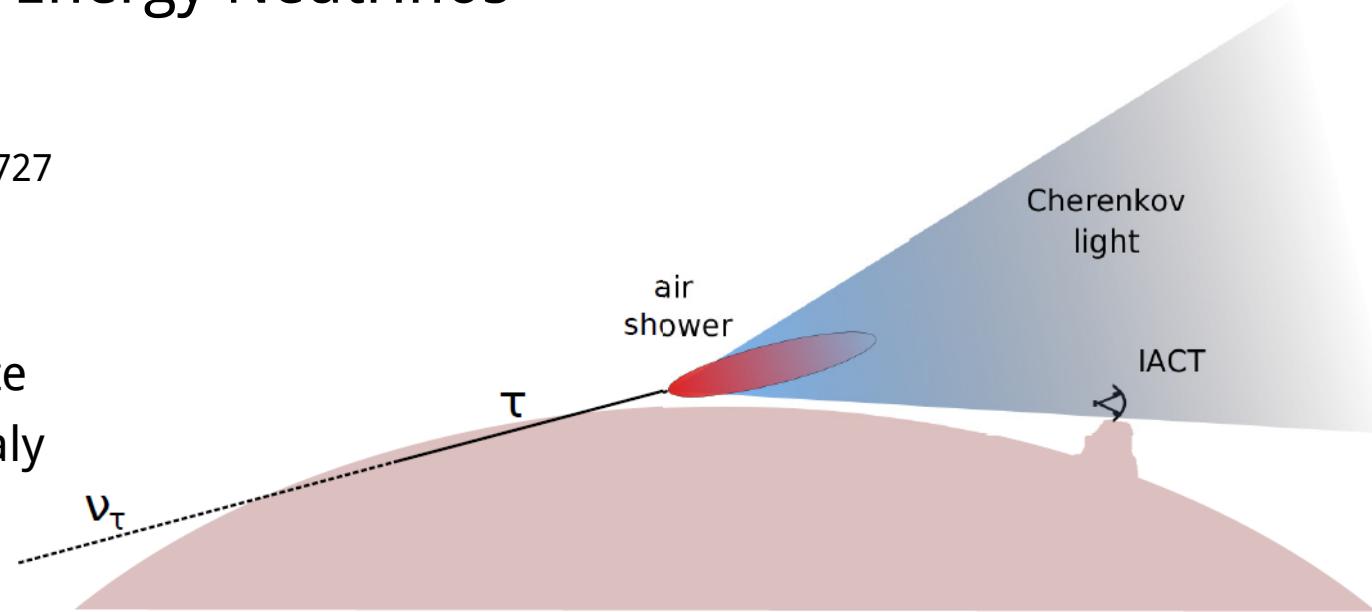
Phys. Rev. D 99, 083012 (2019)  
and  
Astro2020 white paper arXiv:1907.08727

Anthony Brown, Durham, UK

Abraham D. Falcone, Penn state

Mosè Mariotti, Uni. Padova, Italy

Ignacio Taboada, Georgia Tech

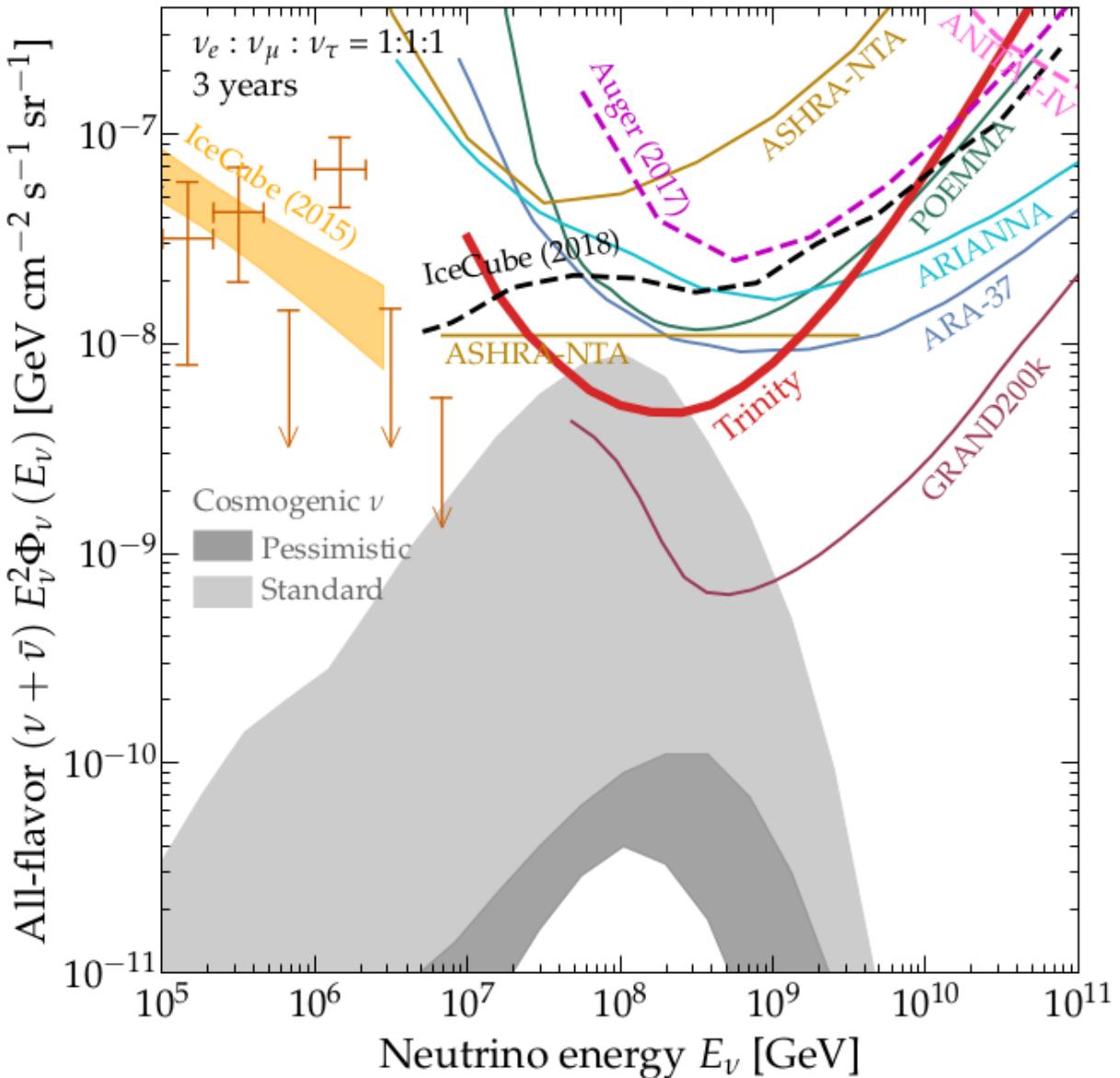


Nepomuk Otte

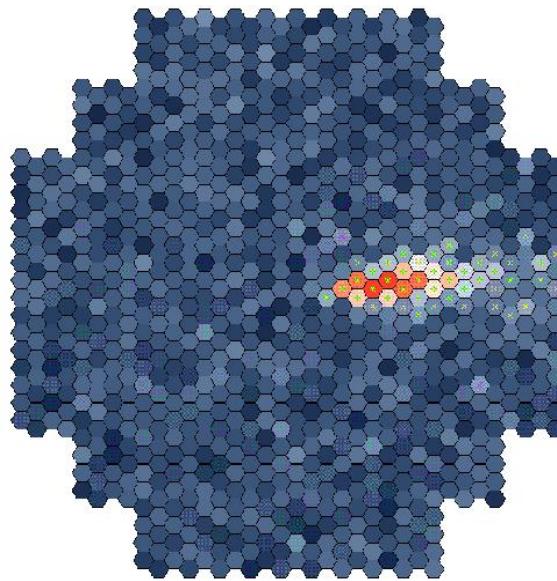
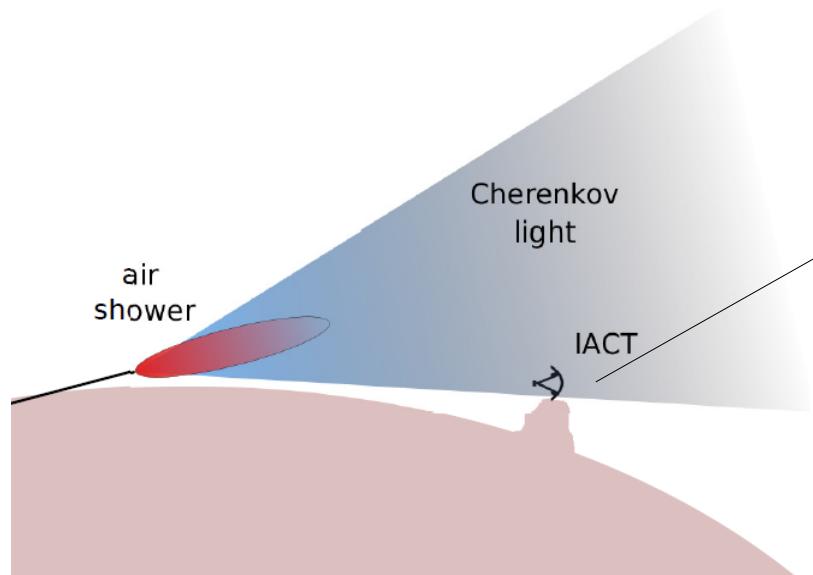
School of Physics  
&  
Center for Relativistic Astrophysics

**Georgia**  
**Tech**  **Physics**  
College of Sciences

Three year sensitivity



# Air-Shower Imaging



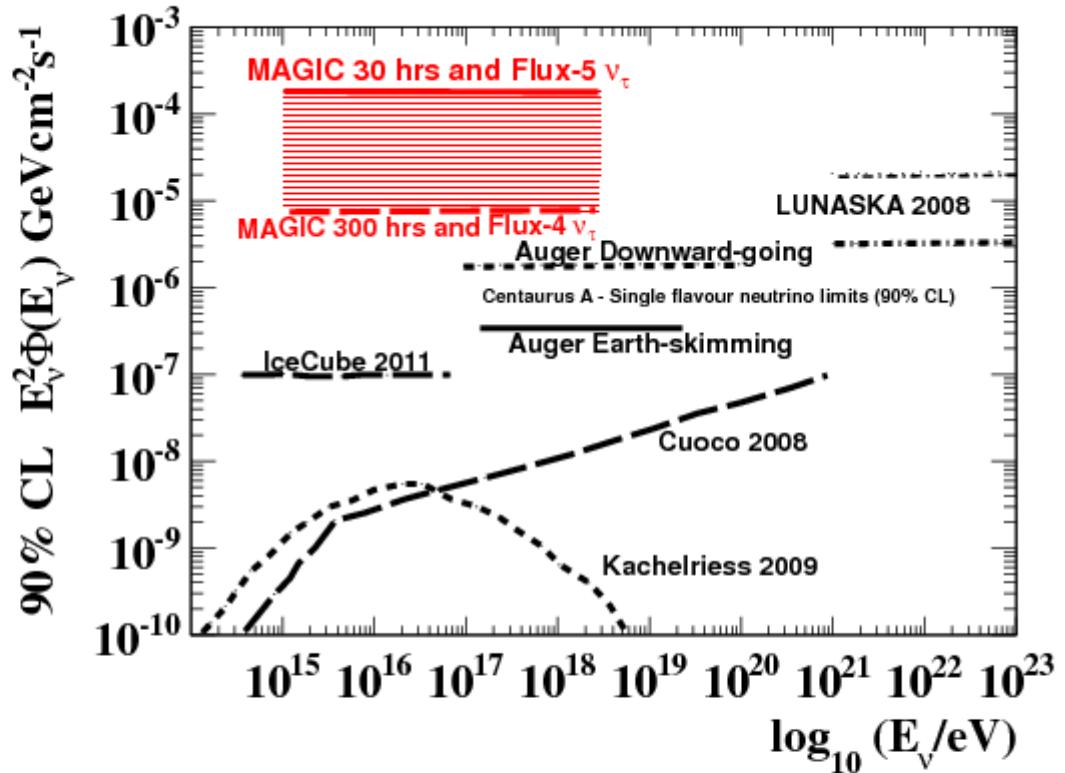
## Proven Technique

- Angular resolution  $\sim 0.1^\circ$
- Energy resolution  $\sim 10\%$
- Excellent background suppression

# UHE Neutrino Searches with Air-Shower Imaging Telescopes



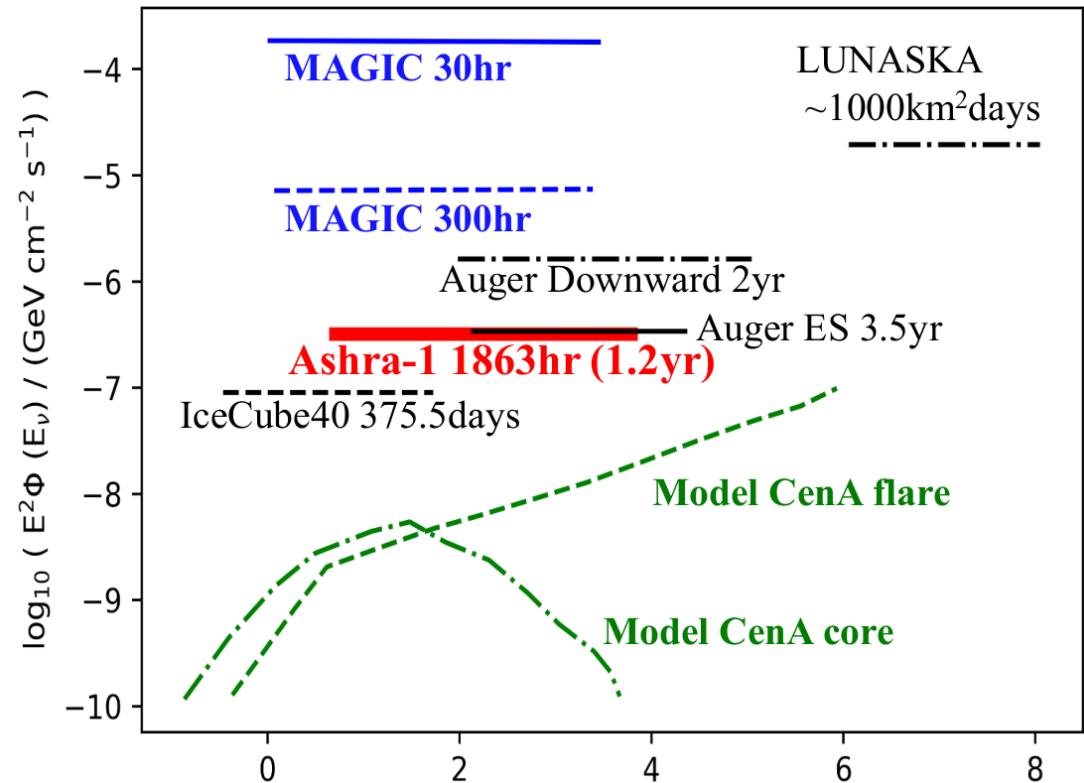
MAGIC (2018), Astropart.Phys.102,77-88.



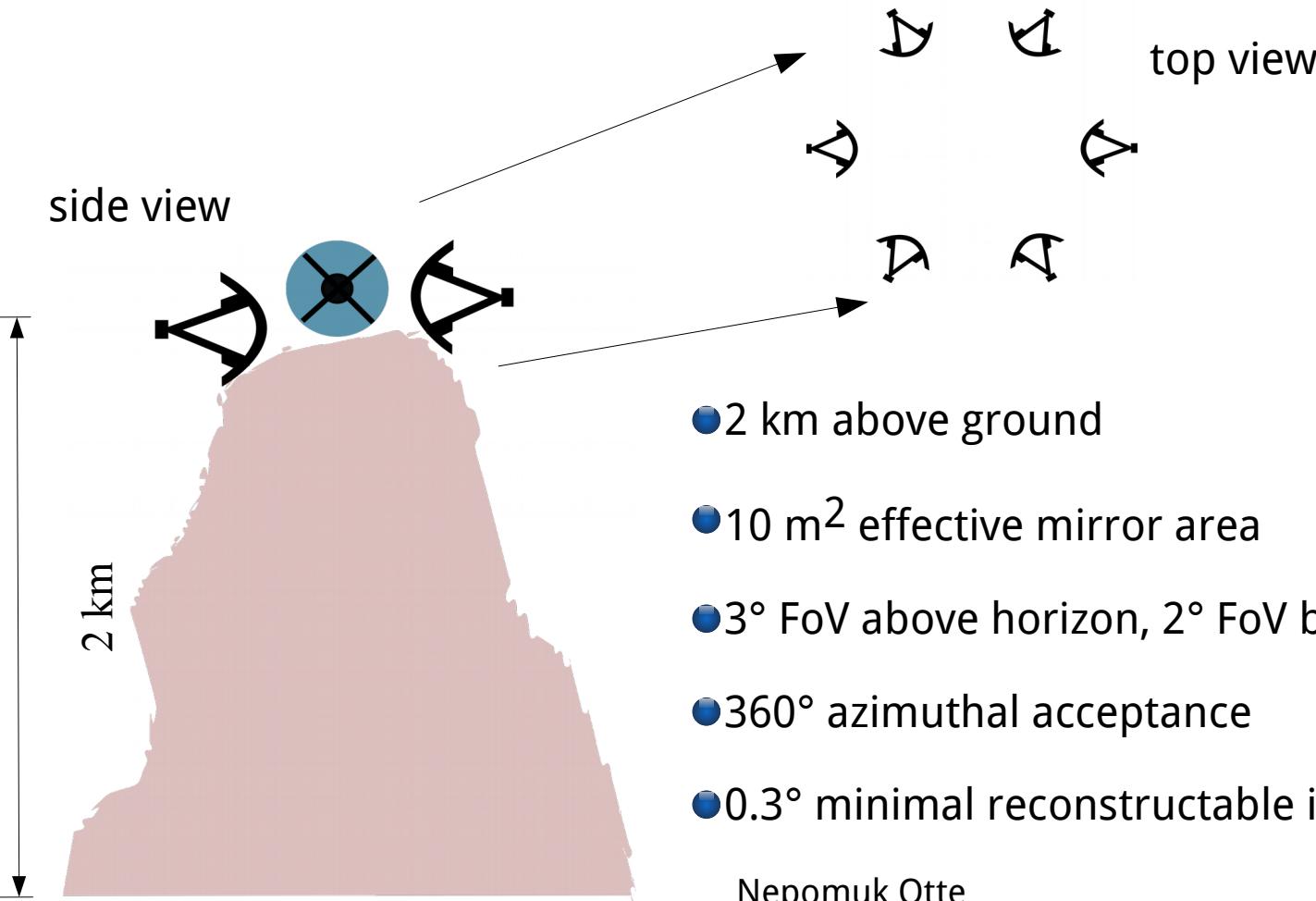
# UHE Neutrino Searches with Air-Shower Imaging Telescopes



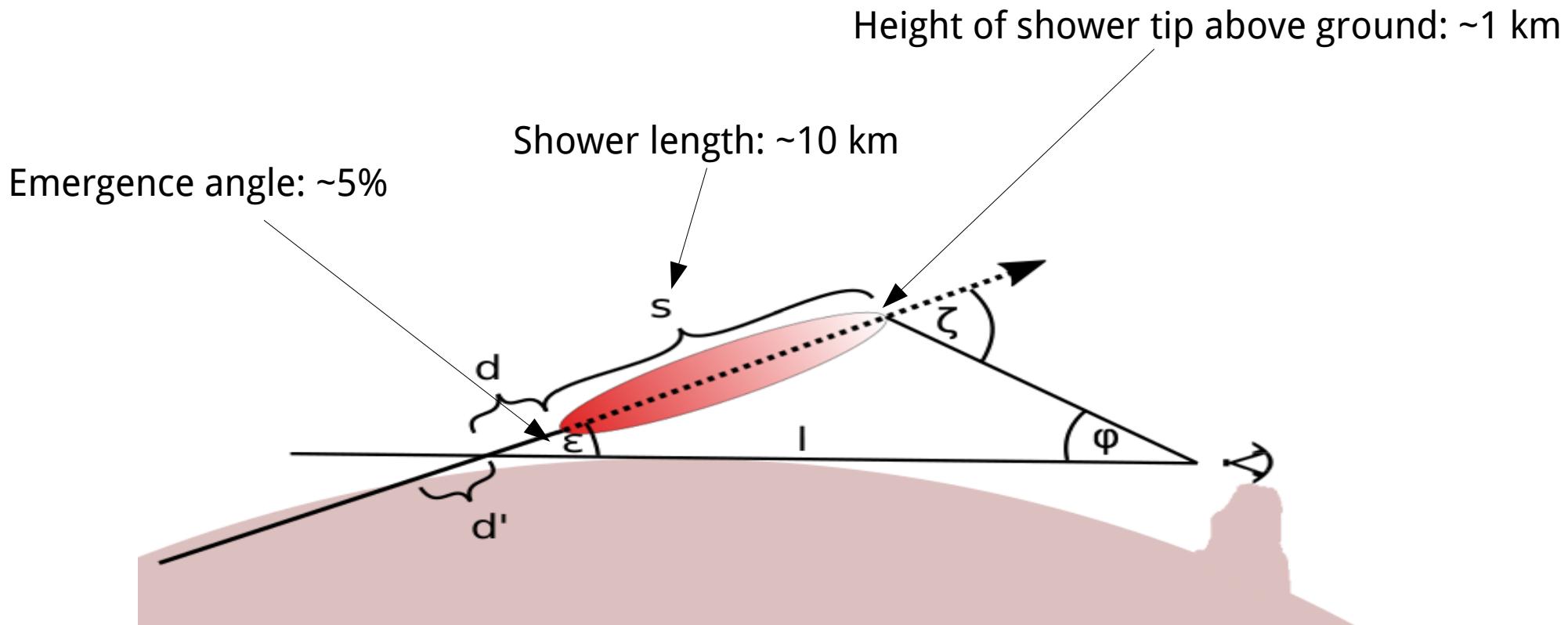
Ashra-1 PoS(ICRC2019)970



# Trinity: Baseline Configuration

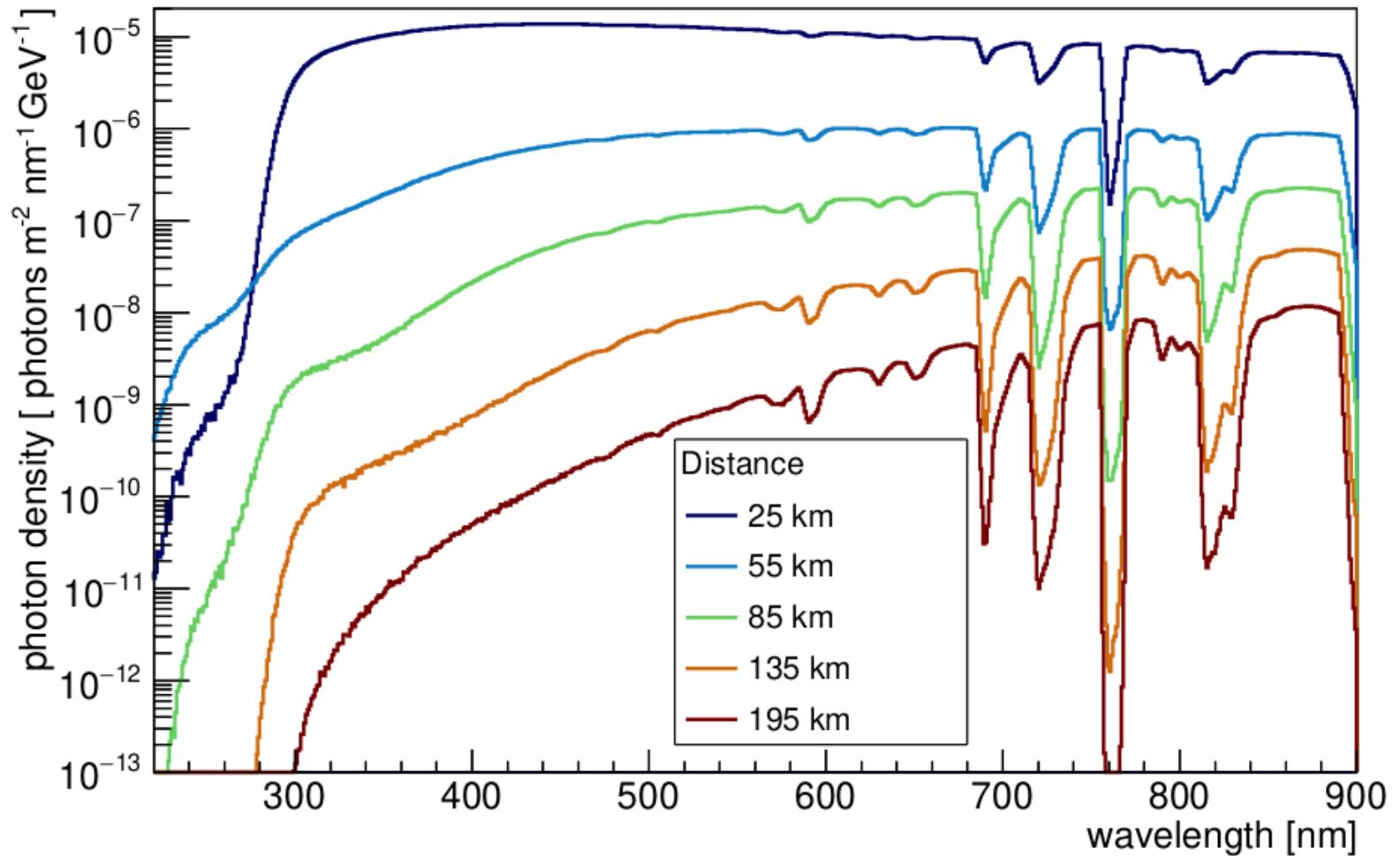


# UHE Tau initiated Air-Shower Fun Facts



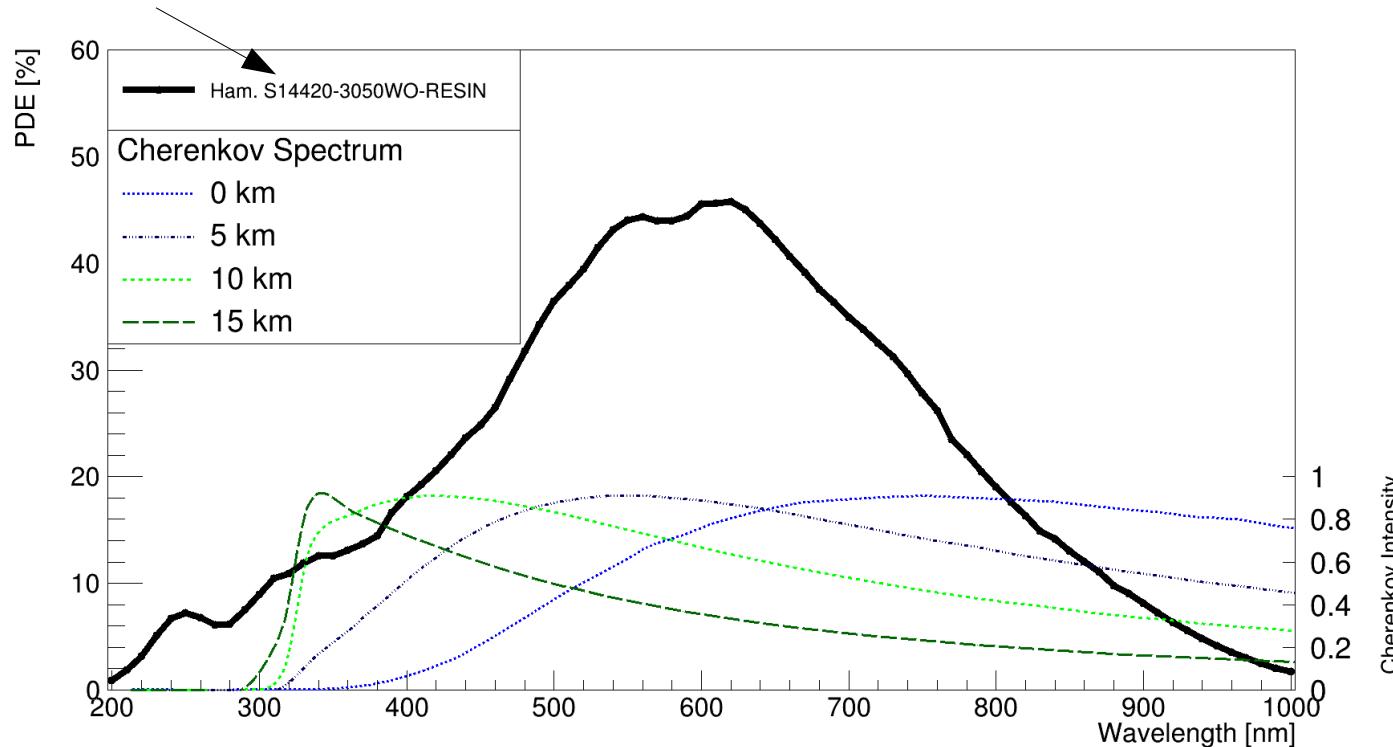


# Air Showers for Lovers



# Red Sensitive Silicon Photomultipliers

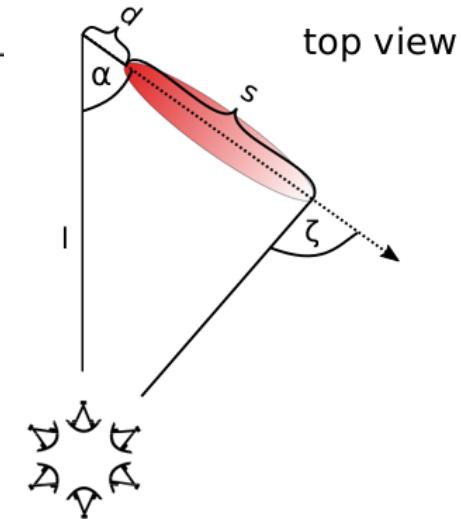
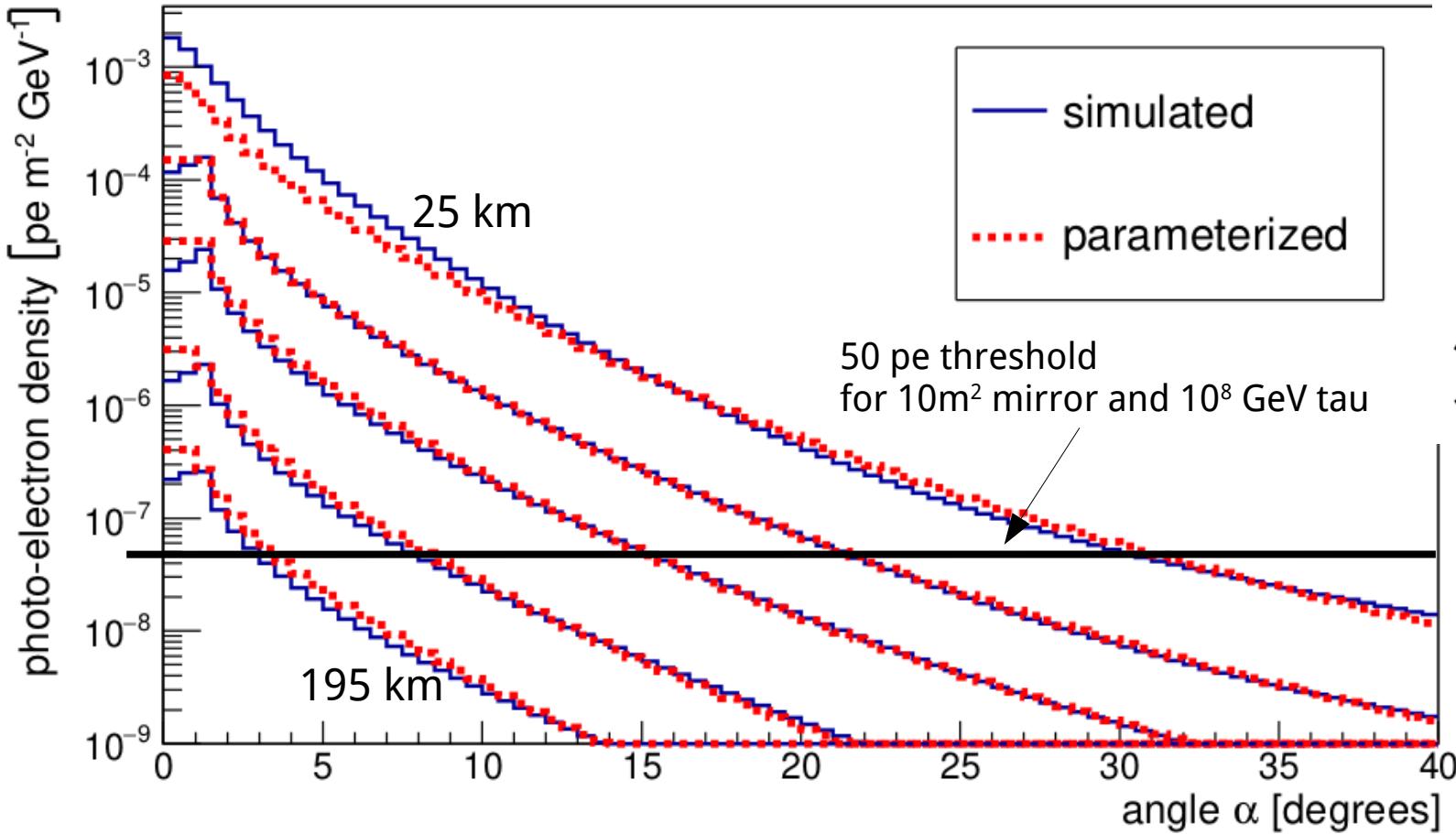
Hamamatsu S14420



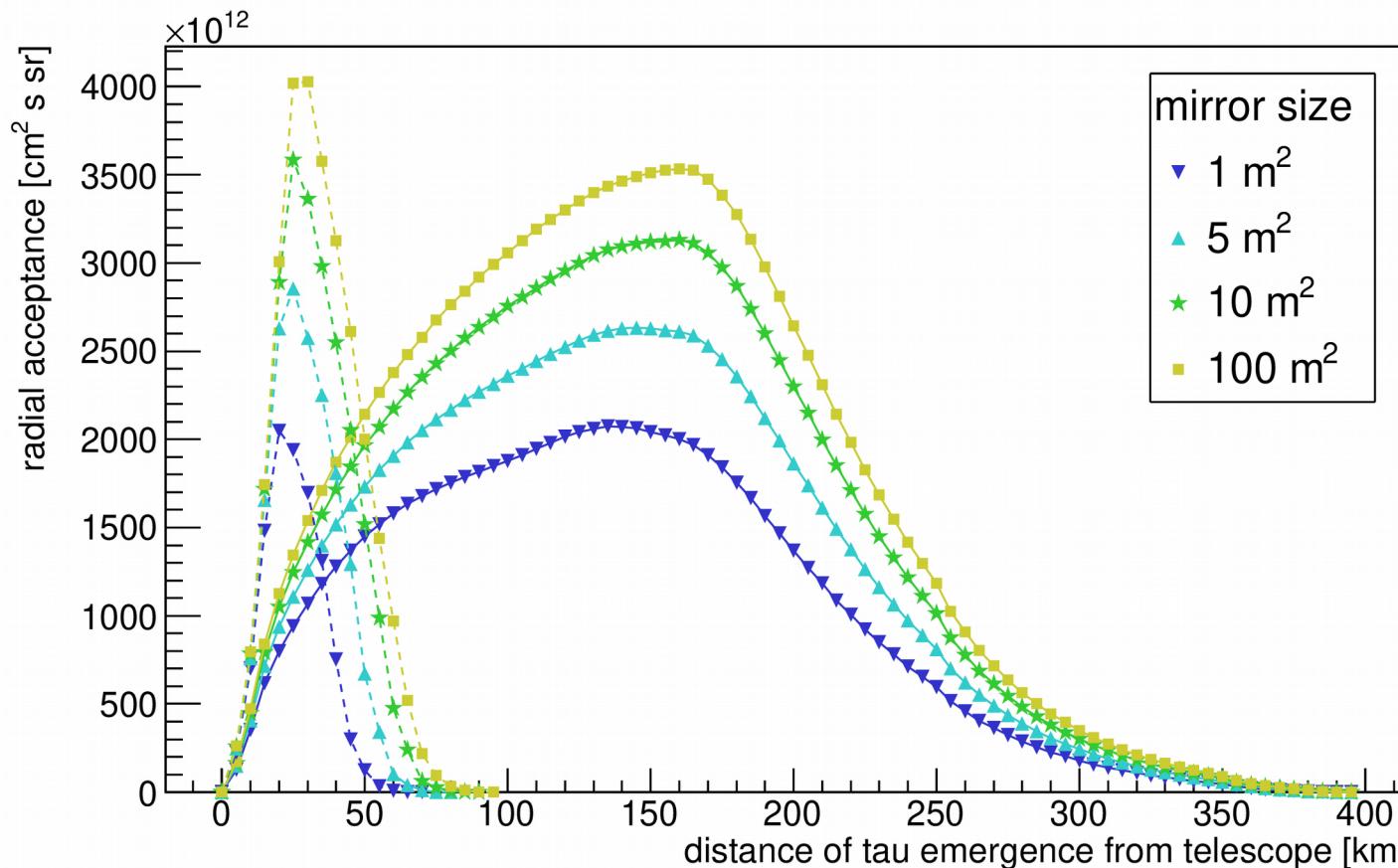
NO et al. EUSO-SPB2 PoS(ICRC2019)977

Nepomuk Otte

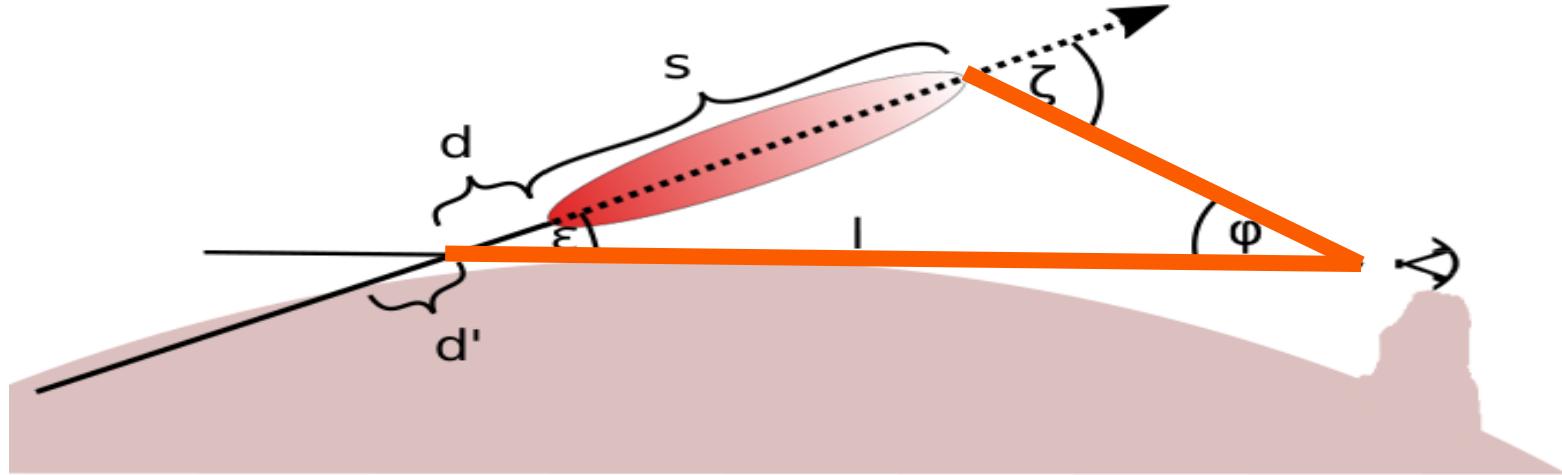
# Detected Cherenkov Light



# Acceptance vs. Light Collection Area

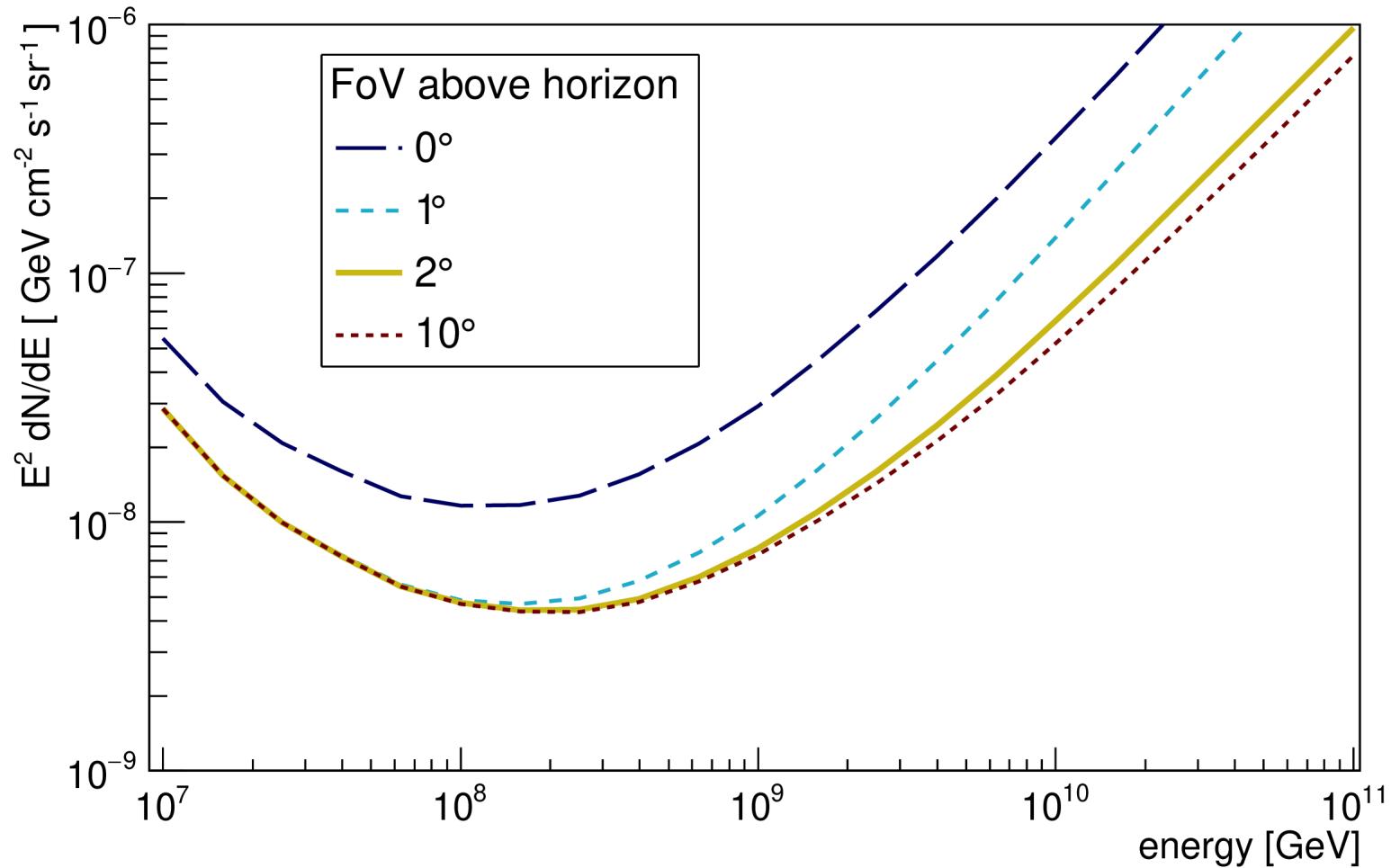


# How much Field of View?

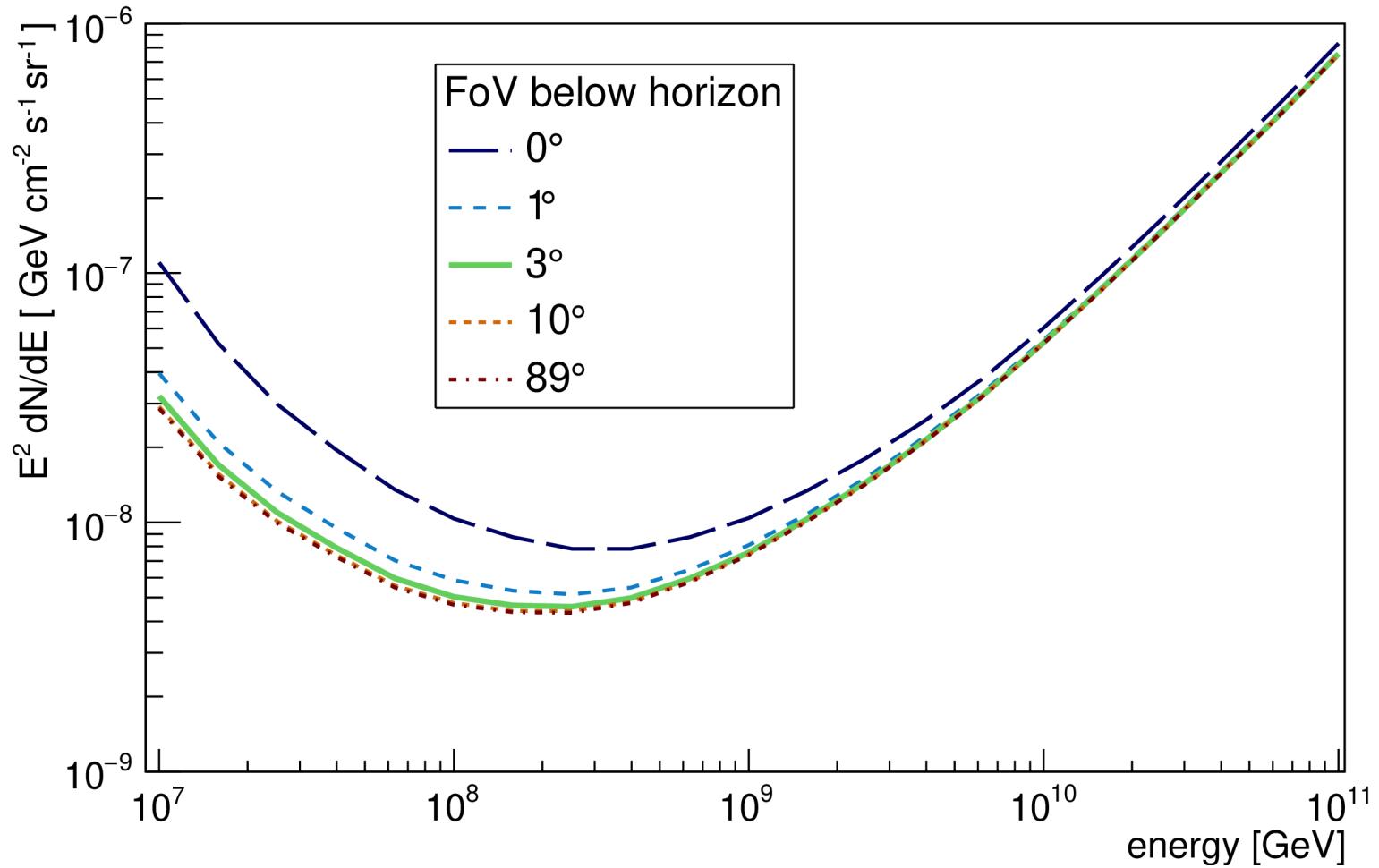


Full image containment is required.

# Sensitivity vs. FoV above Horizon



# Sensitivity vs. FoV below Horizon



PREF

+031.69168° / -110.88498° 8436ft 13:15:52

MAIL

MAP

POSITION - ALTITUDE - TIME



LOG

HORIZON

+00.8°

ANGLE

1° veto region

4° signal region

ELEVATION

-01.9°

ANGLE

AZIMUTH - BEARING

301° N59W 5351mils TRUE

ZERO

A-B

A

W | 300 | 330

LENS

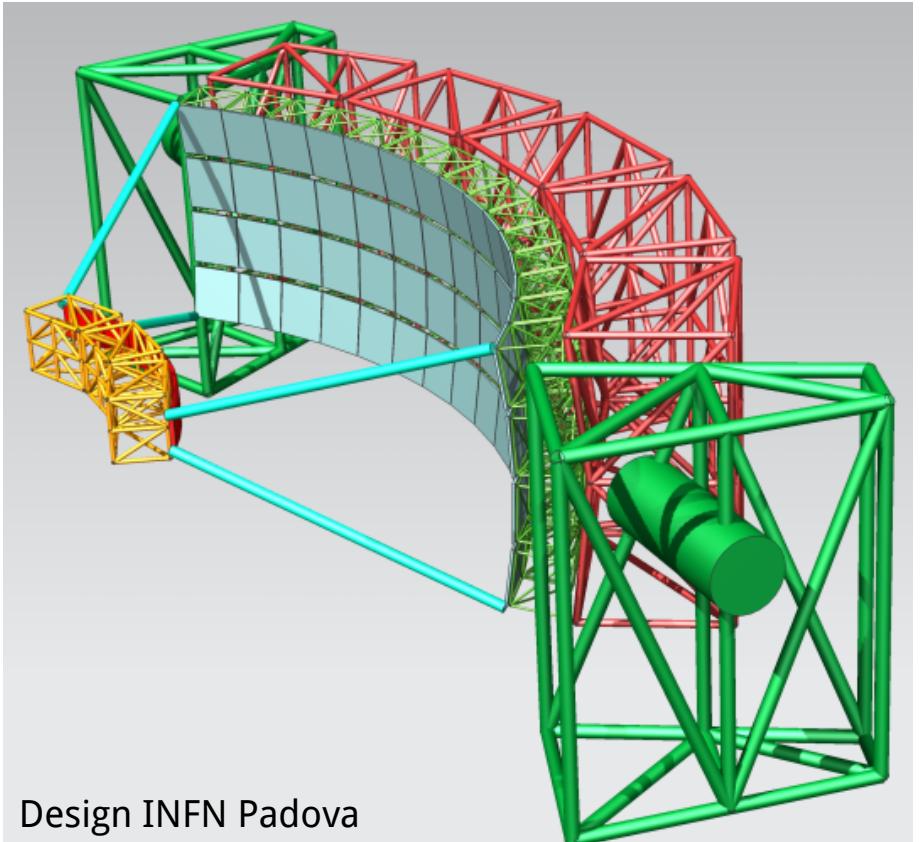
10X

S

# Optics

Based on J. Cortina et al., Astrop. Physics 72 (2016) 46

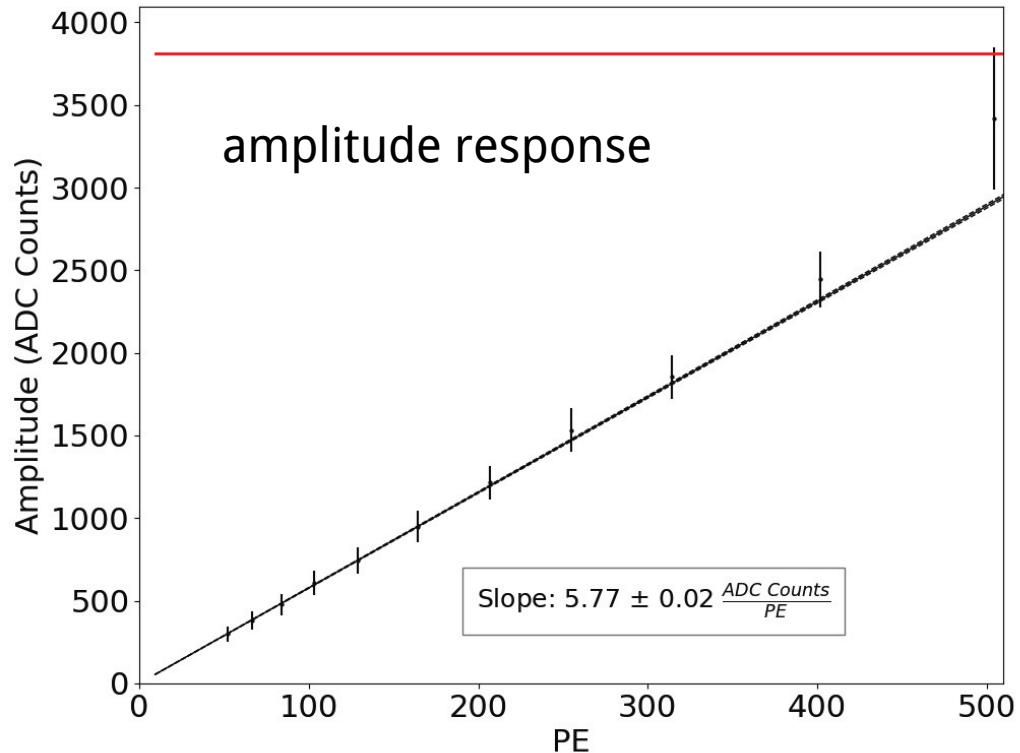
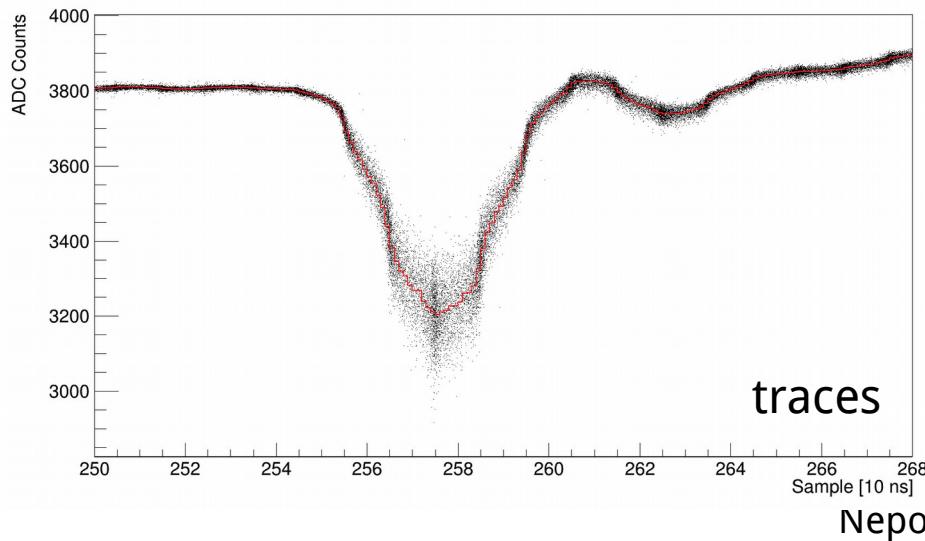
- FoV 5° X 60°.
- 5.6 m focal length.
- 0.3° optical PSF.
- 20 mm Winston cones coupled to **9 mm SiPMs**.
- 3,300 pixel camera.
- 68 m<sup>2</sup> mirror area → **16 m<sup>2</sup>** in any direction.
- Rotates in elevation.
- Thin-glass replica mirror technology ~\$2k/m<sup>2</sup>.
- Implementation based on MAGIC structure.
- \$170k for one telescope (excl. camera).



# Signal Chain

Same as in NO et al. EUSO-SPB2 PoS(ICRC2019)977

- MUSIC: preamp ASIC developed for IACTs.
- AGET digitizer: 100MS/s, switch capacitor readout, 12 bit.
- \$100 per channel → \$330k per camera



Tested with picosecond laser flashing Hamamatsu S14520 SiPM + MUSIC + AGET

# Site

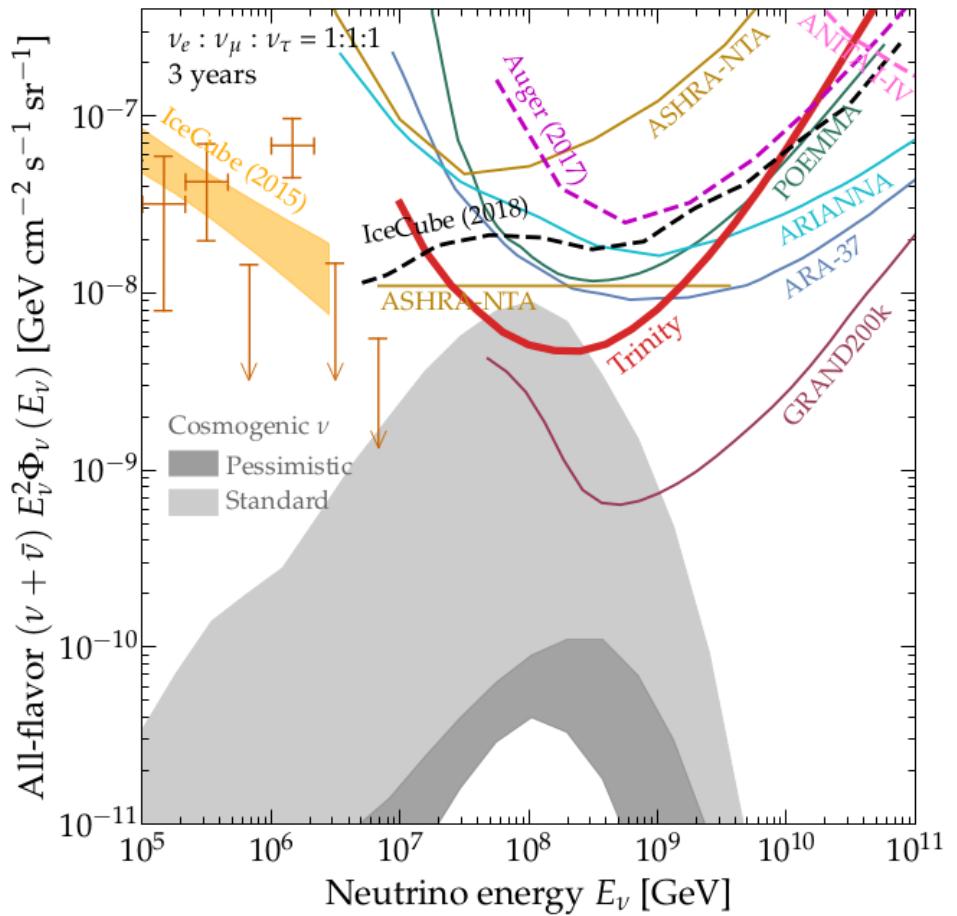
- Needs to oversee a remote area.
- Modest light contamination is ok.
  - Images happen on <100ns timescales.
- Ashra Site; BEACON Test site; Frisco Peak, UT.



View from Frisco Peak, UT

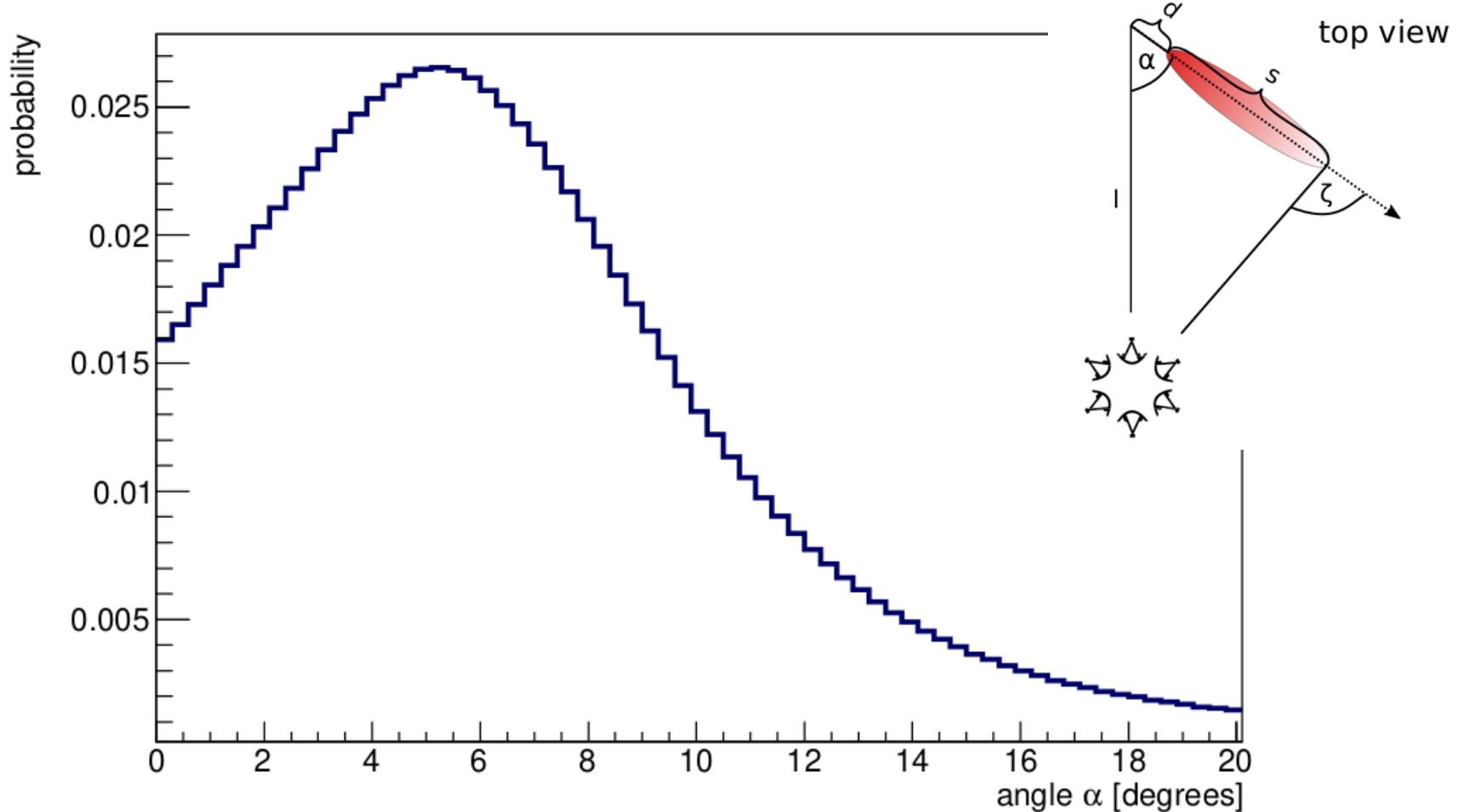
# Conclusions

- Air-shower imaging is a viable technique to search for UHE neutrinos.
- Trinity is optimized to detect distant UHE tau showers.
- Cost effective solutions exist for the optics and the camera (~\$3M for 360°).
- Developing of main components is ongoing.
- Deploy prototype next year?

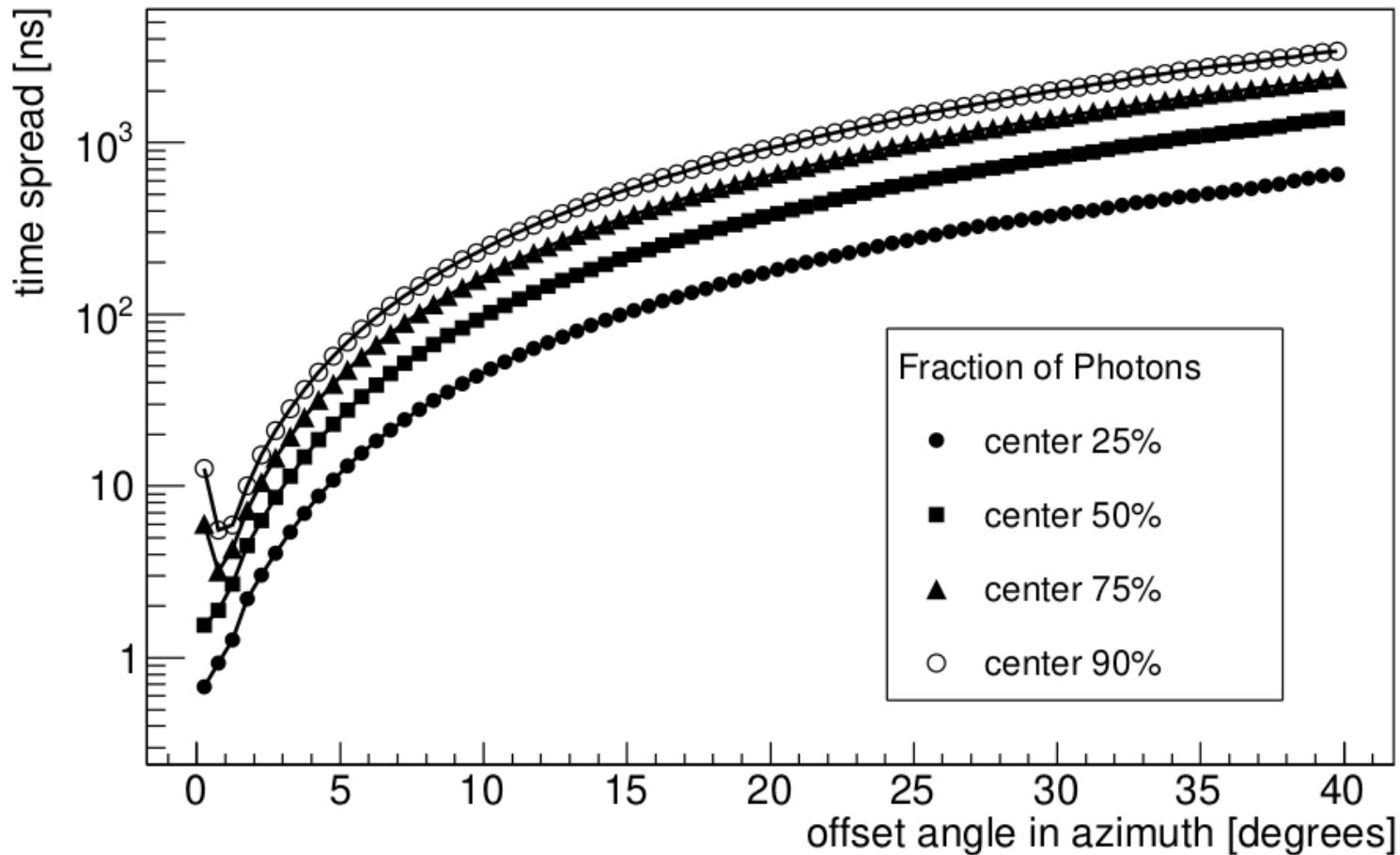


# Backup

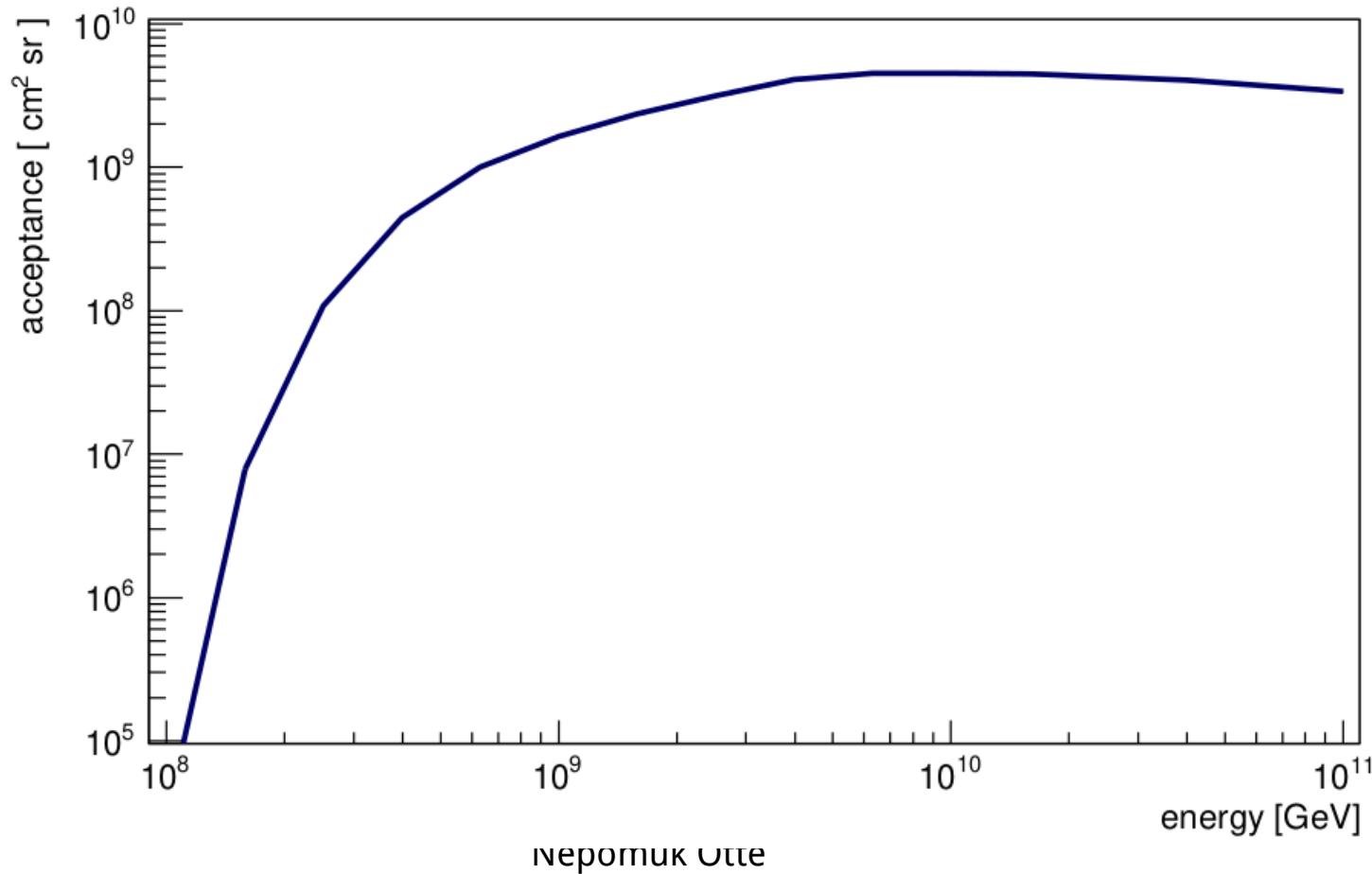
# Triggered Viewing Angles



# Spread of Photon Arrival Times

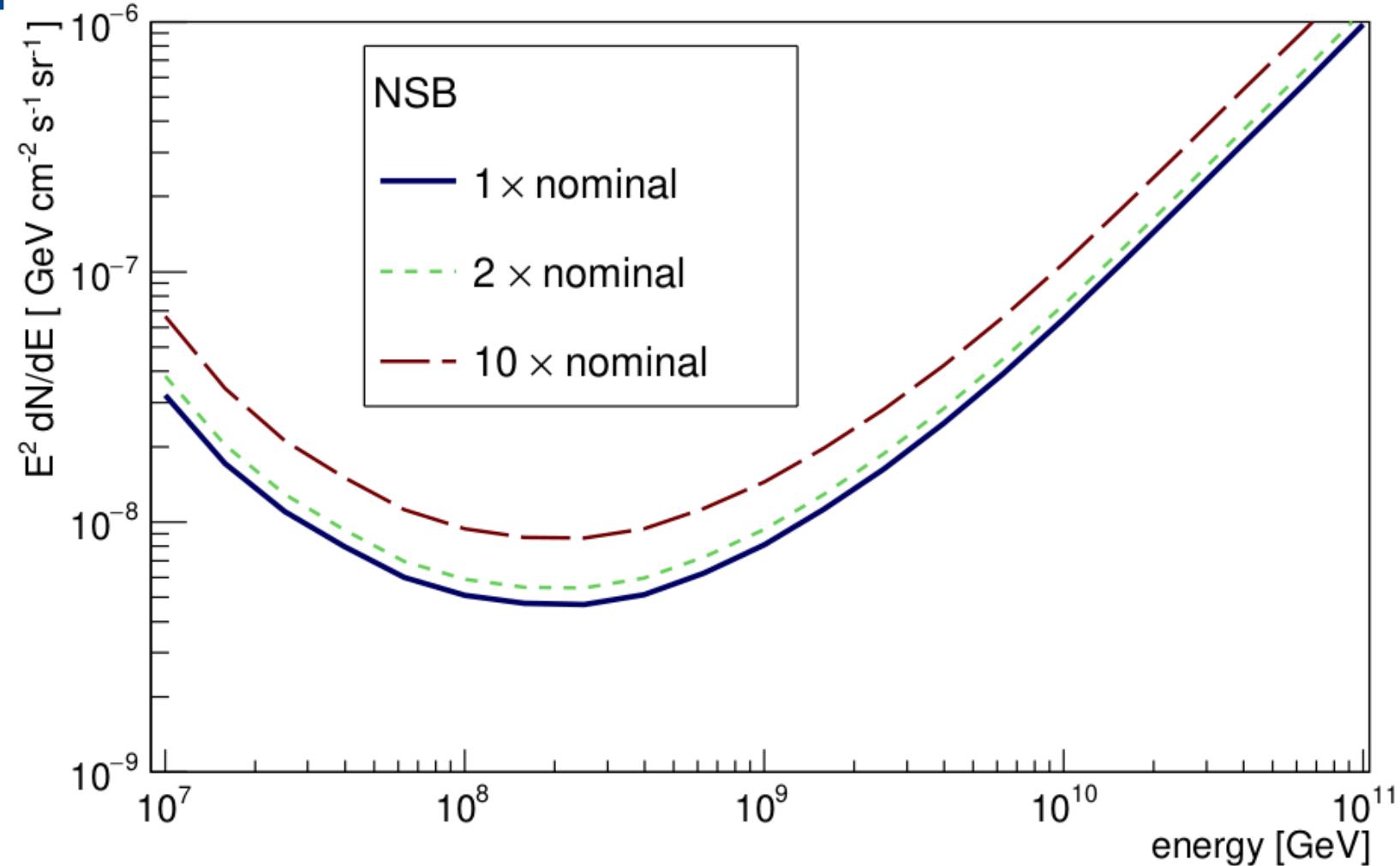


# Acceptance

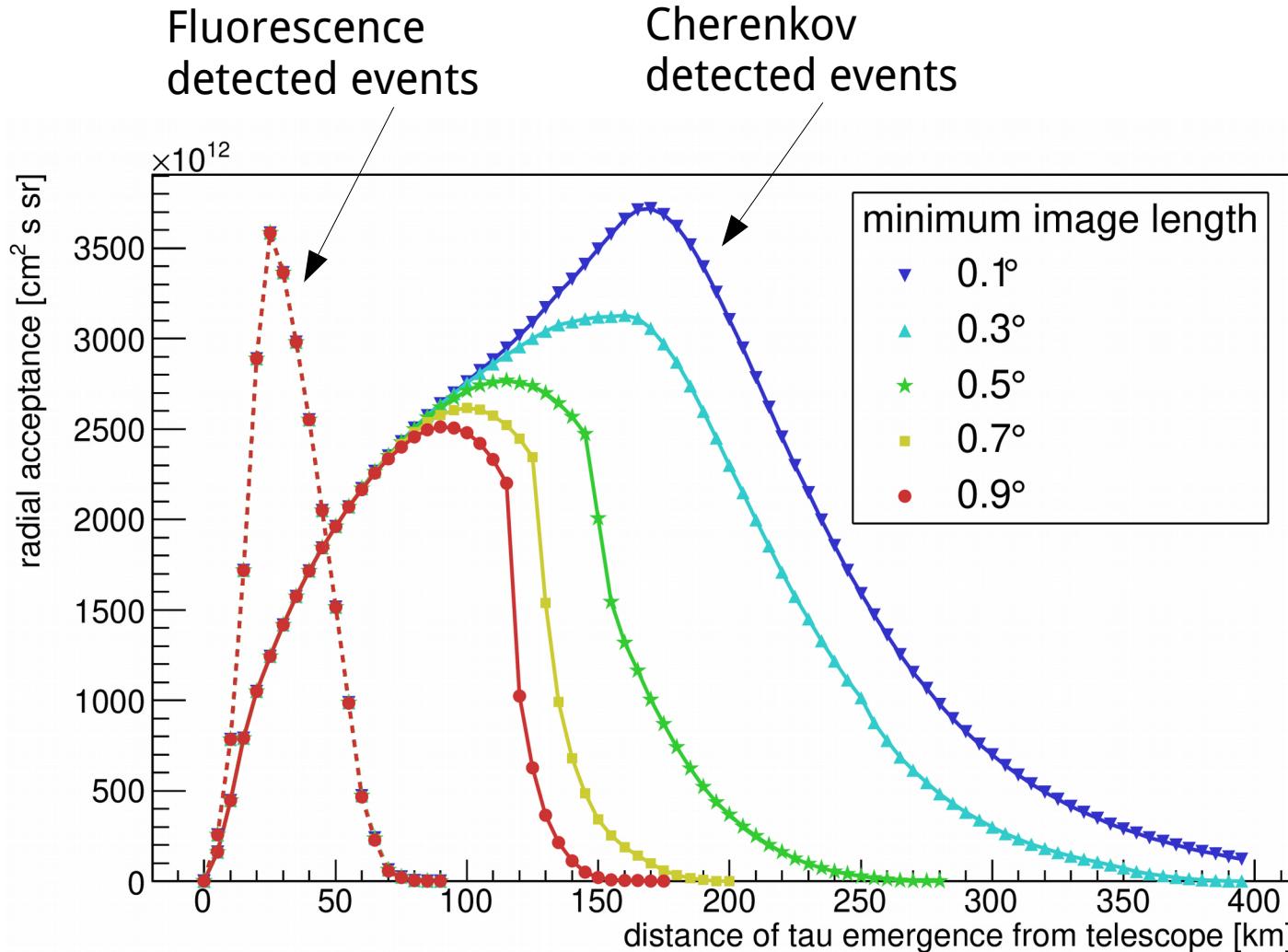


Nepomuk Ulte

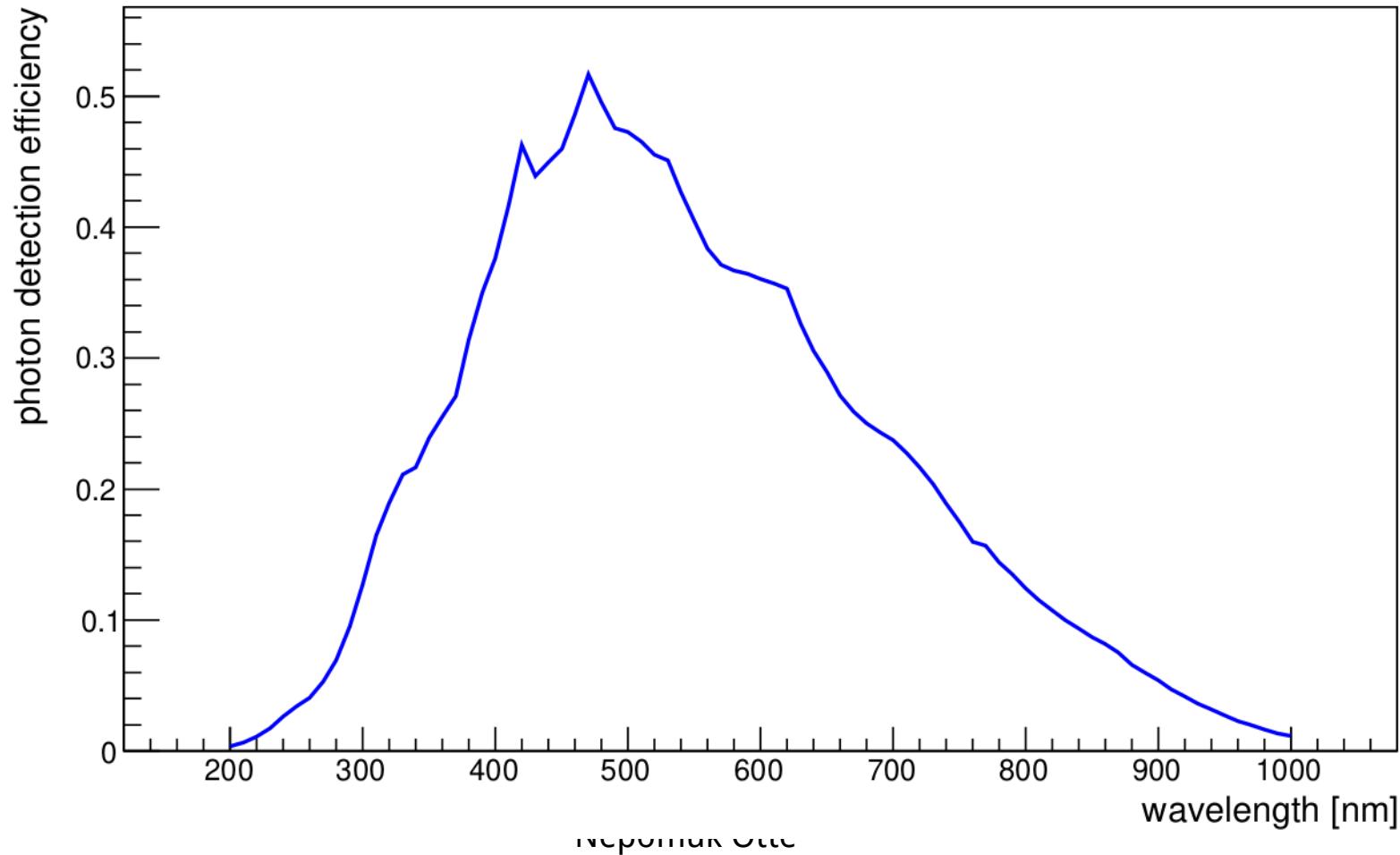
# Impact of Night Sky Background



# Angular Resolution



# PDE Hamamatsu S14520



# Thin-Glass Mirrors



Nepomuk Otte