



中国科学院高能物理研究所
Institute of High Energy Physics
Chinese Academy of Sciences

The Cosmic-ray Anisotropy Observed by YBJ-HA Experiment

Yingying Guo

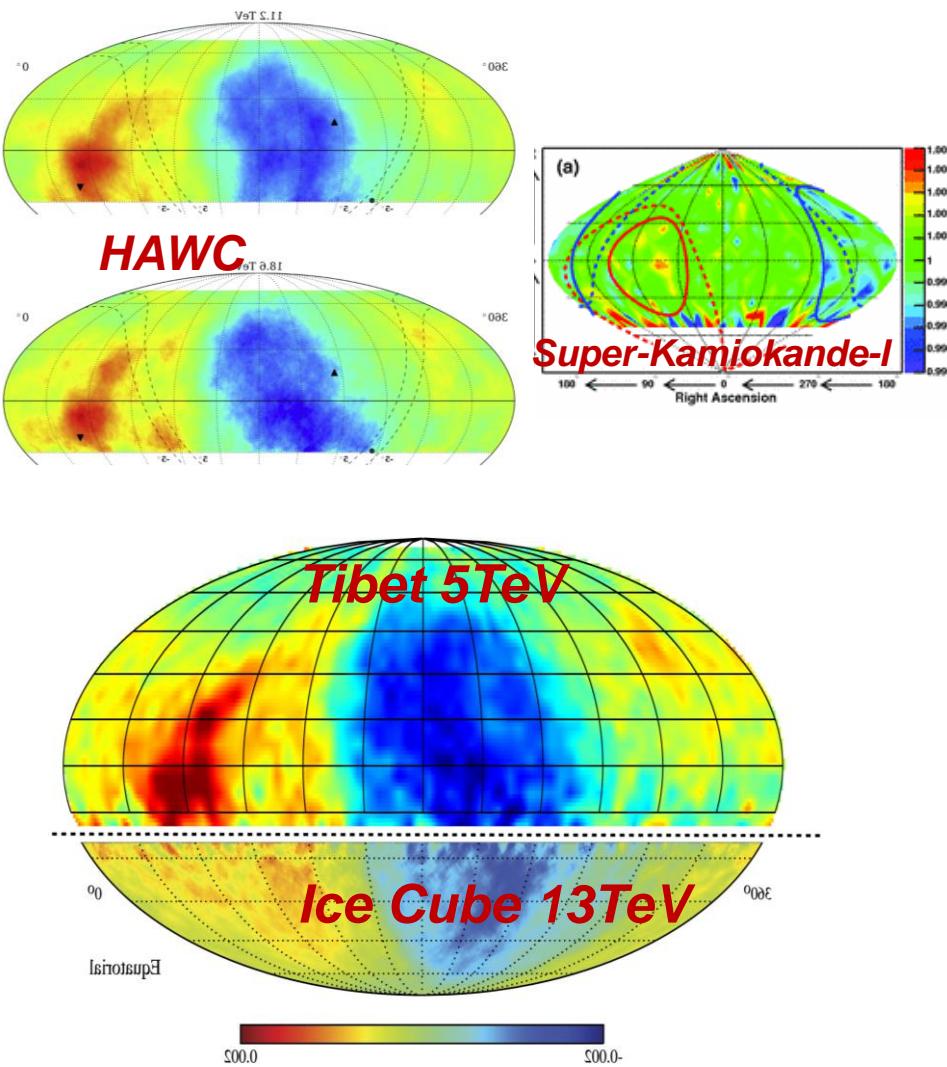
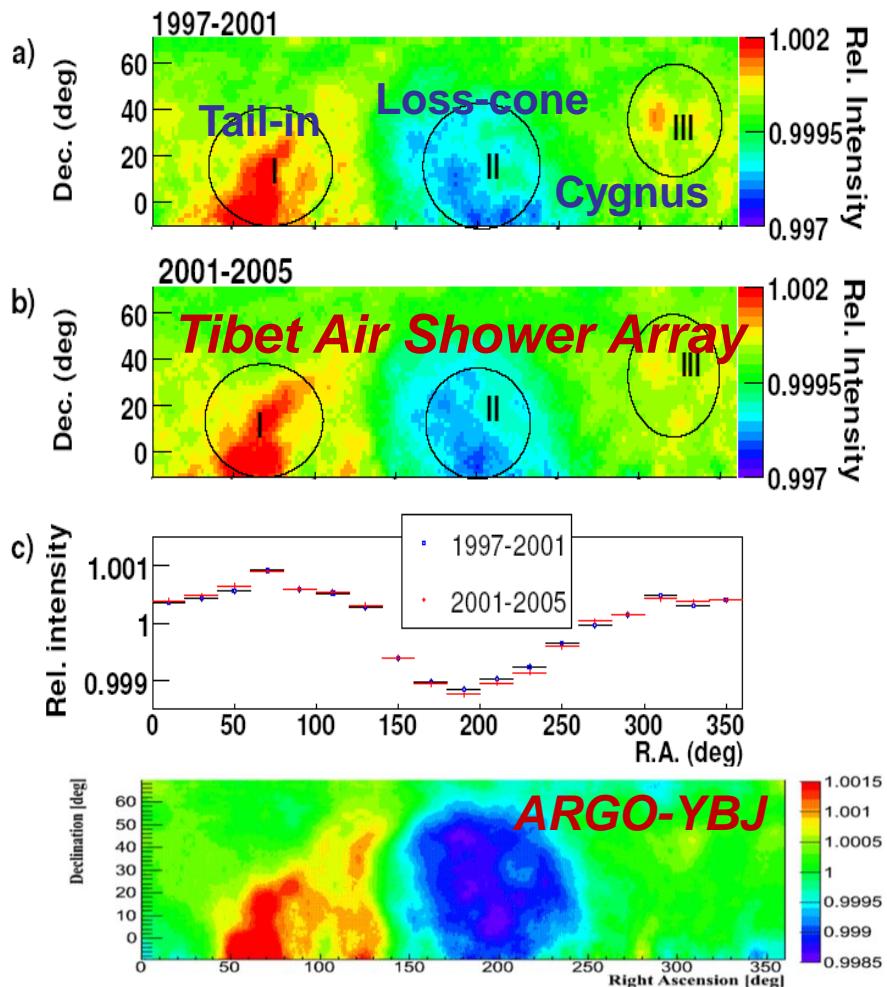
*Institute of High Energy Physics,
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For 2019 ICRC

Outline

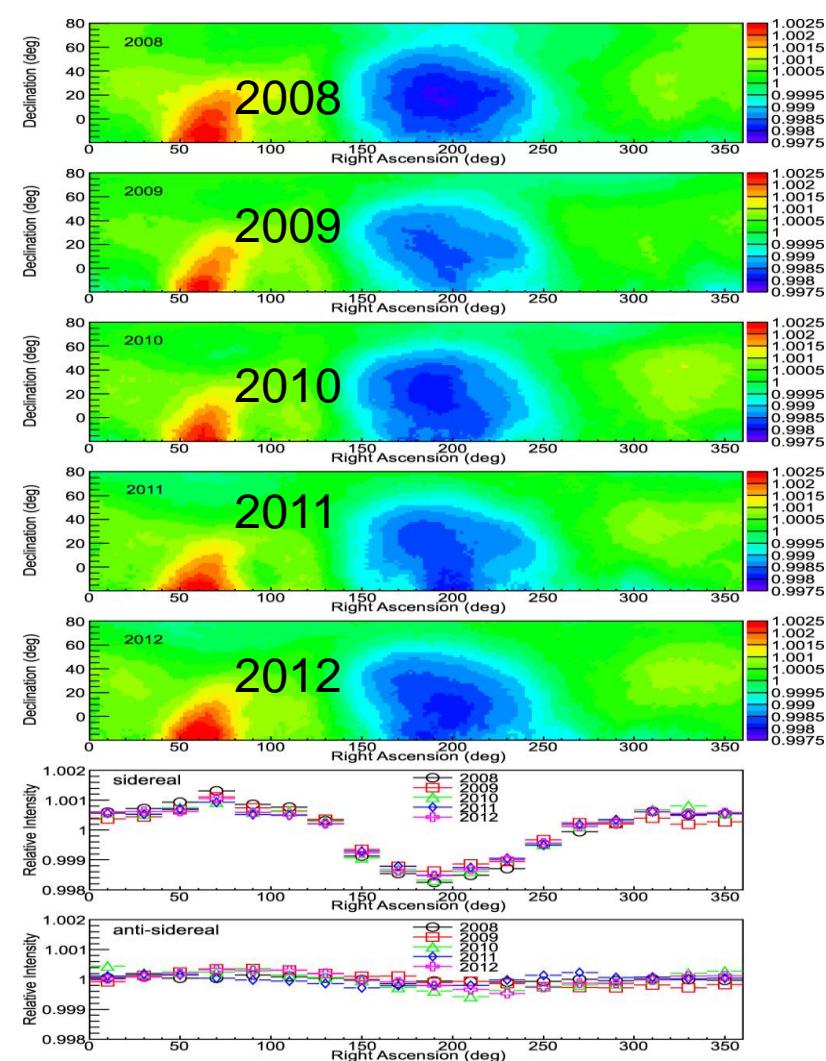
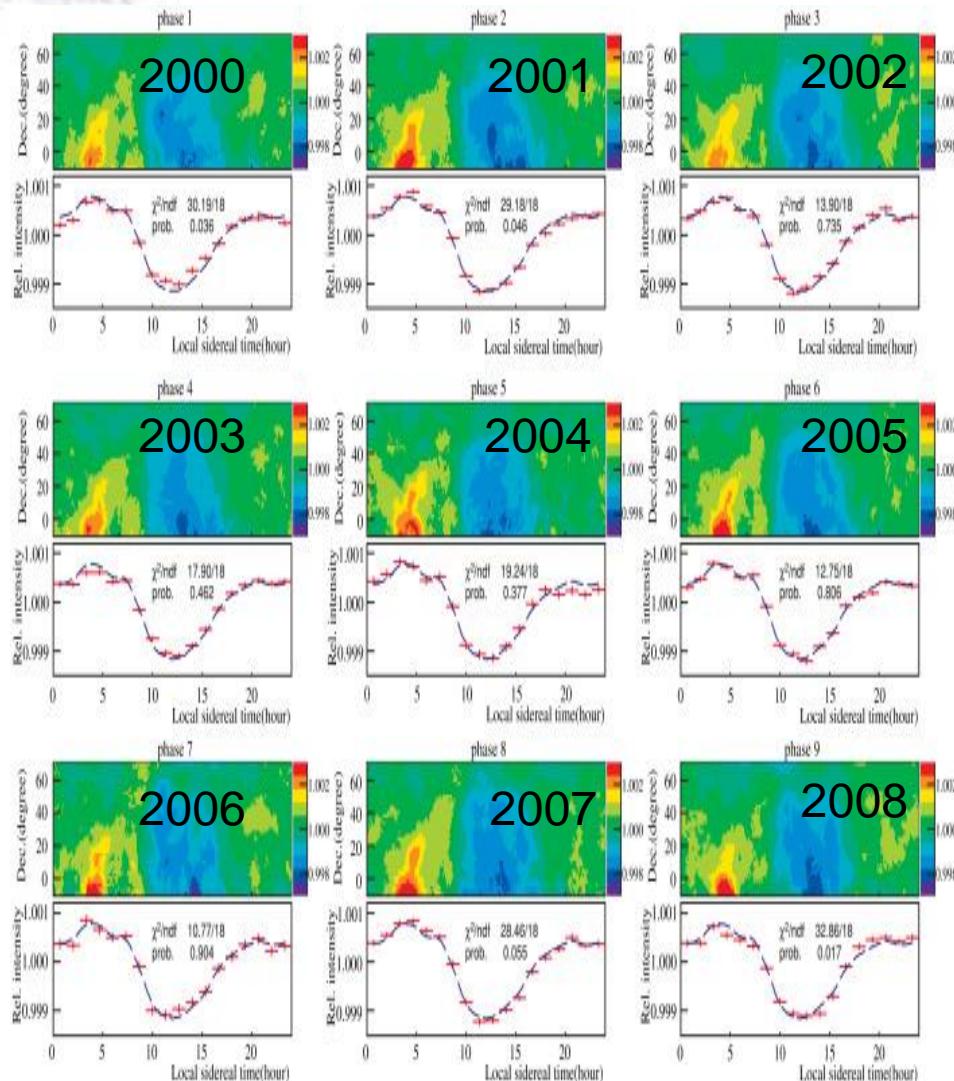
- *Some observations of anisotropy*
- ***YBJ-HA(YangBajing Hybrid Array)***
- *Data selection && Background estimation*
- *Summary*

Experiment Observation



<https://doi.org/10.1126/SCIENCE.1131702> <https://doi.org/10.3847/1538-4357/aad90c>
<https://doi.org/10.1016/j.pppn.2017.01.004> <https://doi.org/10.1103/PhysRevD.75.062003>

No time dependence of Sidereal anisotropy

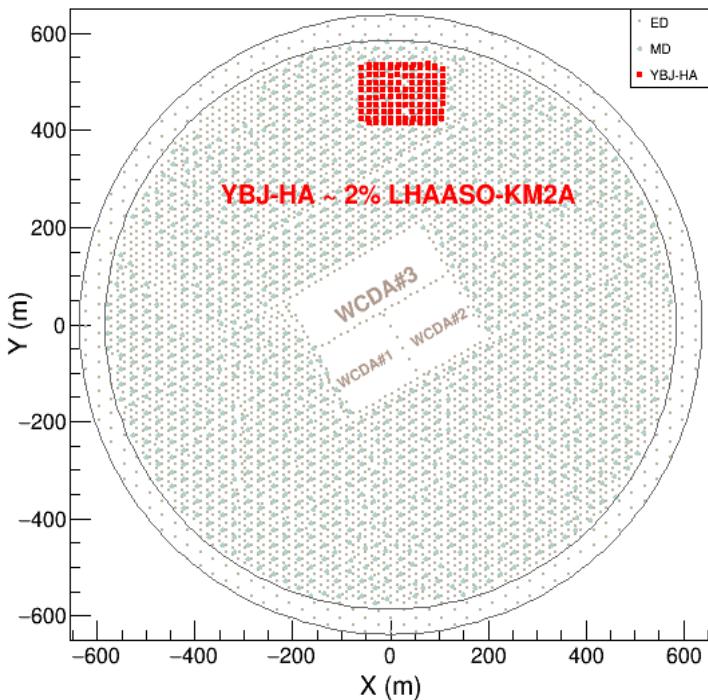


THE TIBET III AIR SHOWER ARRAY

ARGO-YBJ

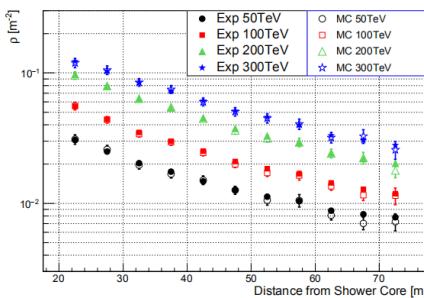
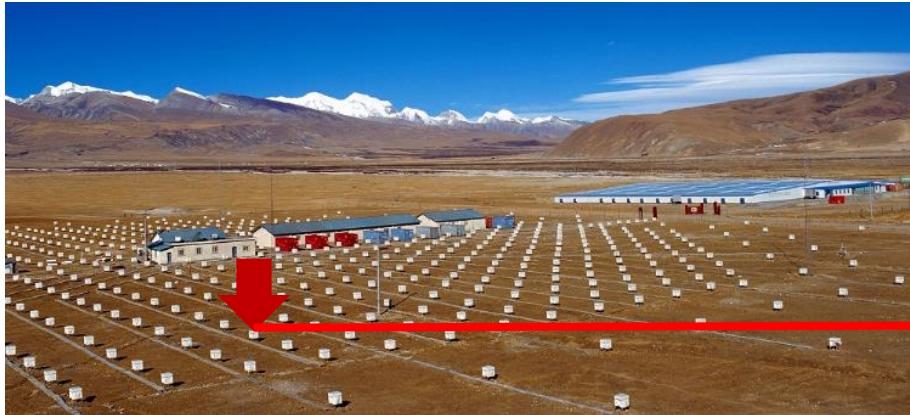


YBJ-HA Array / testing Array of LHAASO-KM2A



	YBJ-HA	LHAASO-KM2A
<i>Altitude</i>	<i>4300m</i>	<i>4400m</i>
<i>Detector type</i>	<i>ED && MD</i>	<i>same</i>
<i>Separation</i>	<i>15m</i>	<i>same</i>
<i>time synchronization system</i>	<i>white rabbit clock</i> 	<i>same</i>
<i>electronics</i>	ED 	
	<i>MD</i> 	<i>same</i>

YangBaJing Hybrid Array(YBJ-HA)

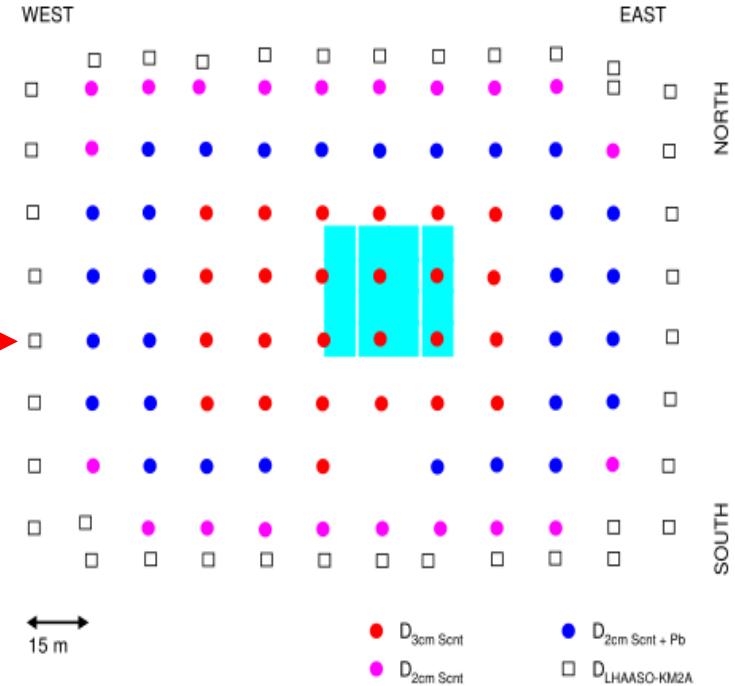


Muons lateral density distribution

Diffuse γ rays upper limits on flux

Data phase and the their respective number of events for analysis

Phase	Time(days)	Number of Events
1	68	1.081×10^8
2	74	1.361×10^8
3	135	2.701×10^8
4	119	2.886×10^8

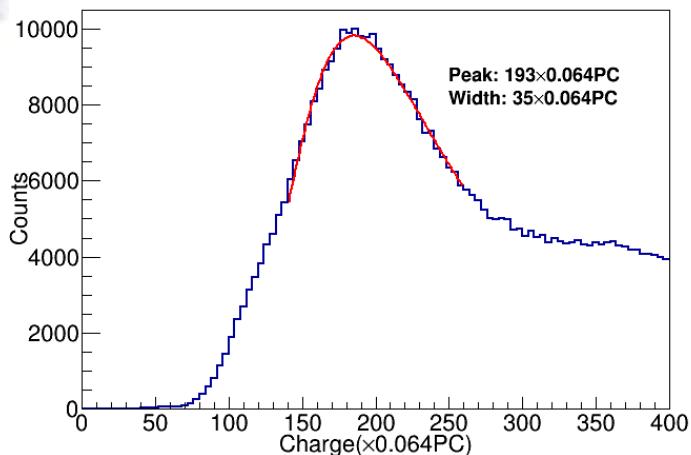


Detector types:(115EDs+16MDs)

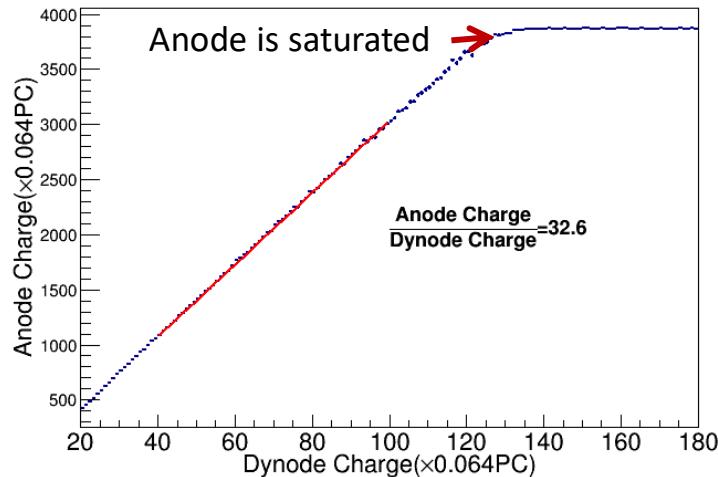
➤ **115 Electromagnetic particle Detectors(EDs)**

➤ **16 Muon Detectos(MDs)**

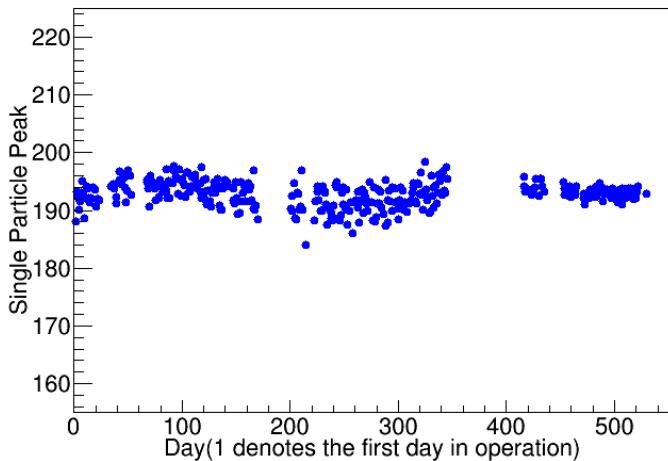
ED charge calibration && electronics stability



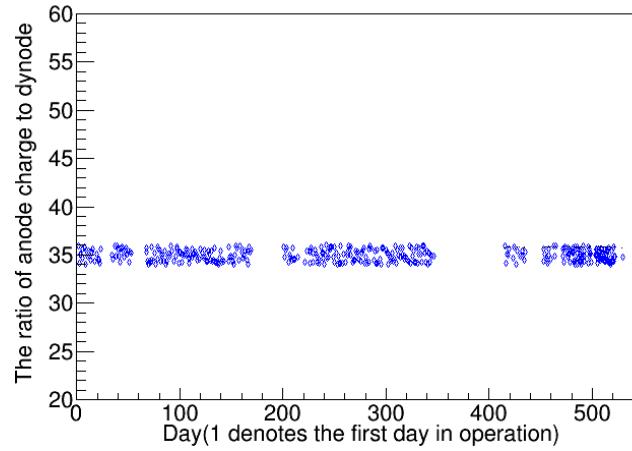
Single particle peak



The ratio of anode charge to dynode

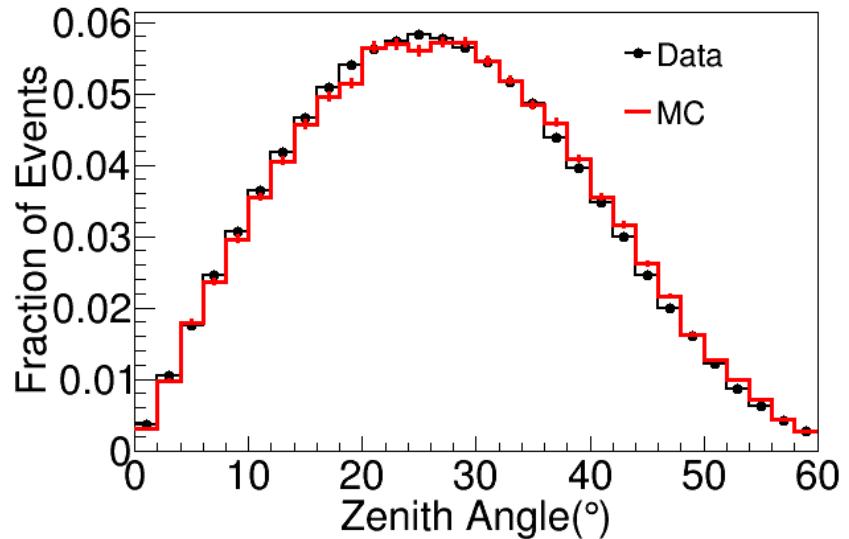
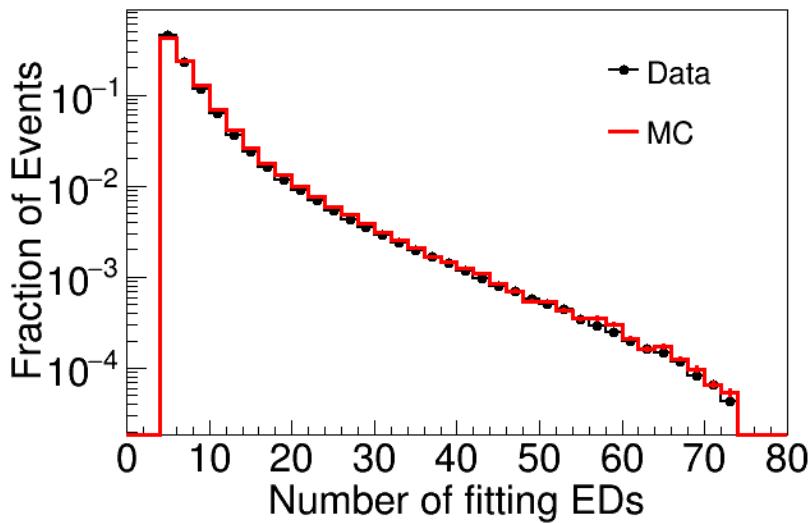
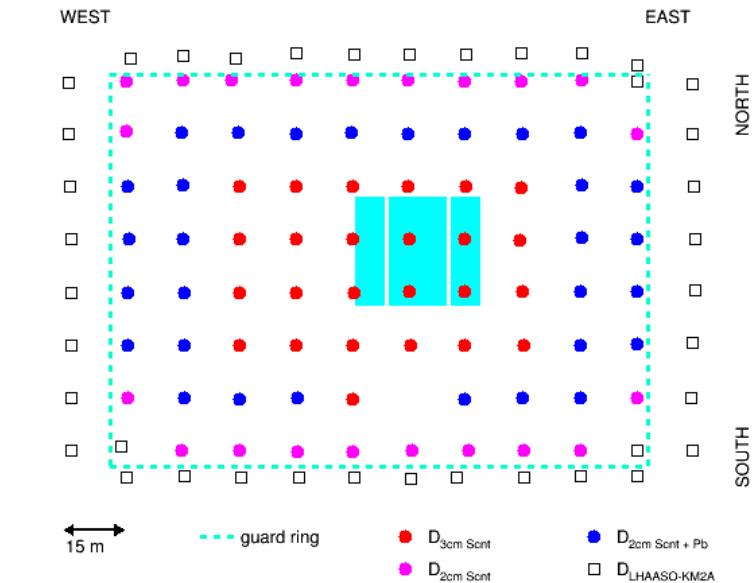


The stability single particle peak



The stability of the ratio

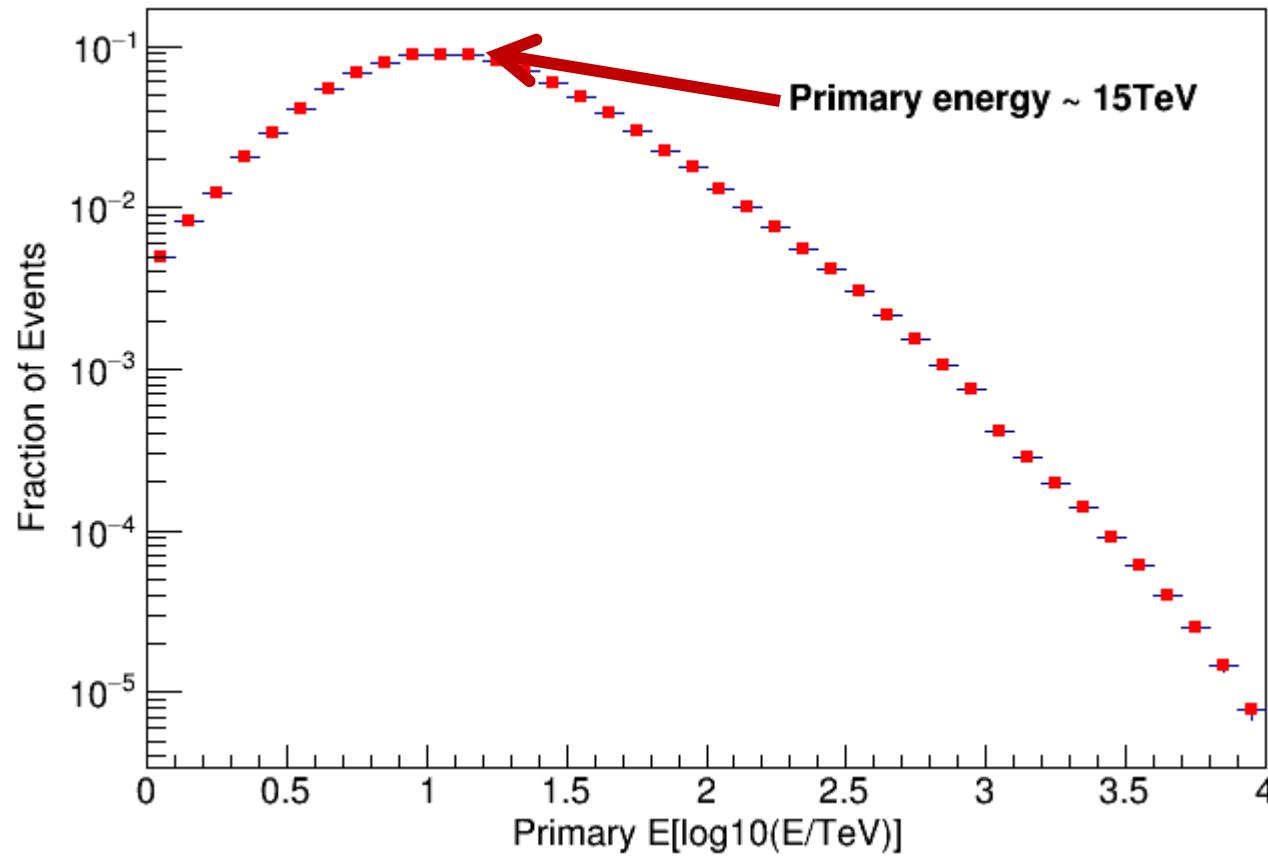
Data selection



- Criteria:
on array
The number of EDs > 4
the zenith angle <60 $^{\circ}$
- Monte Carlo agrees with data



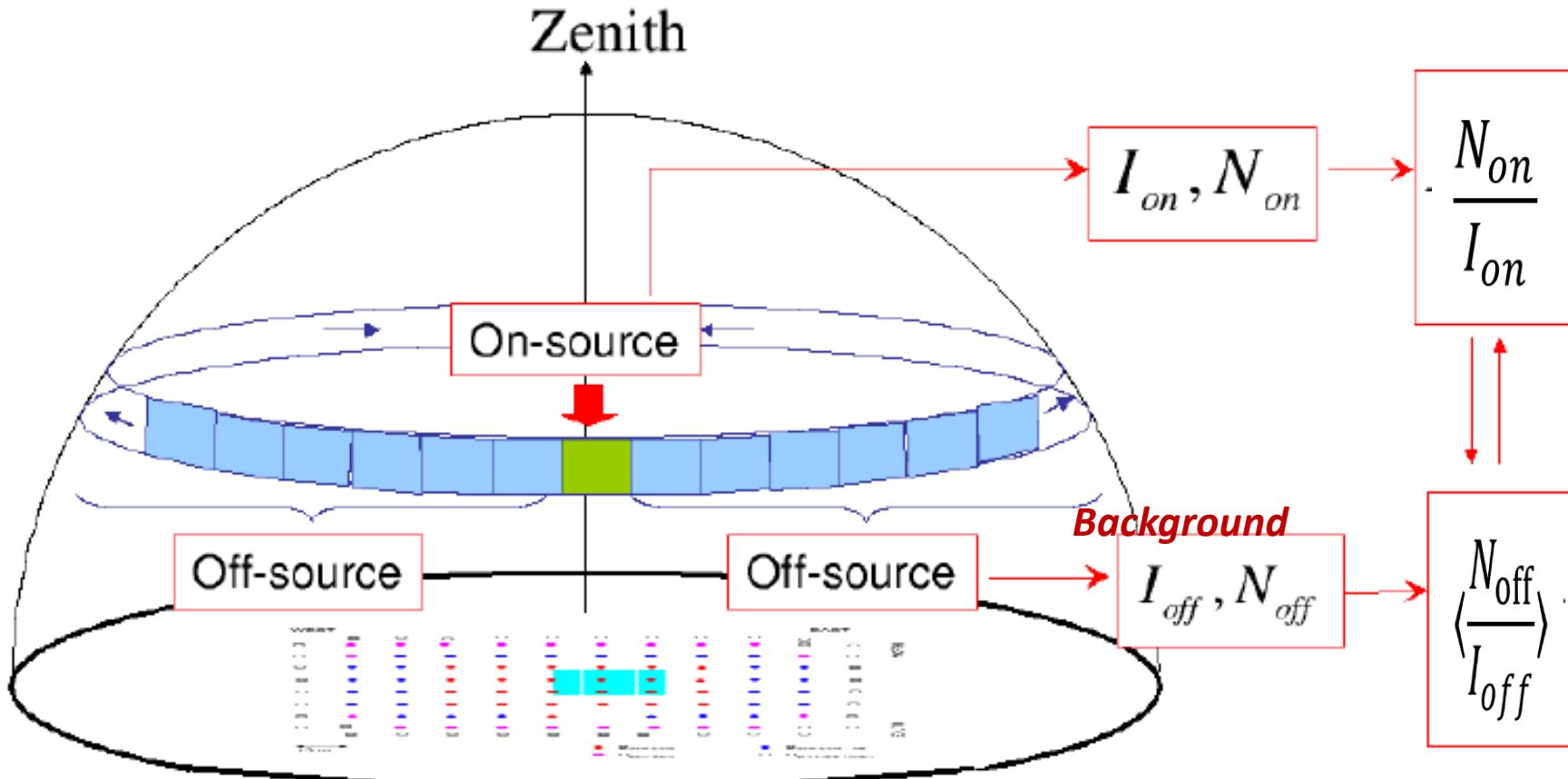
Energy Distribution



The simulated primary energy distribution



Background estimation

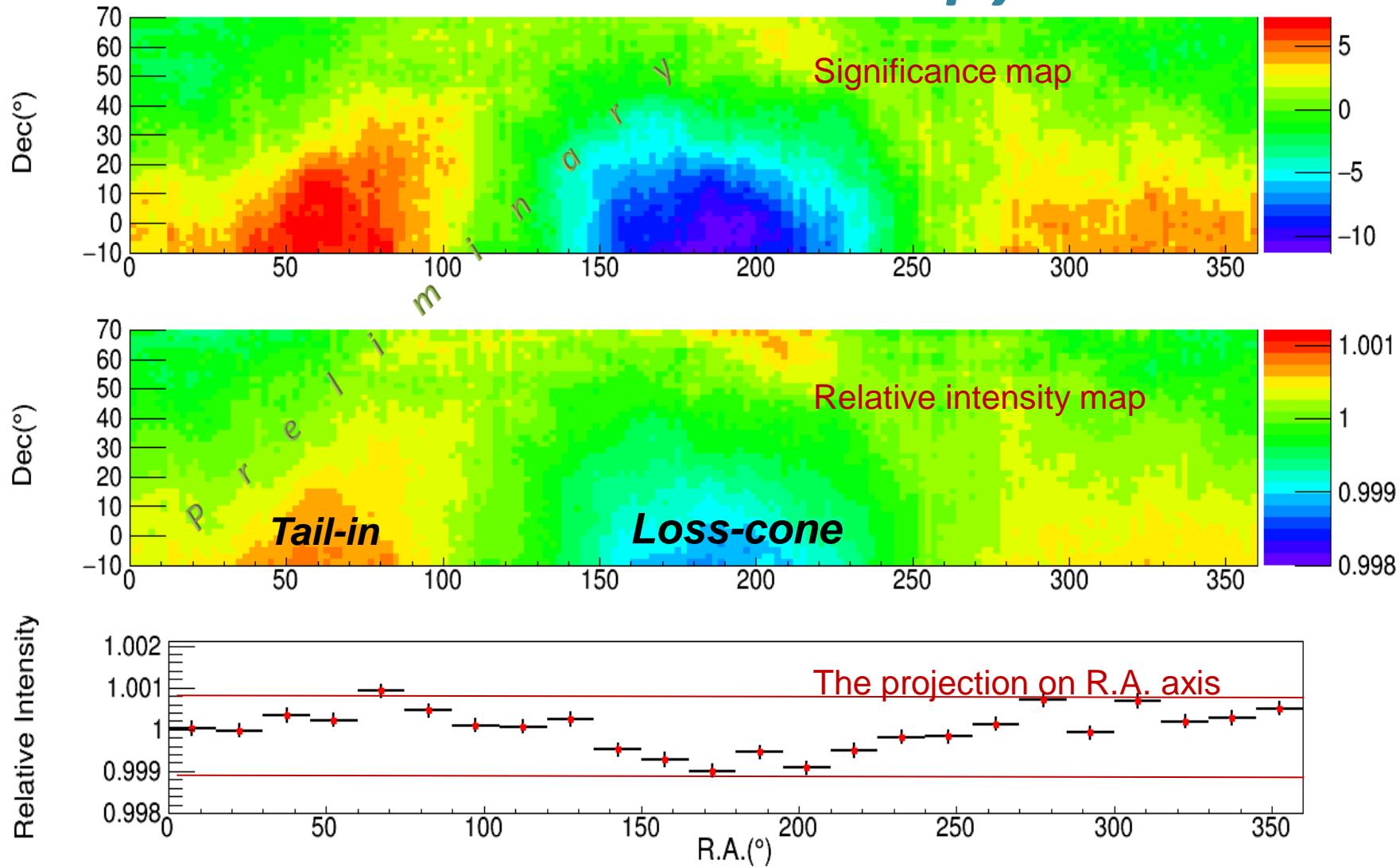


All-Distance Equi-Zenith Angle Method

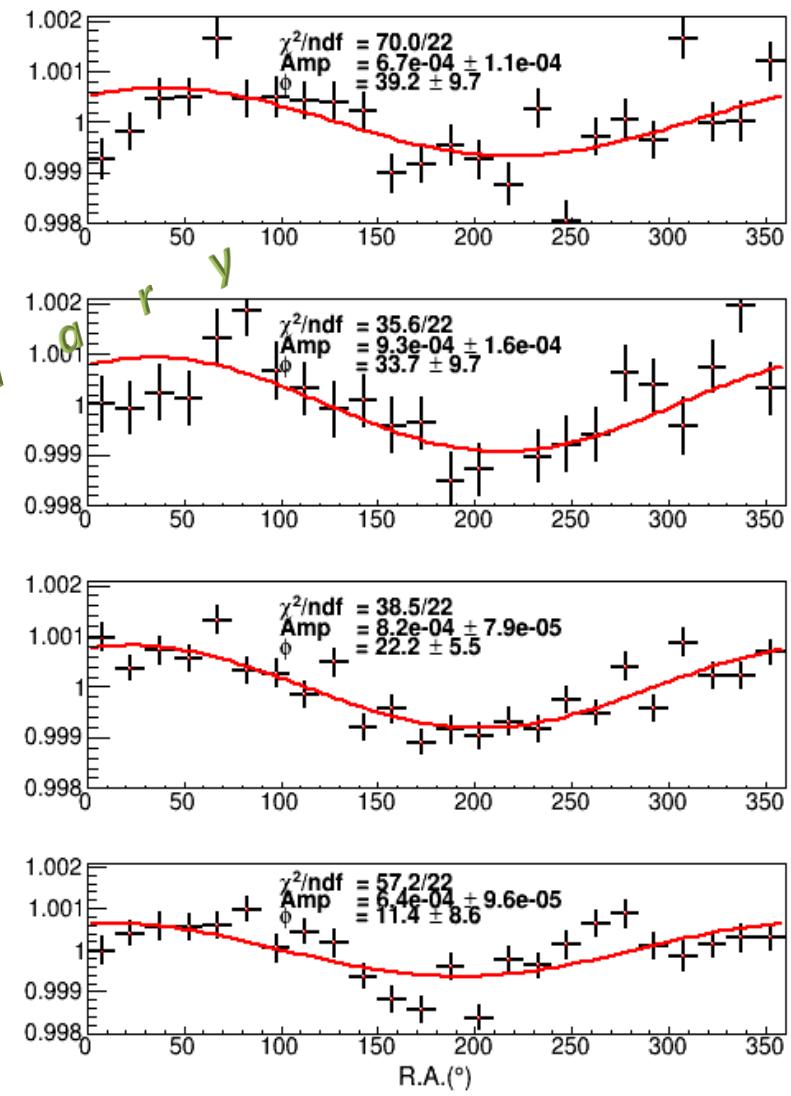
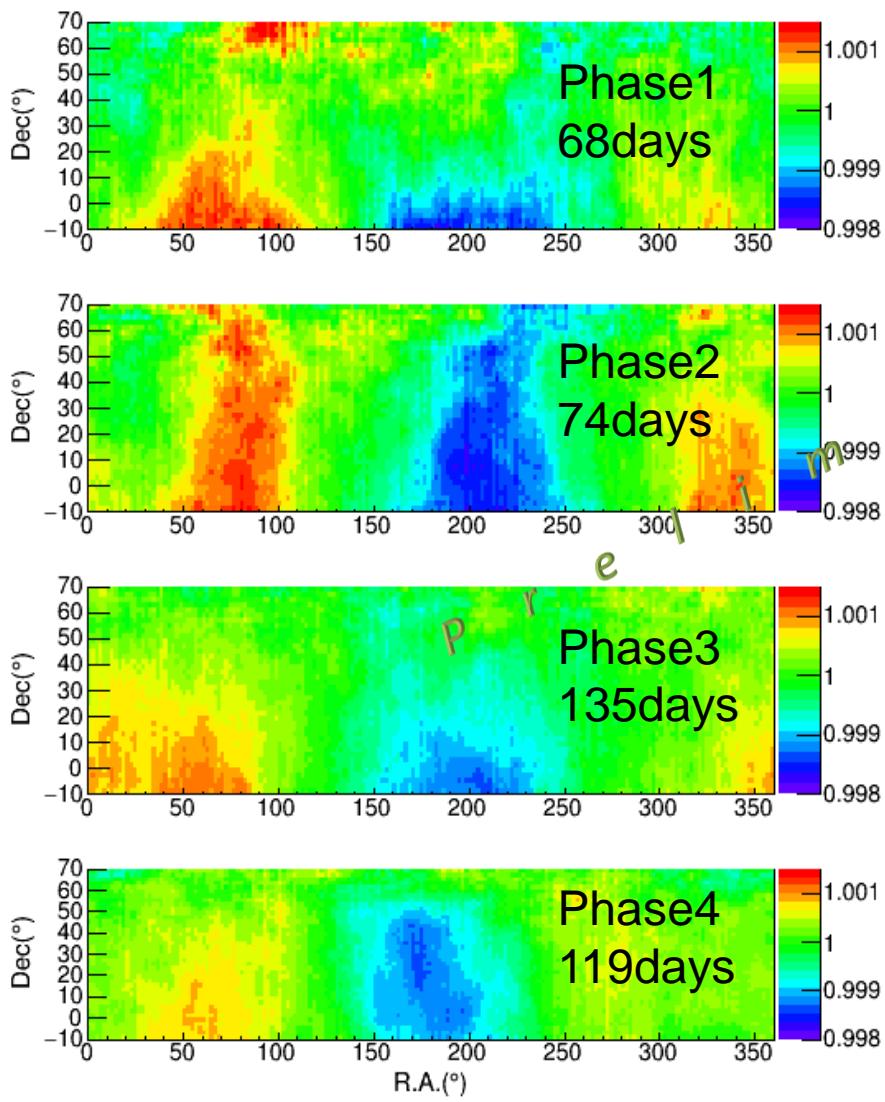
I indicates the relative intensity in some direction

$$Signi_{ij} = \frac{I_{ij} - 1}{\Delta I_{ij}}$$

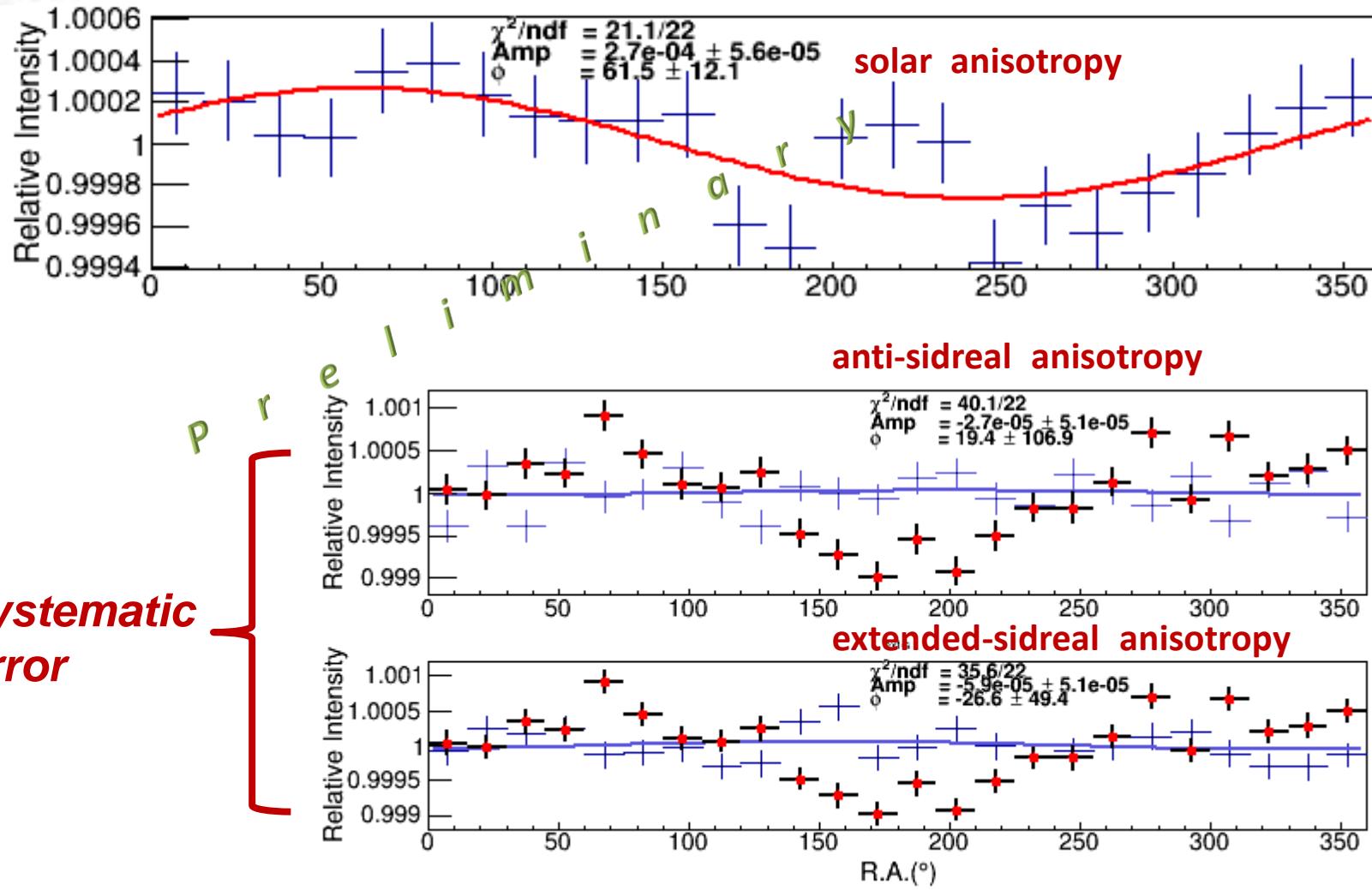
Sidereal Anisotropy



No time dependence of sidereal anisotropy



Systematic checks



summary

- ***We observed the sidereal anisotropy of cosmicray at the median energy of 15 TeV.***
- ***No time-dependence of the Large-structure anisotropy.***
- ***The anti/extended-sidereal anisotropy is $\sim 10^{-5}$, which ensured the reliability of sidereal anisotropy .***
- ***The YBJ-HA is running, LHAASO-KM2A is under construction***